

Stimulated Neutrino Transitions and Rabi Oscillations

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OUTLINE

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 - What are Neutrinos
 - Neutrino Oscillations
 - Why Do Neutrinos Oscillate
2. Matter Effect
 - Matter Interaction
 - MSW Effect
 - Solar Neutrino Problem
 - Stimulated Neutrino Oscillations
3. Stimulated Neutrino Flavor Transitions
 - Hamiltonian and Basis
 - Rabi Oscillations
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 - Basis and Formalism
5. Summary

OVERVIEW

Introduction

What are Neutrinos

Neutrino Oscillations

Why Do Neutrinos Oscillate

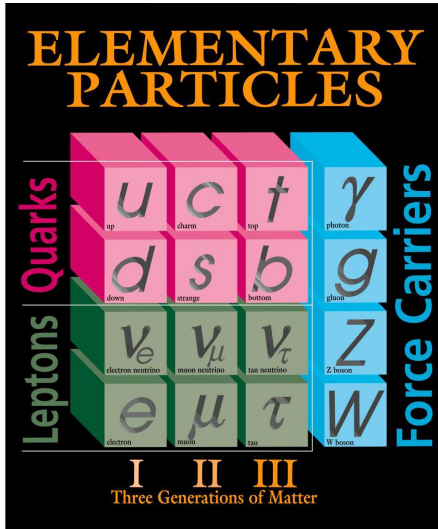
Matter Effect

Stimulated Neutrino Flavor Transitions

Jacobi-Anger Expansion

Summary

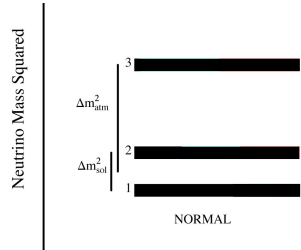
WHAT ARE NEUTRINOS?



Fermilab 95-759

Neutrinos are

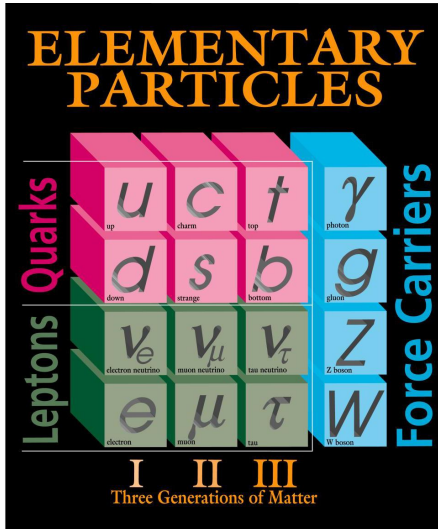
- fermions,
- electrically neutral,
- light.



Adapted from Olga Mena & Stephen Parke (2004)

Table of elementary particles. Source:
Fermilab

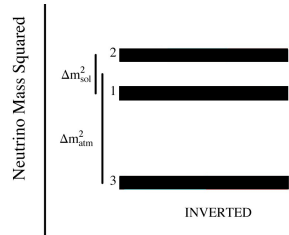
WHAT ARE NEUTRINOS?



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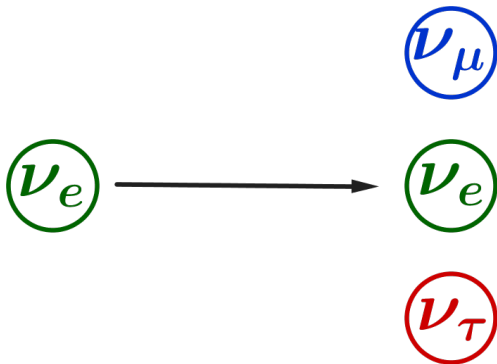
Table of elementary particles. Source:
Fermilab

WHAT IS NEUTRINO OSCILLATION?

Neutrino Oscillation

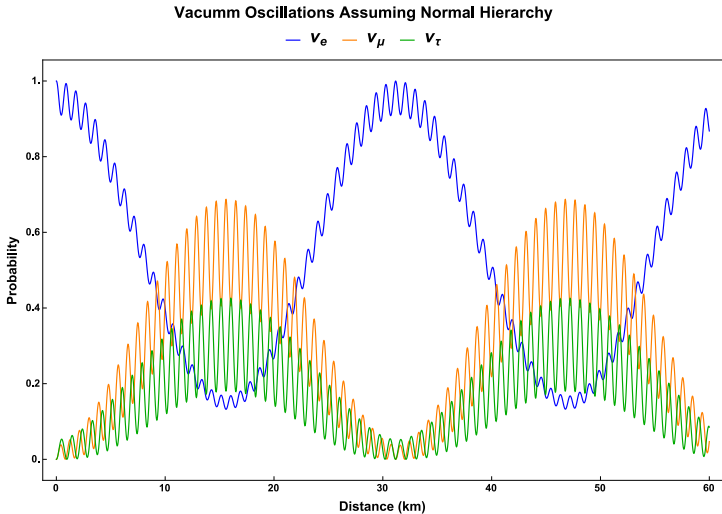
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Neutrino Flavor Conversion



Neutrino Oscillations

WHAT IS NEUTRINO OSCILLATION?



Probabilities of finding neutrinos to be in each flavor.

WHY DO NEUTRINOS OSCILLATE?

Equation of Motion

$$i\partial_x |\Psi\rangle = \hat{\mathbf{H}} |\Psi\rangle$$

- ▶ Basis: Hamiltonian diagonalized basis/mass basis/propagation basis, $\{|\nu_1\rangle, |\nu_2\rangle\}$.

▶

$$\mathbf{H} = -\frac{\omega_v}{2}\sigma_3, \quad \text{where } \omega_v = \frac{\delta m^2}{2E} = \frac{m_2^2 - m_1^2}{2E}.$$

- ▶ The system can be solved given initial condition of the amplitudes of the two eigenstates $(\langle\nu_1|\Psi(0)\rangle, \langle\nu_2|\Psi(0)\rangle)^T$,

$$\begin{pmatrix} \langle\nu_1|\Psi(x)\rangle \\ \langle\nu_2|\Psi(x)\rangle \end{pmatrix} = \begin{pmatrix} \langle\nu_1|\Psi(0)\rangle \exp(i\omega_v x/2) \\ \langle\nu_2|\Psi(0)\rangle \exp(-i\omega_v x/2) \end{pmatrix}$$

WHY DO NEUTRINOS OSCILLATE?

Flavor basis

Neutrino wave function in flavor basis $\{|\nu_e\rangle, |\nu_\mu\rangle\}$ is related to state in energy basis $\{|\nu_1\rangle, |\nu_2\rangle\}$ through

$$\begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \end{pmatrix} = \begin{pmatrix} \cos \theta_v & \sin \theta_v \\ -\sin \theta_v & \cos \theta_v \end{pmatrix} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \end{pmatrix}$$

θ_v : vacuum mixing angle

WHY DO NEUTRINOS OSCILLATE?

