Project 4

Benedicte Allum Pedersen, Emil Helland Broll and Fredrik Oftedal Forr

Abstract

The Ising model in two dimensions has been used to simulate phase transitions in a magnetic system. We have used Monte Carlo and the Metropolis algorithm and our results showes that the number of Monte Carlo cycles wee need to reach equilibrium depends on the spin configuration in the start matrix and also on the temperature. We have estimated the critical temperature for the phase transitions to be 2.78 for a infinitely large lattice.

Contents

1	Introduction	3
	1.1 Phase Transitions	3
2	Method	4
	2.1 The Metropolis Algorithm	5
	2.2 Optimization	5
	2.3 Phase Transitions	6
3	Results	7
	3.1 Number of Monte Carlo Cycles and Equilibration Times	7
	3.2 Probability Distribution	12
	3.3 Phase Transitions	12
4	Discussion	15
5	Conclusion	16
6	Appendix	16
	6.1 Degenerated Energies and Magnetization	16
	6.2 The Partition Function	17
	6.3 Expectation Values for the Energy	
	6.4 Expectation Values for the Magnetization	
	6.5 Heat Capacity	18
	6.6 Susceptibility	19
7	Bibliography	19

1 Introduction

In this project we will study the Ising model in two dimensions. This model is used to simulate phase transitions in materials. A magnetic material will exhibit a phase transition from a magnetic phase to a phase with zero magnetization when the temperature of the system increases. The temperature where this phase transition occurs is called the critical temperature, T_C , where at temperatures above T_C , the average magnetization is zero. We will be studying the spins of electrons in a lattice, which is a binary system because each electron only can take two values, spin up or spin down, which we represent by +1 and -1.

The energy we get from the Ising model without an externally applied magnetic field is given by:

$$E = -J \sum_{\langle kl \rangle}^{N} S_k S_l \tag{1}$$

where $s_k, s_l = \pm 1$ and represents classical spin values of electrons at position k and position l in a grid. N is the total number of spins (electrons) and J is a coupling constant expressing the strength of the interactions between neighboring spins. $\langle kl \rangle$ indicates that we sum over the spins of the nearest neighbors.

In the model, we apply periodic boundary conditions, meaning that for electrons on the edge of the 2D grid, their nearest neighbors on the edge side is the electron by the opposite edge. In addition, we assume that we have a ferromagnetic ordering, meaning that J > 0.

1.1 Phase Transitions

The behavior of physical quantities like the mean magnetization, the heat capacity and the susceptibility can be characterized by a power law behavior when the temperature is near T_C :

$$(T)\rangle \sim (T - T_C)^{\beta},$$

 $C_v(T) \sim |T_C - T|^{\alpha},$
 $\chi(T) \sim |T_C - T|^{\gamma},$

where
$$\beta = 1/8, \alpha = 0$$
 and $\gamma = 7/4$.

The correlation length is another important physical quantity which can be described in the same way. The correlation length, ε , defines the length scale at which the overall properties of a material start to differ from its bulk properties (Jensen, M.). We expect ε to be of the order of the lattice spacing for $T >> T_C$. As a result of more interactions between the spins as T approaches

 T_C , the correlation length increases as we get closer to T_C . This means that the divergent behavior of ε near T_C is:

$$\varepsilon(T) \sim |T_C - T|^{-\nu}.\tag{2}$$

We will always be limited to a finite lattice and ε will be proportional with the size of the lattice. The behavior of a finite lattice can then be related to the behavior of a infinitely large lattice, so the critical temperature will scale as

$$T_C(L) - T_C(L = \infty) = aL^{-1/\nu} \tag{3}$$

where a is a constant. If we set $T = T_C$, the mean magnetization, the heat capacity and the susceptibility will be

$$\langle M(T) \rangle \sim (T - T_C)^{\beta} \to L^{-\beta/\nu},$$

 $C_v(T) \sim |T_C - T|^{\alpha} \to L^{-\alpha/\nu},$
 $\chi(T) \sim |T_C - T|^{\gamma} \to L^{-\gamma/\nu}.$

In this project, we will use a discretised model of the system and run numerical processes to calculate the physical quantities of the system in order to study the phase transitions in the material.

2 Method

In the Appendix, we have described how we calcuate the degenerated energies and magnetization of a 2x2-grid of spins. Table 1 shows the number of possible configurations that result in the same energy and magnetization, for a given number of spins pointing up. In total, 16 different spin configurations are possible for a 2x2 grid but only 3 different energies and 3 different magnetizations.

