

Recap 3-body problem

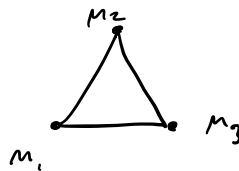
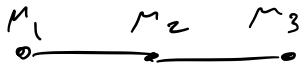
Recall: Assume system w/ All bodies rotate w/ Ω

COM's:
$$(\ddot{x}_k - 2\dot{y}_k \omega - x_k \omega^2) = G \sum_{j \neq k} \frac{m_j}{r_{jk}^3} (x_j - x_k)$$

$$(\ddot{y}_k + 2\dot{x}_k \omega - y_k \omega^2) = G \sum_{j \neq k} \frac{m_j}{r_{jk}^3} (y_j - y_k)$$

$$\ddot{z}_k = G \sum_{j \neq k} \frac{m_j}{r_{jk}^3} (z_j - z_k)$$

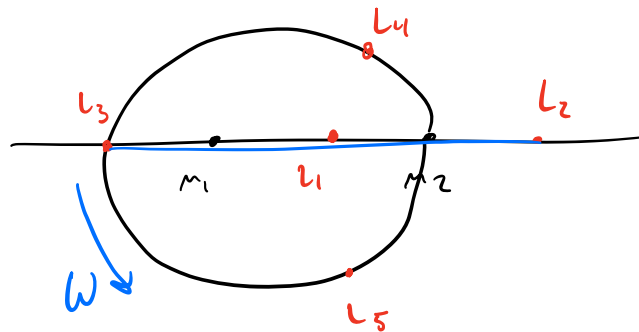
2 solutions:



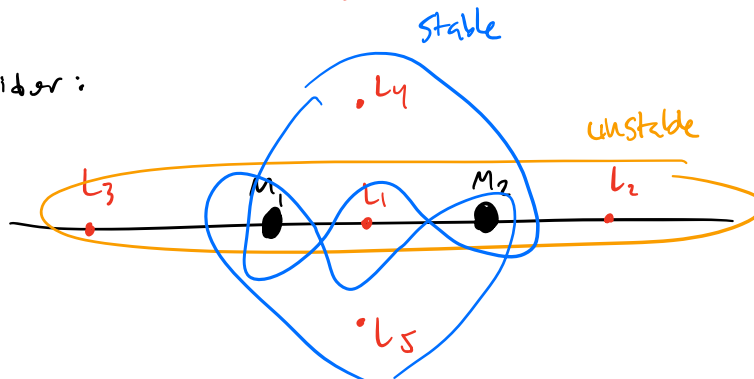
Quintic equation \rightarrow 5 'lagrange points' $L_1 - L_5$

$$M_1 \gg M_2 \gg M_3$$

stable positions for M_3 ?



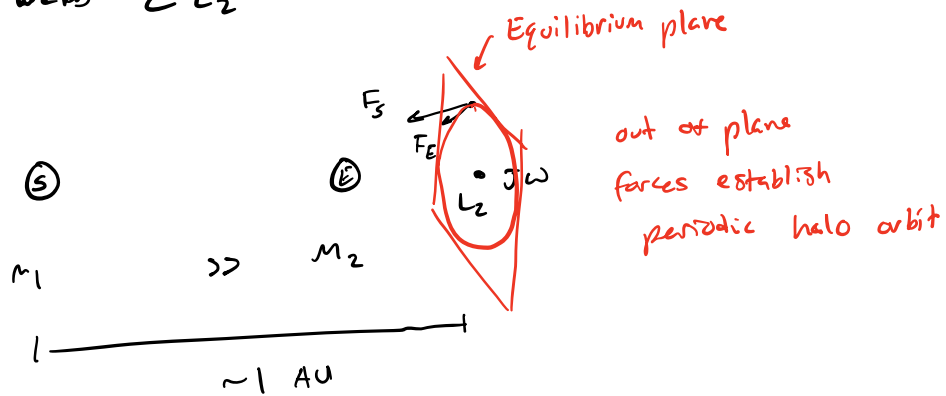
Consider:



$\rightarrow L_1, L_2, L_3 \rightarrow$ unstable

$\rightarrow L_4, L_5 \rightarrow$ stable if $\frac{m_1}{m_2} > 24.96$ (Sun-Earth \checkmark)

James Webb $\in L_2$



We need 3D orbits

$$\mathbf{r} \rightarrow \bar{\mathbf{r}} = r_x \hat{i} + r_y \hat{j} + r_z \hat{k}$$

