

# On Gödel-type Metrics

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joined work with

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- 1 Setting
- 2 Toolkit (part)
- 3 Gödel-type Spacetimes
- 4 Visualization Techniques
- 5 Behavior
- 6 Outlook

## Along the following line:

- 1 Present (a part of) a toolkit for the construction of particular Lorentzian manifolds,
- 2 apply this toolkit in order to construct Gödel-type spacetimes as an interesting laboratory (our methods/ideas are *not* restricted to these spacetimes),
- 3 analyze the behavior of geodesics and light cones for these spacetimes as an example of the visualization techniques.

# Setting

We consider  $(N + 1)$ -dimensional time-oriented smooth Lorentzian manifolds  $(M, g)$  of signature  $(+, -, \dots, -)$ .

Additionally we fix a timelike unit vector field  $X^i$  (observer field).

# Setting

For our analysis the usual decomposition of the covariant derivative of  $X_i$  into irreducible parts will be used:

$$\nabla_k X_i = \omega_{ik} + \sigma_{ik} + \frac{1}{N} \Theta P_{ik} - \dot{X}_i X_k$$

with the antisymmetric part  $\omega_{ik}$  (rotation), the symmetric traceless part  $\sigma_{ik}$  (shear) and the trace  $\Theta$  (expansion) itself.

# Setting

Using the observer field we are able to give particular PDEs relating the metric tensor with the parts of the decomposition before, for example we have

$$\omega_{ik} = \frac{1}{2} (g_{0k} \partial_0 g_{0i} - g_{0i} \partial_0 g_{0k} - \partial_k g_{0i} + \partial_i g_{0k}).$$

# Shear-free spacetimes

	Spatial metric $g_{\alpha\beta}(t) = g_{\alpha\beta}(t, x^\gamma)$
General	$S^2(t)g_{\alpha\beta}(0) + g_{0\alpha}(t)g_{0\beta}(t) - S^2(t)g_{0\alpha}(0)g_{0\beta}(0)$
$\dot{X}_i = 0$	$S^2(t)g_{\alpha\beta}(0) + (1 - S^2(t)) g_{0\alpha}(0)g_{0\beta}(0)$
$\Theta = 0$	$g_{\alpha\beta}(0) + g_{0\alpha}(t)g_{0\alpha}(t) - g_{0\alpha}(0)g_{0\alpha}(0)$
$\Theta = 0, \dot{X}_i = 0$	$g_{\alpha\beta}(0)$

# Starting with the Gödel metric

$$(g_{ik}(0, x^\gamma)) = \begin{pmatrix} 1 & 0 & e^{\sqrt{2}\omega_0 x^1} & 0 \\ 0 & -1 & 0 & 0 \\ e^{\sqrt{2}\omega_0 x^1} & 0 & \frac{1}{2}e^{2\sqrt{2}\omega_0 x^1} & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

Gödel (1949)



# Starting with the Gödel metric

Some properties:

- 5-dim group of isometries which is transitive,
- no imbedded hypersurfaces without boundary which are spacelike everywhere,
- simply connected,
- geodesically complete,
- $\omega \neq 0$ ,  $\sigma = \Theta = 0$ ,
- CTCs, ...

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# Two friends



# Generalized Gödel metric

$$(g_{ik}) = \begin{pmatrix} 1 & 0 & e^{\sqrt{2}\omega_0 x^1} & 0 \\ 0 & -S^2(t) & 0 & 0 \\ e^{\sqrt{2}\omega_0 x^1} & 0 & \frac{1}{2}e^{2\sqrt{2}\omega_0 x^1} (2 - S^2(t)) & 0 \\ 0 & 0 & 0 & -S^2(t) \end{pmatrix}$$

M. Plaue and M. S. (2008)

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# What to do?

