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PROGRAM TB
 PROGRAM CALCULATES BRIGHTNESS TEMP. OF THE JOVIAN PLANETS
C AS A FUNCTION OF WAVELENGTH.
C
C PROGRAM ESPECIALLY FOR Neptune, based upon Sushil/romani's cloud models
C
  L AMMONIA
C f77 -0 -o sus susnep_v2003_orton.for
C ifort -132 susnep 2012 orton.for -o sus.exe
C ./sus.exe
                  to run it
  man ifort to look for compiling options, like -132
C
C
  This program has improved NH3 (joined Spilker and Joiner/Steffes
C This program has H2S de Boer (H2S2, slightly better compared to lab data than
  my older program, improved H2O, and PH3.
C Note: PH3 abundace is absent right now; need to add saturation curves, etc.
C
C
  AS 1, VOL2: JTP.DAT
                         CONTAINS TEMP PRESS DATA IN UPPER ATM.
C
  AS 2, VOL2: JPAR. DAT
                             PARTIAL PRESS DATA ARE CALCUL AND PUT IN THIS
  AS 3, VOL2: JTPZ.DAT
                              TEMP PRESS ALT ARE CALC AND PUT HERE
C
  AS 5, VOL2: JAMMON. DAT
                            INPUT
C AS 6, SPR:
C
  AS 7, VOL2: KAKAR.DAT
                           CONTAINS AMMONIA LINES AND CP FOR H2
  RUN WITH JTP.JOB
  M5 EQ 0 CALCULATE TABLES; NE 0 ASSUMES TABLES EXIST ALLREADY
      dimension Ttab(10), frequ(2428)
      dimension abeh2h2(2428,10),abnh2h2(2428,10)
      dimension abeh2he(2428,10),abnh2he(2428,10)
      dimension abeh2ch4(2428,10),abnh2ch4(2428,10)
      DIMENSION VKAK(16,16), DVKAK(16,16)
      DIMENSION TC(45), CPH2D(45)
      DIMENSION RR(18), CSZA(18), DCSZA(18), STR(19)
      DIMENSION ZT(7000), ZDELZ(7000), ZTPR(7000), ZPH2(7000), ZPH2S(7000),
                ZPHE(7000), ZPNH3(7000), ZPH2O(7000), ZPCH4(7000), ZSOLN(7000),
     +
                ZPCO(7000), ZPCO13(7000), ZPHCN(7000), ZPPH3(7000)
      DIMENSION XWAVE(30000), XTEMP(30000), XNADIR(30000), VSET(30000)
      DIMENSION TAU(7000), FINT(7000), GTAU(5,7000)
      DIMENSION CLH20(7000), CLNH4SH(7000), CLNH3(7000), CLH2S(7000)
      DIMENSION CLCH4(7000), CLAR(7000)
      real*4 F0(311),S0(311),EL(311),aph3,wgths(311)
      real*4 FP0(728),SP0(728),ELP(728)
      REAL*4 wgtS0(40), wgtFGB(40), wgtSB(40)
      real*4 t3,gm,vline,fjk,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,
             deltot, PCO, PCO13, PHCN
      COMMON/orton/Ttab, frequ, abeh2h2, abnh2h2, abeh2he, abnh2he,
                    abeh2ch4,abnh2ch4,nfreq,ntemp
      COMMON/ph3/FP0,SP0,ELP,wgtS0,wgtFGB,wgtSB
      COMMON/h2s/F0,S0,EL,wgths
      COMMON/CL/CLH20, CLNH4SH, CLNH3, CLH2S, CLCH4, CLAR
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
      COMMON/U2/C1, C2, C3, C4
      COMMON/U3/VKAK, DVKAK
      COMMON/U4/T1, T2, T3, T4, T5, T6, T7
      COMMON/U5/SP, SL, SLAM, SLH20, KB, KS
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
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COMMON/V3/QROT,QR1,QR2,S1,S2,S3,S4,P1,P2,P3,P4
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
      COMMON/V6/P5, P6, P7, P8
      COMMON/V7/PR01(20), PR02(20), STOR(20)
      COMMON/V9/TEMP, BETA, SQ, TAU1, TAU2
      COMMON/V10/TAU, FINT, GTAU
      COMMON/W1/FP, REB, RHO(20)
      COMMON/ABUN/AB1, AB2, AB3, AB4, AB5, AB6
      COMMON/ATM/TC, CPH2D
      COMMON/ZP/ZT, ZDELZ, ZTPR, ZPH2, ZPHE, ZPNH3, ZPH20, ZPCH4, ZPH2S,
                 ZPCO, ZPCO13, ZPHCN, ZPPH3
C
      VSET ARE WAVENUMBERS (CM-1).
      DATA CSZA/.9978,.9911,.9798,.9638,.9428,.9165,
        .8844,.8459,.8000,.750,.700,.650,.600,.500,
        .400,.300,.200,.100/
      DATA DCSZA/.10,.0668,.0668,.0667,.0667,.0667,
        .0667,.0666,.0640,.0571,.0492,.0429,.0530,
        .0583,.0440,.0316,.0205,.0126/
      DATA STR/0.0,.10,.1666,.2333,.30,.3667,.4334,
        .50, .5667, .6307, .6878, .737, .7799, .8329,
        .8912,.9352,.9668,.9873,1.0/
C
      OPEN(1,NAME='neptune.paulCO_cloud21_fletcher_best_dry')
      OPEN(10,NAME='nepcloud_C0.cloud21_fletcher_best_dry')
      Open(2,NAME='npar.dat')
      OPEN(3,NAME='ntpz.dat')
      OPEN(5,NAME='nammon.dat')
C
      OPEN(6,NAME='nepco_spectotal_allgases_dry.dat')
      OPEN(6,NAME='test.dat')
      Open(7,NAME='kakar.dat')
      open(11,NAME='orton_H2.tables',form='formatted')
C
      TO SUPPRESS THE LENGHTLY PRINT OUT OF INTERMEDIATE VALUES,
C
      SET NOPR AND NOPR2 TO 0 (NOT EQUAL TO 1).
C
      NOPR3 CONTROLS PRINT OUT OF MIXING RATIOS
\mathsf{C}
      DEN=INTEGRATED CLOUD DENSITY; ALT=ALTITUDE RANGE OF CLOUD
C
      for some reason I thought I calculated H2-CH4 I guess,
                 but this hasn't been done
      CZB=0.0
      DEN=7.2e6
      NOPR=0
      N0PR2=0
      NOPR3=0
      N0PR4=1
C
      READ IN ORTON'S H2 TABLES
C
      read (11,*) ntemp, Tmax, Tmin
      dltemp=(alog(Tmax)-alog(Tmin))/float(ntemp-1)
      do n=1,ntemp
         Ttab(n)=alog(Tmin)+(n-1)*dltemp
      enddo
C
      read (11,*) nfreq
      read (11,*) (frequ(i),i=1,nfreq)
      do n=1,nfreq
         read (11,*) (abeh2h2(n,i),i=1,ntemp)
      enddo
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do n=1,nfreq
         read (11,*) (abnh2h2(n,i),i=1,ntemp)
      enddo
      do n=1,nfreq
         read (11,*) (abeh2he(n,i),i=1,ntemp)
      enddo
      do n=1,nfreq
         read (11,*) (abnh2he(n,i),i=1,ntemp)
      enddo
      do n=1,nfreq
         read (11,*) (abeh2ch4(n,i),i=1,ntemp)
      do n=1,nfreq
         read (11,*) (abnh2ch4(n,i),i=1,ntemp)
      enddo
C
      READ(5,779) MII
      READ(5,779) KQQ
  779 FORMAT(I2)
      READ(5,773) SIZE
C
       WRITE(6,772) SIZE
      SIZE=SIZE/2.0
  772 FORMAT(2X, 'WATER DROPLET SIZE', 2X, E10.5,/)
  773 FORMAT(F10.4)
C
      EQUIL: IQ=2, FP=S1, REB=1.0
      NORMAL: IQ=0; FP=0.25; REB=0
C
C
      INTERM: IQ=1; REB=1.0.; FP=S1
      IQ=2: TEMP FOR EQUIL. CASE
      READ (5,778) IQ, REB, FP
      if(NOPR4.eq.1) WRITE(6,778) IQ,REB,FP
  778 FORMAT(I2,2F10.4)
C
      DO 1357 I=1,18
          RR(I) = ACOS(CSZA(I))
          RR(I)=SIN(RR(I))
          if(NOPR4.eq.1) WRITE(6,1345) CSZA(I),DCSZA(I),RR(I)
          FORMAT(3E15.5)
 1345
 1357 CONTINUE
      D0 2345 J=1,16
          D0 2345 K=1,J
              READ(7,3333) VKAK(J,K)
 2345 CONTINUE
      D0 2346 J=1,16
          D0 2346 K=1,J
              READ(7,2222) DVKAK(J,K)
              IF(NOPR.EQ.1) WRITE(6,4444) J,K,VKAK(J,K),DVKAK(J,K)
              VKAK(J,K)=1.0E-3*VKAK(J,K)
 2346 CONTINUE
 3333 FORMAT(F12.3)
 2222 FORMAT(F8.1)
      READ(7,103) (TC(I), CPH2D(I), I=1,36)
      READ(7,108) (TC(I), CPH2D(I), I=37,38)
  108 FORMAT(2(F3.0,F4.3,1X))
      READ(7,104) (TC(I), CPH2D(I), I=39,44)
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103 FORMAT(9(F3.0,F4.3,1X))
  104 FORMAT(6(F4.0,F4.3,1X))
  105 FORMAT(5(F5.0,1X,F7.3))
      for equilibrium H2
C
      IF(IQ.EQ.2) READ(7,103) (TC(I),CPH2D(I),I=1,24)
C
      WRITE(6,105) (TC(I), CPH2D(I), I=1,44)
C
      read in H2S lines THIS IS OLD
      OPEN(7,NAME='h2s file.list')
C
      read(7,7465) (F0(i),S0(i),EL(i),i=1,311)
c7465 format(F13.4,9x,F7.4,3x,F9.4,32x)
      OPEN(7, NAME='H2S_deboer.LIN')
      read(7,*) (F0(i),S0(i),EL(i),wgths(i),i=1,311)
      close(7)
      OPEN(7,NAME='ph3 file.list')
      read(7,*) (FP0(i),SP0(i),ELP(i),i=1,728)
      close(7)
      OPEN(7,NAME='ph3wgt.dat')
      read(7,*) (k,wgtS0(i),wgtFGB(i),wgtSB(i),i=1,40)
      close(7)
C
 4444 FORMAT(I5, I5, 2X, F15.4, 2X, F15.4)
      READ(5,1041) M5,NL
 1041 FORMAT(I1,1X,I4)
      if(NOPR4.eq.1) WRITE(6,1046) M5,NL
 1046 FORMAT(2X, 'PARAMETER M5 IS EQUAL TO',/,2I5)
 1049 FORMAT(F10.4)
      G=1130.0
      G=1140.0
      GE=G
      if(NOPR4.eq.1) WRITE(6,1050) G
 1050 FORMAT(2X, 'THE GRAVITY G IS ',/,F12.4)
      READ(5,1045) AB1,AB2,AB3,AB4,AB5,AB6
 1045 FORMAT(6E10.3)
      WRITE(6,1047) AB1,AB2,AB3,AB4,AB5,AB6
 1047 FORMAT(2X, 'ABUNDANCES H2, HE, NH3, H20, CH4, H2S ARE', /, 6E12.5)
      IF(M5.EQ.0) CALL ATMOS(NL,SOL)
      if(NOPR4.eq.1) WRITE(6,1051) NL
 1051 FORMAT(2X, 'THE PARAMETER NL AFTER ATMOS IS EQUAL TO', /, I6)
      NL2=NL-1
      REWIND 2
      REWIND 3
      do i=1,NL
          READ(3,23) ZT(I), ZDELZ(I), ZSOLN(I), CLH2O(I), CLNH4SH(I),
                      CLNH3(I),CLH2S(I),CLCH4(I),CLAR(I)
      enddo
  zpco etc are partial pressures
      do i=1,NL
          READ(2,13) ZTPR(I), ZPH2(I), ZPHE(I), ZPNH3(I), ZPH20(I),
                      ZPCH4(I), ZPH2S(I), ZPPH3(i), ZPCO(i), ZPCO13(i), ZPHCN(i)
      enddo
   23 FORMAT(10X,F10.4,8E13.8)
   13 FORMAT(11E13.8)
C
      READ(5,1060) NWN
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```
if(NOPR4.eq.1) WRITE(6,1060) NWN
1060
      FORMAT(I5)
C
      The following could be changed in the future to a read vset(1) and dvset
C
      DO 31 I=1,NWN
   31 READ(5,1061) VSET(I)
 1061 FORMAT(2F10.4)
      WILL PERFORM THE CALCULATIONS FOR THE WAVENUMBERS
C
      read(5,1061) vset(1), vsetlast
      dvset=(vset(1)-vsetlast)/nwn
      do n=2, nwn
          vset(n)=vset(n-1)-dvset
      enddo
\mathsf{C}
      DO 200 N=1,NWN
          VNU=VSET(N)
          FREQ=C*VNU*1.0E9
          VVNU=1.0/VNU
C
       WRITE(6,9) VNU, FREQ, VVNU
   9
          FORMAT(/,/,2X,'V(CM-1),FREQ(HZ) ',1P,3(2X,E10.3))
C
      INITIALIZE THE OPTICAL DEPTH TO ZERO.
          DH2=0.0
          DHE=0.0
          DNH3=0.0
          DH20=0.0
          DCH4=0.0
          DH2S=0.0
          dco=0.0
          dco13=0.0
          dhcn=0.0
          dpph3=0.0
          DOPTD=0.0
C
          SUM OVER THE VARIOUS PRESSURE LEVELS.
          DO 100 M=1,NL2
C
              SPECIFY THE PRESSSURES OF H2, HE, NH3, H20, AND CH4.
              T=ZT(M)
              DELZ=ZDELZ(M)
              PH2=ZPH2(M)
              PHE=ZPHE(M)
              PNH3=ZPNH3(M)
              PH20=ZPH20(M)
              PCH4=ZPCH4(M)
              PH2S=ZPH2S(M)
              PPH3=ZPPH3(M)
              pco=zpco(M)
              pco13=zpco13(M)
              phcn=zphcn(M)
              TPR=ZTPR(M)
              T1=273./T
              T2=300./T
              T3=T2**C2
              T4=T1**C1
              T5=T1**C2
              T6=1.0/(T**3.5)
              T7=4.8/T
              ARGON=tpr*6.0E-14
              wtmol=PH2*2.0158 + PHE*4.0026 + PH20*18.0153 + PCH4*16.04 + PNH3*
```

```
1
                                       + PH2S*34.08 + ARGON*39.06 + PCO*28.01 + PHCN*27.018+ PPH3*
                   33.9976
                              wtmol=wtmol*1.67e-24/tpr
C
                              CALL THE VARIOUS SUBROUTINES WHICH CALCULATE THE ABSORPTION
        COEFFICIENTS.
                              IF(IO.GE.1) CALL PART(T)
                              IF(IQ.GE.1) FP=S1
                              IF(TPR.gt.1.0.and.PNH3.NE.0.0.and.vnu.lt.1.0) CALL NH3TOM(VNU,T,ANH
                              IF(TPR.le.1.0.and.PNH3.NE.0.0.and.vnu.lt.1.0) CALL NH3(VNU,T,ANH3)
                              if(vnu.ge.1.0.and.PNH3.ne.0.0) CALL NH3J0IN(VNU,T,ANH3)
                              IF(PNH3.EQ.0.0) ANH3=0.0
                              CALL H20(VNU,T,AH20)
C
                              AH20=0.0
C
                              ANH3=0.0
C
C
                              H2H2 WITH MASSIE FORMALISM
C
                              H2H2JS WITH J0INER-STEFFES F0RMALISM; not very good
\mathsf{C}
                              H2H2OR with Orton's tables; this is best, but note the range
        inside which it works.
                              if(vnu.le.frequ(1)) CALL H2H2(VNU,T,A,B)
                              if(T.le.Tmin) CALL H2H2(VNU,T,A,B)
                              if(T.ge.Tmax) CALL H2H2(VNU,T,A,B)
                              if(T.gt.Tmin.and.T.lt.Tmax.and.vnu.gt.frequ(1))
                                       CALL H2H2OR(IQ, VNU, T, A, B, CZB)
                              CALL absbr_H2S2(ah2s,t,tpr,vnu)
C
                              Ah2S=0.0
                              CALL absbr_ph3(aph3,t,tpr,vnu,ZPPH3)
С
                              CALCULATE THE OPTICAL DEPTH TAU
C
C
                              ACO = absco(lco,fr0,freq,tpr,T,wtmol)
C
                              the line shape to be used is determined in absco routine.
                              call absx_co(aco,t,tpr,vnu)
C
                              statement above is including all CO lines, VVW only; this is
        useful for continuum measurements
                              AC0=AC0*pco/tpr
                              ****still an error in CO13. Also when simply running absx_co13 --
C
        assumed to be continuum
C
                              I need to test this. Note that fr0 is wrong for CO13 if the CO12
        parameters are used
C
                              AC013=0.0
C
                              ACO13 = absco13(lco,fr0,freq,tpr,T,wtmol)
C
                              call absx co13(aco13,t,tpr,vnu)
C
                              ACO13=ACO13*pco13/tpr
                              write(6,1345) aco,pco,tpr
C
                              A8=0.0
                              goto 9191
C
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        UDSCLOUDSCLOUDSCLOUDS
                              IF(MII.NE.1) GOTO 9191
                              H20 liquid and ice
C
                              IF(T.LT.273.0) THEN
                                       CALL DIEL(2,0,VVNU,T,E1,E2,D1,D2)
                                       X1=SIZE
                                       X2=X1*2.0*3.1415/VVNU
                                       CALL MIE(X2, D1, D2, QABS)
```

```
X=18*1.67E-24
                   Y=SIZE*SIZE*SIZE/8.0E-24
                   X=X*Y
                   XDEN=CLH20(M)/X
                   A8=A8+XDEN*QABS*3.1415*DELZ*size*size
               ENDIF
               IF(T.GE.273.0) THEN
                   CALL DIEL(1,0,VVNU,T,E1,E2,D1,D2)
C
                   X1 IS RADIUS PARTICLE
                   X1=SIZE
                   X2=X1*2.0*3.1415/VVNU
                   CALL MIE(X2, D1, D2, QABS)
                   TAKE MOLEC.SIZE FEW ANGSTROMS; OR ITS RADIUS 2.00E-8
C
C
                   TAKE H20-NH3 SOLUTION CLOUD WITH H20-ICE /WATER PROPERTIES
C
                   X=MASS GR/CM3;
                  X=18*1.67E-24
                   Y=SIZE*SIZE*SIZE/8.0E-24
C
                   WEIGHT OF ONE DROPLET IS X*Y
\mathsf{C}
                  NUMBER OF DROPLETS PER CM**3:
                  XDEN=ZSOLN(M)/X
                   A8=A8+XDEN*QABS*3.1415*DELZ*size*size
               ENDIF
               goto 9191
C
\mathsf{C}
              NH4SH CLOUD
              X1=SIZE
              X2=X1*2.0*3.1415/VVNU
              D1=1.7
              D2=.05
               CALL MIE(X2, D1, D2, QABS)
               X=51*1.67E-24
              Y=SIZE*SIZE*SIZE/8.0E-24
              X=X*Y
              XDEN=CLNH4SH(M)/X
               A8=A8+XDEN*QABS*3.1415*DELZ*size*size
               G0T0 9191
C
              NH3-ICE
              D1=1.3
              D2=.05
              X1=SIZE
              X2=X1*2.0*3.1415/VVNU
               CALL MIE(X2, D1, D2, QABS)
              X=17*1.67E-24
              Y=SIZE*SIZE*SIZE/8.0E-24
              X=X*Y
              XDEN=CLNH3(M)/X
               A8=A8+XDEN*QABS*3.1415*DELZ*size*size
C
              H2S-ICE
              D1=1.3
              D2 = .01
              X1=SIZE
              X2=X1*2.0*3.1415/VVNU
               CALL MIE(X2, D1, D2, QABS)
              X=34*1.67E-24
              Y=SIZE*SIZE*SIZE/8.0E-24
              X=X*Y
```