Table 1: Spinconfigurations grouped by their total energy and magnetization

# spins up	# configurations	E^2	\mathbf{M}
4	1	-8J	4
3	4	0	2
2	4	0	0
2	2	+8J	0
1	4	0	-2
0	1	-8J	-4

We have used the values in Table 1 to calculate the expectation values for the energy and the mean magnetization. These values have then been used to calculate the variance for the two physical quantities. The variance of the energy and the mean magnetization have respectively been used to calculate the heat capacity $c_v = \sigma_E^2/k_B T^2$ and the susceptibility $\chi = \sigma_M^2/k_B T$.

2.1 The Metropolis Algorithm

The Metropolis algorithm uses a propability distribution to obtain a sequence of random samples. In our case it is used to decide if the spins should flip or not. If we have a uniform propability distribution function(PDF), $\omega \in [0, 1]$ this is given by:

$$\omega = \frac{P_i}{P_j} = \frac{P_{new \ state}}{P_{previous \ state}} = \frac{e^{-\beta E_i/z}}{e^{-\beta E_j/z}}$$

$$\omega = e^{-\beta(E_i - E_j)} = e^{-\beta\Delta E}$$

where ΔE is the energy difference between the new and the previous state, E_i and E_j . $\beta = 1/k_BT$ and Z is the partition function described in the Appendix(Section 7.2). If $\omega = 1$ the system will stay at the same place, there will be no change in the spin configuration. If $\omega > 1$ the system will move to a larger propability of spin configurations and if $\omega < 1$ the system move to a smaller propability.

2.2 Optimization

In order to calculate the spin energies for different systems, we implement a discretised Ising model using the Monte Carlo method. This is a mostly random method, where we perform a loop over random positions in the system grid of spins and decide whether or not we should flip spins using the Metropolis algorithm.

The most effective way to run this simulation is to loop over one random position at a time and decide if we should flip that single spin or not. Because we can easily calculate the local energy contribution of that one spin, if we keep track of the current enery of the system before flipping, it is easy to calculate the new total energy, so we don't need to loop over the full grid each time we want to find the total energy.

In the Ising model, we know that when $N \to \infty$, the system should reach equilibrium. For our model, N is the number of Monte Carlo cycles, so we want to keep N as large as possible while still being able to run the calculations in a reasonable amount of time. This means we need to optimize our Monte Carlo loop as much as possible.

In theory, what we need to to within the MC loop is the following:

- Find the spin at a random position
- Calculate the current local energy contribution of that spin

- Calculate the new total energy of the system if the spin is flipped
- Use the Metropolis algorithm to decide if we should flip or not
- If we flip: calculate the new total energy and new total magnetization
- Update the expectation values with the (new) energy and total magnetization

Because we are looking at a 2D system, the energies actually get surprisingly easy to calculate. Because we know we only flip one spin at a time, the surrounding four spins (above, below, left and right) stay the same. The consequence of this is that the local energy contribution can only ever "flip" – if you flip the centre spin, the local energy will be $E_{new}(x,y) = -E_{old}(x,y)$). Since the local energy only ever depends on five spins, the possible local energies are also easily calculated. The only possible energies of a local part of the system are: E(x,y) = -4J, -2J, 0, 2J, 4J, making the only possible transition energies: $\Delta E(x,y) = 8J, 4J, 0, -4J, -8J$

We can use this knowledge to pre-calculate the probabilities of accepting a flip from one energy to another, using the Metropolis algorithm. Our requirement for flipping is:

$$r \le e^{-\beta \Delta E} \tag{4}$$

This, again, can be simplified to say that as long as $\Delta E \leq 0$, accept the flip, to avoid having to deal with the random number at all, saving some time.

These simplifications should make our code run quite fast, in theory, fast enough that we don't need to parallelize the Monte Carlo loop itself. However, if we want to run multiple Monte Carlo loops of millions or billions of cycles, we should definitely parallelize in order to utilze our full computing power.

In this project, we parallelize using Open MPI, dividing the different Monte Carlo loops among the processors we have available, and feed the results back to the main process for analysis.

2.3 Phase Transitions

We want to study phase transitions in lattices of different sizes, and in order to do so, we need to run simulations for the different lattice sizes with multiple temperatures, and compare these to extrapolate the critical temperatures of the systems.

