

Direct Approach to Quantum Tunneling

Anders Andreassen,^{*} David Farhi,[†] William Frost,[‡] and Matthew D. Schwartz[§]

Harvard University, Cambridge, Massachusetts 02138, USA

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The decay rates of quasistable states in quantum field theories are usually calculated using instanton methods. Standard derivations of these methods rely in a crucial way upon deformations and analytic continuations of the physical potential and on the saddle-point approximation. While the resulting procedure can be checked against other semiclassical approaches in some one-dimensional cases, it is challenging to trace the role of the relevant physical scales, and any intuitive handle on the precision of the approximations involved is at best obscure. In this Letter, we use a physical definition of the tunneling probability to derive a formula for the decay rate in both quantum mechanics and quantum field theory directly from the Minkowski path integral, without reference to unphysical deformations of the potential. There are numerous benefits to this approach, from nonperturbative applications to precision calculations and aesthetic simplicity.

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Quantum tunneling is a hallmark of nonclassical physics. In 1D quantum mechanics, the decay of quasistable states can be seen by solving the Schrödinger equation. An example is shown in Fig. 1. A wave function ψ , initially localized in the false-vacuum (FV) region near the point a , will evolve in time to have support in the region labeled R . When the lifetime is long, there is a well-defined decay rate Γ . To see this, we begin with the probability of ψ being found in region R after time T , given by

$$P_R(T) \equiv \int_R dx |\psi(x, T)|^2. \quad (1)$$

This probability for the potential in Fig. 1 is shown in Fig. 2 (computed numerically by solving the Schrödinger equation). Note that there are wiggles in the probability on the short time scale $T_{\text{slosh}} \sim \omega_a^{-1}$, where $m\omega_a^2 = V''(a)$ characterizes the frequency of sloshing around the false vacuum. At intermediate times, the probability is exponential, $P_{\text{FV}}(T) \sim e^{-\Gamma T}$, and we can extract Γ from this regime. At late times, nonlinearities set in, from both the wave function bouncing off the far edge of R and the initial wave function being depleted.

There is no practical way to generalize directly the above procedure, of numerically solving the Schrödinger equation, to compute decay rates in quantum field theory. There are alternative ways to compute the decay rate in 1D [1], such as the WKB approximation, or finding the imaginary part of resonance energies with Gamow-Siegert outgoing boundary conditions [2,3]. However, one cannot easily justify the simplest generalizations of these to field theory either. The only approach that seems to generalize nicely is based on the path integral.

In the approach of Callan and Coleman [4], the decay rate is extracted from the persistence amplitude in Euclidean time:

$$Z(T) = \langle a | e^{-HT} | a \rangle = \sum_n e^{-E_n T} |\psi_n(a)|^2 \xrightarrow{T \rightarrow \infty} e^{-E_0 T}. \quad (2)$$

At large T , the sum is dominated by the ground state energy E_0 , whose imaginary part gives half the decay rate. Writing the matrix element in terms of a path integral, the claim is that

$$\frac{\Gamma}{2} \sim \text{Im} \lim_{T \rightarrow \infty} \frac{1}{T} \ln \int_{x(-T/2)=a}^{x(T/2)=a} \mathcal{D}x e^{-S[x]}. \quad (3)$$

A similar formula results from considering the partition function, which gives a path integral with periodic rather than fixed boundary conditions [5,6].

We write \sim in Eq. (3) because one cannot simply compute the imaginary part: The path integral produces $Z(T)$ of Eq. (2), which is manifestly real. To get an imaginary part, most discussions typically take an example such as the quartic potential $V(x) = x^2 - gx^4$. For $g > 0$, the spectrum for this potential is unbounded from below,

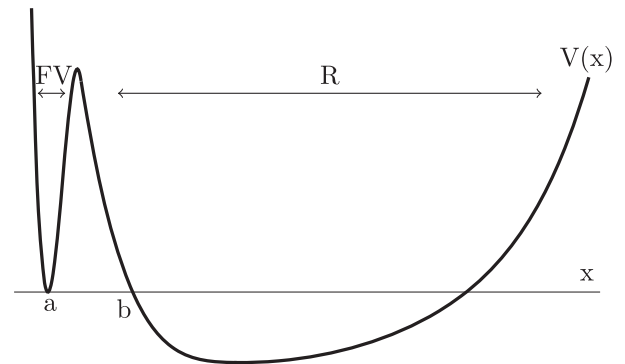


FIG. 1. An example 1D potential exhibiting quantum tunneling from the region FV to the region R .