```
XDEN=CLH2S(M)/X
                                  A8=A8+XDEN*QABS*3.1415*DELZ*size*size
                                  G0T0 9191
CLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSCLOUDSC
          UDSCLOUDSCLOUDSCLOUDS
                                  CALCULATE THE OPTICAL DEPTH TAU
C
  9191
                                  CONTINUE
C
                                  FOR MASSIE FORMALISM USE FOLLOWING EXPRESSIONS
                                  DENH2=PH2/T
                                  DENHE=PHE/T
                                  DENCH4=PCH4/T
                                  A2=DENH2*DENH2*A*5.246E-3*DELZ/2.9979
                                  A3=DENH2*DENHE*B*5.246E-3*DELZ/2.9979
C
                                  For Orton tables (best to use): amagat is p(bar)*269.6/T with
          269.6=273.15/1.013
                                  if(T.gt.Tmin.and.T.lt.Tmax.and.vnu.gt.frequ(1))
                                               A2=DENH2*DENH2*273.*273.*delz*A/1.013/1.013
            +
                                  if(T.gt.Tmin.and.T.lt.Tmax.and.vnu.gt.fregu(1))
                                               A3=DENH2*DENHE*273.*273.*delz*B/1.013/1.013
                                  if(T.gt.Tmin.and.T.lt.Tmax.and.vnu.gt.frequ(1))
                                               A12=DENH2*DENCH4*273.*273.*DELZ*CZB/1.013/1.013
C
                                  AH2HE=(A2+A3+A12)
                                  WRITE(6,1003) T,PH2,A,B,CZB,A2,A3,A12
C
                                  write(6,1004) nfreq,ntemp
C
                                  FORMAT(8(1X,E10.3))
c 1003
c 1004
                                  format(2I10)
                                  Joiner\Steffes formalism:
C
C
                                  AH2HE=AH2H2*DELZ
C
                                  we absorb PH3 absorption into dh2s
                                  A4=ANH3*DELZ
                                  A5=AH20*DELZ
                                  A6=AH2S*DELZ
                                  A9=APH3*DELZ
                                  DH2=DH2+A2
                                  DHE=DHE+A3
                                  DCH4=DCH4+A12
                                  DNH3=DNH3+A4
                                  DH20=DH20+A5
                                  DH2S=DH2S+A6+A9
                                  DC0=DC0+AC0*delz
                                  DHCN=DHCN+AHCN*delz
                                  DC013=DC013+AC013*delz
                                  A7=A4+A5+AH2HE+A8+A6+A9+AC0*DELZ
C
                                  NH3, H20, HCN and 13CO do not make a difference.
                                  DOPTD=DOPTD+A7
C
                                  TAU IS THE OPTICAL DEPTH.
                                  TAU(M)=DOPTD
                                  GTAU(1.M)=DH2
                                  GTAU(2,M)=DHE
                                  GTAU(3,M)=DNH3
                                  GTAU(4,M)=DH20
                                  GTAU(5,M)=DH2S
                                  wt=a7/delz*exp(-tau(m))
C
                                  write weighting function
C
```

```
WRITE(6,262) T,TPR,DELZ,wt
C
            FORMAT (4E15.5)
 262
 100
         CONTINUE
         EVALUATE THE BRIGHTNESS TEMPERATURE OF THE CALCULATED INTENSITY.
C
         DBT=0.0
         DDR=0.0
         DO 2001 JJ=1,18
            CS=CSZA(JJ)
            DC=DCSZA(JJ)
            CALL BRIGHT(VNU, FREQ, NL, CS, BRI)
C
            CALL BRIGHTor(VNU, FREQ, NL, CS, BRI)
            STR(JJ)=BRI
            DBT=DBT+BRI*RR(JJ)*DC
            DDR=DDR+RR(JJ)*DC
2001
         CONTINUE
         DBT=DBT/DDR
         WAVEL=1./VNU
         GHZ=C*VNU
         WRITE(6,225) VNU, WAVEL, GHZ
C
225
         FORMAT(2X, 'V(CM-1)', 4X, 'WAVELENGTH(CM)', 3X, 'FREQ(GHZ)',
    1
               /,1P,E11.4,3X,E11.4,4X,E11.4)
         WRITE(6,226)
C
         FORMAT('
                       1)
226
         WRITE(6,235) wavel,STR(1),DBT
C
235
         FORMAT(' wavel, wave^-1, center disk, DISK-AVERAGED TB',/,1P,6X,3E11.4)
C
         XWAVE(N)=WAVEL
         XTEMP(N)=DBT
         XNADIR(N) = STR(1)
 200 CONTINUE
     DO 290 N=1, NWN
        x=1./xwave(N)
        GHZ=C*VSET(N)
        WRITE(6,5768) GHZ, XWAVE(N), x, XNADIR(N), XTEMP(N)
 290 CONTINUE
5768 FORMAT (5F20.8)
     ST0P
     END
SUBROUTINE BRIGHT(VNU, FREQ, NL, CMU, BRI)
     DIMENSION ZT(7000), ZDELZ(7000)
     DIMENSION TEM(7000), TOPR(7000)
     DIMENSION TAU(7000), FINT(7000), GTAU(5,7000)
     COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
     COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
     COMMON/V10/TAU, FINT, GTAU
     COMMON/ZP/ZT, ZDELZ
     BR IS THE INTENSITY AT THE TOP OF THE ATMOSPHERE.
C
C
     CMU IS THE COSINE OF THE SOLAR ZENITH ANGLE.
C
     FINT(I) IS THE FRACTION OF THE INTENSITY CONTRIBUTED BY LEVEL I.
C
     BR IS GIVEN BY EQUATION 90, PAGE 16 OF CHANDRASEKHAR'S
C
     RADIATIVE TRANSFER BOOK. BR IS DERIVED FROM THE INTENSITY
C
     CALCULATED FROM EQUATION 90. EQUATION 90 ASSUMES THAT
```

```
C
      THE SOURCE FUNCTION IS SIMPLY GIVEN BY THE PLANCK FUNCTION.
      BR=0.0
      YY=FRE0
      A1=2.0*6.626E-9*YY*YY*YY/(C*C)
      A2=6.626*1.0E-11*YY/1.3806
C
     WRITE(6,190)
C190 FORMAT(2X, 'LEVEL, H*FREQ/KT, EXP(-TAU), DTAU, B(T), ',/,
C
     1 2X, 'TAU, DINT, PARTIAL SUM OF I')
      L1=NL
      L2=L1-1
      BR=0.0
      D0 200 L=2,L2
          K=L1+1-L
          K2=K-1
          S1=ZT(K)
          S2=ZT(K2)
          TAVE=0.5*(S1+S2)
          A4=A2/TAVE
          A5=(TAU(K)+TAU(K2))/(2.0*CMU)
          A6=(TAU(K)-TAU(K2))/CMU
          PL=A1/(EXP(A4)-1.00)
          CALL EX(A5,A7)
          DINT=PL*A7*A6
          BR=BR+DINT
          FINT(K)=DINT
C
          if(cmu.gt.0.99) WRITE(6,205) K,Tave,A7,A6,PL,TAU(K),DINT,BR
 205
          FORMAT(2X, I4, 7(2X, E10.3))
200
     CONTINUE
      A8 IS THE BRIGHTNESS TEMPERATURE CORRESPONDING TO BR.
      A7=A1/BR
      A8=A2/(ALOG(1.00+A7))
     WRITE(6,888) FREQ, CMU, BR, A8
 888
     FORMAT(2X, 'FREQ, CMU, BR, T ARE ', /, 4E15.5)
      BRI=A8
      RETURN
SUBROUTINE BRIGHTor(VNU, FREQ, NL, CMU, BRI)
      real*8 BR,DINT,PL,A44,A55,A66,A77,A88,A99,A11,A22,YY
      DIMENSION ZT(7000), ZDELZ(7000)
      DIMENSION TEM(7000), TOPR(7000)
      DIMENSION TAU(7000), FINT(7000), GTAU(5,7000)
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
      COMMON/V10/TAU, FINT, GTAU
      COMMON/ZP/ZT, ZDELZ
      BR IS THE INTENSITY AT THE TOP OF THE ATMOSPHERE.
C
C
      CMU IS THE COSINE OF THE SOLAR ZENITH ANGLE.
C
      FINT(I) IS THE FRACTION OF THE INTENSITY CONTRIBUTED BY LEVEL I.
C
      BR IS GIVEN BY EQUATION 90, PAGE 16 OF CHANDRASEKHAR'S
C
      RADIATIVE TRANSFER BOOK. BR IS DERIVED FROM THE INTENSITY
C
      CALCULATED FROM EQUATION 90. EQUATION 90 ASSUMES THAT
      THE SOURCE FUNCTION IS SIMPLY GIVEN BY THE PLANCK FUNCTION.
      BR=0.0
      YY=FRE0
      A11=2.0*6.626E-9*YY*YY*YY/(C*C)
```

```
A22=6.626*1.0E-11*YY/1.3806
C
     WRITE(6,190)
C190 FORMAT(2X, 'LEVEL, H*FREQ/KT, EXP(-TAU), DTAU, B(T), ',/,
C
     1 2X, 'TAU, DINT, PARTIAL SUM OF I')
      L1=NL
      L2=L1-1
      BR=0.0
      K=NL
      S1=ZT(K)
      TAVE=S1
      A44=A22/TAVE
      PL=A11/(EXP(A44)-1.00)
      determine BR at bottom atm
C
C
      if(cmu.gt.0.99) WRITE(6,204) K, Tave, PL, BR, tau(K)
 204 format(I5,4E15.5)
     D0 200 L=2,L2
          K=L1+1-L
          K2 = K - 1
          S1=ZT(K)
          S2=ZT(K2)
          TAVE=0.5*(S1+S2)
          A44=A22/TAVE
C
          A5=TAU(K2)/CMU
          A55=(TAU(K)+TAU(K2))/(2.0*CMU)
          A66=(TAU(K)-TAU(K2))/CMU
          PL=A11/(EXP(A44)-1.00)
          CALL EX(A66,A77)
C
          DINT=PL*A7*A6
          BR=BR*A77+PL*(1.-A77)
          if(cmu.gt.0.99) WRITE(6,205) K, Tave, BR, A77, A66, PL, TAU(K), TAU(K2)
          FORMAT(2X, I4, 7(2X, E10.3))
 205
200 CONTINUE
      A8 IS THE BRIGHTNESS TEMPERATURE CORRESPONDING TO BR.
      A99=A11/BR
      A88=A22/(DLOG(A99+1.000))
     WRITE(6,888) FREQ, CMU, BR, A8
888 FORMAT(2X, 'FREQ, CMU, BR, T ARE ', /, 4E15.5)
      BRI=A88
      RETURN
      END
SUBROUTINE ATMOS(NN, SOL)
      DIMENSION XP(7000), XT(7000), XG(7000), H2S(7000), CH4(7000),
                co(7000),co13(7000),hcn(7000)
      DIMENSION TEMPA(7000), TEMPW(7000), TEMPP(7000), TEMPHY(7000)
      DIMENSION TEMPHE(7000), TEMPXG(7000), HELIUM(7000), XH(7000)
     DIMENSION TEMPT(7000), TMPCH4(7000), TMPH2S(7000), TSOLN(7000),
                tmpph3(7000),tmpco(7000),tmphcn(7000)
      DIMENSION ZDELZ(7000), TDELZ(7000), SOLN(7000), fosf(7000)
      DIMENSION AMMON(7000), WATER(7000), DRUK(7000), HYDR(7000), fosfien(7000)
      DIMENSION TC(45), CPH2D(45)
      DIMENSION TCLH20(7000), TCLNH4SH(7000), TCLNH3(7000), TCLH2S(7000)
      DIMENSION TCLCH4(7000), TCLAR(7000)
      DIMENSION CLH20(7000), CLNH4SH(7000), CLNH3(7000), CLH2S(7000)
      DIMENSION CLCH4(7000), CLAR(7000)
```

```
COMMON/CL/CLH20, CLNH4SH, CLNH3, CLH2S, CLCH4, CLAR
      COMMON/U2/C1, C2, C3, C4
      COMMON/U5/SP, SL, SLAM, SLH20, KB, KS
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/ABUN/AB1, AB2, AB3, AB4, AB5, AB6
      COMMON/ATM/TC, CPH2D
C
C
      AMMON, WATER MIXING RATIOS NH3, H20
C
      THEY ARE READ IN AT PRESSURE LEVELS OF
\mathsf{C}
      .01,.02,.03,.05,.07,.1,.2,.3,.5,.7,1,2,3,5,7,10 BARS.
      TFLAT=1.0E10
      KOUNT=0
      ZZZ=0.0
      READ(1,19) JMAX
  19 FORMAT(1x, I5)
      write(6,19) JMAX
      S0L=0.0
      ZD=0.0
      zdelz(1)=0.0
      DO 8079 II=1, JMAX
          READ(1,*) XH(II),XT(II),XP(II),HYDR(II),HELIUM(II),CH4(II),
                     AMMON(II), WATER(II), H2S(II), SOLN(II)
C
          read(1,*) s,XH(II),XP(II),XT(II)
\mathsf{C}
          Hydr(ii)=0.831
C
          helium(ii)=0.149
C
          h2s(ii)=0.0
C
          water(ii)=2.0E-4
C
          ammon(ii)=1.0E-4
\mathsf{C}
          ch4(ii)=2.0E-2
          co(ii)=1.0E-6
          if(xp(ii).ge.0.1585) co(ii)=0.0
C
          ***NOTE: check on pco in lats loop; do 50 M=1,----
          co13(ii)=co(ii)*1.0E-2
          hcn(ii)=0.0
          soln(ii)=0.0
 222
          FORMAT(F5.1,2X,F6.2,E13.3,2F7.3,5(E10.3))
 223
          FORMAT(I5, F7.2, E11.3, 2F7.3, 4(E10.3))
          IF(II.GT.1) ZDELZ(II) = (XH(II-1)-XH(II))*1.0E5*900./1130.
C
          IF(II.GT.1) ZDELZ(II)=abs(XH(II-1)-XH(II))*1.0E5
          SOLN(II) = SOLN(II) / 10.0
C
          IF(II.NE.1) SOL=SOL+SOLN(II)*ZDELZ(II)
          IF(SOLN(II).GT.0.0) ZD=ZD+ZDELZ(II)
C
          IF(XT(II).GT.400.0) ammon(ii)=1.94E-4
          fosfie=0.0
          fosfien(ii)=fosfie
C
          CALCULATE VAPOR PRESSURE FOR NH3
C
          Skip steps if Paul's files are used
      goto 8079
      if(ammon(ii).gt.3.0E-6) ammon(ii)=3.0E-6
      IF(XT(II).GT.195.0) GOTO 608
      PNH3=1.342E7*EXP(-3753.6/XT(II))
      PRNH3=PNH3/XP(II)
      IF(PRNH3.GE.AMMON(II)) PNH3=AMMON(II)*XP(II)
      AMMON(II)=PNH3/XP(II)
  608 CONTINUE
C
      calculate vapor pressure for water
```

```
C
      skip step if Paul's profiles are used
      goto 8079
      IF(XT(II).le.273.) goto 9910
      IF(XT(II).gt.273.) goto 9920
 9910 PH20=1.013*4.327E7*EXP(-6194./XT(II))
      PRH20=PH20/XP(II)
      GOTO 9921
 9920 PH20=1.013*1.5350E6*EXP(-5273./XT(II))
      PRH20=PH20/XP(II)
9921 IF(PRH20.GE.water(ii)) PH20=water(ii)*xp(ii)
      water(ii)=PH20/xp(ii)
      TT=XT(ii)*XT(ii)
C
      VPCH4S = 4.425070D0 - 453.92414D0/XT(ii) - 4055.6016D0/TT
C
         + 115352.19D0/XT(ii)**3 - 1165560.7D0/XT(ii)**4
C
       PCH4 = (10.0D0**VPCH4S) * 1.013
C
       PRCH4=PCH4/XP(II)
C
       if(PRCH4.ge.CH4(ii)) PCH4=ch4(II)*XP(II)
C
       if(xt(ii).lt.80.) CH4(ii)=0.0
C
      continue
   skip step if Paul's profiles are used
       goto 8069
C
       xtabun=ammon(jmax)+water(jmax)+H2S(jmax)+ch4(jmax)+co(jmax)
       fhe=helium(jmax)/hydr(jmax)
       fxtab=xtabun/hydr(jmax)
       DO 8089 II=1, JMAX
       xabun=ch4(ii)+ammon(II)+water(ii)+H2S(ii)+co(ii)
       fac=(xabun/xtabun)*fxtab
       hydr(ii)=1.0/(1.0+fhe+fac)
       helium(ii)=hydr(ii)*fhe
 8089
       continue
8069
       continue
C
        D0 i=1,jmax
            j=jmax+1-i
            read(10,*) x,y,z,soln(j),clh2o(j),clnh4sh(j),
C
C
                        clnh3(j),clh2s(j),clch4(j),clar(j)
            soln(j)=0.
            clh2o(i)=0.
            clnh4sh(j)=0
            clnh3(j)=0.
            clh2s(j)=0.
            clch4(j)=0.
            clar(j)=0.
        enddo
        ZDELZ(1)=ZDELZ(2)
        SOL=SOL/ZD
C
C
        WRITE(6,994) SOL,ZD
  994
        FORMAT('DENSITY AND ALTITUDE RANGE ARE', 2E12.4)
        TT=XT(1)
        PR=XP(1)
        JM=JMAX-1
C
        skip all te rest?
C
        goto 6904
        D0 601 II=1,JM
            X=ALOG(XT(II+1)/XT(II))
```

```
Y=ALOG(XP(II+1)/XP(II))
          XG(II)=X/Y
601
      CONTINUE
      write(6,*) xq(jm)
      TEMPT(1)=XT(1)
      TEMPP(1)=XP(1)
      TEMPHY(1)=HYDR(1)
      TEMPHE(1)=HELIUM(1)
      TMPCH4(1)=CH4(1)
      TEMPA(1) = AMMON(1)
      TEMPW(1)=WATER(1)
      TMPph3(1)=fosfien(1)
      TMPco(1)=co(1)
      tmphcn(1)=hcn(1)
      TMPH2S(1)=H2S(1)
      TEMPXG(1)=XG(1)
      TDELZ(1) = ZDELZ(1)
      TSOLN(1)=SOLN(1)
      TCLH20(1)=CLH20(1)
      TCLNH4SH(1)=CLNH4SH(1)
      TCLNH3(1)=CLNH3(1)
      TCLH2S(1)=CLH2S(1)
      TCLCH4(1)=CLCH4(1)
      TCLAR(1) = CLAR(1)
      DO 387 IJ=1,2
          KMAX=JMAX
          JJ=1
          write(6,*) jmax,jm
          D0 102 II=2,JMAX
              D0 101 KK=1,3
                  JJ=JJ+1
                  TEMPXG(JJ)=XG(II)
                  TEMPT(JJ)=KK*(XT(II)-XT(II-1))/3.0 + XT(II-1)
                  TEMPP(JJ)=KK*(XP(II)-XP(II-1))/3.0 + XP(II-1)
                  TEMPHY(JJ)=KK*(HYDR(II)-HYDR(II-1))/3.0 + HYDR(II-1)
                  TEMPHE(JJ)=KK*(HELIUM(II)-HELIUM(II-1))/3.0 + HELIUM(II-1)
                  TMPCH4(JJ)=KK*(CH4(II)-CH4(II-1))/3.0 + CH4(II-1)
                  TEMPA(JJ)=KK*(AMMON(II)-AMMON(II-1))/3.0 + AMMON(II-1)
                  TEMPW(JJ)=KK*(WATER(II)-WATER(II-1))/3.0 + WATER(II-1)
                  TMPH2S(JJ)=KK*(H2S(II)-H2S(II-1))/3.0 + H2S(II-1)
                  TMPco(JJ)=KK*(co(II)-co(II-1))/3.0 + co(II-1)
                  TMPHcn(JJ)=KK*(Hcn(II)-hcn(II-1))/3.0 + hcn(II-1)
                  TMPph3(JJ)=KK*(fosfien(II)-fosfien(II-1))/3.0 + fosfien(II-1)
                  TDELZ(JJ)=ZDELZ(II)/3.0
                  TSOLN(JJ)=kk*(SOLN(II)-soln(ii-1))/3.0 + soln(ii-1)
                  if(soln(ii).eq.0..and.soln(ii-1).gt.0.)
                      tsoln(jj)=(4-kk)*soln(ii-1)/3.0
                  TCLH2O(JJ)=kk*(CLH2O(II)-clh2o(ii-1))/3.0 + clh2o(ii-1)
                  if(clh2o(ii).eq.0..and.clh2o(ii-1).gt.0.)
                      tclh2o(jj)=(4-kk)*clh2o(ji-1)/3.0
                  TCLNH4SH(JJ)=kk*(CLNH4SH(II)-clnh4sh(ii-1))/3.0 + clnh4sh(ii-1)
                  if(clnh4sh(ii).eq.0..and.clnh4sh(ii-1).qt.0.)
                      tclnh4sh(jj)=(4-kk)*clnh4sh(ji-1)/3.0
   +
                  TCLNH3(JJ)=CLNH3(II)
                  TCLH2S(JJ)=CLH2S(II)
```