In order to run effective calculations in this system, we do several tests before beginning to vary the temperatures and gather statistical data. One such test is to determine how many Monte Carlo cycles are needed to reach a good agreement with analytical values, in order to determine when the system has reached the most likely state, equilibrium.

From Equation 3 it is possible to esit mate the critical temperature, T_C for a infinitely large lattice. For this we need to know the value of the constant a, we find this by using T_C for some known lattice sizes, L_1 and L_2

$$T_C(L_1) - T_C(L = \infty) = aL_1^{-1/\nu}$$

 $T_C(L_2) - T_C(L = \infty) = aL_2^{-1/\nu}$

We subtract the two equations above from each other, that gives:

$$a = \frac{T_C(L_1) - T_C(L_2)}{L_1^{-1/\nu} - L_2^{-1/\nu}}. (5)$$

3 Results

In order to test our algorithm and its precision, we compare the analytical values of the expectation values of the energy and magnetization to our program's equivalent values. We ran the test by starting with a low number of MC cycles and increasing the number gradually until we got a good agreement. For a 2x2 lattice, we approximately need 10 000 Monte Carlo cycles in order to achieve a good agreement with the analytical values of the system. Using 10 000 MC cycles, we get $\langle E \rangle = -7.9864J$, which only differ 0.00247J from the analytical value. The values described in the sections above are represented in Table 2, where we have set $J/k_BT = \beta J = 1$. The calculations for the analytical values can be found in the Appendix.

Table 2: Analytical and numerical values for $T=1.0,\,L=2$ and 10^7 Monte Carlo cycles

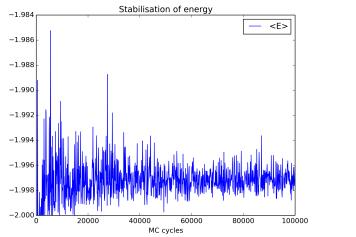
Property	Analytical per spin	Numerical per spin
$\langle E \rangle$	-1.9960	-1.9955
$\left\langle E^{2}\right angle$	15.968	15.9634
C_v	11.9840	11.9817
$\langle M(T) \rangle$	0	-0.0025167
$\langle M(T) \rangle$	0.9987	0.9986
$\langle M(T) ^2 \rangle$	3.9933	3.99323
χ	2.9959	2.9954

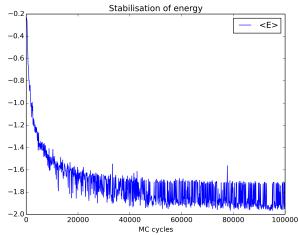
3.1 Number of Monte Carlo Cycles and Equilibration Times

In order to find a more precise measurement of the number of MC cycles needed to reach equilibrium, we run a test with a 20x20 lattice. The test involves logging the mean energy and mean magnetization for each Monte Carlo cycle, for both

T = 1.0 and T = 2.3. Figure 1 shows the results for two different initial spin matrices, one with all spins up and one random matrix for T=1.0.

For a matrix with a random spin configuration at T = 1.0, we reach steady state at around 200 000 Monte Carlo cycles, while for the matrix with all spins up we seemingly only need a few cycles.



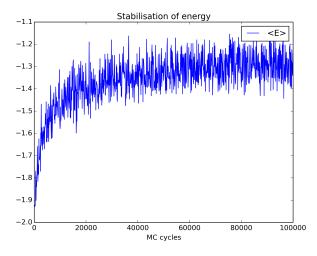


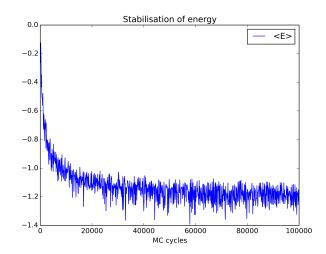
(a) Initial matrix with all spins up.

(b) Initial matrix with random spins.

Figure 1: Energy plotted against number of Monte Carlo cycles when T = 1.0.

When we increase the temperature, we see that we need substantially more Monte Carlo cycles to achive a steady state. Figure 2 shows that we would need around 1 million Monte Carlo cycles for both the random matrix and the matrix with all spins up. An estimate for the actual time it takes to reach equilibrium based on the number of Monte Carlo cycles can be made if we assume that one Monte Carlo cycle equals 1 second. The estimated equilibration times can be found in Table 3.