- Find geodesics  $\gamma$ , solving  $\ddot{\gamma} = 0$  (with initial position and direction),
- 3-dim representation,
- chose convenient coordinates,
- find a parametric representation for the light cone,
- .... use the package *GeodesicGeometry*, T. Schönfeld (2009).

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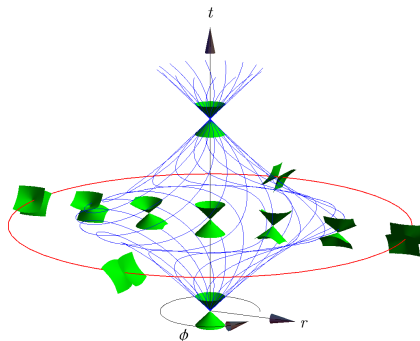
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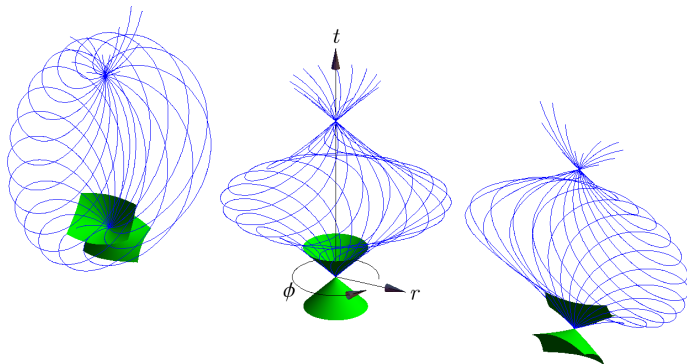
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# Hawking–Ellis (1973)



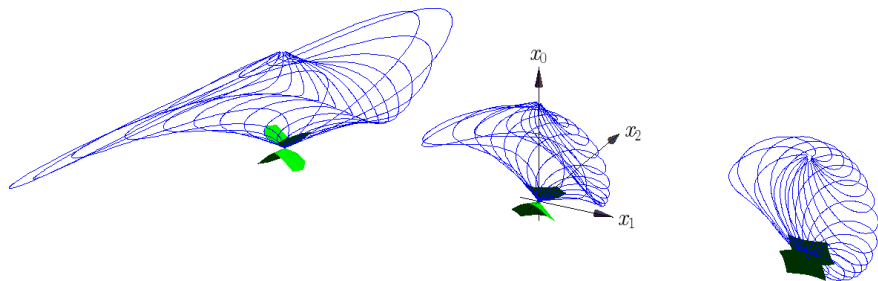
Gödel's spacetime in cylindrical coordinates: A set of geodesics starting from very close to the coordinate origin is shown along with a few light cones which exhibit the tipping effect, showing the existence of CTCs.

# Some more details



Sets of geodesics originating from points with different radial coordinates in Gödel's spacetime. Just as the light cones, the geodesics tip over; the refocusing is not affected.

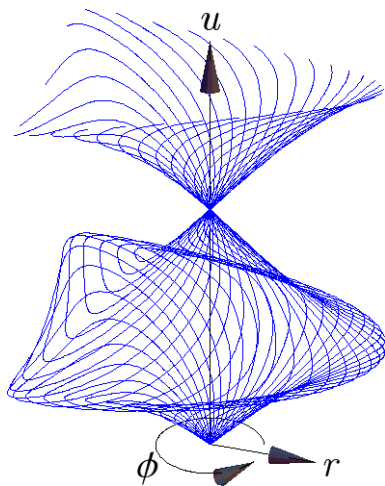




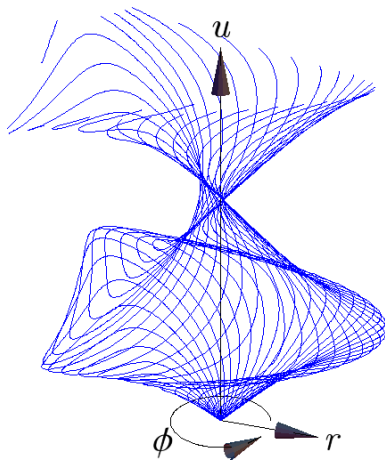
Three sets of geodesics originating from points with different  $x_1$  values in Gödel's spacetime. Just like the light cones, they tip over as one moves to positive  $x_1$  values, and they open up in the  $x_2$  direction as one moves to negative  $x_1$  values.

