

# Exploring The (Metric) Space of Collider Events

Harvard Particle Physics Lunch Talk

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Massachusetts Institute of Technology

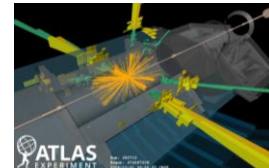
Joint work with Patrick Komiske and Jesse Thaler

[\[1902.02346\]](#)

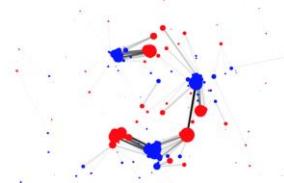
February 13, 2019

# Outline

## Part I Introduction



When are two events similar?

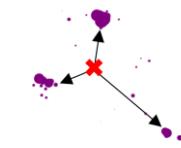


The Energy Mover's Distance

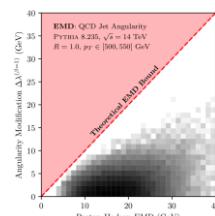


Movie Time

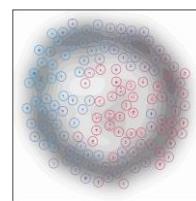
## Part II Applications



Observables



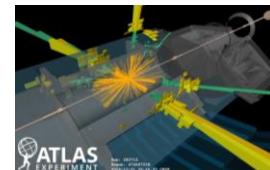
Quantifying event modifications



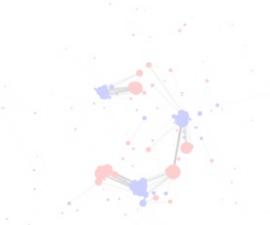
Exploring the Space of Events

# Outline

## Part I Introduction



When are two events similar?



The Energy Mover's Distance

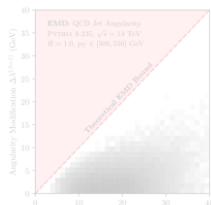


Movie Time

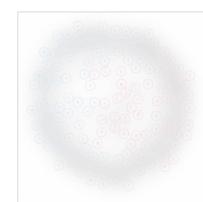
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Observables

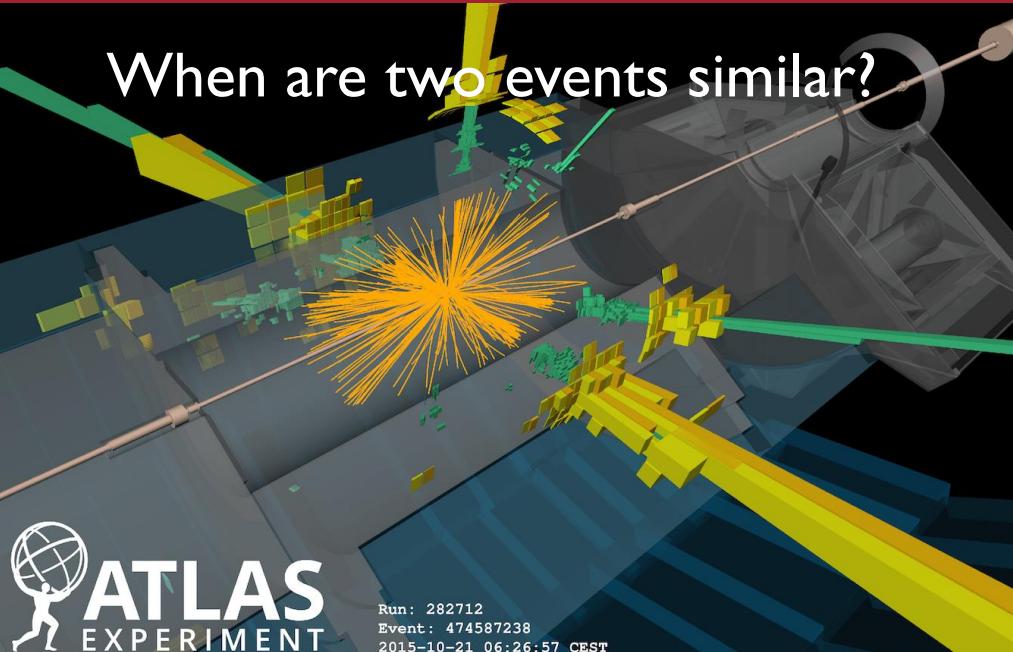


Quantifying event modifications

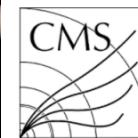


Exploring the Space of Events

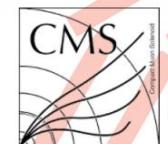
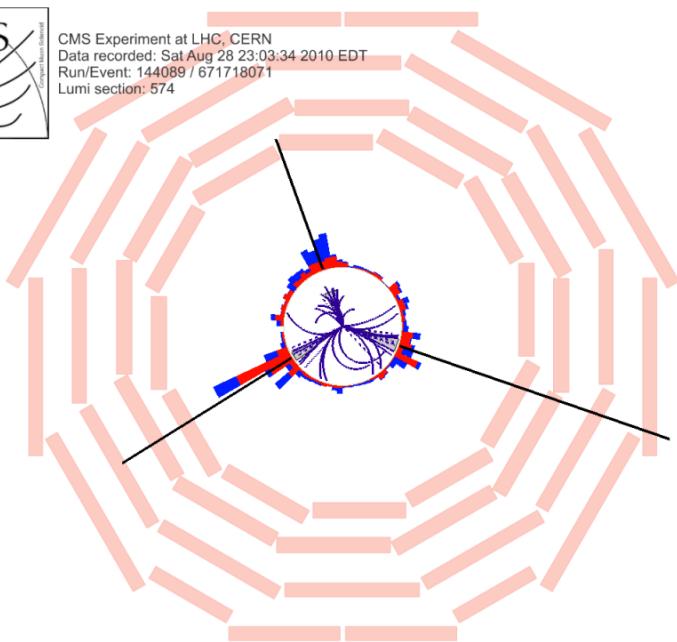
# When are two events similar?



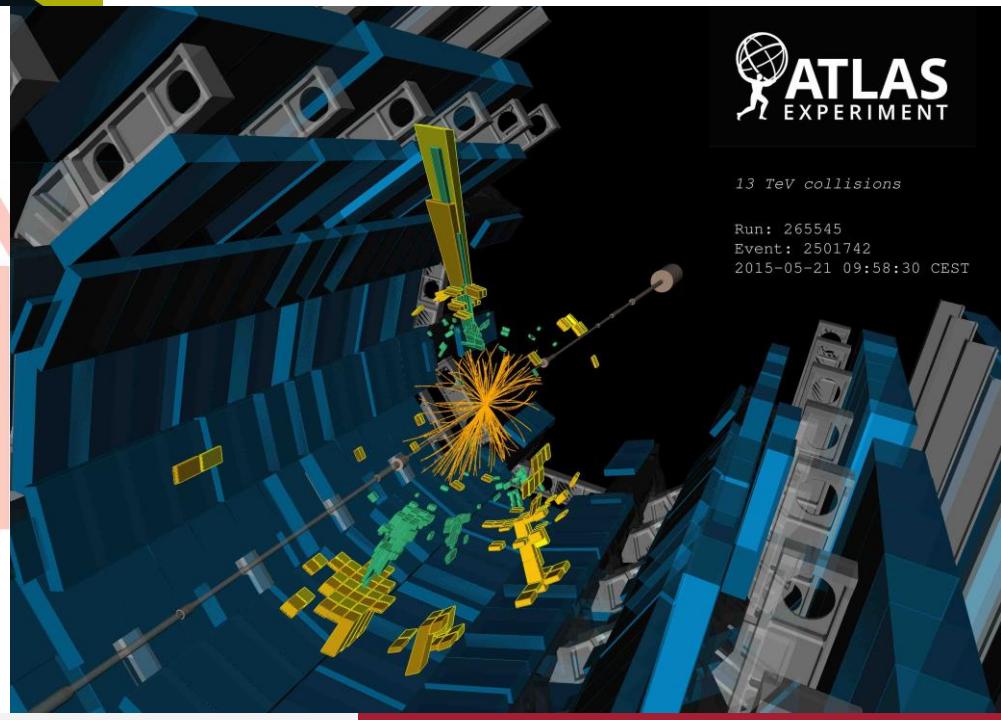
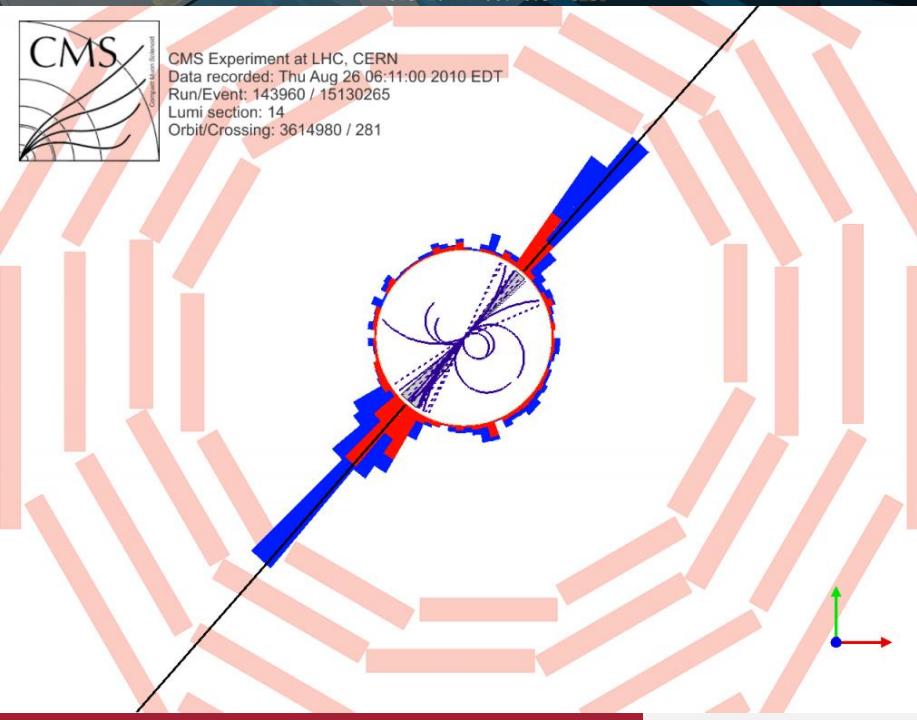
Run: 282712  
Event: 474587238  
2015-10-21 06:26:57 CEST