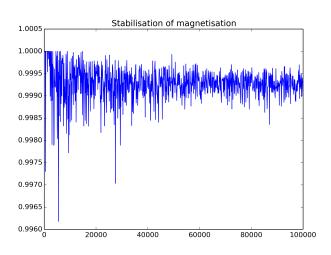


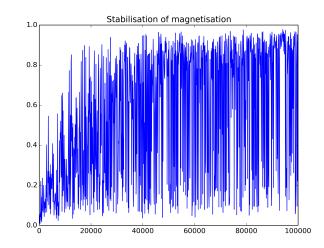
(a) Initial matrix with all spins up.

(b) Initial matrix with random spins.

Figure 2: Energy plotted against number of Monte Carlo cycles when T=2.4.

However when we do the same experimen for the magnetization for the system we get a different result. In Figure 3 we can see the way the magnetization varies over MC cycles for T=1. But there is not so easy to find a sertain location where the system reaches an equilibrium.





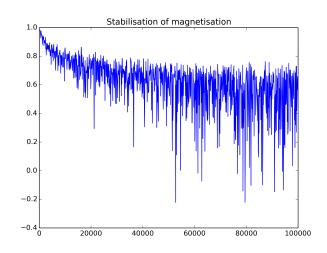
(a) Initial matrix with all spins up.

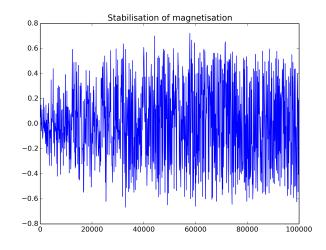
(b) Initial matrix with random spins.

Figure 3: Magnetization plotted against number of Monte Carlo cycles when T=1.

When the temperature increases it does not get easier to judge whether we

get to an equilibrium state or not. This is shown in Figure 4





(a) Initial matrix with all spins up.

(b) Initial matrix with random spins.

Figure 4: Magnetization plotted against number of Monte Carlo cycles when T=2.4.

The estimation of when the system reaches equilibrium for the different initial conditions is found in the following table, 3.

Table 3: Estimated equilibration times.

Temperature	Matrix	Equilibration time (sec)
1.0	Random spins	200 000
1.0	All spins up	≈ 0
2.4	Random spins	1 000 000
2.4	All spins up	1 000 000

Figure 5 and Figure 6 show plots of the total number of accepted configurations as function of the total number of Monte Carlo cycles for respectively an initial matrix with all spins up and an initial matrix with random spins. It is worth mentioning that these plots represent how many flips that are done on the whole lattice per monte carlo cycle. Therefore the values are sub zero.

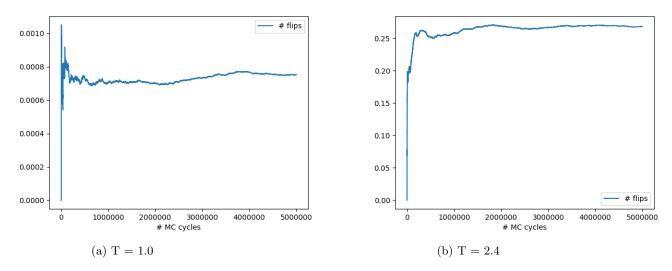


Figure 5: Total number of accepted spin configuration as a function of Monte Carlo cycles, for an initial matrix with all spins up.

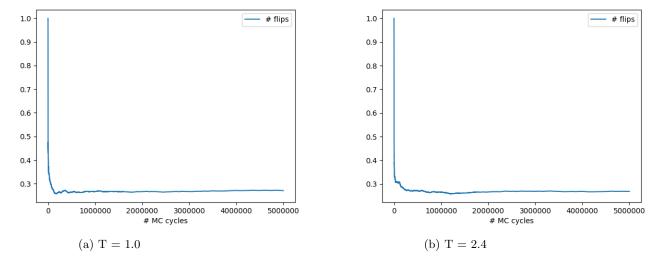


Figure 6: Total number of accepted spin configuration as a function of Monte Carlo cycles, for an initial matrix with all spins up.