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θ_v : vacuum mixing angle

Hamiltonian H

Mass basis

$$\begin{aligned} & \frac{\omega_v}{2} \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix} \\ &= -\frac{\omega_v}{2} \sigma_3 \end{aligned}$$

Flavor basis

$$\begin{aligned} & \frac{\omega_v}{2} \begin{pmatrix} -\cos 2\theta_v & \sin 2\theta_v \\ \sin 2\theta_v & \cos 2\theta_v \end{pmatrix} \\ &= \frac{\omega_v}{2} (-\cos 2\theta_v \sigma_3 + \sin 2\theta_v \sigma_1) \end{aligned}$$

NATURE OF NEUTRINO OSCILLATION

Transition Probability

$$P(|\nu_e\rangle \rightarrow |\nu_\mu\rangle) = \sin^2(2\theta_v) \sin^2(\omega_v x/2)$$

- ▶ $\omega_v = (m_2^2 - m_1^2)/2E$ determines oscillation wavelength.
- ▶ Mixing angle θ_v determines flavor oscillation amplitude.

OVERVIEW

Introduction

Matter Effect

- Matter Interaction

- MSW Effect

- Solar Neutrino Problem

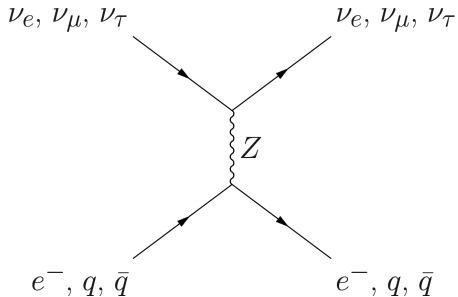
- Stimulated Neutrino Oscillations

Stimulated Neutrino Flavor Transitions

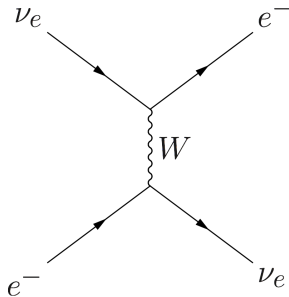
Jacobi-Anger Expansion

Summary

MATTER INTERACTION



Neutral current interaction between
 ν_e, ν_μ, ν_τ ,
and e^- , quarks and antiquarks.



Charged current interaction between
 ν_e and e^-

MATTER INTERACTION

Hamiltonian with matter interaction in flavor basis ($\omega_V = \delta m^2/2E$):

$$\mathbf{H} = \frac{\omega_V}{2} \begin{pmatrix} -\cos 2\theta_V & \sin 2\theta_V \\ \sin 2\theta_V & \cos 2\theta_V \end{pmatrix} \pm \sqrt{2} G_F n_e(x) \begin{pmatrix} 1 & 0 \\ 0 & 0 \end{pmatrix}$$

► Vacuum Hamiltonian

► Matter interaction

MATTER INTERACTION

Hamiltonian with matter interaction in flavor basis ($\omega_v = \delta m^2/2E$):

$$\mathbf{H} = \frac{\omega_v}{2} (-\cos 2\theta_v \sigma_3 + \sin 2\theta_v \sigma_1) + \frac{\lambda(x)}{2} \sigma_3$$

► Vacuum Hamiltonian

► Matter interaction

► $\lambda(x) = \sqrt{2}G_F n_e(x)$

MSW EFFECT

Hamiltonian in Vacuum

$$\mathbf{H}_{\text{vacuum}} = \frac{\omega_v \cos 2\theta_v}{2} \sigma_3 + \frac{\omega_v \sin 2\theta_v}{2} \sigma_1$$

$$\begin{aligned} \mathbf{H} &= \frac{\lambda(x) - \omega_v \cos 2\theta_v}{2} \sigma_3 + \frac{\omega_v \sin 2\theta_v}{2} \sigma_1 \\ &= \frac{\omega_m(x) \cos 2\theta_m(x)}{2} \sigma_3 + \frac{\omega_m(x) \sin 2\theta_m(x)}{2} \sigma_1, \end{aligned}$$

where

$$\begin{aligned} \omega_m(x) &= \sqrt{(\lambda(x) - \omega_v \cos 2\theta_v)^2 + \omega_v^2 \sin^2 2\theta_v}, \\ \tan 2\theta_m(x) &= \frac{\omega_v \sin 2\theta_v}{\omega_v \cos 2\theta_v - \lambda(x)}. \end{aligned}$$

MSW EFFECT

$$\begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \end{pmatrix} = \begin{pmatrix} \cos \theta_v & \sin \theta_v \\ -\sin \theta_v & \cos \theta_v \end{pmatrix} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \end{pmatrix}$$

Constant matter profile λ_0 as an example,

Significance of θ_m

Define matter basis (eigenenergy basis) $\{|\nu_L\rangle, |\nu_H\rangle\}$

$$\begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \end{pmatrix} = \begin{pmatrix} \cos \theta_m & \sin \theta_m \\ -\sin \theta_m & \cos \theta_m \end{pmatrix} \begin{pmatrix} |\nu_L\rangle \\ |\nu_H\rangle \end{pmatrix}$$

In matter basis

$$\mathbf{H}_{\text{matter-basis}} = -\frac{\omega_m}{2} \sigma_3$$

MSW EFFECT

$$\begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \end{pmatrix} = \begin{pmatrix} \cos \theta_v & \sin \theta_v \\ -\sin \theta_v & \cos \theta_v \end{pmatrix} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \end{pmatrix}$$

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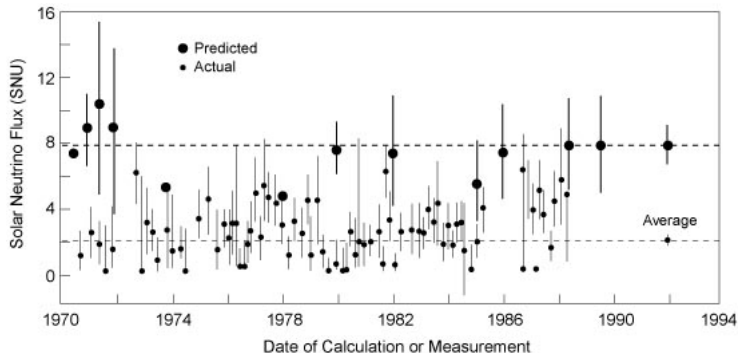
In matter basis

$$\mathbf{H}_{\text{matter-basis}} = -\frac{\omega_m}{2} \sigma_3$$

Transition Probability

$$P(|\nu_e\rangle \rightarrow |\nu_\mu\rangle) = \sin^2(2\theta_m) \sin^2(\omega_m x)$$

SOLAR NEUTRINO PROBLEM



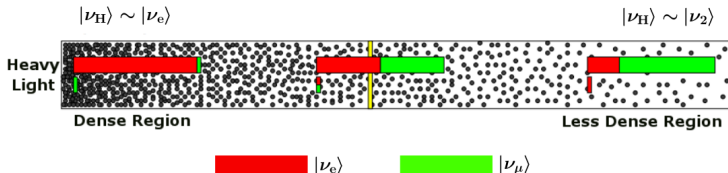
Chlorine detector (Homestake experiment) results and theory predictions.
SNU: 1 event for 10^{36} target atoms per second. Kenneth R. Lang (2010)

MSW EFFECT AND SOLAR NEUTRINOS

$$\mathbf{H} = \frac{\lambda(x) - \omega_v \cos 2\theta_v}{2} \sigma_3 + \frac{\omega_v \sin 2\theta_v}{2} \sigma_1$$

$$\begin{pmatrix} |\nu_L\rangle \\ |\nu_H\rangle \end{pmatrix} = \begin{pmatrix} \cos \theta_m & -\sin \theta_m \\ \sin \theta_m & \cos \theta_m \end{pmatrix} \begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \end{pmatrix}$$

$$\mathbf{H}_{\text{matter-basis}} = -\frac{\omega_m}{2} \sigma_3$$



Yellow bar is the resonance point. Red: $|\nu_e\rangle$. Green: $|\nu_\mu\rangle$. Adapted from Smirnov, 2003.

