Exercise in FYS-3002 — Solutions

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March 3, 2021

Problem 1

Task 1

What is the physical interpretation of χ in eq. (3)?

Equation (3) is what is given in FH (4.35). The parameter χ describe the susceptibility of the medium, that is, the electron population and the ion population in the plasma. In this context the susceptibility is a scalar and its value describe how prone the medium is of being magnetized by an external magnetic field, as such, we get information about the long term behaviour of the plasma through the susceptibility function.

Task 2

Explain what the term $f_{\alpha,1}$ in eq. (2) describe.

This equation describe the perturbation of the phase space function and is the same as FH (4.26) only written out in full and using the Fourier transformed spatial variable and Laplace transformed temporal variable.

Problem 2

Using any one of the expressions for $\langle |n_e(\mathbf{k},\omega)|^2 \rangle$, write a program that calculates the power spectral density. The program should accept a number of input parameters:

- $f_{\rm r}$ Radar frequency
- $n_{\rm e}$ Electron number density
- B Magnetic field strength
- $m_{\rm i}$ Ion mass
- $T_{\rm e}$ Electron temperature
- $T_{\rm i}$ Ion temperature
- θ Aspect angle (the angle between the radar pointing direction and the magnetic field.)

The code should be well commented and included as an appendix.

Explain the code. Some things to consider:

- Where in the code were the different equations solved?
- How was the numerical calculation implemented?

The code itself should be well commented and included as an appendix.

Hint: Before you integrate all the way to infinity: where along the axis of integration does most of the information lie?

Problem 3

We will now look at some specific parameters using our program. Run your program with the parameters given as:

Parameter	Unit	Value
$f_{ m r}$	[Hz]	430×10^{6}
$n_{ m e}$	$[m^{-3}]$	2×10^{10}
B	[T]	3.5×10^{-5}
$m_{ m i}$	[amu]	16
$T_{ m e}$	[K]	200
$T_{ m i}$	[K]	200
θ	[°]	135

for frequencies $f \in [-2 \times 10^6, 2 \times 10^6]$.

Task 1

Where could an experiment with these parameters be done? Make a sketch that includes the radar beam and the magnetic field line. Assume that the radar points directly upwards, i.e., towards zenith.

The radar beam will be scattered at some height, say $h \approx 200 \, \mathrm{km}$. The aspect angle is the angle between the incident wave vector and the magnetic field line at the altitude where the radar beam is scattered.

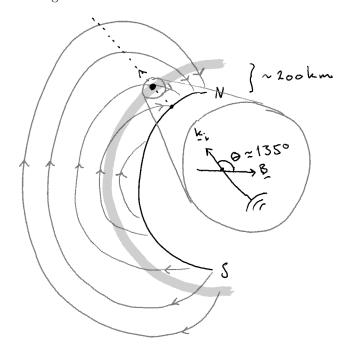


Figure 1: Sketch of a possible geometry based on the parameters given.

Task 2

The spectrum is plotted for frequencies $f \in [-2 \times 10^6, 2 \times 10^6]$; relative to an observer at the radar location, which way does the features found at positive frequencies is the spectrum move?

Structures that move towards the Earth will result in structures in the power spectrum at positive frequencies.

Task 3

Plot the resulting power spectra calculated by the program and explain what the different peaks represent.

The plot from problem 3 should result in something similar to fig. 2.

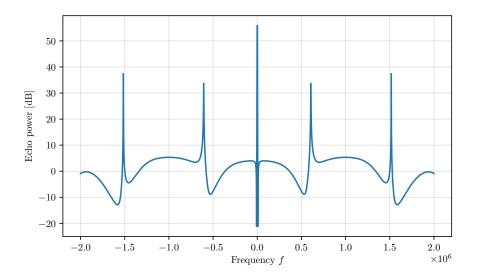


Figure 2: Power spectral density using the parameters above.

The peaks at largest frequency represent the plasma line which describe the electron population, the lines in the middle at frequency $f \approx 0.5\,\mathrm{Hz}$ are gyro lines that appear due to the aspect angle we use (not parallel to the magnetic field line) and the lines at zero frequency are the ion lines.