3.2 Probability Distribution

The probability distribution of the energy for T=1.0 and T=2.4 is shown in figure 7

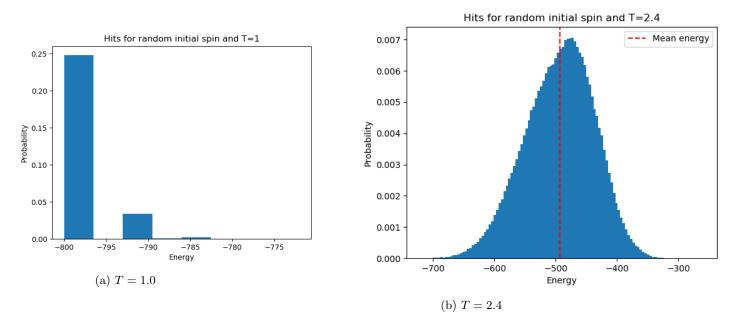


Figure 7: Probability distribution for counted energies.

3.3 Phase Transitions

During our tests runs with the Ising model with various lattice sizes and MC cycles, we gathered information about the time it takes to perform our calculations, which we used to estimate how many cycles we can perform for each of our required temperature steps within a reasonable computation duration.

Running our model for one million Monte Carlo cycles and 20 temperature steps on the same computer using one core takes 4.93 minutes, while using four cores takes 1.59 minutes. This shows that parallelization gives us a significant speed-up for our calculations.

We landed on having 12 temperature steps in the range T = [2.0, 2.3], because we had four cores available for computations, and the way we parallelizes meant each core then had 3 temperature steps to compute, with a total of 10^8 Monte Carlo cycles each. Running this calculation on a four core 2,5 GHz computer took 131.7 minutes to complete.

In Figure 8 and 9 we have plotted the mean energy, the mean magnetization, the heat capacity and the susceptibility as a function of T for different lattice sizes, L=40, L=60, L=80 and L=100. From the simulations and by using

Equation 3 we have estimated the critical temperature, T_C for the system. The critical temperature for the different lattice sizes are shown in Table 4

Table 4: Critical temperatures for different lattice sizes.

Lattice	Critical Temperature
L = 40	2.31
L = 60	2.35
L = 80	2.35
L = 100	2.44

We then use Equation 5 to find $T_C(L=\infty)$ from 3. We use the critical temperature for L = 80 and L = 100 this gives a=-36 and $T_C(L=\infty)=2.8$

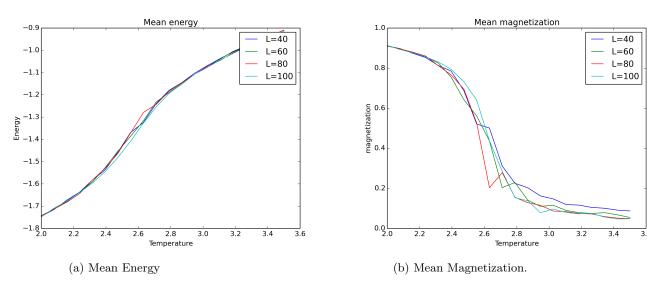


Figure 8: Energy and Magnetization as a function of temperature for different lattice sizes, \mathcal{L} .

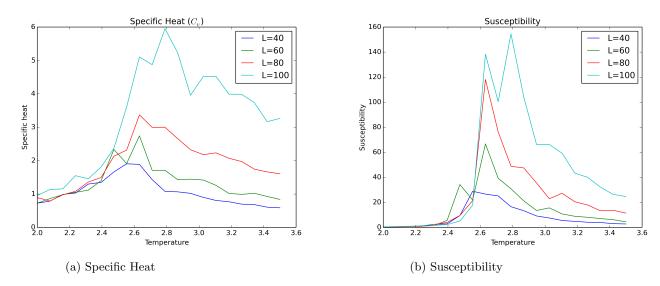


Figure 9: Specific heat and susceptibility as a function of temperature for different lattice sizes, \mathcal{L} .

4 Discussion

The numerical values for a 2x2 lattice is in good agreement with the analytical values. The mean magnetization varies a lot in different calculations, because the magnetization is heavily dependent on each single spin flip. While the energy dependence on each flip is the square value of the product of the neighboring spins, the magnetization depends on the sum of all spins. This might be why we experience different values for the mean magnetization in different calculations.

From Figure 1 and 3 we see that the energy and magnetization needs a lot less time to stabelize when all of the spins in the lattice have the same configuration, when T=1.0. When we increase the temperature, the number of Monte Carlo cycles we need to get good results also increases. For a higher temperature we can also see that the energy in the system increases.

Although the magnetization never seems to stabelize for the initial random matix. As mentioned above, the magnetization is more dependent on the spin configuration in the entire matrix and therefore it will not be as stable as the energy.

