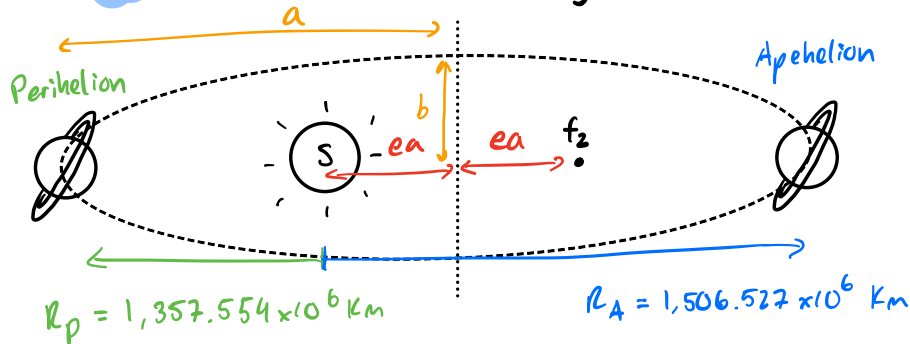


Problem 1 Saturn perihelion = $1,357.554 \times 10^6$ km, aphelion = $1,506.527 \times 10^6$ km

1a) Calculate eccentricity & semi-major axis & compare.



$$R_A = a(1+e) \quad , \quad R_P = a(1-e)$$

$$\frac{R_A}{R_P} = \frac{a(1+e)}{a(1-e)} \rightarrow \left(\frac{R_A}{R_P}\right)(1-e) = 1+e$$

$$\rightarrow \frac{R_A}{R_P} - \frac{R_A}{R_P}e = 1+e \rightarrow \frac{R_A}{R_P} - 1 = e + \frac{R_A}{R_P}e$$

$$\rightarrow \frac{R_A}{R_P} - 1 = e\left(1 + \frac{R_A}{R_P}\right) \rightarrow e = \frac{\left(\frac{R_A}{R_P} - 1\right)}{\left(1 + \frac{R_A}{R_P}\right)}$$

Plug in radii & solve for e:

$$e = 0.05201$$

$$\rightarrow a = \frac{R_A}{(1+e)} = 1432.04 \times 10^6 \text{ km} = a$$

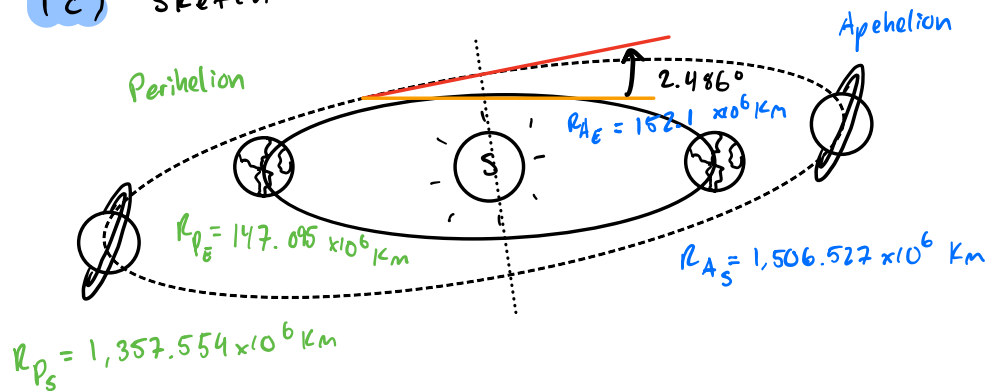
1b) Published values:

From NSSDC.GSFC, NASA.GOV : $e = 0.0520$ ✓

$a = 1,432 \times 10^6$ km ✓

Inclination: 2.486°

1c) sketch



Problem 2 Titan $r_m = 1.22 \times 10^9$ m, $T = 15.95$ days.