MSW EFFECT

Suppose $\omega_v = (m_2^2 - m_1^2)/2E < 0$,

$$\mathbf{H} = -\frac{\omega_v}{2} \begin{pmatrix} -\cos 2\theta_v & \sin 2\theta_v \\ \sin 2\theta_v & \cos 2\theta_v \end{pmatrix} + \sqrt{2}G_F n_e(x) \begin{pmatrix} 1 & 0 \\ 0 & 0 \end{pmatrix}$$

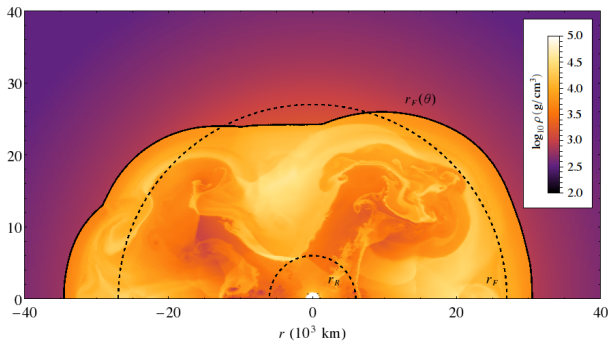


$$\mathbf{H} = \left(\frac{-\omega_v}{2} \cos 2\theta_v + \frac{\lambda(x)}{2} \right) \sigma_3 - \frac{\omega_v}{2} \sin 2\theta_v \sigma_1$$

SUPERNOVA MATTER DENSITY PROFILE

Why Do We Care

Astrophysical environments: supernovae, accretion disks etc



Supernova shock and turbulence. E. Borriello, et al (2014)

$$\Delta n_e(r) = \sum_n c_n \sin(k_n r + \phi_n)$$

STIMULATED NEUTRINO OSCILLATIONS

Matter Profile

$$\lambda(x) = \lambda_0 + \delta\lambda(x)$$

Basis

Background matter basis: Hamiltonian is diagonalized with only background matter profile λ_0 ,

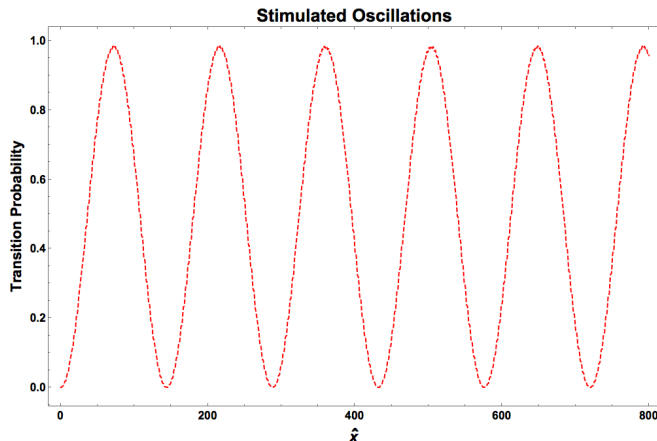
$$H_{\text{background}} = -\frac{\omega_m}{2} \sigma_3.$$

Hamiltonian

$$\mathbf{H} = \frac{1}{2} (-\omega_m + \delta\lambda(x) \cos 2\theta_m) \sigma_3 - \frac{\delta\lambda(x)}{2} \sin \theta_m \sigma_1.$$

STIMULATED NEUTRINO OSCILLATIONS

P. Krastev and A. Smirnov (1989); J. Kneller et al (2013);
K. Patton et al (2014);



Stimulated oscillations. $\lambda(x) = \lambda_0 + A \sin(kx)$ with $\hat{x} = \omega_m x$, $A = 0.1\omega_m$,
 $k = 0.995\omega_m$, $\theta_m = \pi/6$

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Hamiltonian and Basis
Rabi Oscillations

Jacobi-Anger Expansion

Summary

HAMILTONIAN

Hamiltonian in Background Matter Basis

$$\mathbf{H} = \frac{1}{2} (-\omega_{\text{m}} + \delta\lambda(x) \cos 2\theta_{\text{m}}) \sigma_3 - \frac{\delta\lambda(x)}{2} \sin 2\theta_{\text{m}} \sigma_1.$$

Matter profile

$$\lambda(x) = \lambda_0 + A \cos(kx),$$

$$\mathbf{H} = \frac{1}{2} (-\omega_{\text{m}} + \cos 2\theta_{\text{m}} A \cos(kx)) \sigma_3 - \frac{\sin 2\theta_{\text{m}}}{2} A \cos(kx) \sigma_1.$$

RABI OSCILLATIONS

Rabi Oscillation

Hamiltonian

$$-\frac{\omega_0}{2}\sigma_3 - \frac{\alpha}{2} \begin{pmatrix} 0 & e^{ikt} \\ e^{-ikt} & 0 \end{pmatrix}$$

$$E_2 = \frac{\omega_0}{2}$$

$$E_1 = -\frac{\omega_0}{2}$$

Incoming light



Frequency : k

RABI OSCILLATIONS

$$\frac{1}{2} (-\omega_m + \cos 2\theta_m A \cos(kx)) \sigma_3 - \frac{\sin 2\theta_m}{2} A \cos(kx) \sigma_1$$

Rabi Oscillation

Hamiltonian

$$-\frac{\omega_0}{2} \sigma_3 - \frac{\alpha}{2} \begin{pmatrix} 0 & e^{ikt} \\ e^{-ikt} & 0 \end{pmatrix}$$

$$E_2 = \frac{\omega_0}{2}$$

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Incoming light



Frequency : k

The transition probability from low energy to high energy is

$$P_{1 \rightarrow 2} = \frac{\alpha^2}{\alpha^2 + (\omega_0 - k)^2} \sin^2 \left(\frac{\Omega_R}{2} t \right),$$

where

$$\Omega_R = \sqrt{\alpha^2 + (\omega_0 - k)^2}$$

is Rabi frequency.