CMS Experiment at LHC, CERN  
Data recorded: Sat Aug 28 23:03:34 2010 EDT  
Run/Event: 144089 / 671718071  
Lumi section: 574



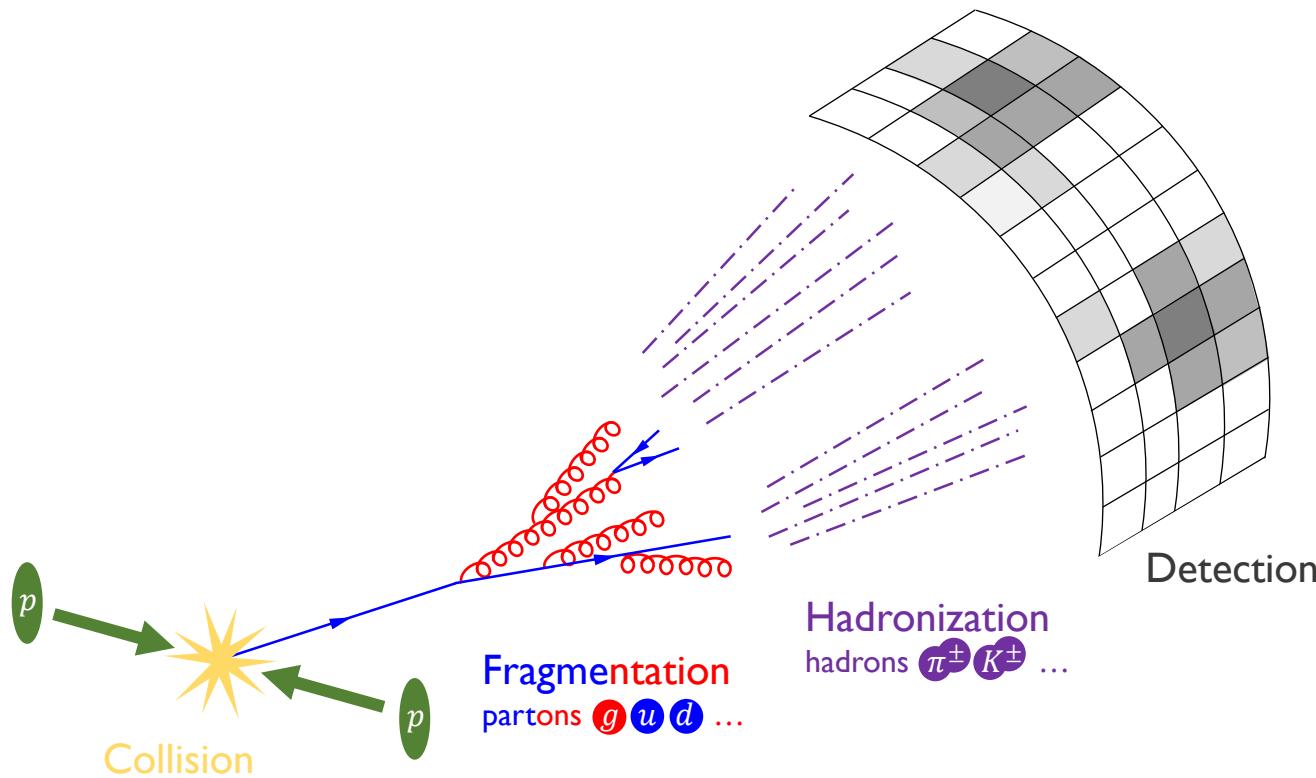
CMS Experiment at LHC, CERN  
Data recorded: Thu Aug 26 06:11:00 2010 EDT  
Run/Event: 143960 / 15130265  
Lumi section: 14  
Orbit/Crossing: 3614980 / 281



13 TeV collisions  
Run: 265545  
Event: 2501742  
2015-05-21 09:58:30 CEST

# When are two collider events similar?

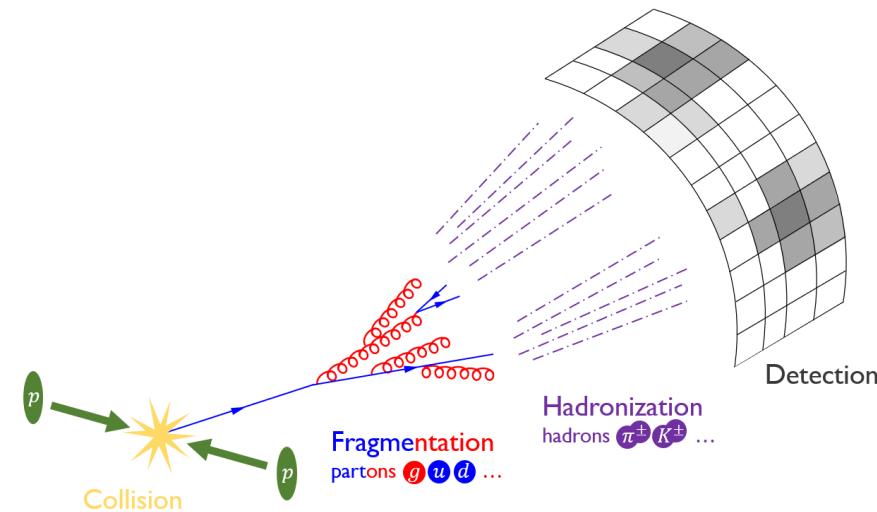
*How an event gets its shape*



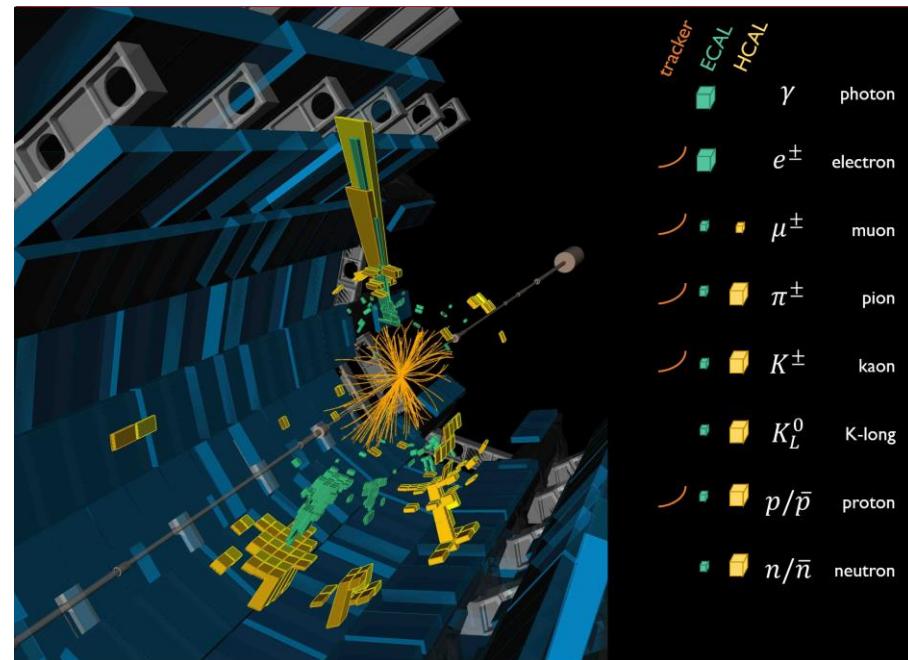
# When are two collider events similar?