Problem 4

In this exercise we will use the parameters:

Parameter	Unit	Value
$f_{ m r}$	[Hz]	933×10^{6}
$n_{ m e}$	$[m^{-3}]$	2×10^{11}
B	[T]	5×10^{-5}
$m_{ m i}$	[amu]	16
$T_{ m i}$	[K]	2000
heta	[°]	180

Task 1

Calculate the power spectral density on $f \in [3.5 \times 10^6, 7 \times 10^6]$ for $T_e = 2000 \,\mathrm{K}, 4000 \,\mathrm{K}, 6000 \,\mathrm{K},$ and 8000 K and plot the power spectra.

The plots from the above should look similar to fig. 3.

Task 2

Explain the changes that can be seen as the electron temperture changes.

This is showing the upshifted plasma line. The equation describing the real part of the plasma line wave resonance frequency is

$$\omega_{\Re,e} = [\omega_{pe}^2 (1 + 3k^2 \lambda_{\rm D}^2) + \Omega_{\rm e}^2 \sin^2 \theta]^{1/2} = [\omega_{pe}^2 + 3k^2 v_{\rm th,e}^2 + \Omega_{\rm e}^2 \sin^2 \theta]^{1/2}$$

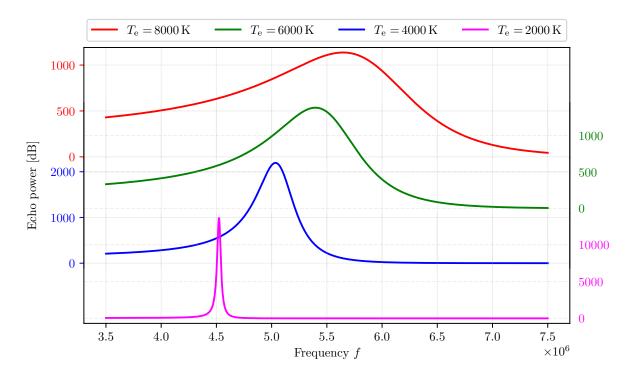


Figure 3: Power spectral density plot obtained with the parameters for problem 4.

and we see that with increasing temperature the thermal velocity will increase, thus increasing the resonance frequency as seen in the fig. 3.

Task 3

Explain what the assumption $k^2 \lambda_D^2 \ll 1$ is referring to. Is this assumption valid for all temperatures? Why/why not?

We also see from the figure that the width of the plasma line gets wider as the temperature is increased. The assumption $k^2\lambda_{\rm D}^2\ll 1$ is usually applied when solving the IS spectrum, implying weak Landau damping. It describe the kind of scattering we are interested in, which is scattering from the larger structures with length scale equal to the Debye length. But with increasing temperature this assumption is no longer valid (the Debye length is proportional to the square root of the temperature) and the Landau damping gets stronger. More power is distributed to the shoulders and therefore the peak power is also decreased to maintain the same power of the plasma line for all temperatures.

Problem 5

You should now be able to experiment a bit for yourself.

Task 1

Explain which parameter(s) that needs to be changed to obtain a similar plot as is shown in fig. 4.

We recognize this as the same as the plot in the compendium by Bjørnå (last page, top panel). If we change the ion temperature and keep the ratio between ion and electron temperature the same we get the desired result. Specifically, we can use $T_i = 100, 200, 300, 400 \,\mathrm{K}$ and a ratio to electron temperature given as $T_e/T_i = 1.5$.

${\bf Task~2}$ Reproduce the plot in fig. 4 using your own program. Include values on the axis and labels to all spectra.