When looking at Figure 5 and Figure 6 we can easier see when the system stabelizes. In the beginning of the calculation we can se that there is some change in number of flips but it evens out fairly fast. When these plots even out we can consider that as a stable configuration since very few proposed flips is accepted. That means we have reached an equilibrium where the energy is as low as possible.

For this calculations we calculated the spins inside a single loop, a result of this is that the flip factor will not vary as much since the total number og flipped spins at the end of the loop does not differ that much. The flip facor could have been a lot more variating if we had calculated the flipping of spins in a different Monte Carlo loop.

From Figure 6 we can see that for a random inital spin matrix the flipfactor increases as the temperature increases. This indicates that the spins flip more often inside a single loop for higher temperatures.

For the probability distribution we have highlighted the mean energy when T=2.4. We coose not to do this for T=1 since this is clearly around -800. This is as expected since for low temperature none of the atoms will be recieving thermal energy from its suroundings. Therfore the expected value for the system should be the lowest possible value (with some exeptions). For T=2.4 we can se that we have a nice gaussian distribution around the mean energy value. When the temperature increases some of the atoms will recieve energy from the surroundings and it will be more random how much energy each of the configurations will have. With this in mind it is likley that we will have a ditribution looking like this. Because of some trouble getting the right value for the variance for the energy this is not highlighted in the plot. If there was good numbers for the variation it should take up around 60% of the total area of the

distribution.

The analytical value for the critical temperature is 2.269, we dont get the expected value and assume that this has somthing to do with the variance that mentioned above does not correspond good to our assumptions.

5 Conclusion

Our results shows that the inital matrix as well as temperature plays a role in the physical values for a magnetic system. We have experienced that a random spin matrix needs more time to reach the equilibrium state for the energy than a initial matrix where all the spins are aligned. From Table ?? we can tell that the estimated equilibration times is strongly dependent on both the temperature and the initial spin matrix. When the temperature increases the spins flip more often.

For a low temperature our expectations for the mean energy corresponds good with our results because the system will lay in the lowest and most stable energy state.

6 Appendix

6.1 Degenerated Energies and Magnetization

When calculating the degenerate energies for a 2x2 lattice, we start with the equation $E_i=-J\sum_{\langle kl\rangle}^2 s_k s_l$

The case for all spins up looks like this: $\uparrow \uparrow \uparrow$

The energy for this system is:

$$E_1 = -J \sum_{\langle kl \rangle}^2 s_k s_l$$

$$= -J((s_1 s_2 + s_1 s_3) + (s_2 s_1 + s_2 s_4) + (s_3 s_1 + s_3 s_4) + (s_4 s_3 + s_4 s_2))$$

$$= -J((1+1) + (1+1) + (1+1) + (1+1))$$

$$E_1 = -8J$$

In this calculation, we can see that the same interactions are counted twice. The reason for this is that the unit cell repeats itself to infinity in both x and y direction, which is our periodic boundary condition. Therefore the spin s_1 will interact with s_2 and s_3 inside of our unit cell, while s_2 and s_3 interact "outside" of the unit cell.

The magnetization is defined as:

$$M_i = \sum_{j=1}^{N} s_j$$

which sums over all spins for a given configuration i. For the same spin configuration as above, this gives us:

$$M_i = \sum_{j=1}^{4} (1+1+1+1) = 4$$

6.2 The Partition Function

When we known the values for all the degenerate energies we can calculate the value of the partition function.

$$z = \sum_{i=1}^{2^{+}} e^{-\beta E_i}$$
 For a 2x2-lattice, $n = 4$.
$$z = \sum_{i=1}^{2^{4}} e^{-\beta E_i}$$

$$z = e^{-\beta E_1} + e^{-\beta E_2} + \dots + e^{-\beta E_1 6}$$

$$z = e^{8\beta J} + 4e^{-\beta \cdot 0} + 2e^{-8\beta J} + 4e^{-\beta \cdot 0} + 4e^{-\beta \cdot 0} + e^{8\beta J}$$

$$z = 2e^{8\beta J} + 2e^{-8\beta J} + 12$$

6.3 Expectation Values for the Energy

The partition function is used to calculate the expectation value of the energy, $\langle E \rangle$, and the terms involving energies that are 0 are omitted.