```
TCLCH4(JJ)=CLCH4(II)
                    TCLAR(JJ)=CLAR(II)
  101
                CONTINUE
                KMAX=KMAX+2
  376
                CONTINUE
  102
            CONTINUE
            JMAX=KMAX
            WRITE(6,19) JMAX
            D0 103 II=1, JMAX
                ZDELZ(II)=TDELZ(II)
                SOLN(II)=TSOLN(II)
                CLH20(II)=TCLH20(II)
                CLNH4SH(II)=TCLNH4SH(II)
                CLNH3(II)=TCLNH3(II)
                CLH2S(II)=TCLH2S(II)
                CLCH4(II)=TCLCH4(II)
                CLAR(II)=TCLAR(II)
                XT(II)=TEMPT(II)
                XP(II)=TEMPP(II)
                HYDR(II)=TEMPHY(II)
                HELIUM(II)=TEMPHE(II)
                CH4(II)=TMPCH4(II)
                AMMON(II)=TEMPA(II)
                WATER(II)=TEMPW(II)
                H2S(II)=TMPH2S(II)
                Hcn(II)=TMPHcn(II)
                co(II)=TMPco(II)
                fosfien(ii)=tmpph3(ii)
                XG(II)=TEMPXG(II)
  388
                CONTINUE
  103
            CONTINUE
        CONTINUE
  387
 if skipping extension of the adiabat, then (there is an error in the below,
    and it is better to use Paul's code)
6904
      continue
       KOUNT=0
CHEREIAMHEREIAM
       D0 50 M=1,JMAX
C
    DELP IS THE DIFFERENCE IN THE PRESSURE LEVELS OF THE
C
   MODEL ATMOSPHERE, IN LOG10 COORDINATES.
      P0=PR
      T0=TT
      IF(M.EQ.1) GOTO 50
      PR=XP(M)
      TT=XT(M)
      PCH4=CH4(M)*PR
      H20MR=WATER(M)
      AMR=AMMON(M)
      PHE=HELIUM(M)*PR
      PH2S=H2S(M)*PR
      PNH3=AMR*PR
      PH20=H20MR*PR
      PH2=HYDR(M)*PR
    PPh3=fosfien(m)*pr
        if(pr.ge.0.1585) co(m)=0.0
        pco=co(m)*pr
```

```
pco13=co13(m)*pr
        phcn=hcn(m)*pr
C
C
   COMPUTE DELZ, DIFFERENCE IN HEIGHT BETWEEN THE TWO
C
  PRESSURE LEVELS PO AND PR
      DELZ=ZDELZ(M)
      ZZZ=ZZZ+DELZ*1.0E-5
      TPR=PR
      WRITE(2,7) TPR, PH2, PHE, PNH3, PH20, PCH4, PH2S, PCO, PCO13, PHCN
      IF(TT.GT.TFLAT)
     +WRITE(3,8) TPR, TFLAT, DELZ, SOLN(M), CLH2O(M), CLNH4SH(M),
     + CLNH3(M), CLH2S(M), CLCH4(M), CLAR(M)
      IF(TT.LE.TFLAT)
     +WRITE(3,8) TPR,TT,DELZ,SOLN(M),CLH2O(M),CLNH4SH(M),
     + CLNH3(M), CLH2S(M), CLCH4(M), CLAR(M)
    7 FORMAT(10E13.8)
    8 FORMAT(E10.5,F10.4,8E13.8)
       KOUNT=KOUNT+1
 50
      CONTINUE
 9999 NN=KOUNT
      M5 = 0
 1049 FORMAT(F10.4)
  110 FORMAT(I1,1X,I4,I4)
      RETURN
      END
      SUBROUTINE MIE(X,A,B,QABS)
      COMPLEX ZM
      ZM=CMPLX(A,-B)
      ZM=4.0*X*((ZM*ZM-1.0)/(ZM*ZM+2.0))
      QABS=-AIMAG(ZM)
C
       A2=2.0*A*B
C
       A3 = A2 * A2
C
       A4=(A*A)-(B*B)+2.0
C
       A5=A4*A4
       QABS=4.0*X*((3.0*B)/(A3+A5))
      RETURN
      END
C
C
      SUBROUTINE EX(TX,S)
C
      SINCE THE COMPUTER DOES NOT LIKE TOO LARGE OR SMALL ARGUMENTS
C
      FOR THE EXPONENTIAL FUNCTION, WE SET THE VALUE TO ZERO IF
C
      NEED BE.
      IF (TX .GT. 75.0) S=0.0
      IF (TX .LT. 75.0) S=EXP(-TX)
      RETURN
      END
      BLOCK DATA
C
      CONSTANTS USED IN THE PROGRAM ARE SPECIFIED HERE.
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/U2/C1, C2, C3, C4
      PI, SPEED OF LIGHT, PLANCK'S CONST/TWO PI, TWO PI C,
C
C
      BOLTZMANN'S CONST, AW, 1 AMAGAT=2.69E19 CM-3.
C
      AW IS 2*PI*PI/3*HB*C
      DATA PI,C,HB,TPC,BOLC,AW,DEN/3.14159,2.99792458E1,1.054,
```

```
1 1.884,1.380622,2.082,2.687E19/
C
      BAR IS 1.0E6 DYNES/CM2.
      DATA BAR/1.0E6/
C
      THESE MOLECULAR CONSTANTS FOR NH3 ARE IN MHZ.
      C1,C2,C3,C4 ARE 13/3,2/3,1/3,30*760*SQRT(293)
      DATA C1,C2,C3,C4/4.333,0.666,0.333,3.902E+05/
C
      SUBROUTINE H20(VNU,T,AH20)
C
      SEE BERGE AND GULKIS PAPER IN GEHRELS JUPITER BOOK, PG 676, 1969.
C
      SEE PAGE 118 OF GOODMAN'S THESIS.
   Updated July 2003 with de Boer's thesis (p. 66), and refs therein
        REAL*4 AON, T, P, F, V
    real*4 gm,vline,fjk,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,PCO,PCO13,PHCN
    REAL*4 QLOG2, SHAPE2
    REAL*4 C, ZH2O, DV, VNU, gamma
        EXTERNAL QLOG2
        REAL*4 DF, DFC, DFD, EL(10), Q0(5), F0(10), FS, FV
        REAL*4 S0(10),S1(10),TCOR,X,X1,Y
    real*4 gh2(10),ghe(10),gh2o(10),xh2(10),xhe(10),xh2o(10)
  S0 stands for A in de Boer's table 3.5
C
         DATA Q0 /2.711, 2.5237, 2.2605, 1.8164,1.3778/ !LOG PF AT
    300,225,150,75,37.5K Values aren't right, I think
        DATA Q0 /2.2507, 2.0645, 1.804, 1.3649,0.9335/ !LOG PF AT
            300,225,150,75,37.5K These are right, I think
     DATA F0(1),S0(1),EL(1),gh2(1),ghe(1),gh2o(1),xh2(1),xhe(1),xh2o(1)
     & /22235.15, 1.0, 644.0, 2.395, 0.67, 10.67, 0.90, 0.515, 0.626/
        DATA F0(2),S0(2),EL(2),gh2(2),ghe(2),gh2o(2),xh2(2),xhe(2),xh2o(2)
       /183310.12, 41.9, 196.0, 2.400, 0.71, 11.64, 0.95, 0.490, 0.649/
        DATA F0(3),S0(3),EL(3),gh2(3),ghe(3),gh2o(3),xh2(3),xhe(3),xh2o(3)
     & /323000.0, 334.4, 1850.0, 2.395, 0.67, 9.59, 0.90, 0.515, 0.420 /
        DATA F0(4),S0(4),EL(4),gh2(4),ghe(4),gh2o(4),xh2(4),xhe(4),xh2o(4)
     & /325153.8, 115.7, 454.0, 2.395, 0.67, 11.99, 0.90, 0.490, 0.619/
        DATA F0(5),S0(5),EL(5),gh2(5),ghe(5),gh2o(5),xh2(5),xhe(5),xh2o(5)
       /380196.8, 651.8, 306.0, 2.390, 0.63, 12.42, 0.85, 0.540, 0.630/
        DATA F0(6),S0(6),EL(6),gh2(6),ghe(6),gh2o(6),xh2(6),xhe(6),xh2o(6)
     & /390000.0, 127.0, 2199.0, 2.395, 0.67, 9.16, 0.90, 0.515, 0.330/
        DATA F0(7), S0(7), EL(7), gh2(7), ghe(7), gh2o(7), xh2(7), xhe(7), xh2o(7)
     & /436000.0, 191.4, 1507.0, 2.395, 0.67, 6.32, 0.90, 0.515, 0.290/
        DATA F0(8), S0(8), EL(8), gh2(8), ghe(8), gh2(8), xh2(8), xh2(8), xh2(8)
     & /438000.0, 697.6, 1070.0, 2.395, 0.67, 8.34, 0.90, 0.515, 0.360 /
        DATA F0(9),S0(9),EL(9),gh2(9),ghe(9),gh2o(9),xh2(9),xhe(9),xh2o(9)
     & /442000.0, 590.2, 1507.0, 2.395, 0.67, 6.52, 0.90, 0.515, 0.332 /
        DATA F0(10),S0(10),EL(10),gh2(10),ghe(10),gh2o(10),xh2(10),xhe(10),xh2o
     & /448000.8, 973.1, 412.0, 2.395, 0.67, 11.57, 0.90, 0.515, 0.510/
       COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
       COMMON/U4/T1,T2,T3,T4,T5,T6,T7,T8,T9,T10,T11
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
      CONVERT VNU(CM-1) TO GHZ. C = 29.97
    NLINES=10
    C=2.99791E1
                                    !FREOU IN MHZ
    F=VNU*C*1.0E3
```

```
V=VNU*C
                         !FREO IN GHZ
    rho=1.0E12*18.0*PH20/(8.34143E7*T) !rho in g/m^3
    PA=0.81*PH2 + 0.35*PHe
    AON=0.0
        D0 N = 1, NLINES
    flc=f0(n)
                                      !line center in MHz
    VLINE=F0(N)/1.0E3
                             !LINE CENTER IN GHZ
        TCOR = SO(N) * exp(-EL(N)/T)
 THE GROSS LINE SHAPE
        X2=(-v*v+vline*vline)
        X2 = X2 * X2
    q1=qh2(N)*PH2*(300.0/T)**xh2(N)
    g2=ghe(N)*PHe*(300.0/T)**xhe(N)
    g3=gh2o(N)*PH20*(300.0/T)**xh2o(N)
    qamma=q1+q2+q3
        X3=gamma*gamma*4**v*v
        FJK=TCOR*gamma/(X2+X3)
 Divide FJK by 10 to test Steffes student's idea that DeBoer is off by 10.
    FJK=FJK/10.0
    AON=AON+FJK
   perhaps multiply by 4/pi
    enddo
C
    AON=AON*4.0/3.14159 !we think that it is all included in deBoer's stuff; It
    X1=(300.0/T)**2.1
    X1=1.08E-11*X1*PA*v*v*rho
    X2=(300.0/T)**2.5
    X2=2.0*v*v*rho*X2*2.3E-5*A0N
    write(6,*) AON, X1, X2
C
    A0N=X1+X2
C
      H20 ABSORPTION COEFFICIENT IN CM-1;
C
       AH20=A0N
C
      RETURN
      SUBROUTINE H20old(VNU,T,AH20)
C This is the old H2O routine
C
      SEE BERGE AND GULKIS PAPER IN GEHRELS JUPITER BOOK, PG 676, 1969.
C
      SEE PAGE 118 OF GOODMAN'S THESIS.
       COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
       COMMON/U4/T1,T2,T3,T4,T5,T6,T7,T8,T9,T10,T11
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
C
      CONVERT VNU(CM-1) TO GHZ. C = 29.97
      V=VNU*C
C
C
      EVALUATE EOUATION 29.
      A2=V/29.97
      A3=(A2-0.74)*(A2-0.74)
      A4=(A2+0.74)*(A2+0.74)
      A5=9.88E-2*T5*((0.81*PH2)+(0.35*PHE))
      A5=A5/1.013
      A6=A5*A5
```

```
A7=(A5/(A3+A6))+(A5/(A4+A6))
C
      H20 ABSORPTION COEFFICIENT IN CM-1
      AH20=PH20*T4*V*V*((9.07E-9*A7)+(1.45E-7*A5))
C
   CORRECTION TERM B(T)/B(293) SEE GOODMAN, P. 117
      AH20=AH20*1.179
      AH20=AH20/1.013
C
      IF (NOPR .EQ. 1) WRITE(6,10) V,A2,A3,A4,A5,A6,A7,AH20
  10 FORMAT(2X, 'H20 V(GHZ), A2, A3, A4, A5, A6, A7, AH20', /,
     1 2X,1P,8(2X,E10.3),/)
      RETURN
      END
C
C
C
      SUBROUTINE NH3TOM(VNU,T,ANH3)
C
      CALCULATE THE AMMONIA ABSORPTION COEFFICIENT (CM-1)
C
      USED BY GULKIS AND POYNTER (1972), AND GIVEN IN BERGE
      AND GULKIS, GEHRELS JUPITER BOOK, PAGE 675.
      DIMENSION VKAK(16,16), DVKAK(16,16)
       COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
      COMMON/U2/C1,C2,C3,C4
      COMMON/U3/VKAK, DVKAK
      COMMON/U4/T1, T2, T3, T4, T5, T6, T7, T8, T9, T10, T11
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
    VKAK(J,K)=CENTER FREQU FOR THE (J,K) TRANSITION ACCORDING TO THE
C
  TABULATED VALUES OF POYNTER AND GULKIS (1975).
C
C
  DVKAK(J,K) ARE THE CORRESPONDING SELF BROADENED LINE WIDTHS
   AS TABULATED BY THEM (RESP. GHZ, MHZ/TORR)
C
      CHANGE VNU(CM-1) TO GHZ.
      V=(C*VNU)
   CHANGE PRESSURES FROM BARS TO ATM, AND BACK AT END OF SUBROUTINE; MARCH89
    PH2=PH2/1.013
    PHE=PHE/1.013
    PH20=PH20/1.013
    PNH3=PNH3/1.013
    PCH4=PCH4/1.013
    TPPR=PH2+PHE+PNH3+PH20+PCH4
C
      CORRECTION FRACTOR IS GIVEN ON PAGE 676 OF THE BERGE AND GULKIS
      PAPER, GEHRELS JUPITER BOOK.
      CORBG=1.0075+((0.0308+(0.0552*(PH2/T)))*(PH2/T))
    CORR=-0.33664+T/110.4-T*T/70600.
    IF(T.GT.320.) CORR=1.1115
    IF(T_LT_180_L) then
    x=(180.*180./70600.) - 0.337
    y=(1./110.4) - (180./35300.)
    corr=x + y*T
    endif
C
    CORR=CORR*CORBG
    IF(TPPR.LE.1.0) CORR=1.0
C
      SUM=0.0
C
      CALCULATE THE INVERSION LINE WAVENUMBERS
```

```
C
      FOR NH3, FROM THE DATA IN POYNTER AND KAKAR
C
      AP J SUPPL SERIES, VOL 29, P87, 1975.
      J AND K ARE QUANTUM NUMBERS.
C
C
C
      NLINE ARE THE NUMBER OF LINES TO BE CALCULATED.
      NLINE=16
C
      DO 10 J=1, NLINE
      L1=J*(J+1)
      L3=0
      D0 20 K=1,J
      L2=K*K
      L3=L3+1
C
      SB IS THE SELF BRODENED LINE WIDTH OF NH3 IN MHZ/TORR.
C
      C4 IS GIVEN IN THE BLOCK DATA STATEMENT.
\mathsf{C}
      SK IS 1.5 OR 3, DEPENDING UPON THE K VALUE.
      SK=1.5
      IF (L3 .NE. 3) GO TO 21
      SK=3.0
      L3=0
  21
     VLINE=VKAK(J,K)
       SB=DVKAK(J,K)
      IF (NOPR .EQ. 1) WRITE(6,22) J,K,SK,VLINE,SB
     FORMAT(2X, 'J, K, SK, VLINE(GHZ), SB(MHZ/TORR)',/,
     1 2X,I3,2X,I3,2X,F4.1,2X,1P,2(2X,E10.3))
C
      CALCULATE A(J,K).
      CALL SAJTOM(T, VLINE, J, L1, L2, SK, SB, GM, GH2, AJK)
      IF(J.E0.1.AND.K.E0.1) ZZV=2.0*AJK*EXP(23.3/T)/3.0
\mathsf{C}
      FJK IS THE FREQUENCY DEPENDENT LINE SHAPE OF BEN REUVEN.
      CALL SFJTOM(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
      SB=AJK*FJK
      SUM=SUM+(CORR*AJK*FJK)
  no correction term in Joiner/Steffes;
      SUM=SUM+(AJK*FJK)
C
  20 CONTINUE
  10 CONTINUE
  CALCULATE ROTATION LINE J=0; K=0
  VLINE=5.73E5 MHZ; CHOSE DVLINE=14. MHZ
C
      L1=2
      L2 = 1
      VLINE=5.73E2
      SB=14.0
      GM=(2.318*T3*PH2)+(0.79*T3*PHE)+(0.75*T2*SB*PNH3)
      GM=GM
      CALL SFJK(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
      ZZV1=ZZV*7.0E-2*T*(1.0-EXP(-28.6/T))
      ZZV2=ZZV*5.25E-2*T*EXP(-23.3/T)*(1.0-EXP(-57.2/T))
      ZZV3=ZZV*7.50E-2*T*EXP(-28.6/T)*(1.0-EXP(-57.2/T))
      SUM=SUM+(ZZV1*FJK)
      L1=2
      L2=1
      VLINE=1169.0
      CALL SFJK(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
```

```
SUM=SUM+(ZZV2*FJK)+(ZZV3*FJK)
C
C
C
      THE AMMONIA ABSORPTION COEFFICIENT IS IN CM-1.
        ANH3=sum
      IF (NOPR .EQ. 1) WRITE(6,30) CORR,SUM,ANH3
      FORMAT(2X, 'CORR, SUM, ANH3', 2X, 1P, 3(2X, E10.3),/)
C
    PH2=PH2*1.013
    PHE=PHE*1.013
    PH20=PH20*1.013
    PNH3=PNH3*1.013
    PCH4=PCH4*1.013
      RETURN
      END
C
      SUBROUTINE NH3(VNU,T,ANH3)
C
      CALCULATE THE AMMONIA ABSORPTION COEFFICIENT (CM-1)
C
      USED BY GULKIS AND POYNTER (1972), AND GIVEN IN BERGE
\mathsf{C}
      AND GULKIS, GEHRELS JUPITER BOOK, PAGE 675.
      DIMENSION VKAK(16,16), DVKAK(16,16)
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
      COMMON/U2/C1, C2, C3, C4
      COMMON/U3/VKAK, DVKAK
      COMMON/U4/T1,T2,T3,T4,T5,T6,T7,T8,T9,T10,T11
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
C
    VKAK(J,K)=CENTER FREQU FOR THE (J,K) TRANSITION ACCORDING TO THE
C
   TABULATED VALUES OF POYNTER AND GULKIS (1975).
  DVKAK(J,K) ARE THE CORRESPONDING SELF BROADENED LINE WIDTHS
   AS TABULATED BY THEM (RESP. GHZ, MHZ/TORR)
C
      CHANGE VNU(CM-1) TO GHZ.
      V=(C*VNU)
C
   CHANGE PRESSURES FROM BARS TO ATM, AND BACK AT END OF SUBROUTINE; MARCH89
    PH2=PH2/1.013
    PHE=PHE/1.013
    PH20=PH20/1.013
    PNH3=PNH3/1.013
    PCH4=PCH4/1.013
C
      CORRECTION FRACTOR IS GIVEN ON PAGE 676 OF THE BERGE AND GULKIS
C
      PAPER, GEHRELS JUPITER BOOK.
C
      CORR=1.0075+((0.0308+(0.0552*(PH2/T)))*(PH2/T))
C
       CORR=SQRT(PH2/T)
C
       CORR=1.0+(0.155*CORR+0.52*SQRT(CORR))*SQRT(CORR)
      IF (NOPR .EQ. 1) WRITE(6,5) V, CORR
      FORMAT(2X, 'NH3 V(GHZ), CORR', 2X, 1P, 2(2X, E10.3))
      IF (NOPR .EQ. 1) WRITE(6,6) C1,C2,C3,C4
      FORMAT(2X, 'CONSTANTS C1, C2, C3, C4', 2X, 1P, 4(2X, E10.3))
C
      SUM=0.0
C
      CALCULATE THE INVERSION LINE WAVENUMBERS
      FOR NH3, FROM THE DATA IN POYNTER AND KAKAR
C
C
      AP J SUPPL SERIES, VOL 29, P87, 1975.
C
      J AND K ARE OUANTUM NUMBERS.
```