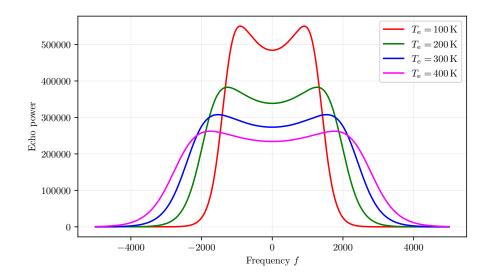


Figure 4: Power spectral density plot

Appendix

```
#!/home/een023/.virtualenvs/py3.9/bin/python
   import sys
   import numpy as np
   import scipy.constants as const
   import scipy.integrate as si
   import matplotlib.pyplot as plt
   import config as cf
10
11
   def isr():
        # Set physical parameters
13
        f = cf.f # 1/s ~ radar frequency
       T e = cf.T e # K ~ electron temperature
1.5
       T_i = cf.T_i \# K \sim ion temperature
16
       n_e = cf.n_e # 1/m^3 ~ electron number density
17
       B = cf.B # T ~ magnetic field strength (towards Earth)
18
       aspect = cf.aspect # degree ~ radar pointing direction to magnetic field line
19
        aspect = np.pi / 180 * aspect
20
       M_amu = cf.M # amu ~ ion mass
21
       M = M_amu * (const.m_p + const.m_n) / 2 # Convert to kq
22
       nu = cf.nu # 1/s ~ collision frequency
24
        # Calculate constants
       k = -4 * np.pi * f / const.c
26
       l_D = debye(T_e, n_e)
       w c = gyro('e', B)
28
       W_c = gyro('i', B, M_amu)
30
        # Susceptibility
       f ax = np.linspace(- 1e4, 1e4, cf.N F) # Frequency axis
32
        # Integration variable of Gordeyev
       y_e = np.linspace(0, 1.5e-4**(1 / cf.ORDER), cf.N_Y)**cf.ORDER
        y_i = np.linspace(0, 1.5e-2**(1 / cf.ORDER), cf.N_Y)**cf.ORDER
        G_e = maxwellian_integrand(y_e, nu, k, aspect, T_e, w_c, const.m_e)
36
        G_i = maxwellian_integrand(y_i, nu, k, aspect, T_i, W_c, M)
37
        Fe = F(f_ax, y_e, nu, G_e)
       Fi = F(f_ax, y_i, nu, G_i)
39
       Xp = (1 / (2 * 1_D**2 * k**2))**(1 / 2)
41
        chi_e = 2 * Xp**2 * Fe
        chi_i = 2 * Xp**2 * Fi
43
        with np.errstate(divide='ignore', invalid='ignore'):
45
            IS = n_e / (np.pi * 2 * np.pi * f_ax) * \
                    (np.imag(Fe) * np.abs(1 + chi_i)**2 + np.imag(Fi) * np.abs(chi_e)**2) / 
47
                    (np.abs(1 + chi e + chi i)**2)
49
        return f_ax, IS
50
   def F(f_ax, y, nu, G):
```

```
# Calculate the F functions that include susceptibility
53
         a = np.array([])
54
         for f in f_ax:
             w = 2 * np.pi * f
56
             sint = my_integration_method(w, y, G)
             a = np.r_[a, sint]
58
59
         func = 1 + (1j * 2 * np.pi * f_ax + nu) * a
60
         return func
61
62
    def maxwellian_integrand(y, nu, k, aspect, T, w_c, m):
63
        G = np.exp(-y * nu -
64
                 k**2 * np.sin(aspect)**2 * T * const.k /
65
                 (m * w_c**2) * (1 - np.cos(w_c * y)) -
66
                 .5 * (k * np.cos(aspect) * y)**2 * T * const.k / m)
67
         return G
69
    def my_integration_method(w, y, G):
71
         val = np.exp(1j * w * y) * G
         sint = si.simps(val, y)
73
         return sint
74
75
    def debye(T, n):
76
         ep0 = 1e-9 / 36 / np.pi
77
         1_D = (ep0 * const.k * T / (n * const.e**2))**(1 / 2)
78
         return 1 D
79
80
    def gyro(p, B, m=16):
81
         if p == 'e':
82
             w = const.e * B / const.m_e
83
         elif p == 'i':
84
             w = const.e * B / (m * (const.m_p + const.m_n) / 2)
86
             sys.exit(f'I do not know what kind of particle {p} is.')
         return w
88
89
    def plot():
90
        x, y = isr()
91
         # y = 10 * np.log10(y)
92
93
        plt.figure()
94
        plt.plot(x, y)
95
        plt.grid(alpha=0.4)
96
        plt.show()
97
98
    if __name__ == '__main__':
99
        plot()
100
```

0.1 References