# Expanding Gödel model (by M. Plaue and M. S.)

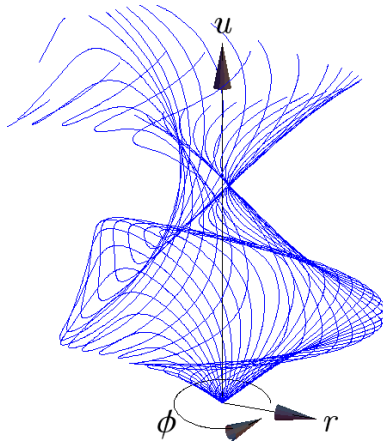
In the following: Six sets of geodesics in the expanding spacetime with identical initial conditions but with different values for the expansion  $\Theta$ .



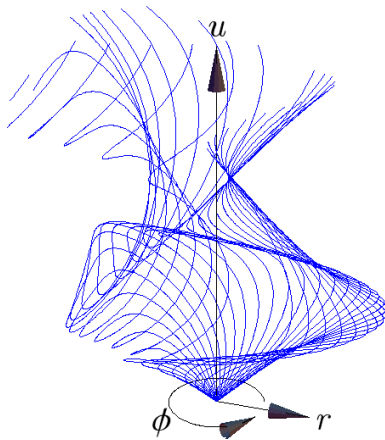
$$\Theta = 0.00$$



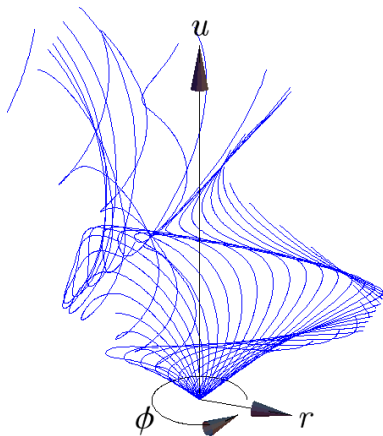
$$\Theta = 0.02$$



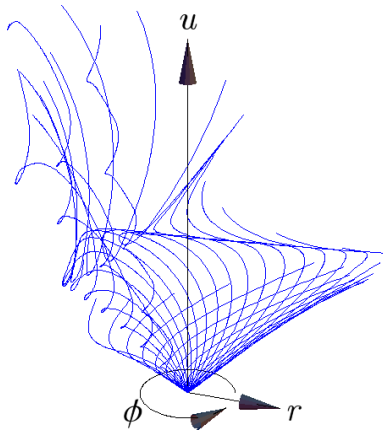
$$\Theta = 0.05$$



$$\Theta = 0.10$$

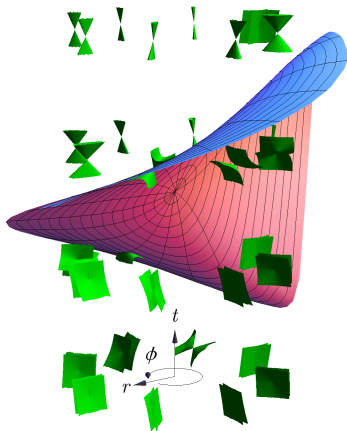


$$\Theta = 0.20$$



$$\Theta = 0.40$$





Causality change in the expanding spacetime with  $\Theta = 0.4$ . The area above the surface is known to be free of closed timelike or null curves.

# Another metric: M. Gürses, M. Plaue and M. S. (2009)

$$g_{ik} = X_i X_k - S^2(t) P_{ik}$$