```
C
C
      NLINE ARE THE NUMBER OF LINES TO BE CALCULATED.
      NLINE=16
C
      D0 10 J=1,NLINE
      L1=J*(J+1)
      L3=0
      D0 20 K=1,J
      L2=K*K
      L3=L3+1
C
      SB IS THE SELF BRODENED LINE WIDTH OF NH3 IN MHZ/TORR.
C
      C4 IS GIVEN IN THE BLOCK DATA STATEMENT.
\mathsf{C}
      SK IS 1.5 OR 3, DEPENDING UPON THE K VALUE.
      SK=1.5
      IF (L3 .NE. 3) GO TO 21
      SK=3.0
      L3=0
  21 VLINE=VKAK(J,K)
       SB=DVKAK(J,K)
      IF (NOPR .EQ. 1) WRITE(6,22) J,K,SK,VLINE,SB
     FORMAT(2X, 'J, K, SK, VLINE(GHZ), SB(MHZ/TORR)',/,
     1 2X,I3,2X,I3,2X,F4.1,2X,1P,2(2X,E10.3))
C
      CALCULATE A(J,K).
      CALL SAJK(T, VLINE, J, L1, L2, SK, SB, GM, GH2, AJK)
      IF(J.E0.1.AND.K.E0.1) ZZV=2.0*AJK*EXP(23.3/T)/3.0
\mathsf{C}
      FJK IS THE FREQUENCY DEPENDENT LINE SHAPE OF BEN REUVEN.
      CALL SFJK(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
      SB=AJK*FJK
      SUM=SUM+(CORR*AJK*FJK)
      IF (NOPR .EQ. 1) WRITE(6,23) AJK,FJK,SB
     FORMAT(2X, 'AJK, FJK, (AJK*FJK)', 2X, 1P, 3(2X, E10.3))
  20
      CONTINUE
  10 CONTINUE
C
C
  CALCULATE ROTATION LINE J=0;K=0
  VLINE=5.73E5 MHZ; CHOSE DVLINE=14. MHZ
      L1=2
      L2 = 1
      VLINE=5.73E2
      SB=14.0
      GM=(2.318*T3*PH2)+(0.79*T3*PHE)+(0.75*T2*SB*PNH3)
      CALL SFJK(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
      ZZV1=ZZV*7.0E-2*T*(1.0-EXP(-28.6/T))
      ZZV2=ZZV*5.25E-2*T*EXP(-23.3/T)*(1.0-EXP(-57.2/T))
      ZZV3=ZZV*7.50E-2*T*EXP(-28.6/T)*(1.0-EXP(-57.2/T))
      SUM=SUM+(ZZV1*FJK)
      L1=2
      L2 = 1
      VLINE=1169.0
      SB=23.8
      CALL SFJK(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
      SUM=SUM+(ZZV2*FJK)+(ZZV3*FJK)
C
C
```

```
C
      THE AMMONIA ABSORPTION COEFFICIENT IS IN CM-1.
      ANH3=SUM
      IF (NOPR .EQ. 1) WRITE(6,30) CORR, SUM, ANH3
      FORMAT(2X, 'CORR, SUM, ANH3', 2X, 1P, 3(2X, E10.3),/)
C
    PH2=PH2*1.013
    PHE=PHE*1.013
    PH20=PH20*1.013
    PNH3=PNH3*1.013
    PCH4=PCH4*1.013
      RETURN
      END
C
C
C
      SUBROUTINE SAJK(T, VLINE, J, L1, L2, SK, SB, GM, GH2, AJK)
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
      COMMON/U4/T1,T2,T3,T4,T5,T6,T7,T8,T9,T10,T11
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
      GM IS GAMMA(J,K) IN GHZ.
C
    2.318 IPV 3.318
      GM=(2.318*T3*PH2)+(0.79*T3*PHE)+(0.75*T2*SB*PNH3)
C
       GM=(4.8*PH2*(T2**0.50))+(0.79*T3*PHE)+(0.75*T2*SB*PNH3)
C
       GM=17.4*T2*PNH3+1.73*T8*PH2+0.40*T8*PHE
      A1=(2.*J+1.)*FLOAT(L2)/FLOAT(L1)
      A5=((2.98*L1)-(1.09*L2))*T7
C
      AJK IS A(J,K), SEE EQN 24 OF BERGE AND GULKIS PAPER, GEHRELS
      JUPITER BOOK, PAGE 675.
      CALL EX(A5,A10)
      AJK=1.23E3*A1*SK*PNH3*T6*A10
      IF (NOPR .EQ. 1) WRITE(6,10) GM,A1,A5,AJK
  10 FORMAT(2X, 'GM, A1, A5, AJK', 2X, 1P, 4(2X, E10.3))
      RETURN
      END
      SUBROUTINE SAJTOM(T, VLINE, J, L1, L2, SK, SB, GM, GH2, AJK)
      COMMON/V1/A1,A2,A3,A4,A5,A6,A7,A8,A9,A10
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
      COMMON/U4/T1,T2,T3,T4,T5,T6,T7,T8,T9,T10,T11
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
      GM IS GAMMA(J,K) IN GHZ.
    PTOT=PH2+PHE+PNH3+PH20+PCH4
   spilker uses equations below
    GNH3=0.74
    GHE=0.46+T/3000.
    R=8.79*EXP(-T/83.)
    X=2.122*EXP(-T/116.8)
    Y = EXP(9.024 - T/20.3)
    Y=(Y-0.9918+PH2)**R
    GH2=2.34*(1.0-X/Y)
   Joanna Joiner uses:
         gh2=1.69
С
         ghe=0.75
С
C
         gnh3=0.60
      GM=(GH2*T3*PH2)+(GHE*T3*PHE)+(GNH3*T2*SB*PNH3)
      A1=(2.*J+1.)*FLOAT(L2)/FLOAT(L1)
```

```
A5=((2.98*L1)-(1.09*L2))*T7
C
      AJK IS A(J,K), SEE EQN 24 OF BERGE AND GULKIS PAPER, GEHRELS
C
      JUPITER BOOK, PAGE 675.
      CALL EX(A5,A10)
      AJK=1.23E3*A1*SK*PNH3*T6*A10
      IF (NOPR .EQ. 1) WRITE(6,10) GM,A1,A5,AJK
      FORMAT(2X, 'GM, A1, A5, AJK', 2X, 1P, 4(2X, E10.3))
      RETURN
      END
C
C
      SUBROUTINE SFJK(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
    real*4 t3,gm,vline,fjk,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,
     + deltot, PCO, PCO13, PHCN
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
       COMMON/U4/T1,T2,T3,T4,T5,T6,T7,T8,T9,T10,T11
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
      FJK IS THE BEN REUVEN LINE SHAPE, SEE EQN 25 OF BERGE AND
C
C
      GULKIS ARTICLE, PAGE 675 OF GEHRELS JUPITER BOOK.
C
      X=PH2+PHE+PH20+PNH3+PCH4
C
   IF X.LT.1 ATM VAN VLECK-WEISKOPF LINE SHAPE CAN BE
  USED; HOWEVER, BEN REUVEN IS STILL BETTER OUT IN THE
  WINGS OF THE AMMONIA LINES.
C
      A2 IS THE COUPLING FACTOR ZETA(J,K) IN GHZ.
C
    G0T0 500
C
 for Gross lineshape, goto 500 (first VVW, then gross
      IF(VLINE.GT.100.0) GOTO 500
      A2=(1.92*PH2*T3)+(0.49*T2*PNH3*SB)+
     1 (0.3*T3*PHE)
C
       A2=(0.49*T2*PNH3*SB)+(0.3*T3*PHE)+
C
      +(4.80*(T2**0.5)-0.6*(T2**2.0))*PH2
C
       A2=(1.73*T8-0.87*T10)*PH2+
C
      + (0.40*T8-0.21*T9)*PHE+
C
      + (17.4*T2-6.0*T11)*PNH3
C
       A2=0.655*GM
C
      A4 IS THE PRESSURE SHIFT DEL (GHZ)
    A4=-0.45*PNH3
C
      SQUARES OF VARIOUS TERMS.
      A5=(VLINE+A4)*(VLINE+A4)
       A5=VLINE*VLINE
C
      A6=GM*GM
      A7=A2*A2
      A8=V*V
C
      NUMERATOR
      A9=((GM-A2)*A8)+((GM+A2)*(A5+A6-A7))
      A10=A8-A5-A6+A7
C
      DENOMINATOR
      A10=(A10*A10)+(4.0*A8*A6)
C
      THE BEN REUVEN LINE SHAPE.
      FJK=2.0*A8*(A9/A10)
      IF (NOPR .EQ. 1) WRITE(6,10) A2,A4,A5,A6,A7,A8,A9,
     1 A10, FJK
      FFF=FJK
      GOTO 501
```

```
500 CONTINUE
 THE VAN VLECK-WEISKOPF LINE SHAPE
C
      Z1=0.755*PH2+0.231*PHE
      Z2=FLOAT(L2)/FLOAT(L1)
      Z1=Z1*(Z2**0.3333)
      Z2=SQRT(Z2)*PNH3*6.23
      Z=(Z1+Z2)*1.0E3/T
      X1=V*V*Z
      X2=(V-VLINE)*(V-VLINE)
      X3=(V+VLINE)*(V+VLINE)
      FJK=X1/(X2+Z*Z)+X1/(X3+Z*Z)
C
       G0T0 501
  502 CONTINUE
  THE GROSS LINE SHAPE
        X2=(-V*V+VLINE*VLINE)
        X2 = X2 * X2
       if(vline.lt.1000.) Z1=123.3551*PH2+60.7441*PHE
    if(vline.ge.1000.) z1=2.885*PH2+1.00*PHE
       Z2=FLOAT(L2)/FLOAT(L1)
C
    Z2=300./T
      Z1=Z1*(Z2**0.3)
    if(vline.lt.1000.) Z=154.1939*PNH3*Z2
    if(vline.ge.1000.) Z=22.989*PNH3*Z2
       Z=Z1+Z
        X3=Z*Z*4*V*V
        X1=v*VLINE*Z*4.
        X1=X1*V*VLINE
        FJK=X1/(X2+X3)
       FFF=FJK
  501 CONTINUE
  10 FORMAT(2X, 'A2, A4, A5, A6, A7, A8, A9, A10, FJK', /,
     1 2X,1P,9(2X,E10.3))
      RETURN
      END
C
      SUBROUTINE SFJTOM(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
    real*4 t3,gm,vline,fjk,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,
     + deltot, PCO, PCO13, PHCN
      COMMON/V1/A1,A2,A3,A4,A5,A6,A7,A8,A9,A10
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
       COMMON/U4/T1,T2,T3,T4,T5,T6,T7,T8,T9,T10,T11
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
C
      FJK IS THE BEN REUVEN LINE SHAPE, SEE EQN 25 OF BERGE AND
C
      GULKIS ARTICLE, PAGE 675 OF GEHRELS JUPITER BOOK.
C
      X=PH2+PHE+PH20+PNH3+PCH4
C
   IF X.LT.1 ATM VAN VLECK-WEISKOPF LINE SHAPE SHOUL BE
  USED; HOWEVER, BEN REUVEN IS STILL BETTER OUT IN THE
  WINGS OF THE AMMONIA LINES. SO OPTION VAN VLECK-WEISKOPF
  IS NEGLECTED BY SETTING IN NEXT LINE:
  IF(X.LT.0.0) INSTEAD OF IF(X.LT.1.0)
C
C
C
        IF(X.LT.1.) GOTO 500
C
```

```
\mathsf{C}
      A2 IS THE COUPLING FACTOR ZETA(J,K) IN GHZ.
c spliker's formalism uses equations below
    ZNH3=0.5
    ZHE=0.28-T/1750.
    ZH2=5.7465+GH2*(-7.7644+GH2*(9.1931+GH2*(-5.6816+1.2307*GH2)))
    if(v.le.30.0) goto 600
c Joanna Joiner's paper gives parameters below
C These parameters are better at freq. (v in GHz) above 40 GHz
C see de Boer's thesis, p. 61-64
        znh3=0.20
       zh2=1.35
        zhe=0.30
\mathsf{C}
      IF(VLINE.GT.100.0) GOTO 500
     A2=(ZH2*PH2*T3)+(ZNH3*T2*PNH3*SB)+(ZHE*T3*PHE)
600
      A4 IS THE PRESSURE SHIFT DEL (GHZ)
C
    A4=-0.45*PNH3
C
      SQUARES OF VARIOUS TERMS.
      A5=(VLINE+A4)*(VLINE+A4)
      A6=GM*GM
      A7 = A2 * A2
      A8=V*V
C
      NUMERATOR
      A9=((GM-A2)*A8)+((GM+A2)*(A5+A6-A7))
      A10=A8-A5-A6+A7
C
      DENOMINATOR
      A10=(A10*A10)+(4.0*A8*A6)
\mathsf{C}
      THE BEN REUVEN LINE SHAPE.
      FJK=2.0*A8*(A9/A10)
      FFF=FJK
      IF(V.LT.29.0) GOTO 501
C
 April 2003: we skip VVW because Spilker's formalism does already
c do the right thing, I believe
    GOTO 501
  500 CONTINUE
C THE VAN VLECK-WEISKOPF LINE SHAPE
      Z1=0.755*PH2+0.231*PHE
      Z2=FLOAT(L2)/FLOAT(L1)
      Z1=Z1*(Z2**0.3333)
      Z2=SQRT(Z2)*PNH3*6.23
      Z=(Z1+Z2)*1.0E3/T
      X1=V*V*Z
      X2=(V-VLINE)*(V-VLINE)
      X3=(V+VLINE)*(V+VLINE)
      FJK=X1/(X2+Z*Z)+X1/(X3+Z*Z)
C
       GOTO 501
  502 CONTINUE
 THE GROSS LINE SHAPE
        X2=(-V*V+VLINE*VLINE)
        X2 = X2 * X2
       if(vline.lt.1000.) Z1=123.3551*PH2+60.7441*PHE
    if(vline.ge.1000.) z1=2.885*PH2+1.00*PHE
C
       Z2=FLOAT(L2)/FLOAT(L1)
    Z2=300 . /T
      Z1=Z1*(Z2**0.3)
    if(vline.lt.1000.) Z=154.1939*PNH3*Z2
```

```
if(vline.ge.1000.) Z=22.989*PNH3*Z2
       Z=Z1+Z
        X3=Z*Z*4.*V*V
        X1=v*VLINE*Z*4.
        X1=X1*V*VLINE
        FJK=X1/(X2+X3)
       FFF=FJK
  501 CONTINUE
  10 FORMAT(2X, 'A2, A4, A5, A6, A7, A8, A9, A10, FJK', /,
     1 2X,1P,9(2X,E10.3))
      RETURN
      END
C
      SUBROUTINE H2H2(VNU,T,A,B)
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V3/QROT, QR1, QR2, S1, S2, S3, S4, P1, P2, P3, P4
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
      COMMON/V6/P5, P6, P7, P8
      COMMON/V7/PR01(20), PR02(20), STOR(20)
      COMMON/W1/FP, REB, RHO(20)
      COMMON/V9/TEMP, BETA, SQ, TAU1, TAU2
      H2 OPACITIES ARE CALCULATED FOLLOWING THE PAPER "ANALYSIS OF
C
C
      THE SHAPE OF THE FAR-INFRARED SPECTRA OF H2-H2 AND H2-HE
C
      COLLISIONS, BY COHEN AND BIRNBAUM, APRIL 1981.
C
      A RELATED PAPER IS GIVEN BY BIRNBAUM AND COHEN, CANADIAN
C
      JOURNAL OF PHYSICS, VOL 54, PAGE 593, 1976.
C
      THIS WORK IS NBSIR 80-2175(R).
C
      A AND B COEFFICIENTS IN THE TRAFTON TRADITION ARE CALCULATED.
      SEE TRAFTON, ASTROPHYSICAL JOURNAL, V 147, PAGE 765, 1967.
C
C
C
      THE H2-H2 COLLISIONS ARE CONSIDERED HERE FOR THE A VALUES.
C
      TEMP=T
C
      BOLTZMANN STATISTICS ARE CALCULATED.
      CALL JROT(TEMP)
C
      EVALUATE SQ, TAU1, TAU2 PARAMETERS.
      CALL PARAM(1)
      IF (NOPR .EQ. 1) WRITE(6,330) SQ,TAU1,TAU2
 330 FORMAT(/,2X,'FOR H2-H2 COLLISIONS',/,2X,'SQ,TAU1,TAU2',
     1 2X,1P,3(2X,E10.3))
      EVALUATE THE A FACTOR.
      CALL FAC(10, AFAC)
      IF (NOPR .EQ. 1) WRITE(6,355) AFAC
      FORMAT(2X, 'AFAC=', 2X, 1P, E10.3)
355
      V=VNU
      IF (NOPR .EQ. 1) WRITE(6,310) V
      FORMAT(2X, V(CM-1) = 1, 1P, E10.3)
310
      OM IS TWO PI TIMES C TIMES THE WAVENUMBER.
C
      FREQ IS SEC-1.
      OM=TPC*V
      FREQ=C*V*1.0E9
      BETA=1.0/(BOLC*TEMP)
      IF (NOPR .EQ. 1) WRITE(6,315) OM, FREQ, BETA
```

```
FORMAT(2X, 'OM(SEC-1), FREQ(SEC-1), BETA(ERG-1)',/,
 315
     1 2X,1P,3(2X,E10.3))
C
      EVALUATE THE 1-EXP TERM.
      TX=BETA*HB*0M
      CALL EX(TX,S)
   AW2 SHOULD BE MULTIPLIED BY 1.0E-36
      AW2 = AW * OM * (1 \cdot 0 - S) * SQ
      IF (NOPR .EQ. 1) WRITE(6,340) AW2
 340
      FORMAT(2X, 'AW2= ', 1P, E10.3)
C
C
      THE SIX TERMS OF EQUATION 7 OF NBS 80 ARE CALCULATED.
\mathsf{C}
      THE FIRST TERM IS EVALUATED.
      0M2=0M
      CALL GAMMA (OM2, GAM)
      IF (NOPR .EQ. 1) WRITE(6,356) TEMP, TAU1, TAU2, OM, BETA, GAM
 356
      FORMAT(2X, 1P, 6(2X, E10.3))
      TERM1=AFAC*GAM
          ZERO AND TWO ARE EVALUATED.
C
      CALL EVAL(0,2,0M,B1,B2)
      TERM2 = (RHO(1)*B1) + (RHO(3)*B2)
C
          ONE AND THREE ARE EVALUATED.
      CALL EVAL(1,3,0M,B1,B2)
      TERM3 = (RHO(2)*B1) + (RHO(4)*B2)
      TERM3 = (9.0/5.0) * TERM3
C
          TWO AND FOUR ARE EVALUATED.
      CALL EVAL(2,4,0M,B1,B2)
      TERM4=(RHO(3)*B1)+(RHO(5)*B2)
      TERM4 = (18.0/7.0) * TERM4
C
          THREE AND FIVE ARE EVALUATED.
      CALL EVAL(3,5,0M,B1,B2)
      TERM5 = (RHO(4)*B1) + (RHO(6)*B2)
      TERM5 = (10.0/3.0) * TERM5
C
   SUM3 SHOULD BE MULTIPLIED BY 1.0E-14
      SUM3=TERM1+TERM2+TERM3+TERM4+TERM5
      IF (NOPR .EQ. 1) WRITE(6,360) TERM1,TERM2,TERM3,TERM4,
     1 TERM5, SUM3
     FORMAT(2X, 'TERM1, TERM2, TERM3, TERM4, TERM5, SUM', /,
        2X,1P,6(2X,E10.3))
C
      NOTE: THE RESULT IS MULTIPLIED BY 1.0E36
      SUM=AW2*SUM3*2.9979E-4
C
C
        NOTICE FACTOR OF 2. NEED TO RESOLVE THIS DISCREPANCY
C
        WITH BIRNBAUM. (FACTOR OF 2 GIVES D(V) VALUES WHICH
C
        AGREE WITH BIRNBAUM'S GRAPHED VALUES.)
C
      A(T) IN THE TRAFTON STYLE IS PUT INTO A
      A=SUM
      IF (NOPR .EQ. 1) WRITE(6,104) V,T,A
      FORMAT(2X, V, T, A = 1, 2X, 1P, 3(2X, E10.3))
 104
C
      THE H2-HE COLLISIONS ARE CONSIDERED HERE FOR THE B VALUES.
C
C
      TEMP=T
C
      BOLTZMANN STATISTICS ARE CALCULATED.
      CALL JROT(TEMP)
```