Hyperion $r_m = 1.48 \times 10^9$ m, Find T_H

Law 3: The square of the period T of the orbit is proportional to the cube of the semi-major axis a $T^2 \propto a^3$

$$T_T^2 = K r_T^3, \quad T_H^2 = K r_H^3$$

$$\hookrightarrow K = \frac{T_T^2}{r_T^3} \rightarrow T_H = \sqrt{\frac{T_T^2}{r_T^3} r_H^3} = \sqrt{\frac{15.95^2}{1.22^3} (1.48^3)}$$

$$= \sqrt{454.18} \rightarrow T_H \approx 21.31 \text{ days}$$

Actual: 21.28 days (NASA Horizons)

Problem 3 Find normal accel. of Earth around sun & Sun around

MW. Around sun: $r = 93 \text{ E}6$ mi, $T = 1 \text{ yr}$

|| MW: $r = 1.9356 \text{ E}17$ mi, $v = 180 \text{ mi/s}$

$$a_{\text{normal}}(2D) = -\omega^2 r$$

$$\text{Sun: } \omega_{\text{sun}} = \frac{2\pi}{T} = \frac{2\pi \text{ rad}}{3.154 \text{ E}7 \text{ sec}} = 1.9921 \text{ E}-7 \text{ rad/s}$$

$$r_{\text{sun}} = 1.4966 \text{ E}11 \text{ m}$$

$$\rightarrow |a_{n, \text{sun}}| = (1.9921 \times 10^{-7})^2 (1.4966 \times 10^{11}) = \frac{0.00594 \text{ m/s}^2}{9.81 \text{ m/s}^2}$$

$$\rightarrow a_{n, \text{sun}} = 0.06 \% \text{ Earth gravity}$$

$$\text{MW: } T_{\text{mw}} = \frac{2\pi r_{\text{mw}}}{v_{\text{mw}}} = \frac{2\pi (1.9356 \times 10^7 \text{ m})}{180 \text{ m/s}} = 6.756 \times 10^5 \text{ sec}$$

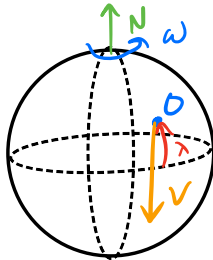
$$\omega_{\text{mw}} = \frac{2\pi}{T_{\text{mw}}} = 9.3 \times 10^{-16} \text{ rad/s}$$

$$a_{n, \text{mw}} = \omega_{\text{mw}}^2 r_{\text{mw}} = (9.3 \times 10^{-16})^2 (3.115 \times 10^{20}) = \frac{2.694 \times 10^{-16} \text{ m/s}^2}{9.81 \text{ m/s}^2}$$

$$\rightarrow a_{n, \text{mw}} = 0.0000000275 \% \text{ Earth gravity}$$

Problem 4

Particle launched in rotating frame

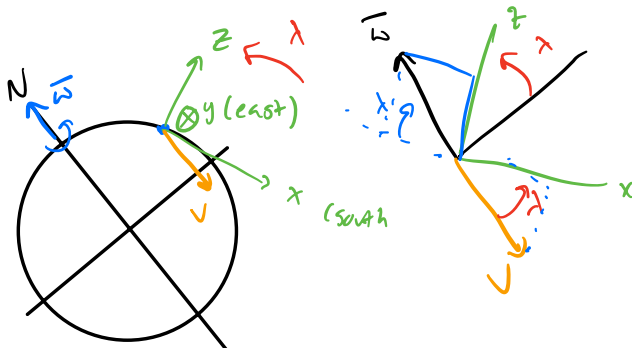


λ = latitude

V = velocity launched at $t=0$

4a) Find magnitude & direction of Coriolis force in terms of V , ω , λ

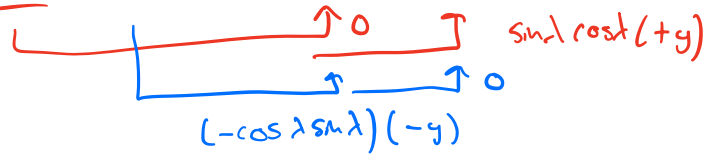
$$\text{Coriolis accel.: } \vec{a}_c = 2 \vec{\omega} \times (\vec{v}_p)_{xyz}$$



$$\vec{\omega} = \omega \sin \lambda \hat{z} - \omega \cos \lambda \hat{x}$$

$$\vec{V} = V \sin \lambda \hat{z} + V \cos \lambda \hat{x}$$

$$\bar{a}_c = 2\omega(\sin\lambda \hat{z} - \cos\lambda \hat{x}) \times v(\sin\lambda \hat{z} + \cos\lambda \hat{x})$$



$$\rightarrow \bar{a}_c = 4\omega V \cos\lambda \sin\lambda (\hat{y}) \rightarrow \text{Eastward}$$

4b) The hemisphere will change the sign of λ . This will not affect the $\cos\lambda$ term, but it will affect the $\sin\lambda$, since $\sin(-\lambda) = -\sin(\lambda)$, flipping the sign and by extension the direction 180° , changing it in this case to westward.