A *collider event* is...

Theoretically: very complicated



Experimentally: very complicated



However:

The *energy flow* (distribution of energy) is the information that is robust to:  
fragmentation, hadronization, detector effects, ...

[\[N.A. Sveshnikov, F.V. Tkachov, 9512370\]](#)

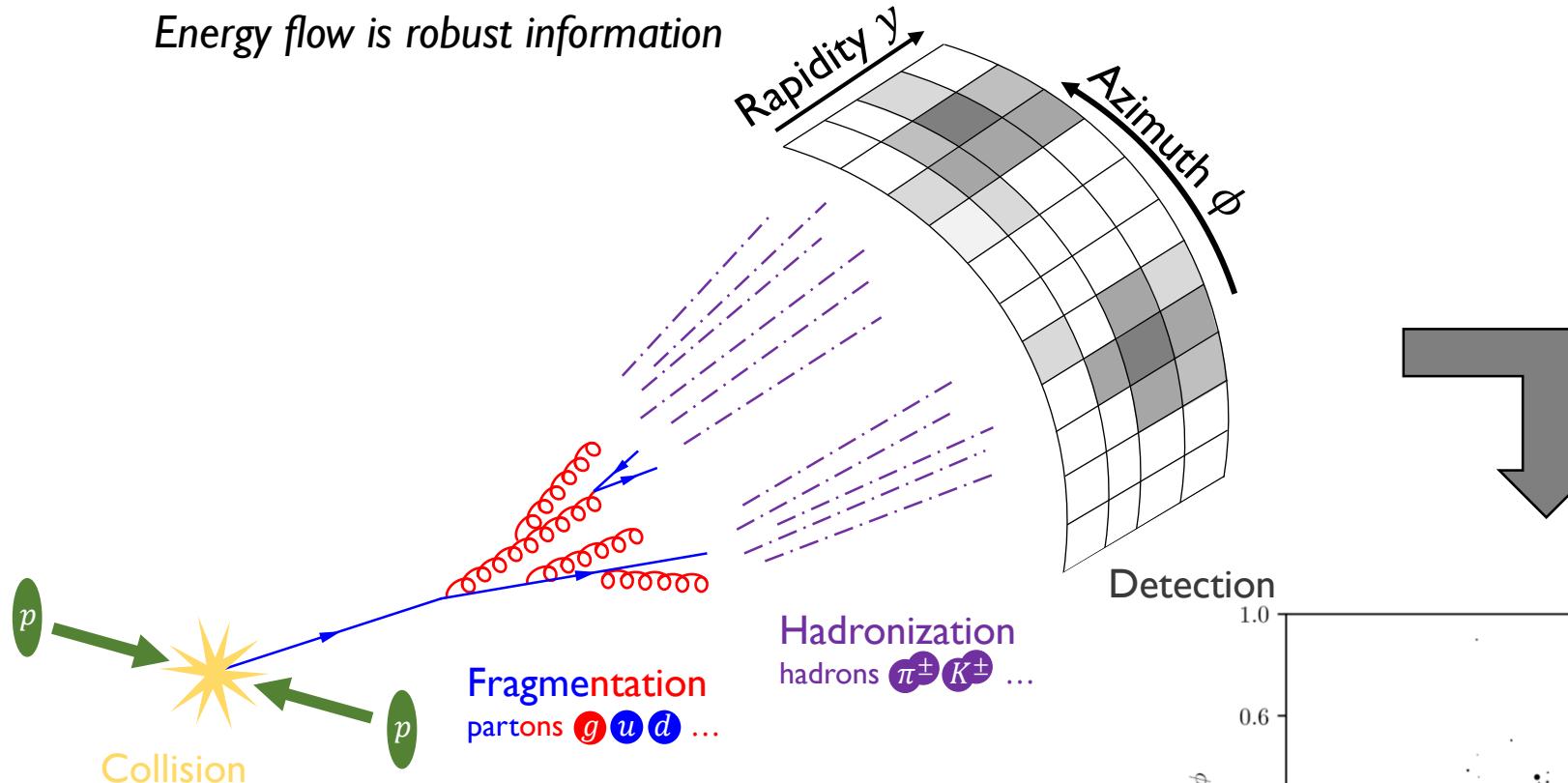
[\[F.V. Tkachov, 9601308\]](#)

[\[P.S. Cherzor, N.A. Sveshnikov, 9710349\]](#)

Energy flow  $\Leftrightarrow$  Infrared and Collinear (IRC) Safe information

# When are two collider events similar?

*Energy flow is robust information*

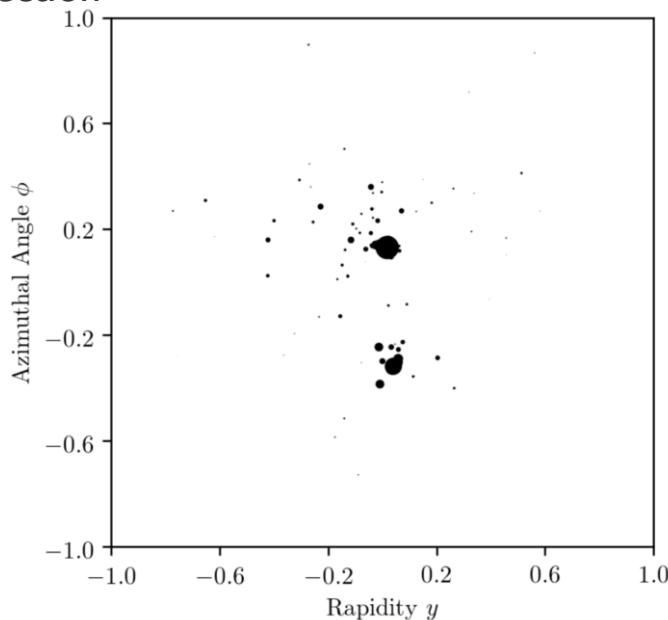


Treat events as distributions of energy:

Ignoring particle flavor, charge, multiplicity...

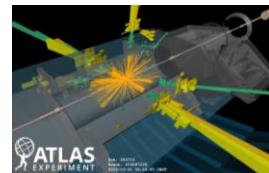
$$\sum_{i=1}^M E_i \delta(\hat{p}_i)$$

↑  
energy      ↑  
                direction



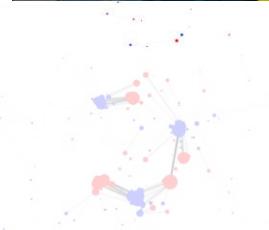
# Outline

## Part I Introduction



When are two events similar?

*When they have similar distributions of energy*



The Energy Mover's Distance

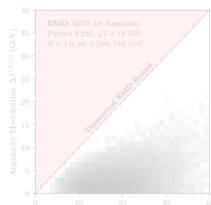


Movie Time

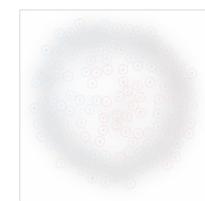
## Part II Applications



Observables



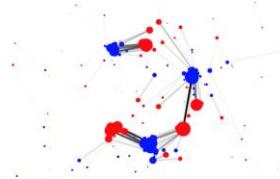
Quantifying event modifications



Exploring the Space of Events

# Outline

## Part I Introduction



When are two events similar?