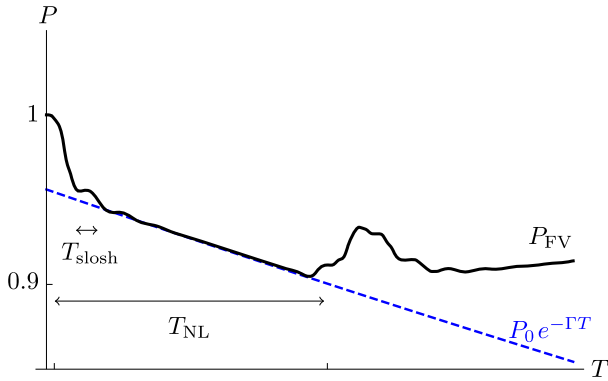


FIG. 2. The probability $P_{\text{FV}}(T)$ of funding a wave function in FV at time t for the toy 1D potential in Fig. 1. This curve is computed by numerically integrating the Schrödinger equation beginning with a Gaussian wave packet localized near a at $t = 0$. The false-vacuum probability falls exponentially, $P_{\text{FV}} \sim e^{-\Gamma T}$, for times intermediate between the false-vacuum sloshing time T_{slosh} and the time when nonlinearities set in and the flux starts flowing back into the FV region.

and $Z(T)$ is infinite. One can nevertheless get a sensible answer for $\text{Im}Z(T)$ with $g > 0$ by an analytic continuation from $g < 0$, where the potential has a stable minimum. The result for this case is that $Z(T)$ has a branch cut for $g > 0$, with $\text{Im}Z(T)$ given by $\pm \frac{1}{2}$ the discontinuity across the cut, which in turn is approximated by the sum over saddle-point expansions of $Z(T)$. These saddle points are finite-action instanton configurations called bounces. Summing over these bounces in the dilute gas approximation, with the extra factor of $\frac{1}{2}$, gives a tunneling rate in quantum mechanics which agrees with the formula from the WKB approximation to at least $\mathcal{O}(\hbar^2)$ [7].

The divergence of $Z(T)$ for the quartic case, following from the unboundedness of the potential, is not the general justification of the required deformation. In fact, all physical potentials are bounded from below. For these, like $V(x) = x^2 - gx^4 + \lambda x^6$ or the one sketched in Fig. 1, there is a ground state E_0 which is real. Then at large T , Eq. (2) picks up the true ground state, associated with the minimum of region R . In the path integral, the dominant path is neither the bounce nor the static solution $x(t) = a$ but rather a path which starts at a , quickly goes to the minimum of region R , stays there for most of T , and then returns to a . This path, which we call *the shot*, has very little to do with the decay but is in fact the result of Eq. (2).

The general procedure to get an imaginary part is to transplant the integration contour from its original trajectory along real paths including the false vacuum, the bounce, and the shot, to a contour along complex paths passing through the false-vacuum saddle point [5,8,9]. Thereby, the contribution of the shot is eliminated. This realignment of the contour agrees with the result of analytic continuation in cases commonly considered, where the potential is unbounded from below, but is different for

physical potentials with ground state energies. Half of these complex paths coincide with paths along the steepest-descent contour passing through the bounce. This procedure results in an algorithm: The decay rate is given by the imaginary part of an asymptotic expansion around the bounce saddle point, multiplied by $\frac{1}{2}$.