```
IF (NOPR .EQ. 1) WRITE(6,200)
200
      FORMAT(/,2X,'FOR H2-HE COLLISIONS')
      EVALUATE SQ, TAU1, TAU2 PARAMETERS.
C
      FOR THE ISOTROPIC S AND TAU FACTORS.
      CALL PARAM(2)
      SOA=SO
      TAU1A=TAU1
      TAU2A=TAU2
      IF (NOPR .EQ. 1) WRITE(6,210) SQA, TAU1A, TAU2A
210 FORMAT(2X, 'ISOTROPIC SQA, TAU1, TAU2A', 2X, 1P, 3(2X, E10.3))
      EVALUATE SQ, TAU1, TAU2 PARAMETERS.
C
C
      FOR THE ANISOTROPIC S AND TAU FACTORS.
      CALL PARAM(3)
      SQB=SQ
      TAU1B=TAU1
      TAU2B=TAU2
      IF (NOPR .EQ. 1) WRITE(6,220) SQB, TAU1B, TAU2B
      FORMAT(2X, 'ATISOTROPIC SQB, TAU1B, TAU2B', 2X, 1P, 3(2X, E10.3))
220
      EVALUATE THE A FACTOR.
      CALL FAC(10, AFAC)
      IF (NOPR .EQ. 1) WRITE(6,355) AFAC
C
      V=VNU
      IF (NOPR .EQ. 1) WRITE(6,310) V
\mathsf{C}
      OM IS TWO PI TIMES C TIMES THE WAVENUMBER.
C
      FREQ IS SEC-1.
      OM=TPC*V
      FRE0=C*V*1.0E9
      BETA=1.0/(BOLC*TEMP)
      IF (NOPR .EQ. 1) WRITE(6,315) OM, FREQ, BETA
C
      EVALUATE THE 1-EXP TERM.
      TX=BETA*HB*0M
      CALL EX(TX,S)
C
      NOTE THAT THE SQ FACTOR IS NOT IN THE AW2 EXPRESSION.
C
      FOR H2-HE, THERE IS AN ADDITIONAL FACTOR OF 2.
\mathsf{C}
      THE SQA AND SQB FACTORS WILL BE INTRODUCED LATER.
      AW2=2.0*AW*OM*(1.0-S)
      IF (NOPR .EQ. 1) WRITE(6,340) AW2
C
C
      THE ISOTROPIC TERM IS CALCULATED HERE.
      S0=S0A
      TAU1=TAU1A
      TAU2=TAU2A
      IF (NOPR .EQ. 1) WRITE(6,430) SQ,TAU1,TAU2
 430
      FORMAT(2X, 'ISOTROPIC SQ, TAU1, TAU2', 2X, 1P, 3(2X, E10.3))
      0M2=0M
      CALL GAMMA (OM2, GAM)
      IF (NOPR .EQ. 1) WRITE(6,356) TEMP, TAU1, TAU2, OM, BETA, GAM
C
      NOTE: THE RESULT IS MULTIPLIED BY 1.0E36
      TERM6=S0*GAM
      IF (NOPR .EQ. 1) WRITE(6,357) SQ,GAM,TERM6
357
      FORMAT(2X, 'ISOTROPIC SQ, GAM, TERM6 ', 1P, 6(2X, E10.3))
C
      THE ANISOTROPIC TERMS ARE CALCULATED HERE.
C
C
      THE SIX TERMS OF EQUATION 7 OF NBS 80 ARE CALCULATED.
      SQ=SQB
```

```
TAU1=TAU1B
      TAU2=TAU2B
      IF (NOPR .EQ. 1) WRITE(6,431) SQ, TAU1, TAU2
      FORMAT(2X, 'ANISOTROPIC SQ, TAU1, TAU2', 2X, 1P, 3(2X, E10.3))
431
      THE FIRST TERM IS EVALUATED.
      0M2=0M
      CALL GAMMA (OM2, GAM)
      IF (NOPR .EQ. 1) WRITE(6,356) TEMP, TAU1, TAU2, OM, BETA, GAM
      TERM1=AFAC*GAM
C
          ZERO AND TWO ARE EVALUATED.
      CALL EVAL(0,2,0M,B1,B2)
      TERM2=(RHO(1)*B1)+(RHO(3)*B2)
C
          ONE AND THREE ARE EVALUATED.
      CALL EVAL(1,3,0M,B1,B2)
      TERM3 = (RHO(2)*B1) + (RHO(4)*B2)
      TERM3 = (9.0/5.0) * TERM3
C
          TWO AND FOUR ARE EVALUATED.
      CALL EVAL(2,4,0M,B1,B2)
      TERM4=(RHO(3)*B1)+(RHO(5)*B2)
      TERM4=(18.0/7.0)*TERM4
C
          THREE AND FIVE ARE EVALUATED.
      CALL EVAL(3,5,0M,B1,B2)
      TERM5 = (RHO(4)*B1) + (RHO(6)*B2)
      TERM5 = (10.0/3.0) * TERM5
C
      SUM3=TERM1+TERM2+TERM3+TERM4+TERM5
C
      THE ANISOTROPIC TERMS ARE MULTIPLED BY SQB
C
      NOTE: THE RESULT IS MULTIPLIED BY 1.0E36
      SUM3=SQ*SUM3
C
      ADD UP THE ISOTROPIC AND ANISOTROPIC TERMS.
      SUM3=SUM3+TERM6
      IF (NOPR .EQ. 1) WRITE(6,460) TERM1,TERM2,TERM3,TERM4,TERM5,
     1 SUM3
460 FORMAT(2X, 'TERM1, TERM2, TERM3, TERM4, TERM5, TOTAL SUM*SQ',/,
     1 2X,1P,7(2X,E10.3))
C
      NOTE: THE RESULT IS MULTIPLIED BY 1.0E36
      SUM=AW2*SUM3*2.9979E-4
C
C
        NOTICE FACTOR OF 2. NEED TO RESOLVE THIS DISCREPANCY
C
        WITH BIRNBAUM. (FACTOR OF 2 GIVES D(V) VALUES WHICH
        AGREE WITH BIRNBAUM'S GRAPHED VALUES.)
C
C
      B(T) IN THE TRAFTON STYLE IS PUT INTO B
      B=SUM
      IF (NOPR .EQ. 1) WRITE(6,404) V,T,B
      FORMAT(2X, V, T, B = (2X, 1P, 3(2X, E10.3), /)
 404
C
C
      RETURN
      END
C Subroutine for H2H2 absorption from Joiner and Steffes
    SUBROUTINE H2H2JS(VNU,T,AH2H2)
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V3/QROT,QR1,QR2,S1,S2,S3,S4,P1,P2,P3,P4
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
```

```
COMMON/V6/P5, P6, P7, P8
      COMMON/V7/PR01(20), PR02(20), STOR(20)
      COMMON/W1/FP, REB, RHO(20)
      COMMON/V9/TEMP, BETA, SQ, TAU1, TAU2
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
C
    X1=(273.0/T)**3.12
    X2=(273.0/T)**2.24
    X3=(273.0/T)**3.34
    X=PH2*X1 + 1.382*PHE*X2 + 9.322*PCH4*X3
      FREQ=C*VNU*1.0E9
      ZL=1.0/VNU
    Y=3.557E-11*PH2/(ZL*ZL)
    AH2H2=Y*X
    RETURN
    END
C Following subroutine is based on Orton's tables
C Following subroutine is based on Orton's tables
    SUBROUTINE H2H2OR(IQ, VNU, T, A, B, CZB)
      dimension Ttab(10), frequ(2428)
      dimension ah2h2e(10), ah2h2n(10), ah2hee(10), ah2hen(10)
      dimension ah2che(10),ah2chn(10)
      dimension abeh2h2(2428,10),abnh2h2(2428,10)
      dimension abeh2he(2428,10),abnh2he(2428,10)
      dimension abeh2ch4(2428,10),abnh2ch4(2428,10)
       real*8 A,B,CZB,ah2e,ah2n,ahee,ahen,ache,achn
C
       integer ntemp, nfreq
С
      COMMON/orton/Ttab, frequ, abeh2h2, abnh2h2, abeh2he, abnh2he,
     + abeh2ch4,abnh2ch4,nfreq,ntemp
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V9/TEMP, BETA, SQ, TAU1, TAU2
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
      COMMON/W1/FP, REB, RHO(20)
C
      algT=alog(T)
C
   CHANGE PRESSURES FROM BARS TO ATM, AND BACK AT END OF SUBROUTINE;
        kp=0
        do 10 ip=1,nfreq
           if(vnu_ge_frequ(ip)) kp=kp+1
           if(vnu.lt.frequ(ip)) goto 11
 10
           continue
C so vnu is in between kp-1 and kp (after next line)
 11
            kp=kp+1
            x=vnu-frequ(kp-1)
           do i=1,ntemp
            ah2h2e(i)=abeh2h2(kp-1,i)
     +
            +x/(frequ(kp)-frequ(kp-1))*(abeh2h2(kp,i)-abeh2h2(kp-1,i))
            ah2h2n(i)=abnh2h2(kp-1,i)
            +x/(frequ(kp)-frequ(kp-1))*(abnh2h2(kp,i)-abnh2h2(kp-1,i))
            ah2hee(i)=abeh2he(kp-1,i)
            +x/(frequ(kp)-frequ(kp-1))*(abeh2he(kp,i)-abeh2he(kp-1,i))
            ah2hen(i)=abnh2he(kp-1,i)
            +x/(frequ(kp)-frequ(kp-1))*(abnh2he(kp,i)-abnh2he(kp-1,i))
     +
            ah2che(i)=abeh2ch4(kp-1,i)
            +x/(frequ(kp)-frequ(kp-1))*(abeh2ch4(kp,i)-abeh2ch4(kp-1,i))
```

```
ah2chn(i)=abnh2ch4(kp-1,i)
            +x/(frequ(kp)-frequ(kp-1))*(abnh2ch4(kp,i)-abnh2ch4(kp-1,i))
           enddo
C
           kp=0
        do 12 ip=1,ntemp
           if(algT_ge_Ttab(ip)) kp=kp+1
           if(algT.lt.Ttab(ip)) goto 13
 12
        continue
C
   so algT is in between kp-1 and kp
 13
                kp=kp+1
                x=algT-Ttab(kp-1)
C
               ah2e=ah2h2e(kp-1)
                 +x/(Ttab(kp)-Ttab(kp-1))*(ah2h2e(kp)-ah2h2e(kp-1))
     +
               ah2n=ah2h2n(kp-1)
                 +x/(Ttab(kp)-Ttab(kp-1))*(ah2h2n(kp)-ah2h2n(kp-1))
     +
               ahee=ah2hee(kp-1)
                 +x/(Ttab(kp)-Ttab(kp-1))*(ah2hee(kp)-ah2hee(kp-1))
               ahen=ah2hen(kp-1)
                 +x/(Ttab(kp)-Ttab(kp-1))*(ah2hen(kp)-ah2hen(kp-1))
               ache=ah2che(kp-1)
                 +x/(Ttab(kp)-Ttab(kp-1))*(ah2che(kp)-ah2che(kp-1))
     +
               achn=ah2chn(kp-1)
                 +x/(Ttab(kp)-Ttab(kp-1))*(ah2chn(kp)-ah2chn(kp-1))
C
        if(IQ.ge.1) A=exp(ah2e)
        if(IQ.ge.1) B=exp(ahee)
        if(IQ.ge.1) CZB=exp(ache)
        if(IQ.eq.0) A=exp(ah2n)
        if(I0.eq.0) B=exp(ahen)
        if(IQ.eq.0) CZB=exp(achn)
C
C
         write(6,667) (Ttab(i), i=1,10)
\mathsf{C}
         write(6,668) vnu,algt,nfreq,ntemp
C
         write(6,669) abeh2h2(128,3), frequ(128)
         write(6,666) T,A,B,CZB,ah2e,ahee,ache
        format(7E15.6)
 666
        format(10f8.3)
 667
 668
        format(2f8.3,2I10)
 669
        format(2E15.5)
    RETURN
    END
C
C
      SUBROUTINE EVAL(L,K,OM,B1,B2)
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
      COMMON/V9/TEMP, BETA, SQ, TAU1, TAU2
C
      CALCULATE THE GAMMA VALUES.
C
      SEE NBS 80.
      CALL EH2(L,E)
      B1=E
      CALL EH2(K,E)
C
      B3 IS W(L,K) IN 2 PI RAD PER SEC.
```

```
B3=TPC*(B2-B1)
      0M2=0M-B3
      IF (NOPR .EQ. 1) WRITE(6,10) L,K,OM,B3
     FORMAT(2X, I3, 2X, 1P, I3, 2X, 1P, 2(2X, E10.3))
      CALL GAMMA (OM2, GAM)
      IF (NOPR .EQ. 1) WRITE(6,15) TEMP, TAU1, TAU2, OM2, BETA, GAM
      FORMAT(2X, 1P, 6(2X, E10.3))
      B1=GAM
      0M2=0M+B3
      CALL GAMMA (OM2, GAM)
      IF (NOPR .EQ. 1) WRITE(6,15) TEMP, TAU1, TAU2, OM2, BETA, GAM
      B2=GAM
      RETURN
      END
C
C
C
      SUBROUTINE PARAM(M)
      COMMON/V1/A1,A2,A3,A4,A5,A6,A7,A8,A9,A10
      COMMON/V9/TEMP, BETA, SQ, TAU1, TAU2
C
      SEE NBS 80 FOR THE SQ, TAU1, TAU2 PARAMETER VALUES.
      IF (M .EQ. 1) GO TO 10
      IF (M .EQ. 2) GO TO 20
      IF (M .EQ. 3) GO TO 30
C
      FOR H2-H2 COLLISIONS.
  10 A1=TEMP/273.15
C
       A2=142.0*EXP(0.26*AL0G(A1))
       A3=4.85*EXP(-0.593*ALOG(A1))
C
\mathsf{C}
       A4=2.17*EXP(-0.523*ALOG(A1))
      A2=281.0*(A1**0.235)
      A3=4.68*(A1**(-0.605))
      A4=2.23*(A1**(-0.607))
      SQ=1.38*A2
      TAU1=A3
      TAU2=A4
      G0 T0 40
C
      FOR H2-HE COLLISIONS, THE ISOTROPIC PARAMETERS.
  20
      CONTINUE
C
      A1=TEMP/273.15
       A2=112.0*EXP(0.93*ALOG(A1))
C
C
       A3=1.74*EXP(-0.54*ALOG(A1))
       A4=3.4*EXP(-0.30*ALOG(A1))
      A1=77.4/TEMP
      A2=33.53/A1
      A3=3.43*SQRT(A1)
      A4=6.56*SQRT(A1)
      S0=1.38*A2
      TAU1=A3
      TAU2=A4
      GO TO 40
C
      FOR H2-HE COLLISIONS, ANISOTROPIC PARAMETERS.
      VALUES ARE FOR 195 DEGREES KELVIN.
C
  30
     CONTINUE
C
      A1=TEMP/273.15
C
       A2=1.914E1
C
       A3=4.13
```

```
C
       A4=2.38
      A1=TEMP/77.4
      A2=12.06*(A1**0.57)
      A3=3.02*(A1**(-0.30))
      A4=8.94*(A1**(-0.60))
      S0=1.38*A2
      TAU1=A4
      TAU2=A3
      G0 T0 40
  40 RETURN
      END
C
C
C
      SUBROUTINE FAC(L, AFAC)
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/V7/PR01(20), PR02(20), STOR(20)
      COMMON/W1/FP, REB, RHO(20)
\mathsf{C}
      SEE NBS 80 FOR THE AFAC FACTOR, EQUATION 8.
      A8=0.0
      D0 10 I=1,L
      J=I-1
      A1=J
      A2=A1+1.00
      A3=2.0*A1
      A4=A3+1.00
      A5=A3-1.00
      A6=A3+3.00
      A7=(A1*A2*A4*RHO(I))/(A5*A6)
      A8 = A8 + A7
  10 CONTINUE
      AFAC=A8
      RETURN
      END
C
C
C
      SUBROUTINE BESSEL(Z,BES)
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
C
      BES IS Z*K1(Z)
C
      SEE PAGES 378-379 OF ABRAMOWITZ AND STEGUN, HANDBOOK OF
C
      MATHEMATICAL FUNCTIONS.
      EQNS. 9.8.3,9.8.4,9.8.7,9.8.8 ARE USED TO CALCULATE Z*K1(Z)
      IF (Z .LE. 2.00) GO TO 10
      IF (Z .GT. 2.00) GO TO 20
C
      A1, A2, ... ARE Z/2 TO THE 2,4, ... POWER. SEE EQN 9.8.7.
      BESI IS I1(Z).
  10 CALL BESS2(Z,BESI)
      A1=(Z*Z)/(2.0*2.0)
      A2=A1*A1
      A3=A2*A1
      A4=A3*A1
      A5 = A4 * A1
      A6 = A5 * A1
      BES=(Z*ALOG(Z/2.0)*BESI)+(1.000)+(.15443144*A1)
      BES=BES-(.67278579*A2)-(.18156897*A3)-(.01919402*A4)
```

```
BES=BES-( 00110404*A5)-( 00004686*A6)
      GO TO 30
C
      A1, A2, ... ARE Z/2 TO THE 1,2, ... POWER. SEE EQN 9.8.8.
  20 A1=2.0/Z
      A2=A1*A1
      A3 = A2 * A1
      A4=A3*A1
      A5 = A4 * A1
      A6 = A5 * A1
      BES=1.25331414+(.23498619*A1)-(.03655620*A2)
      BES=BES+(.01504268*A3)-(.00780353*A4)
      BES=BES+(.00325614*A5)-(.00068245*A6)
      CALL EX(Z,A1)
      BES=BES*SQRT(Z)*A1
  30 RETURN
      END
C
C
\mathsf{C}
      SUBROUTINE BESS2(Z,BESSI)
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
      BESSI IS I1(Z).
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      IF (Z .LE. 3.75) GO TO 10
      IF (Z .GT. 3.75) GO TO 20
C
      A1,A2,.. IS Z/3.75 TO THE 2,4,.. POWER. SEE EQN. 9.8.3.
  10 A1=(Z*Z)/(3.75*3.75)
      A2=A1*A1
      A3 = A2 * A1
      A4=A3*A1
      A5 = A4 * A1
      A6 = A5 * A1
      BESSI=0.5000+(.87890594*A1)+(.51498869*A2)
      BESSI=BESSI+(.15084934*A3)+(.02658733*A4)
      BESSI=BESSI+(.00301532*A5)+(.00032411*A6)
      BESSI=BESSI*Z
      GO TO 30
      A1,A2,.. ARE 3.75/Z TO THE 1,2,.. POWER. SEE EQN. 9.8.4.
  20 A1=3.75/Z
      A2=A1*A1
      A3=A2*A1
      A4=A3*A1
      A5 = A4 * A1
      A6 = A5 * A1
      A7=A6*A1
      A8=A7*A1
      BESSI=.39894228-(.03988024*A1)-(.00362018*A2)
      BESSI=BESSI+(.00163801*A3)-(.01031555*A4)
      BESSI=BESSI+(.02282967*A5)-(.02895312*A6)
      BESSI=BESSI+( 01787654*A7)-( 00420059*A8)
      BESSI=BESSI*(EXP(Z)/SQRT(Z))
      IF ((Z .GT. 150.) .AND. (NOPR .EQ. 1)) WRITE(6,29)
  29 FORMAT(2X, 'Z .GT. 150., EXP PROBLEMS')
  30
      RETURN
      END
C
```