RABI OSCILLATIONS

Rabi Oscillation

Hamiltonian

$$-\frac{\omega_0}{2}\sigma_3 - \frac{\alpha}{2} \begin{pmatrix} 0 & e^{ikt} \\ e^{-ikt} & 0 \end{pmatrix}$$

$$E_2 = \frac{\omega_0}{2}$$

$$E_1 = -\frac{\omega_0}{2}$$

Incoming light



Frequency : k

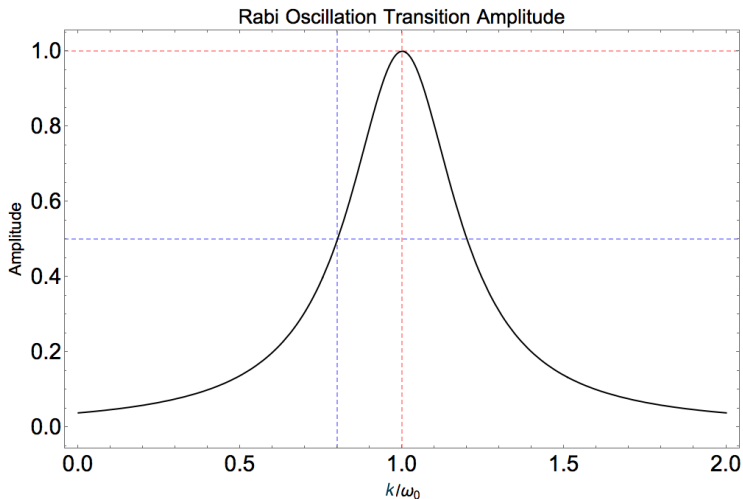
The transition probability from low energy to high energy is

$$P_{1 \rightarrow 2} = \frac{1}{1 + D^2} \sin^2 \left(\frac{\Omega_R}{2} t \right),$$

where

$$D = \left| \frac{\omega_0 - k}{\alpha} \right|.$$

RABI OSCILLATIONS

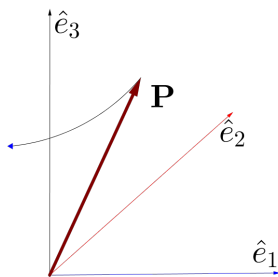


Amplitude of Rabi oscillations for different driving field frequency k

VISUALIZING RABI OSCILLATIONS

$$-\frac{\omega_0}{2}\sigma_3 - \frac{\alpha}{2} \begin{pmatrix} 0 & e^{ikt} \\ e^{-ikt} & 0 \end{pmatrix}$$

$$\begin{aligned}
 & -\frac{\omega_0}{2}\sigma_3 - \frac{\alpha}{2}\cos(kt)\sigma_1 + \frac{\alpha}{2}\sin(kt)\sigma_2 \\
 & = (\alpha\cos(kt) \quad -\alpha\sin(kt) \quad \omega_0) \begin{pmatrix} -\sigma_1/2 \\ -\sigma_2/2 \\ -\sigma_3/2 \end{pmatrix} \\
 & = \vec{H} \cdot (-\vec{\sigma}/2)
 \end{aligned}$$



$$D = \left| \frac{\omega_0 - k}{\alpha} \right|$$

is ratio of the energy gap in corotating frame to width of resonance.

INTERFERENCES OF RABI OSCILLATIONS

$$\begin{aligned}
 & \begin{pmatrix} 0 \\ 0 \\ \omega_0 \end{pmatrix} + \alpha \begin{pmatrix} \cos(kt) \\ -\sin(kt) \\ 0 \end{pmatrix} \\
 \mathbf{H} &= \frac{1}{2} \left(-\omega_m + \cos 2\theta_m A \cos(kx) \right) \sigma_3 - \frac{\sin 2\theta_m}{2} A \cos(kx) \sigma_1 \\
 &\rightarrow -\frac{\omega_m}{2} \sigma_3 - \frac{A \sin 2\theta_m}{2} \cos(kx) \sigma_1
 \end{aligned}$$

$$\vec{H} = \begin{pmatrix} 0 \\ 0 \\ \omega_m \end{pmatrix} + \frac{A \sin 2\theta_m}{2} \begin{pmatrix} \cos(kx) \\ -\sin(kx) \\ 0 \end{pmatrix} + \frac{A \sin 2\theta_m}{2} \begin{pmatrix} \cos(-kx) \\ -\sin(-kx) \\ 0 \end{pmatrix}$$

Two frequencies!

INTERFERENCES OF RABI OSCILLATIONS

$$\vec{H} = \begin{pmatrix} 0 \\ 0 \\ \omega_m \end{pmatrix} + \alpha_1 \begin{pmatrix} \cos(k_1 x) \\ -\sin(k_1 x) \\ 0 \end{pmatrix} + \alpha_2 \begin{pmatrix} \cos(k_2 x) \\ -\sin(k_2 x) \\ 0 \end{pmatrix}$$

Corotating frame of the second frequency,

$$\vec{H} = \begin{pmatrix} 0 \\ 0 \\ \omega_m - k_2 \end{pmatrix} + \alpha_1 \begin{pmatrix} \cos(k_1 - k_2 x) \\ -\sin(k_1 - k_2 x) \\ 0 \end{pmatrix} + \alpha_2 \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$$

Energy gap in this frame becomes the length of the vector

$$\begin{pmatrix} 0 \\ 0 \\ \omega_m - k_2 \end{pmatrix} + \alpha_2 \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$$

INTERFERENCES OF RABI OSCILLATIONS

$$\vec{H} = \begin{pmatrix} 0 \\ 0 \\ \omega_m \end{pmatrix} + \alpha_1 \begin{pmatrix} \cos(k_1 x) \\ -\sin(k_1 x) \\ 0 \end{pmatrix} + \alpha_2 \begin{pmatrix} \cos(k_2 x) \\ -\sin(k_2 x) \\ 0 \end{pmatrix}$$

Corotating frame of the second frequency,

$$\vec{H} = \begin{pmatrix} 0 \\ 0 \\ \omega_m - k_2 \end{pmatrix} + \alpha_1 \begin{pmatrix} \cos(k_1 - k_2 x) \\ -\sin(k_1 - k_2 x) \\ 0 \end{pmatrix} + \alpha_2 \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$$

Energy gap in this frame becomes the length of the vector

$$\sqrt{(\omega_m - k_2)^2 + \alpha_2^2} \rightarrow \omega_m - k_2 + \frac{1}{2} \frac{\alpha_2^2}{\omega_m - k_2}$$

Relative detuning

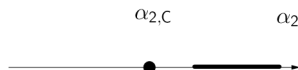
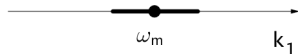
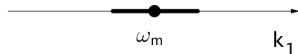
$$D' = \left| \frac{\omega_m - k_1}{\alpha_1} + \frac{\alpha_2^2}{2\alpha_1(\omega_m - k_2)} \right|$$

INTERFERENCES OF RABI OSCILLATIONS

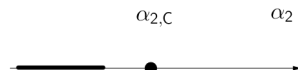
$$D' = \left| \frac{\omega_m - k_1}{\alpha_1} + \frac{\alpha_2^2}{2\alpha_1(\omega_m - k_2)} \right|$$

Two driving frequencies k_1 , and k_2 , with amplitude α_1 , and α_2

Destruction effect: $k_1 = \omega_m$, $|\alpha_2| \gg \sqrt{2\omega_m|\alpha_1(k_2 - \omega_m)|} \equiv \alpha_{2,c}$