$$\begin{split} \langle E \rangle &= \sum_{i}^{2^{n}} \frac{E_{i}e^{-\beta E_{i}}}{z} \\ &= \frac{-8Je^{8\beta J} + 2\left(8Je^{-8\beta J}\right) + \left(-8J\right)e^{8\beta J}}{2e^{8\beta J} + 2e^{-8\beta J} + 12} \\ &= \frac{16J\left(e^{-8\beta J} - e^{8\beta J}\right)}{2\left(e^{8\beta J} + e^{-8\beta J} + 6\right)} \\ &= 8J\frac{e^{-8\beta J} - e^{8\beta J}}{e^{8\beta J} + e^{-8\beta J} + 6} \\ &= 8J\frac{\sinh(8\beta J)}{\cosh(8\beta J) + 3} \end{split}$$

In order to calculate the variance, we need the expectation value of E^2 .

$$\begin{split} \left\langle E^2 \right\rangle &= \sum_{i}^{2^n} \frac{E_i^2 e^{-\beta E_i}}{z} \\ &= \frac{64J^2 e^{8\beta J} + 2\left(64J^2 e^{-8\beta J}\right) + 64J^2 e^{8\beta J}}{2e^{8\beta J} + 2e^{-8\beta J} + 12} \\ &= \frac{128J^2 e^{8\beta J} + 128J^2 e^{-8\beta J}}{2e^{8\beta J} + 2e^{-8\beta J} + 12} \\ &= 64J^2 \frac{e^{8\beta J} + e^{-8\beta J}}{e^{8\beta J} + e^{-8\beta J} + 6} \\ &= 64J^2 \frac{\cosh(8\beta J)}{\cosh(8\beta J) + 3} \end{split}$$

6.4 Expectation Values for the Magnetization

The absolute value of the mean magnetization is given by

$$\langle |M(T)| \rangle = \frac{\sum_{i}^{2^{4}} |M_{i}| e^{-\beta E_{i}}}{\gamma}$$

and can be calculated by using the values for the energy and the magnetization from Table 1:

$$\begin{split} \langle |M(T)| \rangle &= \frac{4e^{8\beta J} + 2 \cdot 4 + 0 \cdot 4 + 0 \cdot 2 + |-2| \cdot 4 + |-4|e^{8\beta J}}{z} \\ \langle |M(T)| \rangle &= \frac{8e^{8\beta J} + 12}{2e^{8\beta J} + 2e^{-8\beta J} + 12} = \frac{4e^{8\beta J} + 8}{e^{8\beta J} + e^{-8\beta J} + 6} \\ &= 8\frac{e^{8\beta J} + 1}{\cosh(8\beta J) + 3} \end{split}$$

We also have that

$$\begin{split} \left\langle |M(T)|^2 \right\rangle &= \frac{\sum_i^{2^4} |M_i^2| e^{-\beta E_i}}{z} \\ \left\langle |M(T)|^2 \right\rangle &= \frac{32 e^{8\beta J} + 32}{2 e^{8\beta J} + 2 e^{-8\beta J} + 12} = 16 \frac{e^{8\beta J} + 1}{e^{8\beta J} + e^{-8\beta J} + 6} \\ &= 16 \frac{e^{8\beta J} + 1}{\cosh(8\beta J) + 3} \end{split}$$

6.5 Heat Capacity

We use the expectation values of the energy to calculate the variance:

$$\sigma_E^2 = \langle E^2 \rangle - \langle E \rangle^2$$

The heat capacity is defined as the variance divided by the energy and temperature squared.

$$C_v = \frac{\sigma_E^2}{k_B T^2}$$

$$= \frac{64J^2}{k_b T^2} \left[\frac{\cosh(8\beta J)}{\cosh(8\beta J) + 3} \left(\frac{\sinh(8\beta J)}{\cosh(8\beta J) + 3} \right)^2 \right]$$

6.6 Susceptibility

The variance of the absolute value of the mean magnetization is given by:

$$\sigma_M^2 = \left<|M(T)^2|\right> - \left<|M(T)|\right>^2$$

We then calculate the susceptibility χ , which is given by $\chi = \sigma_M^2/k_BT$.

$$\chi = \frac{1}{k_B T} \left[16 \frac{e^{8\beta J} + 1}{\cosh(8\beta J) + 3} - \left(8 \frac{e^{8\beta J} + 1}{\cosh(8\beta J) + 3} \right)^2 \right]$$

7 Bibliography

Link to our GitHub repository.