*When they have similar distributions of energy*

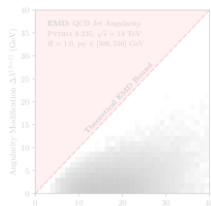
## The Energy Mover's Distance

Movie Time

## Part II Applications



Observables



Quantifying event modifications



Exploring the Space of Events

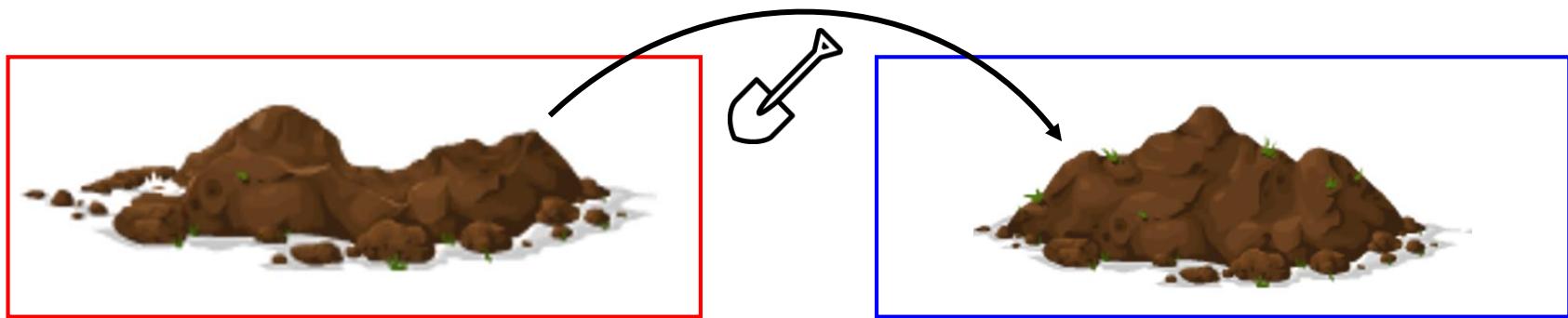
# The Energy Mover's Distance

Review: *The Earth Mover's Distance*

**Earth Mover's Distance:** the minimum “work” ( $\text{stuff} \times \text{distance}$ ) to rearrange one pile of dirt into another

[S. Peleg, M. Werman, H. Rom]

[Y. Rubner, C. Tomasi, and L.J. Guibas]



Metric on the space of (normalized) distributions: *symmetric, non-negative, triangle inequality*

Distributions are close in EMD  $\Leftrightarrow$  their expectation values are close.

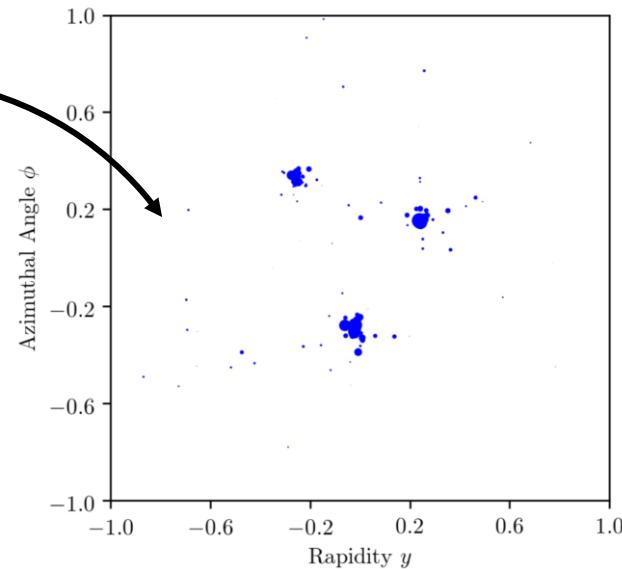
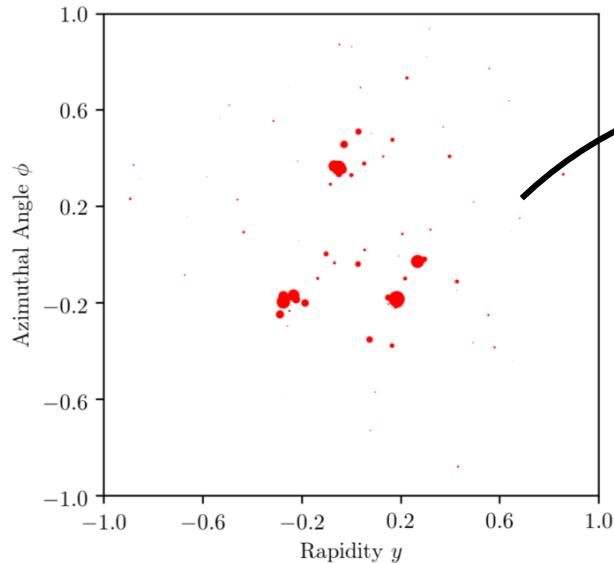
Also known as the 1-Wasserstein metric.

# The Energy Mover's Distance

From Earth to Energy

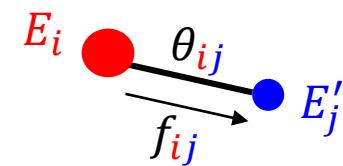
**Energy Mover's Distance:** the minimum “work” (**energy**  $\times$  angle) to rearrange one event (pile of energy) into another

[P.T. Komiske, EMM, J. Thaler, 1902.02346]



$$\text{EMD}(\mathcal{E}, \mathcal{E}') = \min_{\{f\}} \sum_{i=1}^M \sum_{j=1}^{M'} f_{ij} \frac{\theta_{ij}}{R}$$

Difference in  
radiation pattern

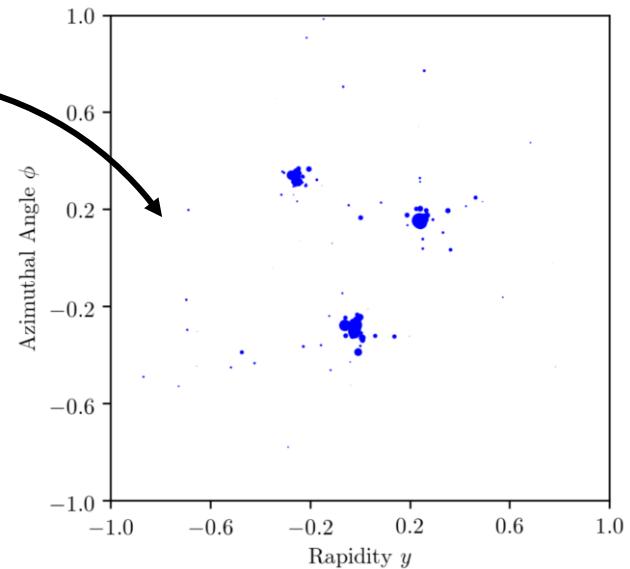
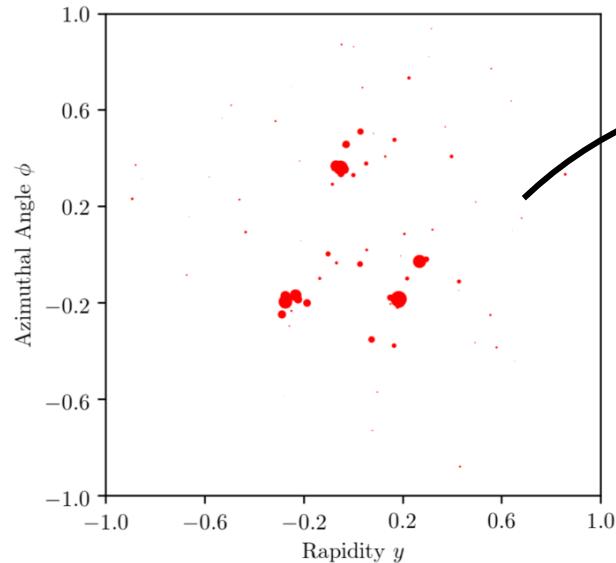


# The Energy Mover's Distance

From Earth to Energy

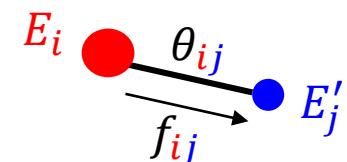
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$$\text{EMD}(\mathcal{E}, \mathcal{E}') = \min_{\{f\}} \sum_{i=1}^M \sum_{j=1}^{M'} f_{ij} \frac{\theta_{ij}}{R} + \left| \sum_{i=1}^M E_i - \sum_{j=1}^{M'} E'_j \right|$$