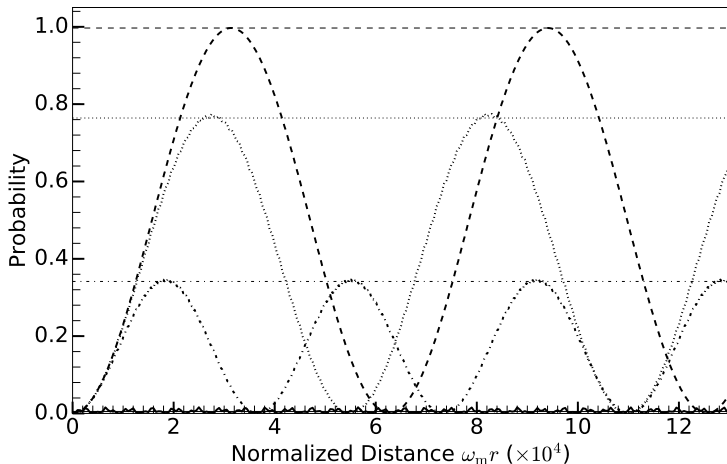


Resonance Destruction



Resonance Survival

INTERFERENCES OF RABI OSCILLATIONS

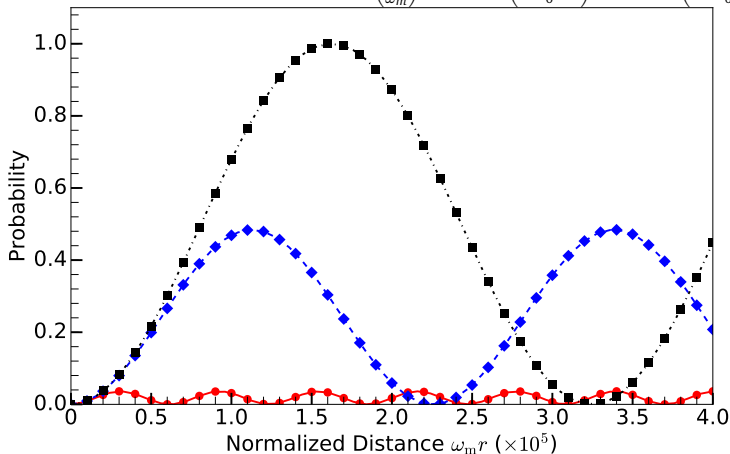


Grid lines: amplitude predicted using $1/(1 + D'^2)$

α_2, k_1 values			
Dashed	dotted	dash-dotted	solid
$10^{-2} \omega_m, 10 \omega_m$	$10^{-2} \omega_m, 10^{-1} \omega_m$	$5.0 \times 10^{-2} \omega_m, 10 \omega_m$	$5 \times 10^{-2} \omega_m, 10^{-1} \omega_m$

RABI FORMULA WORKS

$$\vec{H} = \begin{pmatrix} 0 & \\ 0 & \end{pmatrix} + \frac{A \sin 2\theta_m}{2} \begin{pmatrix} \cos(kx) & \\ -\sin(kx) & 0 \end{pmatrix} + \frac{A \sin 2\theta_m}{2} \begin{pmatrix} \cos(-kx) & \\ -\sin(-kx) & 0 \end{pmatrix}$$



Lines: Rabi formula

Dots, diamonds, triangles, and squares are for $k = \omega_m$, $k = (1 - 2 \times 10^{-5})\omega_m$, and $k = (1 - 10^{-4})\omega_m$ respectively.

$$\alpha_{2,C} = 2\alpha = 2 \times \frac{A \sin 2\theta_m}{2}$$

OVERVIEW

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INTERFERENCES OF RABI OSCILLATIONS

We have been making approximations.

$$\begin{aligned}\mathbf{H} &= \frac{1}{2} \left(-\omega_m + \cos 2\theta_m A \cos(kx) \right) \sigma_3 - \frac{\sin 2\theta_m}{2} A \cos(kx) \sigma_1 \\ &\rightarrow -\frac{\omega_m}{2} \sigma_3 - \frac{A \sin 2\theta_m}{2} \cos(kx) \sigma_1\end{aligned}$$

We need a better basis.

RABI BASIS

Hamiltonian in Background Matter Basis

$$\mathbf{H} = \frac{1}{2} (-\omega_m + \delta\lambda(x) \cos 2\theta_m) \sigma_3 - \frac{\delta\lambda(x)}{2} \sin \theta_m \sigma_1.$$

A Better Basis

Define Rabi basis $\{|\tilde{\nu}_L\rangle, |\tilde{\nu}_H\rangle\}$ is related to background matter basis $\{|\nu_L\rangle, |\nu_H\rangle\}$ through

$$\begin{pmatrix} |\nu_L\rangle \\ |\nu_H\rangle \end{pmatrix} = \begin{pmatrix} e^{-i\eta(x)} & 0 \\ 0 & e^{i\eta(x)} \end{pmatrix} \begin{pmatrix} |\tilde{\nu}_L\rangle \\ |\tilde{\nu}_H\rangle \end{pmatrix},$$

where

$$\eta(x) - \eta(0) = \frac{\cos 2\theta_m}{2} \int_0^x \delta\lambda(\tau) d\tau.$$

SINGLE FREQUENCY MATTER PROFILE

Matter profile

$$\lambda(x) = \lambda_0 + A \sin(kx),$$

Hamiltonian in new basis

$$\tilde{\mathbf{H}} = -\frac{\omega_m}{2}\sigma_3 - \frac{\delta\lambda(x)}{2}\sin 2\theta_m \begin{pmatrix} 0 & e^{2i\eta(x)} \\ e^{-2i\eta(x)} & 0 \end{pmatrix} = -\frac{\omega_m}{2}\sigma_3 + \begin{pmatrix} 0 & h \\ h^* & 0 \end{pmatrix}$$

Hamiltonian in New Basis

$$\begin{aligned} h &\equiv -\frac{\delta\lambda(x)}{2}e^{2i\eta(x)} \\ &= \frac{i}{4} \left[\exp \left(i(k + \omega_m)x + i \cos 2\theta_m \frac{A}{k} \cos(kx) \right) \right. \\ &\quad \left. - \exp \left(i(-k + \omega_m)x + i \cos 2\theta_m \frac{A}{k} \cos(kx) \right) \right] \end{aligned}$$

SINGLE FREQUENCY MATTER PROFILE

Off-diagonal Term in Our System

$$\tilde{\mathbf{H}} = -\frac{\omega_m}{2}\sigma_3 + \begin{pmatrix} 0 & h \\ h^* & 0 \end{pmatrix}$$

$$h \propto \left[\exp \left(i(k + \omega_m)x + i \cos 2\theta_m \frac{A}{k} \cos(kx) \right) - \exp \left(i(-k + \omega_m)x + i \cos 2\theta_m \frac{A}{k} \cos(kx) \right) \right]$$

Jacobi-Anger expansion

$$e^{i\beta \cos(kx)} = \sum_{n=-\infty}^{\infty} i^n J_n(\beta) e^{inkx},$$

where $J_n(\beta)$ are Bessel's functions of the first kind.

SINGLE FREQUENCY MATTER PROFILE

Scaled Quantities

Characteristic scale: ω_m

- ▶ $\hat{A} = A/\omega_m$
- ▶ $\hat{k} = k/\omega_m$
- ▶ $\hat{x} = \omega_m x$
- ▶ $\hat{h} = h/\omega_m$

SINGLE FREQUENCY MATTER PROFILE

Rotation Wave Approximation

The off-diagonal element of Hamiltonian

$$\tilde{\mathbf{H}} = \sum_{n=-\infty}^{\infty} \begin{pmatrix} 0 & \frac{1}{2}\hat{B}_n e^{i(n\hat{k}-1)\hat{x}} \\ \frac{1}{2}\hat{B}_n^* e^{-i(n\hat{k}-1)\hat{x}} & 0 \end{pmatrix}$$

where $\hat{B}_n = -(-i)^n n \hat{k} \tan 2\theta_m J_n(\hat{A} \cos 2\theta_m / \hat{k})$.