Unfortunately, there seems to be no proof in the literature that the final algorithm will always give the decay rate. What seems clear is that there is a mathematically consistent way to define the imaginary part for the real quantity $Z(T)$ in the $T \rightarrow \infty$ limit. However, there is a surprising lack of commentary on the connection between this imaginary quantity and the physical decay rate. Partly, this may be because much of the interest in the path integral formulation of tunneling is related to unraveling non-perturbative elements of quantum mechanics and quantum field theory. (For example, although the quartic potential with $g > 0$ is stable, the ground state energy as a function of g is nonanalytic, with zero radius of convergence. The finite-action instanton configurations for negative g describe the poles in the Borel transform of the positive g series.) Partly, it may be because in quantum mechanics the final prescription can be proven equivalent to WKB at next-to-leading order (NLO), and there has been no reason to doubt the prescription in quantum field theory and no need for ultraprecise calculations. However, for questions about gauge invariance of the decay rate [10–13] or for situations where the potential is itself generated by quantum effects, having a cleaner and more precise derivation may be helpful. Moreover, since the derivation is so far removed from the physical problem, one cannot help but wonder if there might be a less byzantine connection between the path integral and the decay rate. It is the purpose of this Letter to provide such a connection.

From our initial discussion of the decay, as in Fig. 2, we see that there are two characteristic times for physical potentials: a sloshing time T_{slosh} associated with oscillations around the false vacuum and a nonlinearity time T_{NL} associated with the structure of the potential around the true vacuum, on which time scales the picture of probability flowing linearly out of FV breaks down. The decay rate, which characterizes the linear fall during the intermediate time, can be extracted from Eq. (1) through the double limit

$$\Gamma \equiv - \lim_{\substack{T/T_{\text{NL}} \rightarrow 0 \\ T/T_{\text{slosh}} \rightarrow \infty}} \frac{1}{P_{\text{FV}}(T)} \frac{d}{dT} P_{\text{FV}}(T). \quad (4)$$

The limits in this definition are mutually consistent only if $T_{\text{NL}} \gg T_{\text{slosh}}$; i.e., the time scale for the system to leave the destination region and/or reenter the false-vacuum region must be much longer than the time scales of sloshing around in the false vacuum. If this is not the case, then there is no use in characterizing any part of the probability function $P_{\text{FV}}(T)$ with a “decay rate”.

Formula (4) easily generalizes to the case where there are multiple directions for decay. In that situation, we might be interested in the total decay rate Γ as well as the decay rate Γ_R to a particular region R of configuration space, which is not necessarily the entire complement of the false-vacuum region. Then we should look not at the fractional linear falloff of P_{FV} but rather the fractional linear rise of P_R , the probability to find the system in R . Then

$$\Gamma_R \equiv \lim_{\substack{T/T_{\text{NL}}^R \rightarrow 0 \\ T/T_{\text{slosh}} \rightarrow \infty}} \frac{1}{P_{\text{FV}}(T)} \frac{d}{dT} P_R(T), \quad (5)$$

where now T_{NL}^R is the time scale for the system to leave R (back to FV or to any other wells that may be present). The above definition can be recast in terms of the conserved quantum mechanical probability flux if desired.

For potentials of interest, where $T_{\text{NL}}^R \gg T_{\text{slosh}}$, the exact energy eigenstates that have substantial support in the FV region are in bands around resonance energies E_i with widths Γ_i . The higher-energy bands will decay exponentially faster than the lower-energy ones, so we will focus on the lowest energy band. Any wave function with support on this band will have the same decay rate. Thus, for simplicity, we can begin with a position eigenstate localized at exactly the minimum of the false-vacuum well: $\psi(x, 0) = \delta(x - a)$. For $t > 0$, the wave function is simply the propagator, and so

$$P_R(T) = \int_R dx_f |\mathcal{N}D(a, 0; x_f, T)|^2 \quad (6)$$

$$= \int_R dx_f \left| \mathcal{N} \int_{x(0)=a}^{x(T)=x_f} \mathcal{D}x e^{iS[x]} \right|^2. \quad (7)$$

For later convenience, we have written the propagator as $\mathcal{N}D$ rather than simply D , where \mathcal{N} is the path integral normalization defined by the second line.