```
C
C
      SUBROUTINE GAMMA(OM2,GAM)
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
      COMMON/V9/TEMP, BETA, SQ, TAU1, TAU2
C
      SEE EQNS 17 OF BIRNBAUM AND COHEN (CAN J PHYS, 54, PG 593, 1976)
C
      ALSO EQUATION 5 OF NBS 80.
      A1=TAU1
      A2=TAU2
      A3=0.5*BETA*HB*0M2
      A4=1.00+(0M2*0M2*A1*A1*1.0E-6)
      A5=(A2*A2*1.0E-6)+((BETA*BETA*HB*HB)/4.0)
      GAM = (A1/(PI*A4))*EXP((A2/A1)+A3)
        A6=GAM
      Z=SQRT(A4*A5)*1.0E3/A1
      CALL BESSEL(Z,BES)
   GAM SHOULD BE MULTIPLIED BY 1.0E-14
      GAM=GAM*BES
      IF (NOPR .EQ. 1) WRITE(6,10) A1,A2,A3,A4,A5,A6,Z,GAM
     FORMAT(2X, 'TAU1, TAU2, A3, A4, A5, A6, Z, GAM', /,
     1 2X,1P,8(2X,E10.3))
      RETURN
      END
      SUBROUTINE EH2(J,E)
      COMMON/V1/A1,A2,A3,A4,A5,A6,A7,A8,A9,A10
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
C
      ENERGY VALUE IN WAVENUMBERS (CM-1) FOR THE
      JTH ROTATIONAL LEVEL OF H2, SEE NBS 80, EQUATION 2.
      A1=J
      A2=A1+1.00
      A3=A1*A1
      A4=A2*A2
      A5 = A3 * A1
      A6=A4*A2
      E=(59.3392*A1*A2)-(0.04599*A3*A4)+(5.2E-5*A5*A6)
      RETURN
      END
C
C
C
      SUBROUTINE JROT(TEMP)
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
      COMMON/V7/PR01(20), PR02(20), STOR(20)
      COMMON/W1/FP, REB, RHO(20)
C
      THE BOLTZMANN STATISTICS FOR H2 ARE CALCULATED
      CALL PART(TEMP)
C
       RHO(L) GIVES THE FRACTION OF H2 IN THE
      K=L-1 ROTATIONAL LEVEL, DIVIDED BY 2J+1.
C
C
      AND MULTIPLIED BY FP FOR THE NORMAL CASE, AND
C
      BY 1.00 IN THE EQUIL CASE (SEE SUBROUTINE
C
      PART FOR DEFINITION OF STOR).
      FRACT IS THE FRACTION OF H2 IN THE K=L-1 LEVEL.
C
C
      NOTICE SEPARATE EQUIL AND NORMAL DEFINITIONS OF FRACT.
C
      EQUIL CASE, KL=0 AND A2=1.00
```

```
NORMAL CASE, KL=10 AND A2=FP.
345 FORMAT(/,2X,'BOLTZMANN STATISTICS FOR H2',/,
     1 2X, 'KL IS EQUAL TO ', I3)
      FOR NORMAL CASE, FRACTION OF PARA IS 0.25
C
        PARA LEVELS J=0,2,4..ARE SPECIFIED.
      D0 20 L=1,5
C
   K=1,3,5,7,9
      K=((L-1)*2)+1
      J=K-1
      A1=(2*(K-1))+1
C
      POP=(REB*STOR(K))+((1.0-REB)*FP*STOR(K+10))
      RHO(K)=POP/A1
 350
      FORMAT(2X, 'J, FRACT, PR01 ', 2X, I3, 2X, 1P, 2(2X, E10.3))
  20
      CONTINUE
\mathsf{C}
        ORTHO LEVELS J=1,3,5.. ARE SPECIFIED.
      D0 21 L=1,5
      K=((L-1)*2)+2
      J=K-1
      A1=(2*(K-1))+1
      POP=(REB*STOR(K))+((1.0-REB)*(1.0-FP)*STOR(K+10))
      RHO(K)=POP/A1
  21 CONTINUE
      RETURN
      END
      SUBROUTINE PART(TEMP)
      COMMON/V3/QROT, QR1, QR2, S1, S2, S3, S4, P1, P2, P3, P4
      COMMON/V6/P5, P6, P7, P8
      COMMON/V7/PR01(20), PR02(20), STOR(20)
      DIMENSION CC(20), CC2(20)
C
      SEE TATUM, AP J SUPPL SER, VOL 14, PAGE 21, 1967
\mathsf{C}
      FOR THE EXPRESSIONS FOR THE BOLTZMANN STATISTICS
C
      THAT APPLY TO THE GROUND STATE OF H2.
      HERZBERG, MOLECULAR SPECT AND MOL STRUCT, VOL 4, CONSTS
C
C
      OF DIATOMIC MOLECULES IS USED FOR THE BV, DE VALUES.
      HCK=1.4388
      BV=60.853-(3.062*0.5)
      DE=4.71E-2
      SNUC=0.5
      CI1=SNUC/((2.0*SNUC)+1)
      CI2=(SNUC+1)/((2.0*SNUC)+1)
      CONST = ((2.0*0)+1)*((2.0*SNUC)+1)*((2.0*SNUC)+1)
      SUM=0.0
      SUM2=0.0
      L=0
      L1=1
      LR=5
      D0 31 J=1,LR
      DG1=(2.0*L)+1.0
      DG2=(2.0*L1)+1.0
      J1=L*(L+1)
      J2=L1*(L1+1)
      F=(BV*J1)-(DE*J1*J1)
      F2=(BV*J2)-(DE*J2*J2)
      CALL EH2(L,E)
      F=E
```

```
CALL EH2(L1,E)
      F2=E
      TX=HCK*(F/TEMP)
      CALL EX(TX,S)
      CC(J)=S*CONST*CI1*DG1
      SUM=SUM+CC(J)
      TX2=HCK*(F2/TEMP)
      CALL EX(TX2,S2)
      CC2(J)=S2*CONST*CI2*DG2
      SUM2=SUM2+CC2(J)
      L=L+2
      L1=L1+2
  31 CONTINUE
      QROT=SUM+SUM2
      QR1=SUM
      QR2=SUM2
      S1=0.0
      S2=0.0
      S3=0.0
      S4=0.0
C
      S1=PROB, TOTAL PARA, EQUILIBRIUM
C
      S2=PROB, TOTAL ORTHO, EQUILIBRIUM
C
      S3=PROB, TOTAL PARA, NORMAL
C
      S4=PROB, TOTAL ORTHO, NORMAL
C
      P1=PR0B, T0TAL ORTHO/PARA, EQUILIBRIUM
C
      P2=PR0B,ORTHO J=1/PARA J=0,EQUILIBRIUM
C
      P3=PR0B, T0TAL ORTHO/PARA, NORMAL
C
      P4=PR0B, ORTHO J=1/PARA J=0, NORMAL
C
      P5=PR0B,ORTHO J=1,EQUILIBRIUM
C
      P6=PR0B, PARA J=0, EQUILIBRIUM
C
      P7=PR0B, ORTHO J=1, NORMAL
\mathsf{C}
      P8=PR0B, PARA J=0, NORMAL
      D0 32 J=1,LR
      S1=S1+(CC(J)/QROT)
      S2=S2+(CC2(J)/QR0T)
      S3=S3+(CC(J)/QR1)
      S4=S4+(CC2(J)/QR2)
  32 CONTINUE
      P1=S2/S1
      P2=CC2(1)/CC(1)
      P3=S4/S3
      P4=(CC2(1)/CC(1))*(QR1/QR2)
      P5=CC2(1)/QR0T
      P6=CC(1)/QROT
      P7=CC2(1)/QR2
      P8=CC(1)/QR1
      DO 40 J=1,LR
      PR01(J)=CC(J)/QR0T
      PR02(J)=CC2(J)/QR0T
      PR01(J+10)=CC(J)/QR1
      PR02(J+10)=CC2(J)/QR2
  40
      CONTINUE
      STOR(I) I=1,2,3,4.. FOR JROT=0,1,2,3,...
C
C
      EQUIL DISTRIBUTION FOR I=1 TO 10
C
      NORMAL DISTRIBUTION FOR I=11 TO 20
      D0 50 J=1,LR
```

```
L=((J-1)*2)+1
      L1=L+1
C
      AA1=1.00, AA2=1.00 FOR STANDARD BOLTZMANN STATISTICS
C
      WHERE THE DEGENERACY IS UPPSTAIRS AND DOWNSTAIRS.
C
      AA1=.25*L, AA2=.75*L1 FOLLOWS EQUATION 9B OF
      BIRNBAUM AND COHEN(WRONG?)
      AA1=1.00
      AA2=1.00
      STOR(L)=PR01(J)/AA1
      STOR(L1)=PRO2(J)/AA2
      STOR(L+10)=PR01(J+10)/AA1
      STOR(L1+10) = PRO2(J+10)/AA2
  50 CONTINUE
      RETURN
      END
      SUBROUTINE DIEL(I,NPR,W,TK,E1,E2,D1,D2)
    COMPLEX ZM, ZZ
    REAL*4 LF, LF0, LF1
    DIMENSION EIMAG(9), FR(9)
   SUBROUTINE TO CALCULATE DIELECTRIC CONST/REFRACTIVE INDICES FOR
C WATER, BASED UPON ULABY ET AL, 1981.
    F=30.0E9/W
      IF(I.EQ.1) GOTO 100
      IF(I.EQ.2) GOTO 200
  100 TEMP=TK-273.0
      A1=TEMP
      A2 = A1 * A1
      A3 = A2 * A1
    RELT=(1.1109E-10)-A1*(3.824E-12)+A2*(6.938E-14)-
     + A3*(5.096E-16)
      E0=88.045-0.4147*A1+A2*(6.295E-4)+A3*(1.075E-5)
    if(E0.LT.0.0) E0=0.0
      EINF=4.9
      E1=(E0-EINF)/(1+(F*RELT)*(F*RELT))
    E1=EINF+E1
    E2=F*RELT*(E0-EINF)
    E2=E2/(1+(F*RELT)*(F*RELT))
    IF(E2.LT.0.0) E2=0.0
      G0T0 900
  200 CONTINUE
    TEMP=TK-273.0
    FR(1)=1.0E8
    FR(2)=3.0E8
    FR(3)=1.0E9
    FR(4)=2.0E9
    FR(5)=3.0E9
    FR(6)=5.0E9
    FR(7)=1.0E10
    FR(8)=3.0E10
    FR(9)=1.0E11
    EIMAG(1)=8.0E-3
    EIMAG(2)=1.5E-3
    EIMAG(3) = 8.0E - 4
    EIMAG(4)=1.0E-3
    EIMAG(5)=1.2E-3
    EIMAG(6)=1.5E-3
```

```
EIMAG(7)=3.0E-3
    EIMAG(8) = 8.0E - 3
    EIMAG(9)=2.0E-2
    E1=3.15
    LF=L0G10(F)
    IF(F.LE.FR(1)) THEN
    E2=8.0E-3
    G0T0 900
    ENDIF
    IF(F.GE.FR(9)) THEN
    E2=2.0E-2
        G0T0 900
    ENDIF
    IF(F.GT.FR(1).AND.F.LE.FR(2)) J=1
    IF(F.GT.FR(2).AND.F.LE.FR(3)) J=2
    IF(F.GT.FR(3).AND.F.LE.FR(4)) J=3
    IF(F.GT.FR(4).AND.F.LE.FR(5)) J=4
    IF(F.GT.FR(5).AND.F.LE.FR(6)) J=5
    IF(F.GT.FR(6).AND.F.LE.FR(7)) J=6
    IF(F.GT.FR(7).AND.F.LE.FR(8)) J=7
    IF(F.GT.FR(8).AND.F.LE.FR(9)) J=8
    LF0=ALOG10(FR(J))
    LF1=ALOG10(FR(J+1))
    DLF=(LF-LF0)/(LF1-LF0)
    X0=ALOG10(EIMAG(J))
    X1=ALOG10(EIMAG(J+1))
    DX=X0+DLF*(X1-X0)
    E2=10**DX
  900
        CONTINUE
С
    ZM = CMPLX(E1, -E2)
    ZZ=SQRT(ZM)
C
C
    D1=REAL(ZZ)
    D2=-AIMAG(ZZ)
      D1=(E1+SQRT((E1*E1)+(E2*E2)))/2.0
      D1=SQRT(D1)
      IF(D1.EQ.0.0) GOTO 910
      D2=E2/(2.0*D1)
      G0T0 9999
  910 D2=0
 9999 RETURN
      END
C
C
C
       FUNCTION SHAPE2(P,T,F,FLC)
        implicit none
        REAL*4 SHAPE2
        real*4 DL,P,T,F,FLC,DFC,DL2
        DL=F-FLC
    DL2=F+FLC
   line width is taken as water line width
    DFC = 0.08*P*((273/T)**0.666)
    DFC = DFC * 30.0
                             !LINEWIDTH IN MHZ
       SHAPE2=((F*F)/(FLC*FLC))*(DFC/(DL*DL+DFC*DFC)
     + +DFC/(DL2*DL2+DFC*DFC))
    SHAPE2=SHAPE2/3 1415926536
                                     !DIVIDE BY PI
```

```
RETURN
       END
C
       REAL*4 FUNCTION QLOG2(Q0,T)
       IMPLICIT NONE
       REAL*4 Q0(5),T
       REAL*4 SLOPE, TLOG
C LOG(T) FOR T=300,225,150,75, 37.5K
       REAL*4 TLOG0(5) /2.47712,2.35218,2.17609,1.8751, 1.5740/
       TLOG=ALOG10(T)
       IF(TLOG.GT.TLOG0(1)
                                                     )GO TO 4
       IF(TLOG.LE.TLOGO(1) .AND. TLOG.GT. TLOGO(2))GO TO 5
       IF(TLOG.LE.TLOG0(2) .AND. TLOG.GT. TLOG0(3))GO TO 6
       IF(TLOG.LE.TLOG0(3) .AND. TLOG.GT. TLOG0(4))GO TO 7
       IF(TLOG.LE.TLOGO(4) .AND. TLOG.GT. TLOGO(5))GO TO 8
       IF(TLOG.LE.TLOG0(5)
                                                    )GO TO 9
4
       QL0G2=Q0(1)
                      !USE 300 K VALUE NEEDS TO BE FIXED
       GO TO 10
5
       SLOPE=(00(2)-00(1))/(TLOG0(2)-TLOG0(1))
       0L0G2=Q0(1)+SL0PE*(TL0G-TL0G0(1))
       G0 T0 10
       SLOPE=(Q0(3)-Q0(2))/(TLOG0(3)-TLOG0(2))
6
       QLOG2=Q0(2)+SLOPE*(TLOG-TLOG0(2))
       G0 T0 10
7
       SLOPE=(00(4)-00(3))/(TLOG0(4)-TLOG0(3))
       0L0G2=00(3)+SL0PE*(TL0G-TL0G0(3))
       G0 T0 10
 8
       SLOPE=(Q0(5)-Q0(4))/(TLOG0(5)-TLOG0(4))
       QLOG2=Q0(4)+SLOPE*(TLOG-TLOG0(4))
       G0 T0 10
9
                                 !USE 37.5 K VALUE -NEEDS TO BE FIXED
       QLOG2=Q0(5)
10
       CONTINUE
       RETURN
       END
C
\mathsf{C}
      SUBROUTINE SAJph3(N,T,T3,GM)
        REAL*4 T,P,F,V
    real*8 x,ajk,fjk,tcor
    real*4 T3,qm,vline,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,deltot,PCO,PCO13,PHCN
    real*4 FP0(728), SP0(728), ELP(728)
    REAL*4 wgtS0(40),wgtFGB(40),wgtSB(40)
    integer*2 N
        COMMON/ph3/FP0,SP0,ELP,wqtS0,wqtFGB,wqtSB
        COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
C
      GM IS GAMMA(N) IN GHZ.
    GH2=3.293
    GHE=1.6803
    Gph3=4.2157
    if(N.qt.40) goto 40
        GM=(GH2*T3*PH2)+(GHE*T3*PHE)
    GM=GM*wqtfqb(N)
    GM=GM+(GPH3*(300.0/T)*PPh3*wgtSB(N))
    goto 50
       GM=(GH2*T3*PH2)+(GHE*T3*PHE)
    GM=GM+(GPH3*(300 \ 0/T)*PPh3)
```

```
50
     continue
     RETURN
     END
C
     SUBROUTINE SFJPH3(N,T,V,VLINE,GM,FJK)
       REAL*4 T,P,F,V
    real*8 x,ajk,fjk,tcor,y,z
        real*4 t3,gm,vline,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,deltot,PCO,PCO13,PHCN
    real*4 FP0(728),SP0(728),ELP(728)
    REAL*4 wgtS0(40),wgtFGB(40),wgtSB(40)
    integer*2 N
        COMMON/ph3/FP0,SP0,ELP,wgtS0,wgtFGB,wgtSB
        COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
C
    zeta=0.
    delta=0.0
C
     FJK IS THE BEN REUVEN LINE SHAPE, FROM DEBOER AND STEFFES
    x = (qm+zeta)*((vline+delta)*(vline+delta) + GM*GM - zeta*zeta)
    x=x+(gm-zeta)*v*v
        y = (v*v - (vline+delta)*(vline+delta) - gm*gm + zeta*zeta)
    z=y*y + 4.0*v*v*GM*GM
        FJK = 29.97925*2.0*v*v*x/3.14159/z/vline/vline
     RETURN
      END
C
      SUBROUTINE ABSBR_h2s2(AON,T,P,VNU)
C
C THIS SUBROUTINE RETURNS QUANTITY AON, WHICH WHEN MULTIPLIED
C BY THE VOLUME MIXING RATIO, GIVES THE ABSORPTION COEFFICIENT
C OF H2S, IN UNITS OF 1/cm
C
C DEFINITIONS
   AON ABSORPTION COEFFICIENT/VOLUME MIXING RATIO 1/cm
C
C
       TEMPERATURE
                                     K
C
           PRESSUME
                                   BARS
C
       FREQUENCY
                               MHZ
C
 ADDITIONAL FUNCTIONS REQUIRED
   0L0G
           BASE TEN LOGARITHM OF THE RATIO OF PARTITION
C
C
        FUNCTION AT TEMPERATURE T TO PARTITION FUNCTION
C
       AT T=300K
C
    SHAPE(F,P) LINE SHAPE FACTOR
IMPLICIT NONE
        REAL*4 AON, T, P, F, V
    real*4 t3,gm,vline,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,deltot,PCO,PCO13,PHCN
        real*4 DL,FLC,DL2
    REAL*4 C, DV, VNU, coef, hck
    real*8 x,ajk,fjk,tcor
        INTEGER*2 N, NLINES
        PARAMETER (NLINES=311) !NUMBER OF SPECTRAL LINES
    real*4 F0(311),S0(311),EL(311),Q0(5),wgths(311)
    REAL*4 QLOG2, SHAPE2
        EXTERNAL QLOG2
```

```
DATA Q0 /2.711, 2.5244, 2.2619, 1.8164, 1.3778/ !LOG PF AT
            300,225,150,75,37.5K
        COMMON/h2s/F0,S0,EL,wqths
         COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
C
    C=2.997925E1
    coef=7.244E21
    hck=1.438396
                     !hc/k
    V=VNU*C
                         !FREQ IN GHZ
    T3=(296./T)**0.6667
    A0N=0 . 0
        DO N = 1, NLINES
    VLINE=F0(N)
                     !LINE CENTER IN GHZ
        TCOR = EXP(-(hck*EL(N)*(1./T - 3.3784E-03)))
    call sajh2s2(N,T,T3,GM)
        ajk=S0(N)*TCOR
        CALL SFJh2s2(N,T,V,VLINE,GM,FJK)
       AON=AON + ajk*FJK
       ENDD0
    x=(296.0/T)**3.5
    AON=7.244E21*x*AON*Ph2s/296.0 !2.4163E3 is from deboer
C
       RETURN
       END
\mathsf{C}
      SUBROUTINE SAJh2s2(N,T,T3,GM)
        REAL*4 T,P,F,V
    real*8 x,ajk,fjk,tcor
    real*4 T3,gm,vline,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,deltot,PC0,PC013,PHCN
    real*4 F0(311),S0(311),EL(311),wgths(311)
    integer*2 N
       COMMON/h2s/F0,S0,EL,wgths
        COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
C
      GM IS GAMMA(N) IN GHZ.
    GH2=1.96
    GHE=1.20
    GH2S=wqths(N)
      GM=(GH2*T3*PH2)+(GHE*T3*PHE)+(GH2S*T3*PH2S)
      RETURN
      END
C
      SUBROUTINE SFJh2s2(N,T,V,VLINE,GM,FJK)
        REAL*4 T,P,F,V
    real*8 x,ajk,fjk,tcor,y,z
        real*4 t3,gm,vline,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,deltot,PCO,PCO13,PHCN
    real*4 F0(311),S0(311),EL(311),wgths(311)
    integer*2 N
        COMMON/h2s/F0,S0,EL,wgths
        COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
C
    zeta=qm
    delta=1.28*ph2s
      FJK IS THE BEN REUVEN LINE SHAPE, FROM DEBOER AND STEFFES
C
    x = (gm+zeta)*((vline+delta)*(vline+delta) + GM*GM - zeta*zeta)
    x=x+(qm-zeta)*v*v
```