SINGLE FREQUENCY MATTER PROFILE

Transition Probability

$$P_{L \rightarrow H}^{(n)} = \frac{|\hat{B}_n|^2}{|\hat{B}_n|^2 + (n\hat{k} - 1)^2} \sin^2 \left(\frac{q^{(n)}}{2} x \right),$$

where

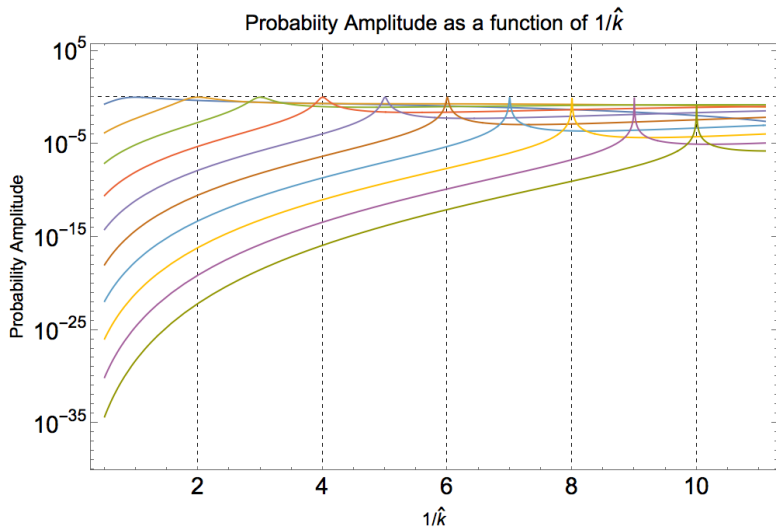
$$q^{(n)} = \sqrt{|\Gamma^{(n)}|^2 + (n\hat{k} - 1)^2}, \quad \text{frequency of oscillations}$$

$$\Gamma^{(n)} = |\hat{B}_n|, \quad \text{width of resonance } (n\hat{k} \text{ as parameter})$$

Resonance conditions

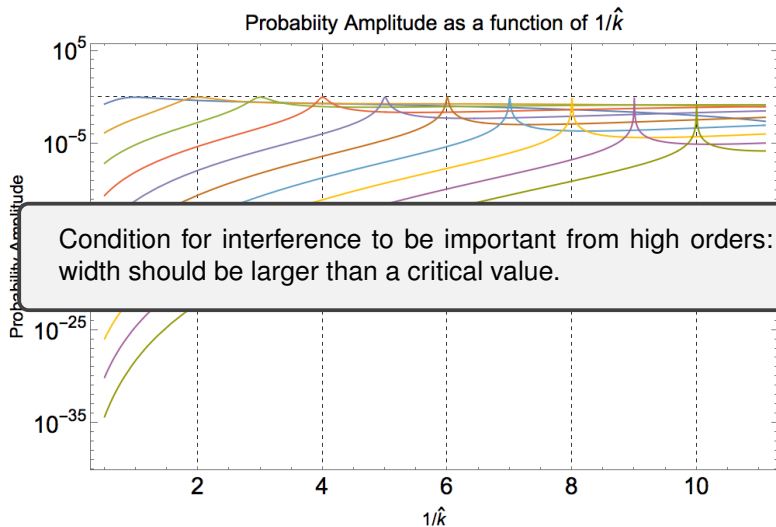
$$\hat{k} \sim \frac{1}{n}$$

SINGLE FREQUENCY MATTER PROFILE



Resonances of different $n = 1/\hat{k}$. Width becomes extremely narrow for high orders.

SINGLE FREQUENCY MATTER PROFILE



Resonances of different $n = 1/\hat{k}$. Width becomes extremely narrow for high orders.

SINGLE FREQUENCY MATTER PROFILE REVISITED

$k_1 = \omega_m$			
n	D	D'_1	$2\pi\omega_m/\Omega_n$
1	0	-	3.2×10^5
-1	10^5	4.8×10^{-6}	3.1
2	1.1×10^9	2.1×10^{-14}	6.3
-2	3.4×10^9	6.9×10^{-15}	2.1

SINGLE FREQUENCY MATTER PROFILE REVISITED

$k_1 = (1 - 2 \times 10^{-5})\omega_m$			
n	D	D'_1	$2\pi\omega_m/\Omega_n$
1	1	-	2.2×10^5
-1	10^5	1	3.1
2	1.1×10^9	1	6.3
-2	3.4×10^9	1	2.1

SINGLE FREQUENCY MATTER PROFILE REVISITED

$$k_1 = (1 - 10^{-4})\omega_m$$

n	D	D'_1	$2\pi\omega_m/\Omega_n$
1	5.2	-	6.2×10^4
-1	10^5	5.2	3.1
2	1.1×10^9	5.2	6.3
-2	3.4×10^9	5.2	2.1

CASTLE WALL MATTER PROFILE

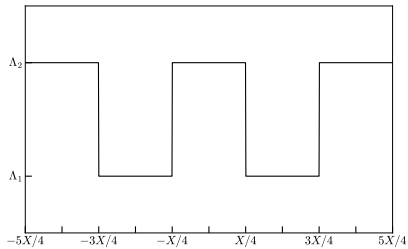
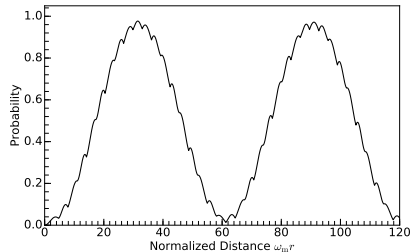


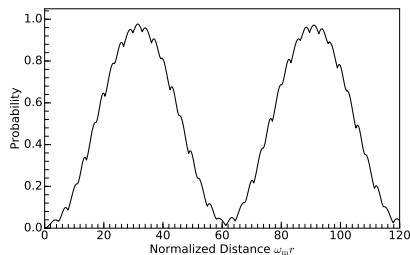
Figure: Castle wall matter profile



CASTLE WALL MATTER PROFILE

Table: Relative detuning of each frequency.

$\{n_1, n_2\}$	D	$D'_{\{1,0\}}$
$\{1, 0\}$	0	-
$\{-1, 0\}$	48	1.0×10^{-2}
$\{0, 1\}$	1.5×10^2	1.1×10^{-3}
$\{2, 0\}$	2.4×10^2	2.0×10^{-4}



OVERVIEW

Introduction

Matter Effect

Stimulated Neutrino Flavor Transitions

Jacobi-Anger Expansion

Summary

SUMMARY

- ▶ The fact that neutrino flavor states are not mass states causes vacuum oscillations.
- ▶ MSW resonance happens when matter potential cancels out the vacuum diagonal elements of the Hamiltonian.
- ▶ Even matter profile doesn't match MSW requirement, variation in matter profile can cause resonances.
- ▶ Single frequency perturbations in matter profile is a combination of many Rabi oscillations.
- ▶ Rabi oscillations with two driving fields of different frequencies: large width to destroy the resonance.

BACKUP SLIDES

BACKUP SLIDES

RABI OSCILLATIONS

The coupling strength is calculated as

$$\alpha = \langle 1 | \mathbf{d} \cdot \mathbf{E} | 2 \rangle$$

where the electric field is

$$\mathbf{E} = \mathbf{E}_0 \sin(kt).$$

and \mathbf{d} is the dipole moment.