Difference in radiation pattern      Difference in total energy

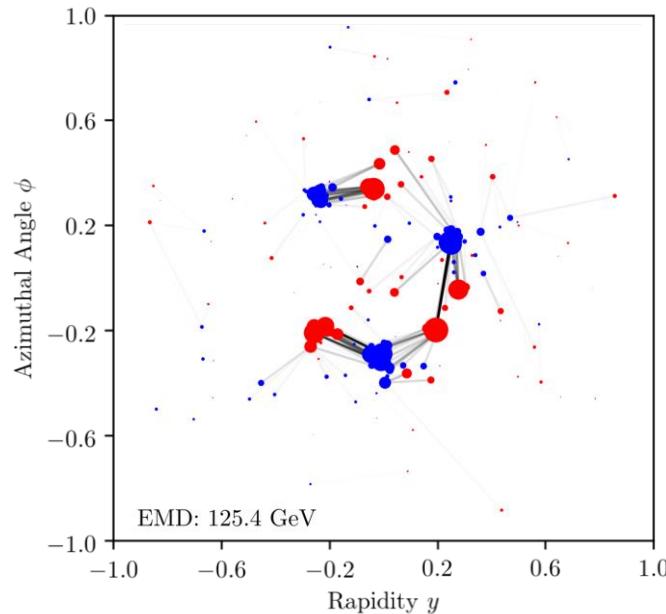


# The Energy Mover's Distance

From Earth to Energy

**Energy Mover's Distance:** the minimum “work” (**energy**  $\times$  angle) to rearrange one event (pile of energy) into another

[P.T. Komiske, EMM, J. Thaler, 1902.02346]



EMD has dimensions of energy

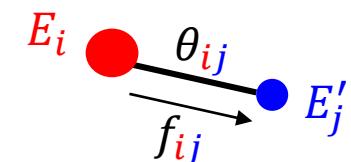
True metric as long as  $R \geq \frac{1}{2} \theta_{\max}$   
 $R \geq$  the jet radius, for conical jets

Solvable via Optimal Transport problem.

~1ms to compute EMD for two jets with 100 particles.

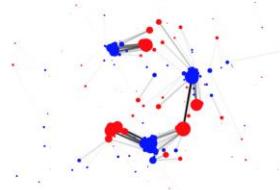
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Difference in radiation pattern      Difference in total energy



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When are two events similar?

*When they have similar distributions of energy*

## The Energy Mover's Distance

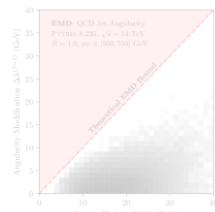
*Work to rearrange one event into another.*

Movie Time

## Part II Applications



Observables



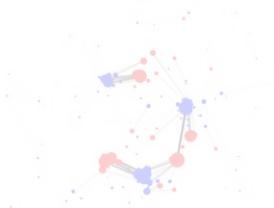
Quantifying event modifications



Exploring the Space of Events

# Outline

## Part I Introduction



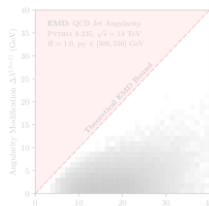
When are two events similar?  
*When they have similar distributions of energy*

## The Energy Mover's Distance

*Work to rearrange one event into another.*

## Movie Time

## Part II Applications



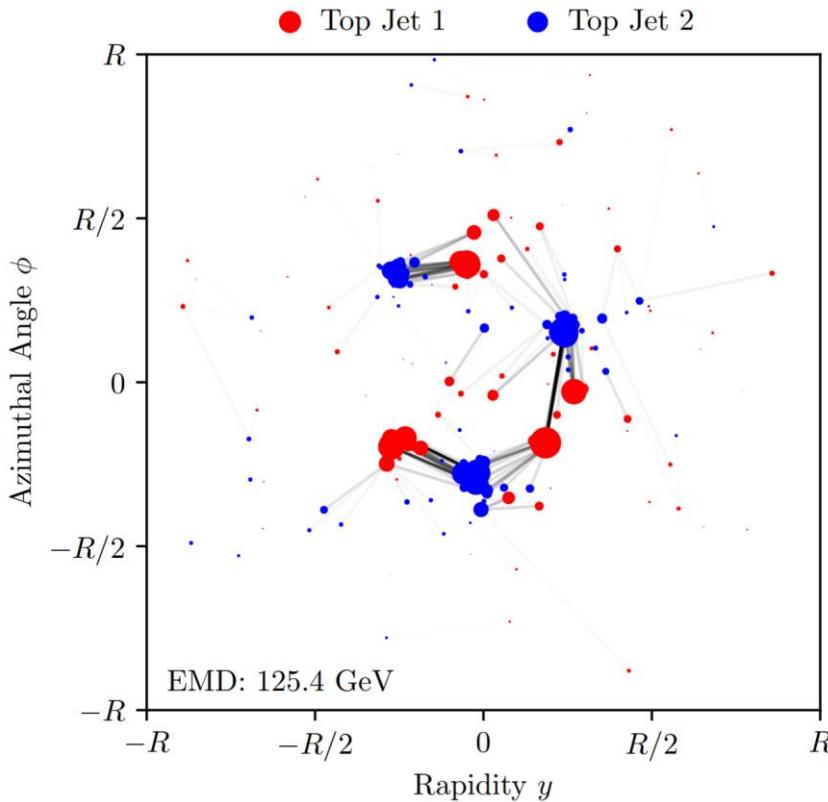
## Observables

## Quantifying event modifications

## Exploring the Space of Events

# Movie Time: Visualizing the EMD

*Taking a walk in the space of events*

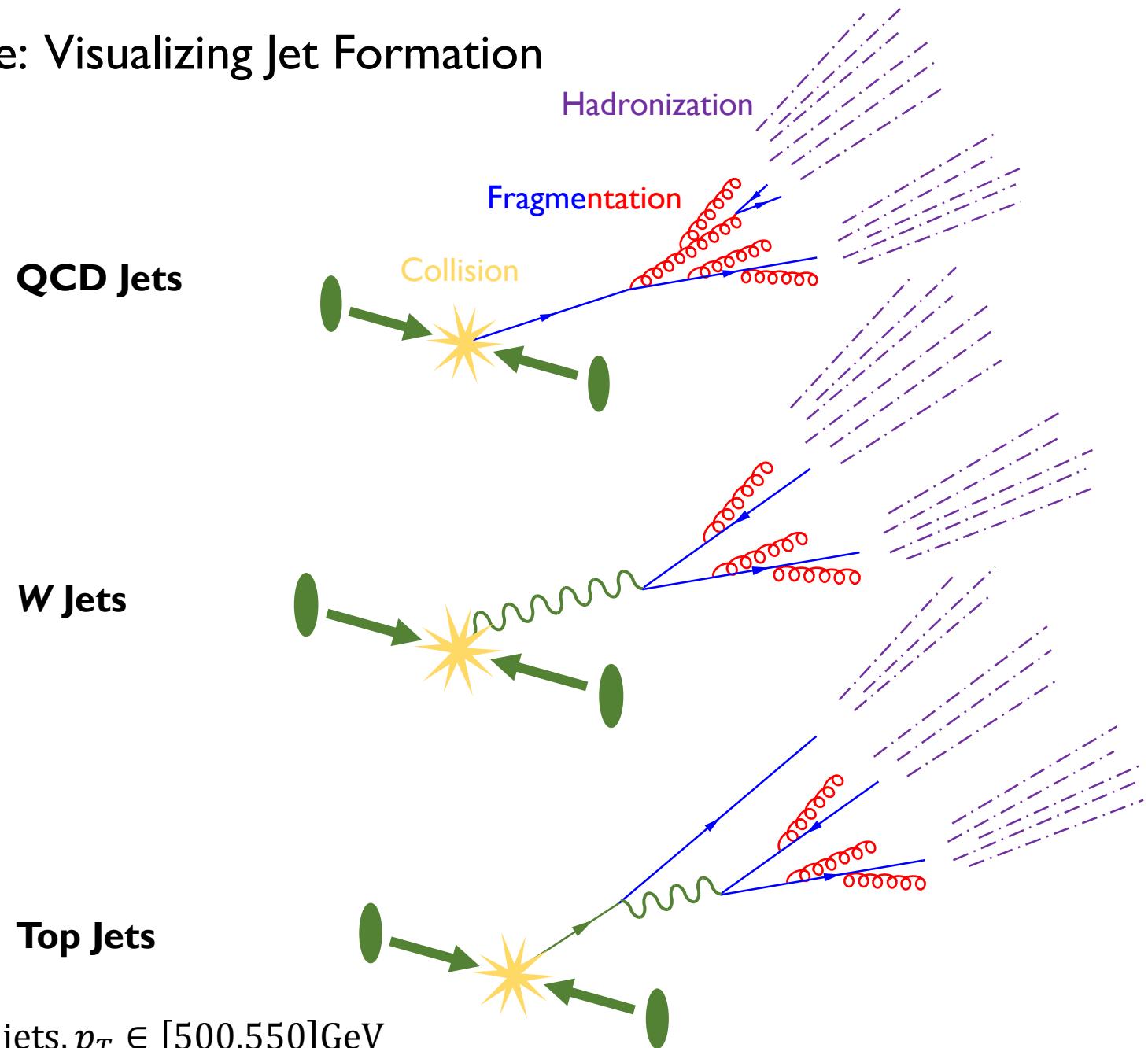


EMD is the cost of an optimal transport problem.