```
y = (v*v - (vline+delta)*(vline+delta) - gm*gm + zeta*zeta)
    z=y*y + 4.0*v*v*GM*GM
        FJK = 29.97925*2.0*v*v*x/3.14159/z/vline/vline
      RETURN
      END
C
      SUBROUTINE NH3JOIN(VNU,T,ANH3)
C
      CALCULATE THE AMMONIA ABSORPTION COEFFICIENT (CM-1)
      USED BY GULKIS AND POYNTER (1972), AND GIVEN IN BERGE
C
C
      AND GULKIS, GEHRELS JUPITER BOOK, PAGE 675.
      DIMENSION VKAK(16,16), DVKAK(16,16)
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
      COMMON/U2/C1, C2, C3, C4
      COMMON/U3/VKAK, DVKAK
      COMMON/U4/T1, T2, T3, T4, T5, T6, T7, T8, T9, T10, T11
      COMMON/V2/PI, HB, C, TPC, BOLC, AW, DEN, BAR, G
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
C
    VKAK(J,K)=CENTER FREQU FOR THE (J,K) TRANSITION ACCORDING TO THE
  TABULATED VALUES OF POYNTER AND GULKIS (1975).
  DVKAK(J,K) ARE THE CORRESPONDING SELF BROADENED LINE WIDTHS
  AS TABULATED BY THEM (RESP. GHZ, MHZ/TORR)
C
C
      CHANGE VNU(CM-1) TO GHZ.
      V=(C*VNU)
   CHANGE PRESSURES FROM BARS TO ATM, AND BACK AT END OF SUBROUTINE; MARCH89
    PH2=PH2/1.013
    PHE=PHE/1.013
    PH20=PH20/1.013
    PNH3=PNH3/1.013
    PCH4=PCH4/1.013
      CORRECTION FRACTOR IS GIVEN ON PAGE 676 OF THE BERGE AND GULKIS
C
C
      PAPER, GEHRELS JUPITER BOOK.
C
    CORR=1.0
      IF (NOPR .EQ. 1) WRITE(6,5) V,CORR
     FORMAT(2X, 'NH3 V(GHZ), CORR', 2X, 1P, 2(2X, E10.3))
      IF (NOPR .EQ. 1) WRITE(6,6) C1,C2,C3,C4
      FORMAT(2X, 'CONSTANTS C1, C2, C3, C4', 2X, 1P, 4(2X, E10.3))
C
      SUM=0.0
C
      CALCULATE THE INVERSION LINE WAVENUMBERS
C
      FOR NH3, FROM THE DATA IN POYNTER AND KAKAR
C
      AP J SUPPL SERIES, VOL 29, P87, 1975.
C
      J AND K ARE QUANTUM NUMBERS.
C
C
      NLINE ARE THE NUMBER OF LINES TO BE CALCULATED.
      NLINE=16
C
      D0 10 J=1,NLINE
      L1=J*(J+1)
      L3=0
      D0 20 K=1,J
      L2=K*K
      L3=L3+1
C
      SB IS THE SELF BRODENED LINE WIDTH OF NH3 IN MHZ/TORR.
C
      C4 IS GIVEN IN THE BLOCK DATA STATEMENT.
```

```
C
      SK IS 1.5 OR 3, DEPENDING UPON THE K VALUE.
      SK=1.5
      IF (L3 .NE. 3) G0 T0 21
      SK=3.0
      L3=0
  21 VLINE=VKAK(J,K)
       SB=DVKAK(J,K)
      IF (NOPR .EQ. 1) WRITE(6,22) J,K,SK,VLINE,SB
     FORMAT(2X, 'J, K, SK, VLINE(GHZ), SB(MHZ/TORR)',/,
     1 2X,I3,2X,I3,2X,F4.1,2X,1P,2(2X,E10.3))
C
      CALCULATE A(J,K).
      CALL SAJKJOIN(T, VLINE, J, L1, L2, SK, SB, GM, GH2, AJK)
      IF(J.EQ.1.AND.K.EQ.1) ZZV=2.0*AJK*EXP(23.3/T)/3.0
C
      FJK IS THE FREQUENCY DEPENDENT LINE SHAPE OF BEN REUVEN.
      CALL SFJKJOIN(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
      SB=AJK*FJK
      SUM=SUM+(CORR*AJK*FJK)
      IF (NOPR .EQ. 1) WRITE(6,23) AJK,FJK,SB
  23 FORMAT(2X,'AJK,FJK,(AJK*FJK)',2X,1P,3(2X,E10.3))
  20 CONTINUE
  10 CONTINUE
\mathsf{C}
  CALCULATE ROTATION LINE J=0;K=0
C
  VLINE=5.73E5 MHZ; CHOSE DVLINE=14. MHZ
      L1=2
      L2 = 1
      VLINE=5.73E2
      SB=14.0
      GM=(1.89*T3*PH2)+(0.75*T3*PHE)+(0.60*T2*SB*PNH3)
      CALL SFJKJOIN(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
      ZZV1=ZZV*7.0E-2*T*(1.0-EXP(-28.6/T))
      ZZV2=ZZV*5.25E-2*T*EXP(-23.3/T)*(1.0-EXP(-57.2/T))
      ZZV3=ZZV*7.50E-2*T*EXP(-28.6/T)*(1.0-EXP(-57.2/T))
      SUM=SUM+(ZZV1*FJK)
      L1=2
      L2=1
      VLINE=1169.0
      SB=23.8
      CALL SFJKJOIN(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
C
    multiply fjk by 4.0 to test sensitivity
C
    FJK=4.0*FJK
      SUM=SUM+(ZZV2*FJK)+(ZZV3*FJK)
C
C
C
      THE AMMONIA ABSORPTION COEFFICIENT IS IN CM-1.
      ANH3=SUM
       IF (NOPR .EQ. 1) WRITE(6,30) CORR,SUM,ANH3
C
  30 FORMAT(2X, 'CORR, SUM, ANH3', 2X, 1P, 3(2X, E10.3), /)
C
    PH2=PH2*1.013
    PHE=PHE*1.013
    PH20=PH20*1.013
    PNH3=PNH3*1.013
    PCH4=PCH4*1.013
      RETURN
```

```
C
      SUBROUTINE SAJKJOIN(T, VLINE, J, L1, L2, SK, SB, GM, GH2, AJK)
    real*4 t3,gm,vline,fjk,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,deltot,PC0,PC013,PHCN
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
      COMMON/U4/T1,T2,T3,T4,T5,T6,T7,T8,T9,T10,T11
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
      GM IS GAMMA(J,K) IN GHZ.
      GM=(1.89*T3*PH2)+(0.75*T3*PHE)+(0.60*T2*SB*PNH3)
      A1=(2.*J+1.)*FLOAT(L2)/FLOAT(L1)
      A5=((2.98*L1)-(1.09*L2))*T7
      AJK IS A(J,K), SEE EQN 24 OF BERGE AND GULKIS PAPER, GEHRELS
C
\mathsf{C}
      JUPITER BOOK, PAGE 675.
      CALL EX(A5,A10)
      AJK=1.23E3*A1*SK*PNH3*T6*A10
       IF (NOPR .EQ. 1) WRITE(6,10) GM,A1,A5,AJK
  10 FORMAT(2X, 'GM, A1, A5, AJK', 2X, 1P, 4(2X, E10.3))
      RETURN
      END
C
      SUBROUTINE SFJKJOIN(T,L1,L2,V,VLINE,GM,GH2,SB,FJK)
    real*4 t3,gm,vline,fjk,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,deltot,PC0,PC013,PHCN
      COMMON/V1/A1, A2, A3, A4, A5, A6, A7, A8, A9, A10
      COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
       COMMON/U4/T1, T2, T3, T4, T5, T6, T7, T8, T9, T10, T11
      COMMON/V4/NOPR, NOPR2, NOPR3, NOPR4
C
C
      FJK IS THE BEN REUVEN LINE SHAPE, SEE EQN 25 OF BERGE AND
C
      GULKIS ARTICLE, PAGE 675 OF GEHRELS JUPITER BOOK.
C
      X=PH2+PHE+PH20+PNH3+PCH4
C
   IF X.LT.1 ATM VAN VLECK-WEISKOPF LINE SHAPE CAN BE
  USED; HOWEVER, BEN REUVEN IS STILL BETTER OUT IN THE
C
C
  WINGS OF THE AMMONIA LINES.
      A2 IS THE COUPLING FACTOR ZETA(J,K) IN GHZ.
С
    G0T0 500
C
   for Gross lineshape, goto 500 (first VVW, then gross
      IF(VLINE.GT.100.0) GOTO 500
      A2=(1.35*PH2*T3)+(0.30*T2*PNH3*SB)+
     1 (0<sub>2</sub>*T3*PHE)
C
      A4 IS THE PRESSURE SHIFT DEL (GHZ)
    A4=-0.45*PNH3
C
      SQUARES OF VARIOUS TERMS.
      A5=(VLINE+A4)*(VLINE+A4)
C
       A5=VLINE*VLINE
      A6=GM*GM
      A7 = A2 * A2
      A8=V*V
C
      NUMERATOR
      A9=((GM-A2)*A8)+((GM+A2)*(A5+A6-A7))
      A10=A8-A5-A6+A7
C
      DENOMINATOR
      A10=(A10*A10)+(4.0*A8*A6)
C
      THE BEN REUVEN LINE SHAPE.
      FJK=2.0*A8*(A9/A10)
```

```
IF (NOPR .EQ. 1) WRITE(6,10) A2,A4,A5,A6,A7,A8,A9,
    1 A10, FJK
    goto 501
  500 CONTINUE
 THE VAN VLECK-WEISKOPF LINE SHAPE
C
     Z1=0.755*PH2+0.231*PHE
     Z2=FLOAT(L2)/FLOAT(L1)
     Z1=Z1*(Z2**0.3333)
     Z2=SQRT(Z2)*PNH3*6.23
     Z=(Z1+Z2)*1.0E3/T
     X1=V*V*Z
     X2=(V-VLINE)*(V-VLINE)
     X3=(V+VLINE)*(V+VLINE)
     FJK=X1/(X2+Z*Z)+X1/(X3+Z*Z)
      G0T0 501
C
  THE GROSS LINE SHAPE
       X2=(-V*V+VLINE*VLINE)
       X2 = X2 * X2
       if(vline.lt.1000.) Z1=123.3551*PH2+60.7441*PHE
    if(vline.ge.1000.) z1=2.885*PH2+1.00*PHE
\mathsf{C}
      Z2=FLOAT(L2)/FLOAT(L1)
    Z2=300./T
     Z1=Z1*(Z2**0.3)
    if(vline.lt.1000.) Z=154.1939*PNH3*Z2
    if(vline.ge.1000.) Z=22.989*PNH3*Z2
      Z=Z1+Z
       X3=Z*Z*4*V*V
       X1=v*VLINE*Z*4.
       X1=X1*V*VLINE
        FJK=X1/(X2+X3)
 501
       FFF=FJK
  10 FORMAT(2X, 'A2, A4, A5, A6, A7, A8, A9, A10, FJK', /,
    1 2X,1P,9(2X,E10.3))
     RETURN
     END
function absco(j,fr0,f,p,t,wtmol)
C
     this function calculates the absorption coeff of CO
C
  f, fr0 in Hz; p in bar, but the routine expects:
C f, fr0 in Hz, and p in microbar
     normalized by the CO mixing ratio (xco)
     the formulae were taken from Clancy's PhD thesis
C
C
     if(j.eq.0) then
         absco = 0.0
          return
     endif
     p=p*1.0E6
C
     pi = 3.14159265308
C
     offset frequency from line center
C
     fpm = f - fr0
      fpp = f + fr0
```

```
C
      pressure broadened line width in Hz
C
 this is the same as in the SHAPE function (IdP)
 note that microbar = dynes/cm^2
C + this number should be the half-width of CO in H2-broadened gas
 which is about 0.071 cm-1/atm
C or 2100 instead of old 3300
C for HCN that would be 2 times larger
 the following line is pretty consistent with the Menger et al 2000
C results for 2-1 transition
       wp = 2100.0 * p * (300.0/t)**0.7
      doppler width (e-folding ... not FWHM) in Hz
C
      wd = fr0 / 3.00e10 * sqrt(2.0*1.38e-16*t/wtmol)
C in his thesis wd is smaller by f. 1.20
C
      number density of atmosphere (ideal gas p = dkt)
С
      d = p / 1.38e-16 / t
C
      a = wp/wd
      x = fpm/wd
C
C
      use the Voigt function line shape
      shape = voigt(x,a) / wd
C
      VanVleck-Weisskopf shape (for future use at large offsets)
C
      vvwshape =
     & 1.0 / pi * f*f/fr0/fr0 * (wp/(fpm*fpm+wp*wp)+wp/(fpp*fpp+wp*wp))
C
C
      Lorentz Shape
      vl = wp / pi / (fpm*fpm + wp*wp)
C
C
        shape=shape*1.0E6
        v1=v1*1.0E6
        vvwshape=vvwshape*1.0E6
        fshape=shape
        if(v1.qt.fshape) fshape=v1
   The following is correct, I think -- 15 April 2011
C
        if(vvwshape.gt.fshape) fshape=vvwshape
   I think pin = p/1.013E6 so P in is in atm; it was written as *1.013E6
        pin = p / 1.01325E6
         xshape=fshape
C
          xshape=vvwshape
        call absx_coline(j,aon,t,pin,f,xshape)
 xshape: p in millibar, f in MHz (p/f); I don't think we need to use this
C vvwshape: p in microbar; f in Hz
C We multiply shape by 1E6 to correct answer to come out in bars and MHz
C
         absco = aon
        write(6,1003) j,aon,absco,shape,f,x,a,wp,wd
C
       write(6,1004) p,xshape,vvwshape,v1,shape
c 1003 format(I5,8E12.4)
c 1004 format(5E12.4)
      p=p/1.0E6
   if other gases are considered, add them here as well I guess; e.g.
```

```
C
  SO2, H2SO4 vapor. VVW line shape is probably OK for lower layers
     return
     end
function absco13(j,fr0,f,p,t,wtmol)
C
     this function calculates the absorption coeff of CO
C
C
  f, fr0 in Hz; p in bar, but the routine expects:
C f, fr0 in Hz, and p in microbar
     normalized by the CO mixing ratio (xco)
     the formulae were taken from Clancy's PhD thesis
C
C
     if(i.eq.0) then
         absco13 = 0.0
         return
     endif
     p=p*1.0E6
C
     pi = 3.14159265308
     offset frequency from line center
C
     fpm = f - fr0
     fpp = f + fr0
C
     pressure broadened line width; take same parameters as for CO12
C
     wp = 2100.0 * p * (300.0/t)**0.7
     doppler width (e-folding ... not FWHM)
C
     wd = fr0 / 3.00e10 * sqrt(2.0*1.38e-16*t/wtmol)
     number density of atmosphere (ideal gas p = dkt)
C
     d = p / 1.38e-16 / t
C
     a = wp/wd
     x = fpm/wd
C
     use the Voigt function line shape
C
     shape = voigt(x,a) / wd
C
C
     VanVleck-Weisskopf shape (for future use at large offsets)
     vvwshape =
    & 1.0 / pi * f*f/fr0/fr0 * (wp/(fpm*fpm+wp*wp)+wp/(fpp*fpp+wp*wp))
C
     Lorentz Shape
C
     vl = wp / pi / (fpm*fpm + wp*wp)
C
       shape=shape*1.0E6
       v1=v1*1.0E6
       vvwshape=vvwshape*1.0E6
       fshape=shape
       if(v1.gt.fshape) fshape=v1
  The following is correct, I think -- 15 April 2011
       if(vvwshape.gt.fshape) fshape=vvwshape
  I think pin = p/1.013E6 so P in is in atm; it was written as *1.013E6
```

```
pin = p / 1.01325E6
       xshape=fshape
C
         xshape=shape
        call absx_co13line(j,aon,t,pin,f,xshape)
C xshape: p in millibar, f in MHz (p/f)
C vvwshape: p in microbar; f in Hz
C We multiply shape by 1E6 to correct answer to come out in bars and MHz
      absco13 = aon
       write(6,1003) j,aon,absco13,shape,f,x,a,wp,wd
C
       write(6,1004) p,xshape,vvwshape,v1,shape
C
c 1003 format(I5,8E12.4)
c 1004 format(5E12.4)
     p=p/1.0E6
 if other gases are considered, add them here as well I guess; e.g.
 SO2, H2SO4 vapor. VVW line shape is probably OK for lower layers
C
     return
     end
function voigt(xx,a)
  Need to know: T, P, But it is all passed along in xx, a
C
C
     voigt function subroutine obtained from D. van Blerkom
С
C
     This is not the ONE TRUE Voigt funtion ...
C
     The normal definition of voigt function is:
C
            H(a,b) = a/pi * INTEGRAL ( exp(-y*y) / ((b-y)**2 + a**2) dy )
C
           where a = wp / wd
C
                 b = (f - fr0) / wd
C
     We compute:
C
C
            voigt = 1/sqrt(pi) * H(a,b)
C
     Function requires further normalization by doppler width.
C
C
     dimension c(31)
     data ic/0/,epsiln/2.5e-12/
C
C
       write(6,1003) xx,a
C
C 1003 format(2E12.4)
     if(ic.ne.0) goto 20
  10 k=-16
     pie = 3.1415926535898
     sqp = sqrt(pie)
     sqpp = 1./sqp
     pie3 = 3.0*pie
     pie32= pie3*pie
     piec = 0.5/(pie*sqp)
     do 15 n=1,31
        k=k+1
        c(n)=piec*exp(-float(k*k)/9.)
  15
     ic=1
C
  20 x = abs(xx)
C
```

```
if(a.gt.1000.0) go to 40
     if(a.lt.0.001) go to 45
C
     a1=3.0*a
     a2=a*a
     if(a.lt.0.1) go to 25
     x1=-pie3*a
     voigt=0.0
     rz=0.0
     b2=0.0
     if(x1.lt.-675.0) go to 30
     y1=pie3*x
     e1=exp(x1)
     rz=e1*cos(y1)
     b2=-e1*sin(y1)
     go to 30
  25 x1=pie3*x
     y1=pie3*a
     e1=exp(y1)
     e2=1.0/e1
     rz=0.5*(e1+e2)*cos(x1)
     b2=-0.5*(e2-e1)*sin(x1)
     qin=a2-x*x
     voigt=0.0
     if(qin.gt.-675.0) voigt=sqpp*exp(qin)*cos(2.0*a*x)
  30 b1=(1.0-rz)*a*1.5
     s=-8.0-1.5*x
     t=s*s+2.25*a2
     c2=-a*pie32
     do 35 n=1,31
        t=t+s+0.25
        s=s+0.5
        b1=a1-b1
        b2 = -b2
        con=c2
        if(t.gt.epsiln) con=(b1+b2*s)/t
  35
        voigt=voigt+con*c(n)
     return
C
     large a limit
C
C
  40 qin=x*x/a+a
     voigt=1.0/(pie*qin)
     return
C
     small a limit
С
C
  45 voigt=0.0
     x2=x*x
     if(x2.lt.675.0) voigt=sqpp*exp(-x2)
     return
     end
SUBROUTINE ABSX_C013(AON,T,P,vnu)
  Note: This subroutine is for 13CO.
```

```
C THIS SUBROUTINE RETURNS QUANTITY AON, WHICH WHEN MULTIPLIED
C BY THE VOLUME MIXING RATIO, GIVES THE ABSORPTION COEFFICIENT
C OF CO13, IN UNITS OF 1/KM
C
C DEFINITIONS
C
   AON ABSORPTION COEFFICIENT/VOLUME MIXING RATIO 1/KM
C
       TEMPERATURE
                                     K
C
           PRESSUME
                                   MILLIBARS
C
      FREQUENCY
                               MHZ
C
C ADDITIONAL FUNCTIONS REQUIRED
           BASE TEN LOGARITHM OF THE RATIO OF PARTITION
C
C
        FUNCTION AT TEMPERATURE T TO PARTITION FUNCTION
C
       AT T=300K
C
   SHAPE(F,P) LINE SHAPE FACTOR
C
  the QO, FO, SO, EL data are from Poynter and Picket, JPL catalogue
IMPLICIT NONE
        REAL*4 AON, T, P, F, V, wtmol
    REAL*4 QLOG2, SHAPE2, SHAPE
    REAL*4 C, ZCO, DV, VNU
        EXTERNAL QLOG2, SHAPE2, SHAPE
        INTEGER*2 N, NLINES
        PARAMETER (NLINES=9)
                              !NUMBER OF SPECTRAL LINES
        REAL*4 DF, DFC, DFD, EL(NLINES), Q0(5), F0(NLINES), FS, FV
        REAL*4 S0(NLINES), S1(NLINES), TCOR, X, X1, Y
        DATA Q0 /2.0564, 1.9317, 1.7565, 1.4579, 1.1620/ !LOG PF AT
            300,225,150,75,37.5K
        DATA F0(1),S0(1),EL(1)
    & /110201.3672, -5.0614, 0.0000/
       DATA F0(2),S0(2),EL(2)
      /220398.7031, -4.1698, 3.6759/
       DATA F0(3), S0(3), EL(3)
    & /330588.0937, -3.6607, 11.0276/
       DATA F0(4),S0(4),EL(4)
    & /440765.4375, -3.3127, 22.0549/
       DATA F0(5),S0(5),EL(5)
    & /550926.7500, -3.0564, 36.7572/
       DATA F0(6),S0(6),EL(6)
    & /661068.0625, -2.8609, 55.1341/
       DATA F0(7), S0(7), EL(7)
    & /771185.3125, -2.7098, 77.1850/
       DATA F0(8),S0(8),EL(8)
       /881274.5625, -2.5933, 102.9090/
       DATA F0(9),S0(9),EL(9)
    & /991331.8125, -2.5049, 132.3051/
    P=P*1000.0/1.013
                          !MILLIBAR
    C=2.99791E1
    F=vnu*C*1.0E3
                                  !FREQU IN MHZ
    A0N=0.0
      DO N = 1, NLINES
      TCOR = EXP(1.4388*EL(N)*(3.33356E-03-1./T))
    \{ *(1.-EXP(-4.7993E-05*F0(N)/T) \}
```