INTERFERENCES OF RABI OSCILLATIONS

$$-\frac{\omega_0}{2}\sigma_3 - \frac{\alpha}{2} \begin{pmatrix} 0 & e^{ikt} \\ e^{-ikt} & 0 \end{pmatrix}$$

$$\begin{aligned} \mathbf{H} &= \frac{1}{2} \left(-\omega_m + \cos 2\theta_m A \cos(kx) \right) \sigma_3 - \frac{\sin 2\theta_m}{2} A \cos(kx) \sigma_1 \\ &\rightarrow -\frac{\omega_m}{2} \sigma_3 - \frac{A \sin 2\theta_m}{2} \cos(kx) \sigma_1 \\ &= -\frac{\omega_m}{2} \sigma_3 - \frac{A \sin 2\theta_m}{2} \frac{1}{2} \begin{pmatrix} 0 & e^{ikx} \\ e^{-ikx} & 0 \end{pmatrix} - \frac{A \sin 2\theta_m}{2} \frac{1}{2} \begin{pmatrix} 0 & e^{i(-k)x} \\ e^{-i(-k)x} & 0 \end{pmatrix} \end{aligned}$$

PARAMETERS USED FOR VACUUM OSCILLATIONS

$$\theta_{12} = 33.36/180\pi; \theta_{13} = 8.66/180\pi; \theta_{23} = 40/180 * \pi; \delta_{cp} = 0;$$
$$m_1^2 = 0.01; m_2^2 = m_1^2 + 0.000079; E = 1\text{MeV}$$

SINGLE FREQUENCY MATTER PROFILE

Why Does It Work?

$$J_n(n \operatorname{sech} \alpha) \sim \frac{e^{-n(\alpha - \tanh \alpha)}}{\sqrt{2\pi n \tanh \alpha}}, \quad \text{for large } n$$

\Rightarrow

$$\Gamma \propto \hat{B}_n \propto \frac{e^{-n(\alpha - \tanh \alpha)}}{\sqrt{2\pi n \tanh \alpha}}$$

Small perturbation \Rightarrow Small $\hat{A} \Rightarrow$ Large $\alpha \Rightarrow$ Drops fast at large n .

TWO-FREQUENCY MATTER PROFILE

Matter Profile

$$\lambda(x) = \lambda_0 + \delta\lambda(x), \quad \delta\lambda(x) = A_1 \sin(k_1 x) + A_2 \sin(k_2 x).$$

TWO-FREQUENCY MATTER PROFILE $\hat{h} = \sum_{n=-\infty}^{\infty} \frac{1}{2} \hat{B}_n e^{i(n\hat{k}-1)\hat{x}}$,

Hamiltonian Off-diagonal Element

Apply Jacobi-Anger expansion,

$$\hat{h} = \sum_{n_1=-\infty}^{\infty} \sum_{n_2=-\infty}^{\infty} \frac{1}{2} \hat{B}_{n_1, n_2}(\hat{k}_1, \hat{k}_2) e^{i(n_1 \hat{k}_1 + n_2 \hat{k}_2 - 1)\hat{x}},$$

where

$$\begin{aligned} & \hat{B}_{n_1, n_2}(\hat{k}_1, \hat{k}_2) \\ &= -(-i)^{n_1+n_2} (n_1 \hat{k}_1 + n_2 \hat{k}_2) J_{n_1} \left(\frac{\hat{A}_1 \cos 2\theta_m}{\hat{k}_1} \right) J_{n_2} \left(\frac{\hat{A}_2 \cos 2\theta_m}{\hat{k}_2} \right) \end{aligned}$$

Which terms are important?

TWO-FREQUENCY MATTER PROFILE

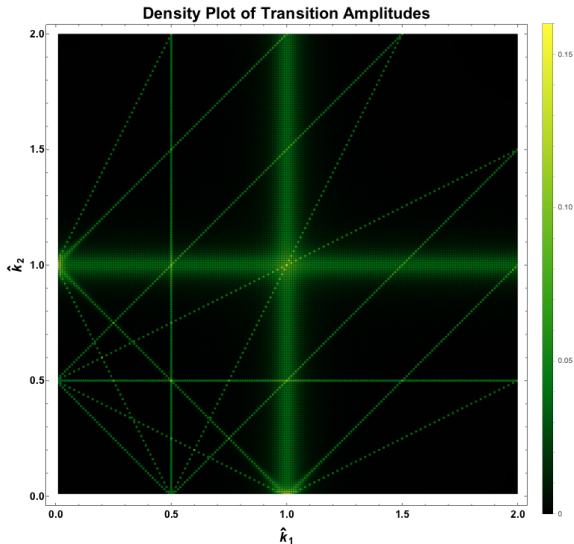
Resonance Lines

There are still resonances, i.e., (almost) zero phases, on lines

$$n_{1,0}\hat{k}_1 + n_{2,0}\hat{k}_2 - 1 = 0$$

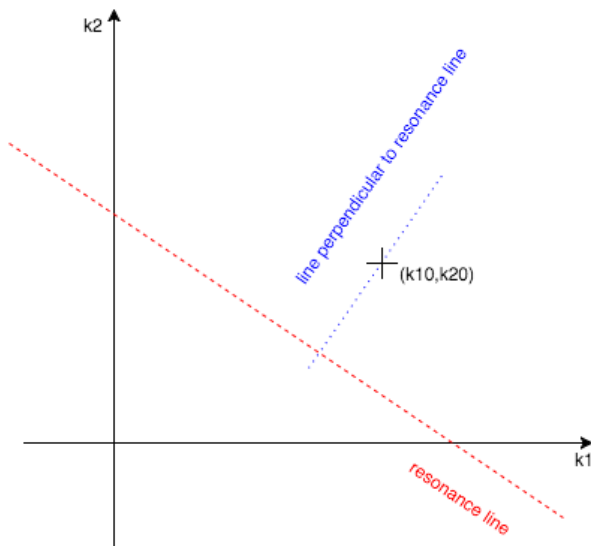
in $\{\hat{k}_1, \hat{k}_2\}$ plane. \Rightarrow Resonance width for each point on resonance lines.

TWO-FREQUENCY MATTER PROFILE $\hat{h} \equiv \sum_{n_1} \sum_{n_2} \frac{1}{2} \hat{B}_{n_1, n_2}(\hat{k}_1, \hat{k}_2) e^{i(n_1 \hat{k}_1 + n_2 \hat{k}_2 - 1)\hat{x}}$



Density plot of transition amplitudes calculated using only one term out of the whole summation in Hamiltonian. $n_1, n_2 \in [-2, 2]$

TWO-FREQUENCY MATTER PROFILE



Resonance line, distance to resonance, and width

TWO-FREQUENCY MATTER PROFILE

Width

$$\Gamma_2 = \frac{\hat{B}_{n_1, n_2}(\hat{k}_{1, \text{intercept}}, \hat{k}_{2, \text{intercept}})}{\sqrt{n_1^2 + n_2^2}}.$$