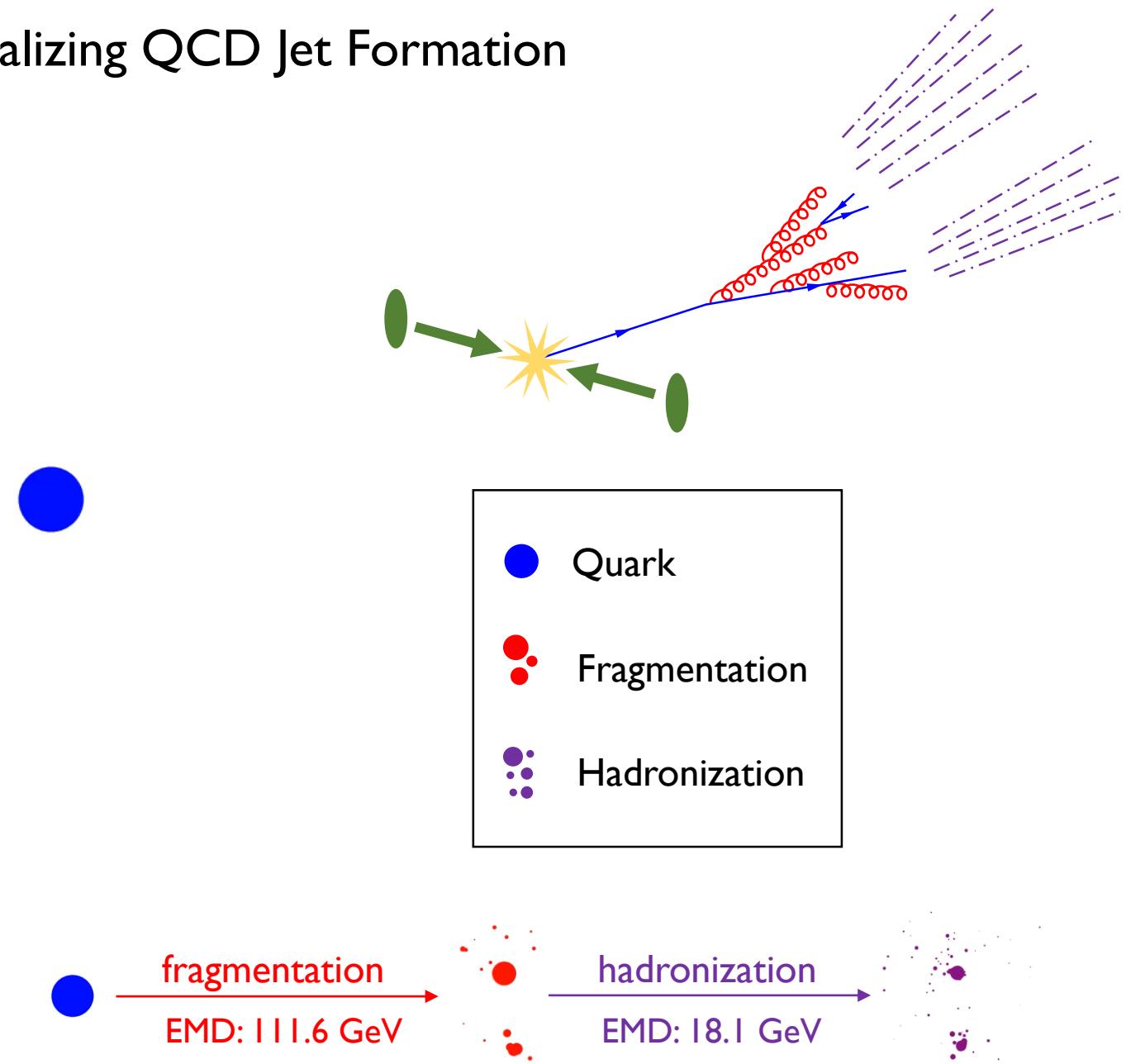
We also get the *shortest path* between the events.

Interpolate along path to visualize!