```
\{(1,-EXP(-1.5998E-07*F0(N)))\}
      AON=AON+TCOR*SHAPE(P,T,F,F0(N))*(10.**S0(N))
      enddo
      AON=AON*10**(Q0(1)-QLOG2(Q0,T)) !T DEPENDENCE OF PARTITION FUNCTION
      AON=0.72435E5*AON*P/T !YIELD ABSORPTION COEFFICIENT IN 1/cm
C
                                  WHEN MULTIPLIED BY MIXING RATIO
C
   P=P*1.013/1000.0
      RETURN
      END
C
SUBROUTINE ABSX_CO13line(n,AON,T,P,f,xshape)
C
 Note: This subroutine is for 13CO.
C
C
C THIS SUBROUTINE RETURNS QUANTITY AON, WHICH WHEN MULTIPLIED
C BY THE VOLUME MIXING RATIO, GIVES THE ABSORPTION COEFFICIENT
C OF CO13, IN UNITS OF 1/KM
C
C DEFINITIONS
   AON ABSORPTION COEFFICIENT/VOLUME MIXING RATIO 1/KM
C
C
       TEMPERATURE
                                   K
C
           PRESSUME
                                 MILLIBARS
C
   F FREQUENCY
                              MHZ
C ADDITIONAL FUNCTIONS REQUIRED
          BASE TEN LOGARITHM OF THE RATIO OF PARTITION
C
C
       FUNCTION AT TEMPERATURE T TO PARTITION FUNCTION
C
       AT T=300K
C
   SHAPE(F,P) LINE SHAPE FACTOR
C
  the QO, FO, SO, EL data are from Poynter and Picket, JPL catalogue
IMPLICIT NONE
       REAL*4 AON, T, P, F, V, wtmol
   REAL*4 QLOG2,SHAPE2,xshape
   REAL*4 C, ZCO, DV, VNU
       EXTERNAL QLOG2, SHAPE2
       INTEGER*2 N, NLINES
       PARAMETER (NLINES=9)
                             !NUMBER OF SPECTRAL LINES
       REAL*4 DF, DFC, DFD, EL (NLINES), Q0(5), F0(NLINES), FS, FV
       REAL*4 S0(NLINES), S1(NLINES), TCOR, X, X1, Y
       DATA Q0 /2.0564, 1.9318, 1.7565, 1.4579, 1.1620/ !LOG PF AT
           300,225,150,75, 37.5K
        DATA F0(1),S0(1),EL(1)
    & /110201.3672, -5.0614, 0.0000/
       DATA F0(2),S0(2),EL(2)
    & /220398.7031, -4.1698, 3.6759/
       DATA F0(3),S0(3),EL(3)
    & /330588.0937, -3.6607, 11.0276/
       DATA F0(4),S0(4),EL(4)
    & /440765.4375, -3.3127, 22.0549/
       DATA F0(5),S0(5),EL(5)
    & /550926.7500, -3.0564, 36.7572/
```

```
DATA F0(6),S0(6),EL(6)
    & /661068.0625, -2.8609, 55.1341/
       DATA F0(7), S0(7), EL(7)
    & /771185.3125, -2.7098, 77.1850/
       DATA F0(8),S0(8),EL(8)
    & /881274.5625, -2.5933, 102.9090/
       DATA F0(9),S0(9),EL(9)
    & /991331.8125, -2.5049, 132.3051/
    P=P*1000.0/1.013
                          !MILLIBAR
    C=2.99791E1
       F=F/1.0E6
                                  !FREQU IN MHZ
C
    F=vnu*C*1.0E3
    A0N=0.0
\mathsf{C}
        xshape=SHAPE2(P,T,F,F0(N))
      TCOR = EXP(1.4388*EL(N)*(3.33356E-03-1./T))
     *(1.-EXP(-4.7993E-05*F0(N)/T)) 
    \{(1.-EXP(-1.5998E-07*F0(N)))\}
      AON=AON+TCOR*xshape*(10***S0(N))
      AON=AON*10**(Q0(1)-QLOG2(Q0,T)) !T DEPENDENCE OF PARTITION FUNCTION
      AON=0.72435E5*AON*P/T
                               !YIELD ABSORPTION COEFFICIENT IN 1/cm
C
                                    WHEN MULTIPLIED BY MIXING RATIO
\mathsf{C}
    P=P*1.013/1000.0
       F=F*1.0E6
      RETURN
      END
C
C*******
C
      SUBROUTINE ABSX_CO(AON,T,P,vnu)
C
C THIS SUBROUTINE RETURNS QUANTITY AON, WHICH WHEN MULTIPLIED
C BY THE VOLUME MIXING RATIO, GIVES THE ABSORPTION COEFFICIENT
C OF CO, IN UNITS OF 1/cm
C
C DEFINITIONS
C
   AON ABSORPTION COEFFICIENT/VOLUME MIXING RATIO
C
       TEMPERATURE
C
           PRESSUME
                                   atm
C
       FREQUENCY
                               MHZ
C
C
 ADDITIONAL FUNCTIONS REQUIRED
C
           BASE TEN LOGARITHM OF THE RATIO OF PARTITION
C
        FUNCTION AT TEMPERATURE T TO PARTITION FUNCTION
C
       AT T=300K
C
    SHAPE(F,P) LINE SHAPE FACTOR
C
C
  the QO, FO, SO, EL data are from Poynter and Picket, JPL catalogue
IMPLICIT NONE
        REAL*4 AON, T, P, F, V, wtmol
    REAL*4 QLOG2, SHAPE2, SHAPE
    REAL*4 C, ZCO, DV, VNU
        EXTERNAL QLOG2, SHAPE2, SHAPE
        INTEGER*2 N, NLINES
```

```
PARAMETER (NLINES=8)
                               !NUMBER OF SPECTRAL LINES
        REAL*4 DF, DFC, DFD, EL(NLINES), Q0(5), F0(NLINES), FS, FV
        REAL*4 S0(NLINES), S1(NLINES), TCOR, X, X1, Y
        DATA Q0 /2.0369, 1.9123, 1.7370, 1.4386, 1.1429/ !LOG PF AT
            300,225,150,75,37.5K
        DATA F0(1),S0(1),EL(1)
     & /115271.2018, -5.0105, 0.0000/
        DATA F0(2),S0(2),EL(2)
     & /230538.0000, -4.1197, 3.8450/
        DATA F0(3),S0(3),EL(3)
       /345795.9899, -3.6118, 11.5350/
        DATA F0(4),S0(4),EL(4)
     & /461040.7682, -3.2657, 23.0695/
        DATA F0(5),S0(5),EL(5)
     & /576267.9305, -3.0118, 38.4481/
        DATA F0(6),S0(6),EL(6)
     & /691473.0763, -2.8193, 57.6704/
        DATA F0(7), S0(7), EL(7)
     & /806651.8060, -2.6715, 80.7354/
        DATA F0(8),S0(8),EL(8)
     & /921799.7000, -2.5590, 107.6424/
                           !MILLIBAR
    P=P*1000.0/1.013
    C=2.99791E1
    F=VNU*C*1.0E3
                                   !FREQU IN MHZ
        A0N=0.0
       DO N = 1, NLINES
       TCOR = EXP(1.4384*EL(N)*(3.3784E-03-1./T))
      *(1.-EXP(-4.7993E-05*F0(N)/T)) 
     \{(1.-EXP(-1.5998E-07*F0(N)))\}
       AON=AON+TCOR*SHAPE(P,T,F,F0(N))*(10.**S0(N))
       enddo
       AON=AON*10**(Q0(1)-QLOG2(Q0,T)) !T DEPENDENCE OF PARTITION FUNCTION
       AON=0.72435E5*AON*P/T
                                   !YIELD ABSORPTION COEFFICIENT IN 1/cm
C
                                    WHEN MULTIPLIED BY MIXING RATIO
       P=P*1.013/1000.0
C
       write(6,1003) n,aon,xz,xz2
 1003
       format(I5,3E12.4)
C
       RETURN
       END
C
C*******
C
       SUBROUTINE ABSX_COline(n,AON,T,P,f,xshape)
C THIS SUBROUTINE RETURNS QUANTITY AON, WHICH WHEN MULTIPLIED
C BY THE VOLUME MIXING RATIO, GIVES THE ABSORPTION COEFFICIENT
C OF CO, IN UNITS OF 1/cm
C
C DEFINITIONS
C
    AON ABSORPTION COEFFICIENT/VOLUME MIXING RATIO
                                                      1/cm
C
       TEMPERATURE
                                      K
C
            PRESSUME
                                    atm
    F
                                MHZ
C
       FREQUENCY
C ADDITIONAL FUNCTIONS REQUIRED
```

```
C
           BASE TEN LOGARITHM OF THE RATIO OF PARTITION
C
       FUNCTION AT TEMPERATURE T TO PARTITION FUNCTION
C
       AT T=300K
C
    SHAPE(F,P) LINE SHAPE FACTOR
C
C
  the QO, FO, SO, EL data are from Poynter and Picket, JPL catalogue
IMPLICIT NONE
        REAL*4 AON, T, P, F, V, wtmol
    REAL*4 QLOG2, SHAPE2, SHAPE, xshape
    REAL*4 C, ZCO, DV, VNU
        EXTERNAL QLOG2, SHAPE2, SHAPE
        INTEGER*2 N, NLINES
        PARAMETER (NLINES=8)
                             !NUMBER OF SPECTRAL LINES
        REAL*4 DF, DFC, DFD, EL(NLINES), Q0(5), F0(NLINES), FS, FV
        REAL*4 S0(NLINES), S1(NLINES), TCOR, X, X1, Y
        DATA Q0 /2.0369, 1.9123, 1.7370, 1.4386, 1.1429/ !LOG PF AT
            300,225,150,75, 37.5K
        DATA F0(1),S0(1),EL(1)
    & /115271.2018, -5.0105, 0.0000/
       DATA F0(2), S0(2), EL(2)
    & /230538.0000, -4.1197, 3.8450/
       DATA F0(3),S0(3),EL(3)
    & /345795.9899, -3.6118, 11.5350/
       DATA F0(4),S0(4),EL(4)
    & /461040.7682, -3.2657, 23.0695/
       DATA F0(5),S0(5),EL(5)
    & /576267.9305, -3.0118, 38.4481/
       DATA F0(6),S0(6),EL(6)
    & /691473.0763, -2.8193, 57.6704/
       DATA F0(7), S0(7), EL(7)
    & /806651.8060, -2.6715, 80.7354/
       DATA F0(8),S0(8),EL(8)
    & /921799.7000, -2.5590, 107.6424/
    P=P*1000.0/1.013
                          !MILLIBAR
    C=2.99791E1
    F=f/1.0E6
                          !FREO IN MHz
C
                                  !FREQU IN MHZ
    F=VNU*C*1.0E3
       A0N=0.0
C
        xshape=SHAPE(P,T,F,F0(N))
      TCOR = EXP(1.4384*EL(N)*(3.3784E-03-1./T))
     *(1.-EXP(-4.7993E-05*F0(N)/T)) 
    \{(1.-EXP(-1.5998E-07*F0(N)))\}
      AON=AON+TCOR*xshape*(10.**SO(N))
      AON=AON*10**(Q0(1)-QLOG2(Q0,T)) !T DEPENDENCE OF PARTITION FUNCTION
      AON=0.72435E5*AON*P/T
                                  !YIELD ABSORPTION COEFFICIENT IN 1/cm
                                   WHEN MULTIPLIED BY MIXING RATIO
C
      P=P*1.013/1000.0
       write(6,1003) n,aon,xz,xz2
C 1003 format(I5,3E12.4)
      f=f*1.0E6
C
      RETURN
      END
C
```

```
SUBROUTINE ABSX_HCN(AON,T,P,vnu)
C
C THIS SUBROUTINE RETURNS QUANTITY AON, WHICH WHEN MULTIPLIED
C BY THE VOLUME MIXING RATIO, GIVES THE ABSORPTION COEFFICIENT
C OF Hcn, THIOFORMALDEHYDE, IN UNITS OF 1/KM
C
C DEFINITIONS
C
    AON ABSORPTION COEFFICIENT/VOLUME MIXING RATIO
                                                    1/CM
C
       TEMPERATURE
                                     K
C
           PRESSUME
                                   MILLIBARS
C
    F
       FREQUENCY
                               MHZ
C
 ADDITIONAL FUNCTIONS REQUIRED
C
           BASE TEN LOGARITHM OF THE RATIO OF PARTITION
       FUNCTION AT TEMPERATURE T TO PARTITION FUNCTION
C
C
       AT T=300K
C
    SHAPE(P,T,F,FLC) LINE SHAPE FACTOR
C
C
  the Q0, F0, S0, EL data are from Poynter and Picket, JPL catalogue
IMPLICIT NONE
        REAL*4 AON, T, P, F, V
    REAL*4 QLOG2, SHAPE
    REAL*4 C, ZHCN, DV, VNU
        EXTERNAL QLOG2, SHAPE
        INTEGER*2 N, NLINES
                                !NUMBER OF SPECTRAL LINES
        PARAMETER (NLINES=11)
        REAL*4 DF, DFC, DFD, EL(NLINES), Q0(5), F0(NLINES), FS, FV
        REAL*4 S0(NLINES), S1(NLINES), TCOR, X, X1, Y
        DATA Q0 /2.6277,2.5031,2.3276,2.0286,1.7317/ !LOG PF AT
            300,225,150,75,37.5K
        DATA F0(1),S0(1),EL(1)
    & /177262.0156, -3.8432,
                                2.9564/
       DATA F0(2),S0(2),EL(2)
                                  0.000/
    & /88630.4141, -2.9588,
       DATA F0(3),S0(3),EL(3)
    & /88631.8437, -2.7370,
                                  0.000/
        DATA F0(4),S0(4),EL(4)
    & /88633.9375, -3.4359,
                                  0.000/
        DATA F0(5), S0(5), EL(5)
    & /177259.6719, -2.6671,
                                  2.9564/
        DATA F0(6),S0(6),EL(6)
    & /177259.9219, -2.5422,
                                  2.9564/
       DATA F0(7), S0(7), EL(7)
    & /177261.1094, -2.1900,
                                  2.9564/
       DATA F0(8), S0(8), EL(8)
                                  2.9564/
    & /177261.2188, -1.9189,
       DATA F0(9),S0(9),EL(9)
    & /177263.4375, -2.6671,
                                  2.9564/
        DATA F0(10),S0(10),EL(10)
    & /265886.1875, -1.0751,
                                  8.8692/
       DATA F0(11),S0(11),EL(11)
    & /354505.4687, -0.7220, 17.7382/
    P=P*1000.0/1.013
                           !MILLIBAR
    C=2.99791E1
```

```
F=VNU*C*1.0E3
                                  !FREOU IN MHZ
   A0N=0.0
      D0 N = 1.NLINES
      TCOR = EXP(1.4388*EL(N)*(3.33356E-03-1./T))
    &*(1.-EXP(-4.7993E-05*F0(N)/T))
    \{(1.-EXP(-1.5998E-07*F0(N)))\}
      AON=AON + TCOR*SHAPE(P,T,F,F0(N))*(10.**S0(N))
      ENDD0
      AON=AON*10**(Q0(1)-QLOG2(Q0,T)) !T DEPENDENCE OF PARTITION FUNCTION
      AON=0.72435E5*AON*P/T !YIELD ABSORPTION COEFFICIENT IN 1/cm
C
                                    WHEN MULTIPLIED BY MIXING RATIO
\mathsf{C}
   P=P*1.013/1000.0
      RETURN
      END
C*****
\mathsf{C}
      FUNCTION SHAPE(P,T,F,FLC)
   implicit none
   REAL*4 SHAPE
   real*4 DL,P,T,F,FLC,DFC,DL2
 HPFW for collisional broadening from Clancy's thesis, for CO
C P in mbar, F and DFC in MHz when called from absx_co routines
C CO 2-1: 2.1*P*((300.0/T)**0.7)
       DFC=188.57*P/(T**0.75)
      DFC=2.1*P*((300.0/T)**0.7)
      DL=F-FLC
      DL2=F+FLC
      SHAPE=((F*F)/(FLC*FLC))*(DFC/(DL*DL+DFC*DFC)
    + +DFC/(DL2*DL2+DFC*DFC))
   SHAPE=SHAPE/3.1415926536 !DIVIDE BY PI
      RETURN
      END
C
      SUBROUTINE ABSBR_PH3(AON,T,P,VNU,zpph3)
C
C THIS SUBROUTINE RETURNS QUANTITY AON, WHICH WHEN MULTIPLIED
C BY THE VOLUME MIXING RATIO, GIVES THE ABSORPTION COEFFICIENT
C OF PH3, IN UNITS OF 1/cm
C
C DEFINITIONS
C
   AON ABSORPTION COEFFICIENT/VOLUME MIXING RATIO
                                                   1/cm
C
       TEMPERATURE
                                     K
C
           PRESSUME
                                   BARS
C
   F FREQUENCY
                               MHZ
C
C
 ADDITIONAL FUNCTIONS REQUIRED
C
          BASE TEN LOGARITHM OF THE RATIO OF PARTITION
C
       FUNCTION AT TEMPERATURE T TO PARTITION FUNCTION
C
       AT T=300K
   SHAPE(F,P) LINE SHAPE FACTOR
C
IMPLICIT NONE
       REAL*4 AON, T, P, F, V
    real*4 t3,gm,vline,ph2,phe,ph2s,pnh3,pch4,ph2o,pph3,
```

```
+ deltot, PCO, PCO13, PHCN
        real*4 DL,FLC,DL2,zpph3
    REAL*4 C,DV,VNU,coef,hck
    real*8 x,ajk,fjk,tcor
        INTEGER*2 N, NLINES
        PARAMETER (NLINES=320)
                                !NUMBER OF SPECTRAL LINES
    real*4 FP0(728),SP0(728),ELP(728),Q0(5)
    REAL*4 wgtS0(40), wgtFGB(40), wgtSB(40)
    REAL*4 QLOG2, SHAPE2
        EXTERNAL QLOG2
       DATA Q0 /2.905, 2.722, 2.4606, 2.0132, 1.5705/ !LOG PF AT 300,225,150,75,
           37.5K
        COMMON/ph3/FP0,SP0,ELP,wgtS0,wgtFGB,wgtSB
        COMMON/U1/PH2, PHE, PNH3, PH20, PCH4, PH2S, PPH3, PCO, PCO13, PHCN
    C=2.997925E1
    coef=7.244E21
    hck=1.438396
                    !hc/k
    V=VNU*C
                         !FREQ IN GHZ
    T3=(300/T)**0.6667
    AON=0.0
        DO N = 1, NLINES
    VLINE=FP0(N)
                        !LINE CENTER IN GHZ
        TCOR = EXP(-(hck*ELP(N)*(1./T - 3.3333E-03)))
    call sajph3(N,T,T3,GM)
    if (N.gt.40) goto 40
    ajk=SP0(N)*wgtS0(N)*TCOR
    goto 50
  40
        ajk=SP0(N)*TCOR
  50
       continue
        CALL SFJph3(N,T,V,VLINE,GM,FJK)
       AON=AON + ajk*FJK
       ENDD0
    x=(300.0/T)**3.5
    AON=2.4163E3*x*AON*PPh3/300.0 !2.4163E3 is from deboer
C
       RETURN
       END
C
```

C