To proceed, we will use the identity that, for any point b between a and x_f , we can label the *first* time the path hits b by t_0 and write

$$D(a, 0; x_f, T) \equiv \int_0^T dt_0 \bar{D}(a, 0; b, t_0) D(b, t_0; x_f, T), \quad (8)$$

where

$$\bar{D}(a, 0; b, t_0) = \int_{x(0)=a} \mathcal{D}x e^{iS[x]} \delta(t_b[x] - t_0) \quad (9)$$

with $t_b[x]$ the functional that returns the first time a path $x(t)$ hits the point b and $\delta(t_b[x] - t_0) = 0$ if $x(t)$ never hits b . Equivalently,

$$\bar{D}(a, 0; b, t_0) = \int_{\substack{x(0)=a \\ x(t_0)=b \\ x(t)<b}} \mathcal{D}x e^{iS[x]} |\dot{x}(t_0)|. \quad (10)$$

Inserting Eq. (8) twice into Eq. (6) gives

$$\begin{aligned} P_R(T) &= \mathcal{N}\mathcal{N}^* \int dt_0 \int dt'_0 \\ &\times \int_R dx_f \bar{D}(a, 0; b, t_0) \bar{D}^*(a, 0; b, t'_0) \\ &\times D(b, t_0; x_f, T) D^*(b, t'_0; x_f, T). \end{aligned} \quad (11)$$

Now, recall the condition $T \ll T_{\text{NL}}^R$ in Eq. (5). In the limit $T/T_{\text{NL}}^R \rightarrow 0$, no flux enters back into the FV region. Note that x_f appears only in the propagators to and from b , so if we choose b to be the classical turning point, where $V(a) = V(b)$, we can use the condition $T \ll T_{\text{NL}}^R$ to justify replacing the integral of x_f over R with an integral of x_f over all of configuration space; the added probability from the propagator to points outside of R is negligible, since it is suppressed by extra factors of the WKB penetration factor. With the restriction on R removed, we can use the completeness of $|x_f\rangle$ to combine the two propagators in the third line of Eq. (11) into a single $D(b, t_0; b, t'_0)$, which can then be recombined with the remaining two \bar{D} factors using Eq. (8) in reverse. This leads to

$$P_R(T) = \mathcal{N}\mathcal{N}^* \int_0^T dt_0 D(a, 0; b, t_0) \bar{D}^*(a, 0; b, t_0) + \text{c.c.} \quad (12)$$

Finally, we find Γ_R by differentiating with respect to T , dividing by P_{FV} , and taking $T \rightarrow \infty$:

$$\Gamma_R = \lim_{T \rightarrow \infty} \frac{D(a, 0; b, T) \bar{D}^*(a, 0; b, T) + \text{c.c.}}{\int dx |D(a, 0; x, T)|^2}. \quad (13)$$

Taking $T \rightarrow \infty$ enforces the other limit $T \gg T_{\text{slosh}}$ in Eq. (5), which is not in conflict with the $T \ll T_{\text{NL}}^R$ limit, since we have already enforced that limit to get to Eq. (12). This is an exact nonperturbative definition of the decay rate independent of the saddle-point approximation.

To make sense of Eq. (13), let us check that it agrees with existing results at NLO. First, we Wick rotate to imaginary time. This provides a further simplification, since the two remaining propagators can then be combined [9]:

$$\Gamma_R = 2\text{Im} \lim_{T \rightarrow \infty} \left(\frac{\int \mathcal{D}x e^{-S_E[x]} \delta(\tau_b[x])}{\int \mathcal{D}x e^{-S_E[x]}} \right)_{T \rightarrow iT}, \quad (14)$$

where both path integrals have boundary condition $x(\pm T) = a$. The functional $\tau_b[x]$ is identical to $t_b[x]$, but we have renamed it as a reminder that x is now a function of Euclidean time τ . The “Im” arises because the δ or \dot{x} in \bar{D} in Eq. (13), with dimensions of time, acquires an i , which causes a relative sign between the two complex conjugate terms.

To proceed with the NLO expansion, we locate the extrema of the Euclidean action. Because of the δ function, all paths must hit b at $\tau = 0$. Thus, the $x(\tau) = a$ extremum is removed, but the bounce $\bar{x}(\tau)$ remains. [Although the shot is removed as well, there is a modified shot $x_s(\tau)$ which is identical to the bounce until it hits b and then shoots into R , eventually returning to a . This modified shot has Euclidean action $S_E[x_s] \approx -T|E_0| + S_1$ with S_1 coming from when the shot is moving fast. Although $S_E[x_s] \ll S_E[\bar{x}]$ at large real T , after $T \rightarrow iT$, the contribution of the modified shot is exponentially suppressed, since $S_1 > S_E[\bar{x}]$. Thus, the bounce is the dominant contribution to the rate.] A minor pleasing feature of this derivation is that approximate solutions, such as time-shifted bounces or multiple bounces, play no role.