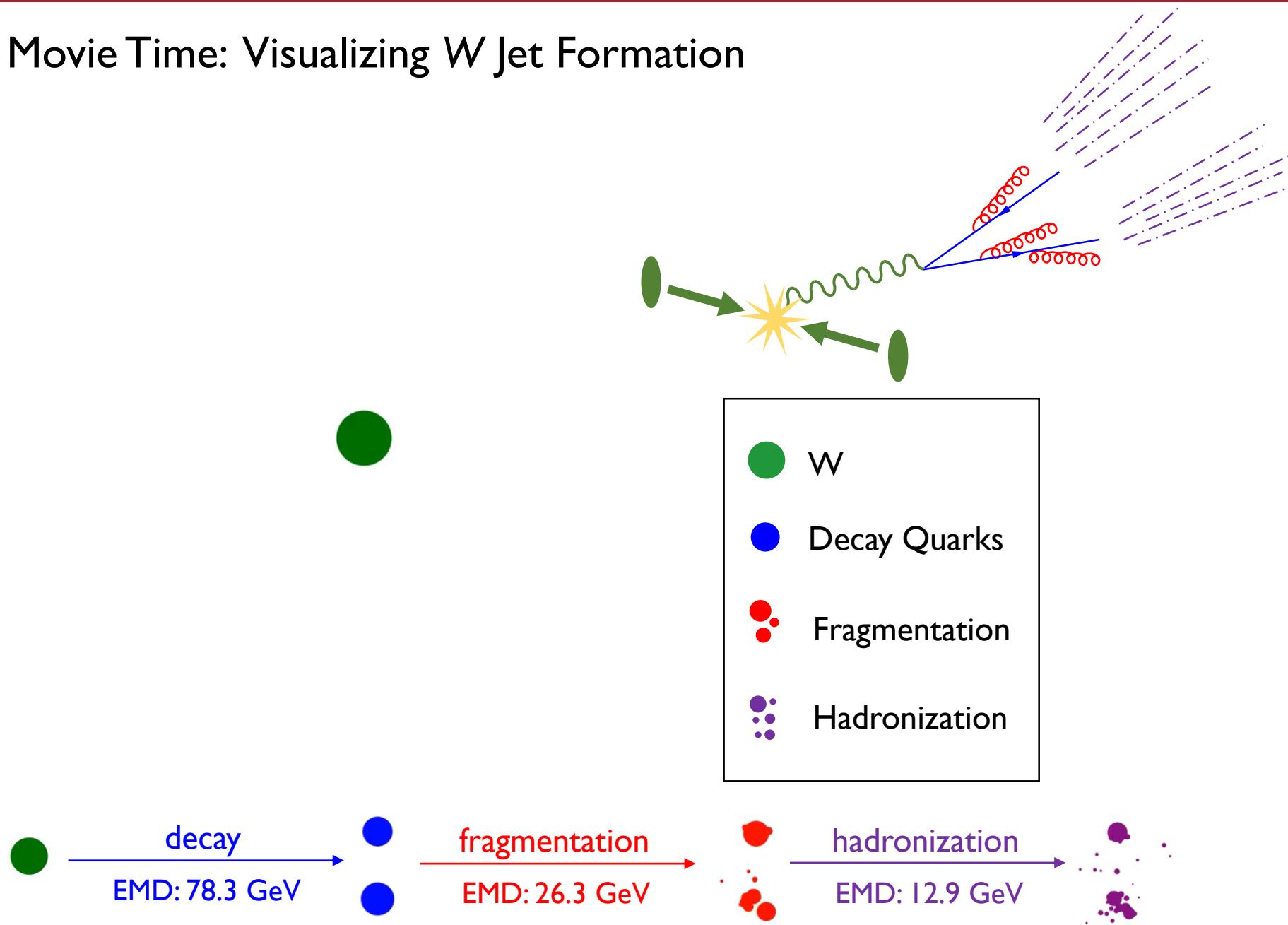
# Movie Time: Visualizing Jet Formation



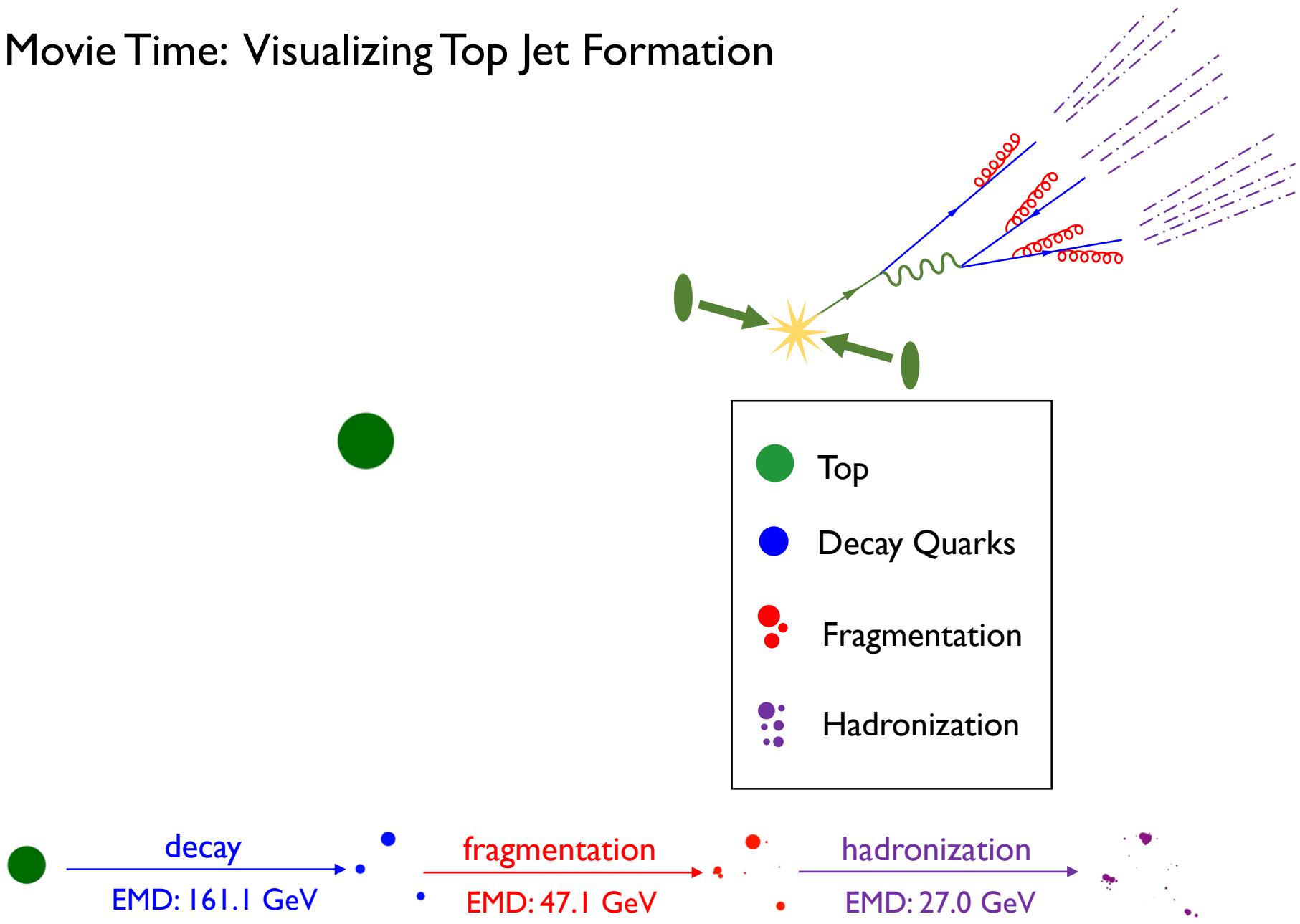
# Movie Time: Visualizing QCD Jet Formation



# Movie Time: Visualizing W Jet Formation

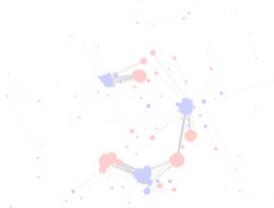


# Movie Time: Visualizing Top Jet Formation



# Outline

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When are two events similar?

*When they have similar distributions of energy*

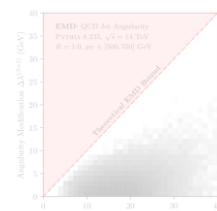
## The Energy Mover's Distance

*Work to rearrange one event into another.*

## Movie Time

*Visualize energy movement and jet formation.*

## Observables



## Part II Applications

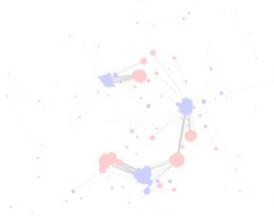
## Quantifying event modifications



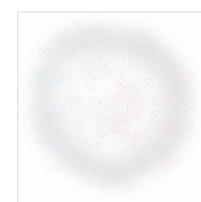
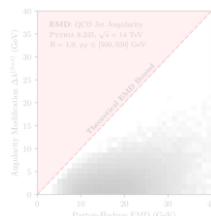
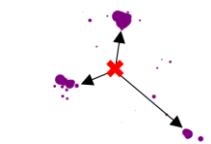
## Exploring the Space of Events

# Outline

## Part I Introduction



## Part II Applications



When are two events similar?

*When they have similar distributions of energy*

The Energy Mover's Distance

*Work to rearrange one event into another.*

Movie Time

*Visualize energy movement and jet formation.*

Observables

Quantifying event modifications

Exploring the Space of Events

# Observables

$N$ -subjettiness:

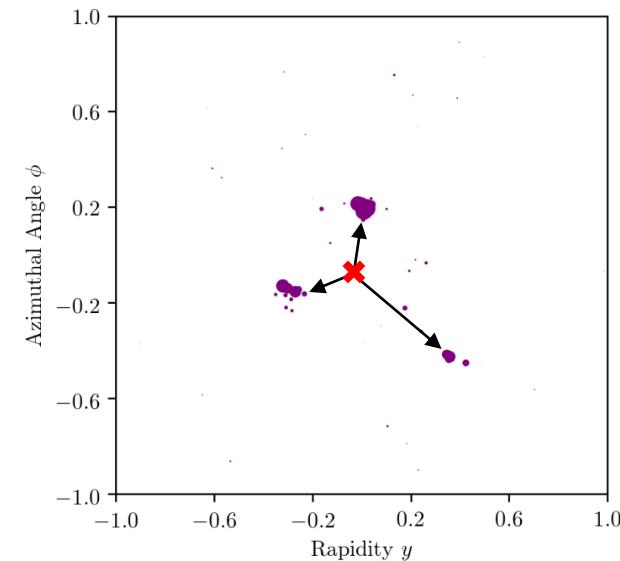
$$\tau_N^{(\beta)} = \sum_{i=1}^M E_i \min_{N \text{ axes}} \{\theta_{1,k}^\beta, \theta_{2,k}^\beta, \dots, \theta_{N,k}^\beta\}$$

measures how well jet energy is aligned into  $N$  subjets

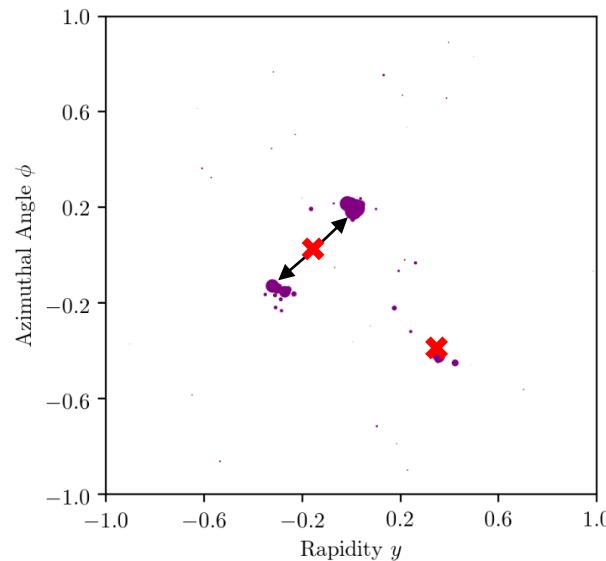
[\[J. Thaler, K. Van Tilburg, 1011.2268\]](#)