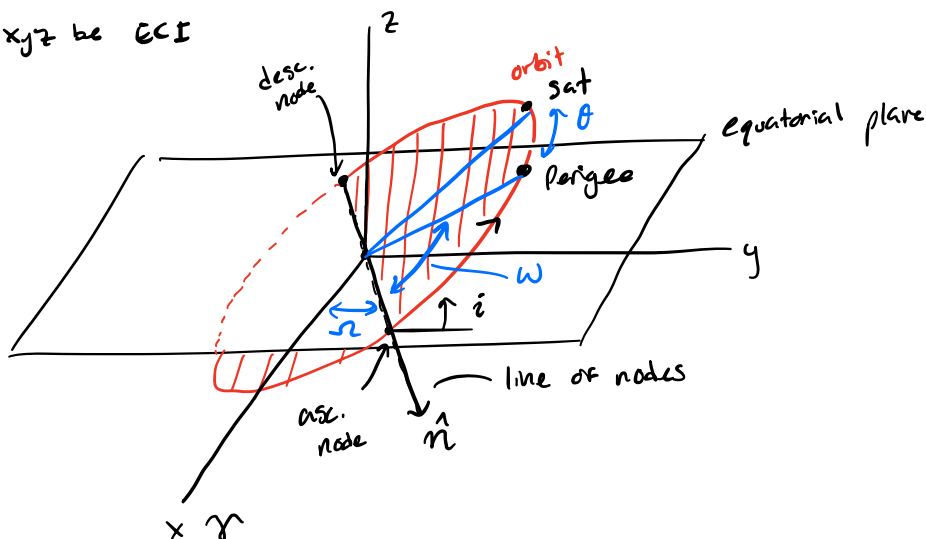
$$\bar{\mathbf{v}} = v_x \hat{i} + v_y \hat{j} + v_z \hat{k}$$

but no orbital elements

\rightarrow define orbits w/ 6 component "state" (3D)

First set: "modified classic set"

Let x, y, z be ECI



Set 1

- a - semi-major axis
- e - eccentricity
- i - inclination
- θ - true anomaly
- Ω - right ascension (\angle from x axis on eq. plane, ccw)
- ω - argument of perigee (\angle from \hat{n})

Set 2

Cartesian x, y, z, v_x, v_y, v_z

Set 3

ADIBARV

- α = right ascension
- δ = declination
- β = flight path angle
- A = angle of \vec{v} from true north
- r = radius
- v = speed

set 4

203ARV

- $\alpha \rightarrow \lambda = \text{longitude}$
 $\lambda = \alpha - \alpha_G$
 $G = \text{greenwich}$

Vectors

$$|\vec{r}| = \sqrt{\vec{r} \cdot \vec{r}}$$

$$u_r = \frac{\bar{r} \cdot \bar{v}}{r}$$

Example 1: construct MCS from \bar{r} and \bar{v}

$$\overline{r}, \overline{v} \rightarrow (a, e, i, \theta, \Omega, \omega$$

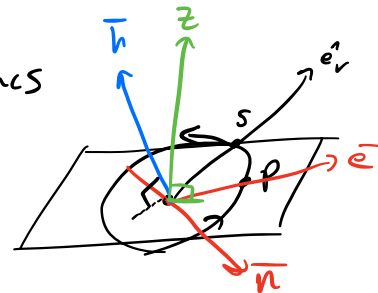
pro: easy to track / predict orbit in MCS

step 1: approaching or leaving perigee
perigee = closest approach

use radial speed $v_r = \frac{\vec{r} \cdot \vec{v}}{r}$

$\bar{v}_r > 0$ moving away

$\bar{v}_r < 0$ moving towards ρ



$$r = \sqrt{\bar{r} \cdot \bar{r}}$$

Step 2: need \bar{h} , \bar{n} , \bar{e} where \bar{n} is dir. of node line
 \bar{e} locates periapsis

$$\text{recall } \bar{h} = \bar{r} \times \bar{v}$$

$$= \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ x & y & z \\ v_x & v_y & v_z \end{vmatrix} \quad h = \sqrt{\bar{h} \cdot \bar{h}}$$

$$\text{then, } \bar{n} = \hat{k} \times \bar{h} \Rightarrow n = \sqrt{\bar{n} \cdot \bar{n}}$$

then use trajectory eqns to show

$$\bar{e} = \left(\frac{v^2}{\mu} - \frac{1}{r} \right) \bar{r} - \frac{\bar{r} \cdot \bar{v}}{\mu} \bar{v} \quad e = \sqrt{\bar{e} \cdot \bar{e}}$$

Step 3: i , Ω , ω

i from \bar{n} & \hat{k}

$$i = \cos^{-1} \left(\frac{h_z}{h} \right)$$

Ω based on \bar{n}

$$\Omega = \cos^{-1} \left(\frac{n_x}{n} \right)$$

ω " " \bar{n}, \bar{e}

$$\omega = \cos^{-1} \left(\frac{\bar{n} \cdot \bar{e}}{|\bar{n}| |\bar{e}|} \right)$$

Step 4: a , semi-major axis

$$\text{trajectory: } a = \frac{h^2}{\mu(1-e^2)}$$

Step 5: find θ

example w/ #5

$$\text{given } \bar{r} = -6045 \hat{i} - 3490 \hat{j} + 2500 \hat{k} \quad \text{km}$$

$$\bar{v} = -3.457 \hat{i} + 6.618 \hat{j} + 2.533 \hat{k} \quad \text{km/s}$$

Find location/ dir in MCS