To integrate over the δ function carefully, we use the collective coordinate τ_0 . Fluctuations around the bounce are parameterized by τ_0 and $\{\zeta_i\}_{i>0}$, as $x(\tau, \tau_0, \zeta_i) = \bar{x}(\tau - \tau_0) + \sum_{i>0} \zeta_i x_i(\tau - \tau_0)$ so that $\tau_b[x] = \tau_0 + \tau_b[x(\tau, 0, \zeta_i)]$. At NLO, the Jacobian generated by changing to $\{\tau_0, \zeta_i\}$ is $J = \sqrt{S_E(\bar{x})/m}$ [4].

The action is independent of τ_0 ; hence, the integral over τ_0 is equivalent to multiplication by J if the path $x(\tau, 0, \zeta_i)$ hits b and 0 otherwise. This leaves the remaining Gaussian fluctuations with the extra boundary condition that all paths must hit b . Around $\bar{x}(t)$, which hits b at its maximum, this restriction adds $\Theta[\zeta_i x_i(0)]$ to our path integral. At NLO, the action is symmetric in the ζ fluctuations, so this restriction to half of the domain can be removed if a factor of $\frac{1}{2}$ is added. This leaves a product of Gaussian integrals over the ζ , whose result is

$$\Gamma_R = e^{-S[\bar{x}]} \sqrt{\frac{S_E[\bar{x}]}{2\pi}} \sqrt{\frac{\det\{-\partial_\tau^2 + V''[a]\}}{-\det'\{-\partial_\tau^2 + V''[\bar{x}(\tau)]\}}}. \quad (15)$$

Here \det' is the usual Fredholm determinant with the zero mode removed, as in Ref. [4]. The minus sign comes from the negative-eigenvalue fluctuation which is still present and contributes to the determinant without ado. This formula is identical to the NLO formula following from Eq. (3) and agrees with the WKB approximation.

The above derivation generalizes naturally to multiple-dimensional quantum mechanics and to field theory. The only difference is that one needs to integrate b over a surface Σ that bounds the destination region R (where Σ and

R are now regions in field configuration space). In field theory, Eq. (14) becomes

$$\frac{\Gamma_R}{V} = \frac{1}{V} 2\text{Im} \frac{\int \mathcal{D}\phi e^{-S_E[\phi]} \delta(\tau_\Sigma[\phi])}{\int \mathcal{D}\phi e^{-S_E[\phi]}}. \quad (16)$$

Here the path integrals have boundary conditions $\phi(\vec{x}, T = \pm\infty) = \phi_a$, where ϕ_a is the false vacuum, and the same analytic continuation and limit as in Eq. (14) are understood. The functional $\tau_\Sigma[\phi(\vec{x}, \tau)]$ returns the first time τ at which the configuration $\phi(\vec{x}, \tau)$ hits the surface Σ .

Often the natural choice for Σ is the *turning point surface*: the set of configurations $\phi(\vec{x})$ satisfying $U[\phi(\vec{x})] = U[\phi_a]$ where the classical energy functional is

$$U[\phi(\vec{x})] = \int d^3x \left[\frac{1}{2} (\vec{\nabla}\phi)^2 + V(\phi) \right]. \quad (17)$$

Solving for Σ will usually indicate one or more regions R disconnected from FV where Eq. (16) can be used.

An important case where Σ cannot be computed directly from V is when tunneling is due to radiative corrections [14]. For example, in the standard model, the leading-order Higgs potential $V(h) \sim \lambda h^4$ with $\lambda > 0$ at the weak scale. The $U = 0$ surface is then simply $h = 0$, and there is no tunneling. But at high energy $\lambda < 0$, and there is tunneling. This tunneling can be seen if the effective potential V_{eff} is used instead of V in Eq. (17). However, using V_{eff} to compute tunneling rates is incorrect: One cannot compute the path integral using V_{eff} without double counting. The formulation we described provides an alternative: One can use Eq. (16) with the classical potential V to compute the tunneling rate.

