Cosmology PSET 6:

Problem 1: Dark Matter and Baryon Density Growth

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In[2]:= vMap = ResourceFunction["ViridisColor"];
   ln[3]:= ode = D[\epsilon[t], \{t, 2\}] + (4/(3t)) D[\epsilon[t], t] == 0;
            \epsilonSol = DSolveValue[ode, \epsilon[t], t]
  Out[4]= -\frac{3 \mathbb{C}_1}{+1/3} + \mathbb{C}_2
  ln[5]:= params = \left\{ \Omega d \rightarrow \frac{5}{6}, \Omega b \rightarrow \frac{1}{6} \right\};
  ln[6]:= odeD = D[\delta_d[t], \{t, 2\}] + \frac{4}{3t}D[\delta_d[t], t] - \frac{2}{3t^2}(\Omega d \delta_d[t] + \Omega b \delta_b[t]) = 0
  Out[6]= -\frac{2 (\Omega b \delta_b[t] + \Omega d \delta_d[t])}{3t^2} + \frac{4 \delta_d'[t]}{3t} + \delta_d''[t] == 0
   In[7]:= odeD /. params
  \text{Out}[7] = -\frac{2\left(\frac{\delta_b[t]}{6} + \frac{5\delta_d[t]}{6}\right)}{2+2} + \frac{4\delta_d'[t]}{3t} + \delta_d''[t] = 0
   ln[8]:= odeB = D[\delta_b[t], \{t, 2\}] + \frac{4}{3t}D[\delta_b[t], t] - \frac{2}{3t^2}(\Omega d \delta_d[t] + \Omega b \delta_b[t]) = 0
  \text{Out[8]=} \ -\frac{2 \ (\Omega b \ \delta_b[t] + \Omega d \ \delta_d[t])}{3 \ t^2} \ + \frac{4 \ \delta_b{'}[t]}{3 \ t} \ + \delta_b{''}[t] = 0
   In[9]:= odeB /. params
  Out[9]= -\frac{2\left(\frac{\delta_{b}[t]}{6} + \frac{5\delta_{d}[t]}{6}\right)}{3+2} + \frac{4\delta_{b}'[t]}{3+} + \delta_{b}''[t] == 0
 In[10]:= sysSolution = NDSolveValue[
                 {odeD, odeB, \delta_d[1] = 1*^-3, \delta_b[1] = 1*^-4, \delta_d'[1] = 0, \delta_b'[1] = 0} /. params,
                 \{\delta_{d}, \delta_{b}\}, \{t, 1, 500\}]
Out[10]=
             \begin{array}{c|c} \textbf{InterpolatingFunction} \Big[ & \blacksquare & \_ & \texttt{Domain: } \{ \texttt{\{1., 500. \}\}} \\ & \texttt{Output: scalar} \\ \end{array} \Big] \Big\}
```

 $\label{eq:deltapprox} $$ \inf[11] = \delta DPlot = Plot[sysSolution[1]][t], \{t, 1, 500\}, ScalingFunctions \rightarrow \{"Log", "Log"\}, $$ PlotStyle \rightarrow vMap[0.6], Frame \rightarrow True, FrameLabel \rightarrow \{"t", "\delta"\}] $$$

Out[11]=

0.010
0.005
0.001

10

In[12]:= δ BPlot = Plot[sysSolution[[2]][t], {t, 1, 500}, ScalingFunctions → {"Log", "Log"}, PlotStyle → vMap[0.2], Frame → True, FrameLabel → {"t", " δ "}]

100

500

50

Out[12]=

0.010
0.005
0.001
5×10⁻⁴
1×10⁻⁴
1 5 10 50 100 500

In[13]:= Show[δ BPlot, δ DPlot]

Out[13]=

0.010
0.005
0.001
5×10⁻⁴
1×10⁻⁴
1 5 10 50 100 500

```
In[14]:= ratio = sysSolution[1][t]
sysSolution[2][t];
 In[15]:= ratioPlot = Plot[{ratio, 1.1}, {t, 1, 500},
            ScalingFunctions \rightarrow {"Log", "Log"}, PlotStyle \rightarrow {vMap[0.8], vMap[0.1]},
            PlotRange \rightarrow All, Frame \rightarrow True, FrameLabel \rightarrow {"t", "\delta_d/\delta_b"}]
Out[15]=
            10
                                                           100
```

In[16]:= tPerturbation = (t /. FindRoot[ratio == 1.1, {t, 1, 500}])
$$\times$$
 380 000 out[16]= 3.17432×10^{7}

Problem 2: Matter Growth with Dark Energy

```
ln[17] = ode = a^2 (1 - \Omega M0 + \Omega M0 / (a^3)) f''[a] + (3a/2)
           (2-2\Omega M0 + \Omega M0 / (a^3)) f'[a] - (3/2) (\Omega M0 / (a^3)) f[a] = 0
Out[17]=
           \frac{3 \; \Omega M0 \; f \left[\, a\,\right]}{2 \; a^3} \; + \; \frac{3}{2} \; a \; \left(2 \; - \; 2 \; \Omega M0 \; + \; \frac{\Omega M0}{a^3} \;\right) \; f' \left[\, a\,\right] \; + \; a^2 \; \left(1 \; - \; \Omega M0 \; + \; \frac{\Omega M0}{a^3} \;\right) \; f'' \left[\, a\,\right] \; = \; 0
 ln[18]:= params2 = {\Omega M0 \rightarrow 0.3};
         ai = 1/3600;
 ln[20]:= sol = NDSolveValue[{ode /. params2, f[ai] == 2*^-3, f'[ai] == f[ai] / ai}, f, {a, ai, 5}]
Out[20]=
         In[21]:= data = Table[{a, sol[a]}, {a, ai, 0.5, 0.001}];
         model = NonlinearModelFit[data, x0 + \alpha a, \{x0, \alpha\}, a]["BestFit"]
         0.034241 + 6.922 a
```

Problem 3: Scalar Field (Inflaton) in Expanding Universe