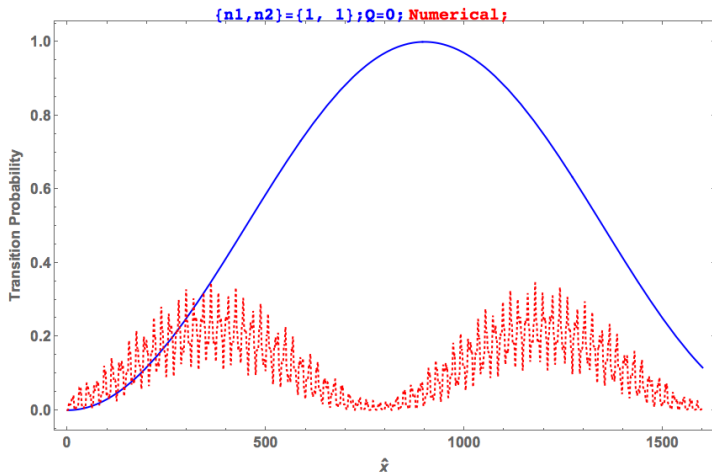
Distance to Resonance Line

$$d = \frac{|n_1 \hat{k}_{10} + n_2 \hat{k}_{20} - 1|}{\sqrt{n_1^2 + n_2^2}}.$$

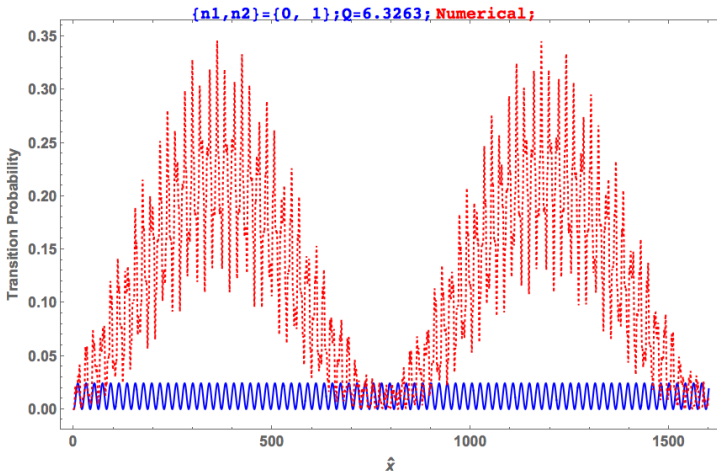
Distance to Resonance Width Ratio

$$Q_2 = \frac{d}{\Gamma_2}.$$

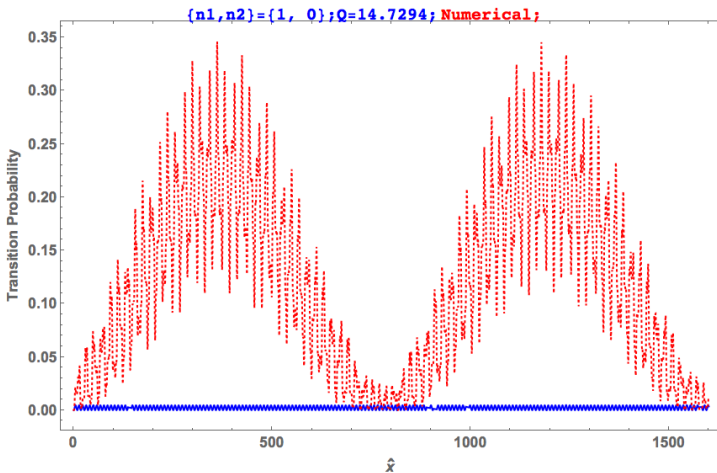
TWO-FREQUENCY MATTER PROFILE



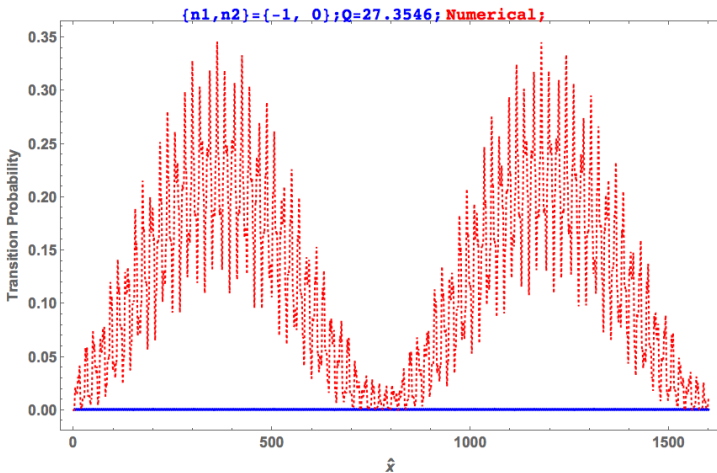
TWO-FREQUENCY MATTER PROFILE



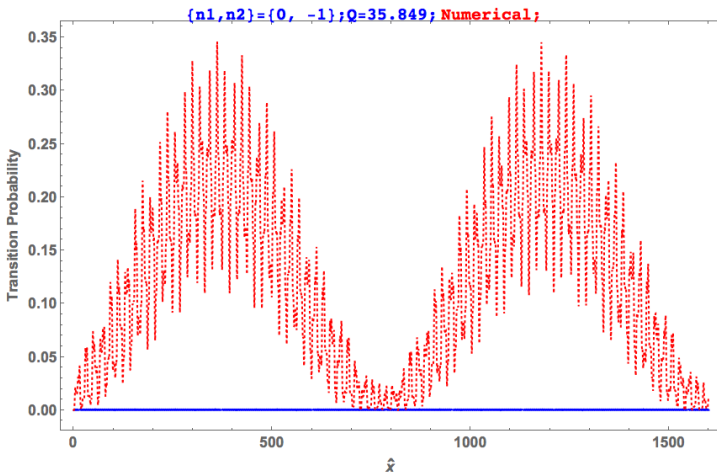
TWO-FREQUENCY MATTER PROFILE



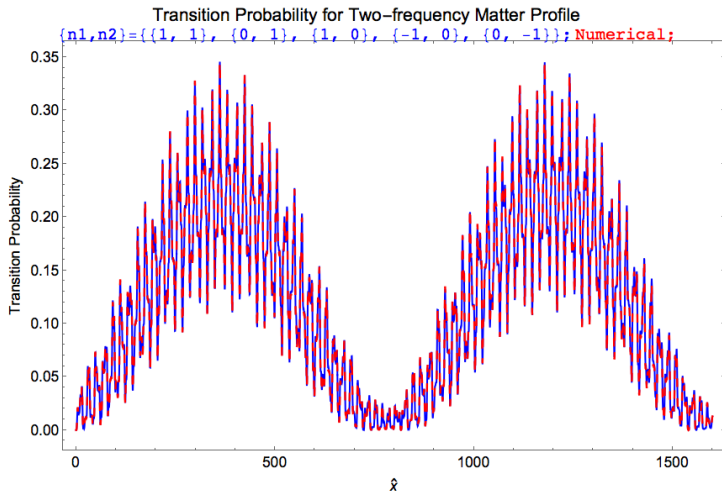
TWO-FREQUENCY MATTER PROFILE



TWO-FREQUENCY MATTER PROFILE



TWO-FREQUENCY MATTER PROFILE



BESSEL'S FUNCTION

$$J_n(\beta) = \sum_{m=0}^{\infty} \frac{(-1)^m}{m! \Gamma(m+n+1)} \left(\frac{\beta}{2}\right)^{2m+n}$$

REFERENCES I