In conclusion, let us summarize some distinctions between the direct approach described here and the conventional potential-deformation method. There are many small technical differences, such as the irrelevance of approximate bounce solutions, multiple bounces, and the dilute gas approximation, but also some larger conceptual differences.

First, in our derivation, the role of the two relevant time scales for the decay to be well defined, T_{NL} and T_{slosh} , is undeniable. Taking $T \ll T_{\text{NL}}$ prevents tunneling back into the false-vacuum region. The limit lets us approximate the integral over R as an integral over all space and is an exponentially small unitarity-violating approximation morally equivalent to the Gamow-Siegert outgoing boundary conditions. In the potential-deformation method, T_{NL} is never invoked. It does, however, also play a critical role there as well: Because T_{NL} is finite, one must analytically continue the potential so that the FV is absolutely stable, then take $T \rightarrow \infty$, and then analytically continue back. This is how the shot contribution is avoided and why one *must* analytically continue the potential to get the decay rate from the partition function.

A second important difference is that in our approach the only analytic continuation used is the usual Wick rotation to imaginary time. When the potential is analytically continued, the sum over instantons gives the discontinuity across the cut, which differs from the naive imaginary part by a factor of $\frac{1}{2}$. In our case, the factor of $\frac{1}{2}$ arises because only half of the fluctuations around the bounce reach b .

Not having to analytically continue the potential implies that one could also, in principle at least, evaluate the decay rate numerically in a quantum field theory. For example, one could use Eq. (13), in principle, to compute the rate for α decay from first principles in QCD on the lattice.

Finally, it is worth recalling that the generalization of the Callan-Coleman approach to the multidimensional case is often done with a leap of faith. Partly, this is because it is meant to agree with WKB and the generalization of WKB to decays in more than one dimension is complicated [15]; partly, it is because studying the analytic continuation of an energy functional of fields with multiple stable regions is grossly complicated. In contrast, the direct derivation described here is fundamentally path-integral based so the multiple-dimensional case and field theory versions are essentially identical.

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*anders@physics.harvard.edu

†farhi@physics.harvard.edu

‡wfrost@physics.harvard.edu

§schwartz@physics.harvard.edu

- [1] M. Razavy, *Quantum Theory of Tunneling* (World Scientific, Singapore, 2013).
- [2] G. Gamow, *Z. Phys.* **51**, 204 (1928).
- [3] A. J. Siebert, *Phys. Rev.* **56**, 750 (1939).
- [4] C. G. Callan, Jr. and S. R. Coleman, *Phys. Rev. D* **16**, 1762 (1977).
- [5] J. Zinn-Justin, *Int. Ser. Monogr. Phys.* **113**, 1 (2002).
- [6] M. Mariño, *Instantons and Large N: An Introduction to Non-Perturbative Methods in Quantum Field Theory* (Cambridge University Press, Cambridge, England, 2015).
- [7] U. D. Jentschura and J. Zinn-Justin, *Ann. Phys. (Berlin)* **326**, 2186 (2011).
- [8] E. J. Weinberg, *Classical Solutions in Quantum Field Theory*, Cambridge Monographs on Mathematical Physics (Cambridge University Press, Cambridge, England, 2015).
- [9] A. Andreassen, D. Farhi, W. Frost, and M. D. Schwartz, [arXiv:1604.06090](https://arxiv.org/abs/1604.06090).
- [10] D. Metaxas and E. J. Weinberg, *Phys. Rev. D* **53**, 836 (1996).
- [11] A. Andreassen, W. Frost, and M. D. Schwartz, *Phys. Rev. Lett.* **113**, 241801 (2014).
- [12] A. Andreassen, W. Frost, and M. D. Schwartz, *Phys. Rev. D* **91**, 016009 (2015).
- [13] A. D. Plascencia and C. Tamarit, *J. High Energy Phys.* **10** (2016) 099.
- [14] S. R. Coleman and E. J. Weinberg, *Phys. Rev. D* **7**, 1888 (1973).
- [15] T. Banks, C. M. Bender, and T. T. Wu, *Phys. Rev. D* **8**, 3346 (1973).