[\[J. Thaler, K. Van Tilburg, 1108.2701\]](#)

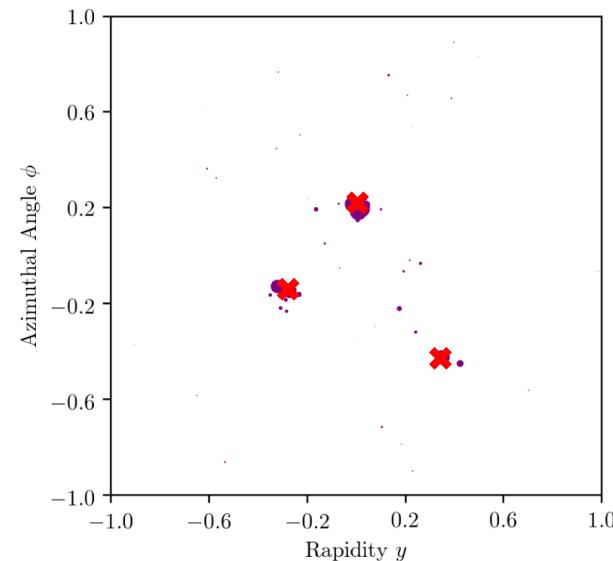
$$\tau_1/E \gg 0$$



$$\tau_1/E > \tau_2/E \gg 0$$



$$\tau_3/E \simeq 0$$



# Observables

$N$ -subjettiness:

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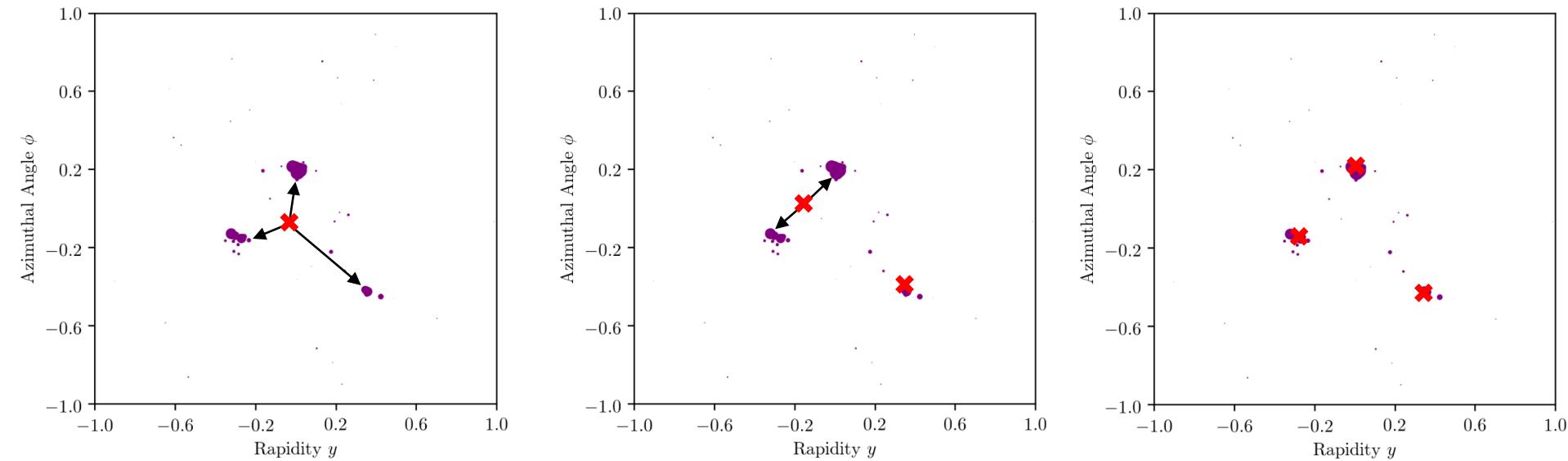
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$$\tau_1/E \gg 0$$

$$\tau_1/E > \tau_2/E \gg 0$$

$$\tau_3/E \simeq 0$$



$N$ -subjettiness is the EMD between the event and the closest  $N$ -particle event.

$$\tau_N(\mathcal{E}) = \min_{|\mathcal{E}'|=N} \text{EMD}(\mathcal{E}, \mathcal{E}').$$

$\beta \neq 1$  corresponds to other  $p$ -Wasserstein distances with  $p = \beta$ .

# Observables

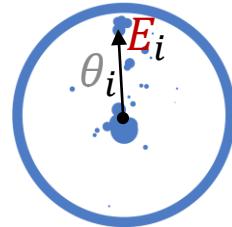
*Getting quantitative*

Take any additive IRC-safe observable:  $\mathcal{O}(\mathcal{E}) = \sum_{i=1}^M E_i \Phi(\hat{p}_i)$

e.g. jet angularities:  $\lambda^{(\beta)} = \sum_{i=1}^M E_i \theta_i^\beta$

[\[C. Berger, T. Kucs, and G. Sterman, 0303051\]](#)

[\[A. Larkoski, J. Thaler, and W. Waalewijn, 1408.3122\]](#)



# Observables

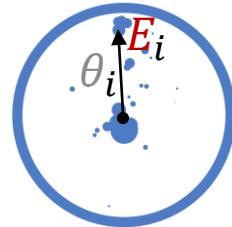
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[C. Berger, T. Kucs, and G. Sterman, 0303051]

[A. Larkoski, J. Thaler, and W. Waalewijn, 1408.3122]



Via the Kantorovich-Rubinstein dual formulation of EMD:

Earth Mover's  
Distance

$$\text{EMD}(\mathcal{E}, \mathcal{E}') \geq \frac{1}{RL} \left| \sum_{i=1}^M E_i \Phi(\hat{p}_i) - \sum_{j=1}^{M'} E'_j \Phi(\hat{p}'_j) \right| = \frac{1}{RL} |\mathcal{O}(\mathcal{E}) - \mathcal{O}(\mathcal{E}')|$$

Difference in  
observable values

“Lipschitz constant” of  $\Phi$   
i.e. bound on its derivative

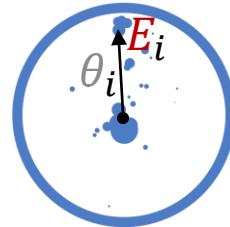
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Difference in  
observable values

“Lipschitz constant” of  $\Phi$   
i.e. bound on its derivative

For  $\beta \geq 1$  jet angularities,  $L = \beta/R$  over the jet cone, so:

$$|\lambda^{(\beta)}(\mathcal{E}) - \lambda^{(\beta)}(\mathcal{E}')| \leq \beta \text{EMD}(\mathcal{E}, \mathcal{E}')$$

The EMD provides a robust upper bound to any modifications of these observables.

# Observables

**Key idea:** Energy-weighted angular structures contain all the IRC-safe information.

$$\frac{1}{RL} \left| \sum_{i=1}^M E_i \Phi(\hat{p}_i) - \sum_{j=1}^{M'} E'_j \Phi(\hat{p}_i) \right| \leq \text{EMD}(\mathcal{E}, \mathcal{E}')$$

**Theorem:** Any infrared and collinear safe observable  $\mathcal{O}$  can be approximated arbitrarily well as:

$$\mathcal{O}(p_1, \dots, p_M) = F \left( \sum_{i=1}^M E_i \vec{\Phi}(\hat{p}_i) \right)$$

for some  $\Phi: \mathbb{R}^2 \rightarrow \mathbb{R}^\ell$  and  $F: \mathbb{R}^\ell \rightarrow \mathbb{R}$  and sufficiently large  $\ell$ .

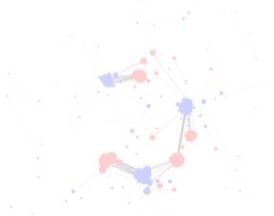
[\[M. Zaheer, et al., 1703.06114\]](#)

[\[P.T. Komiske, EMM, J. Thaler, 1810.05165\]](#)

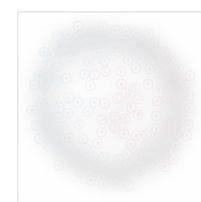
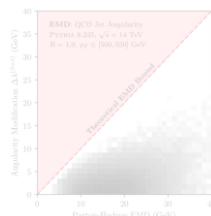
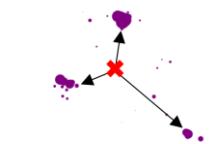
Events close in EMD are close in all infrared and collinear safe information!

# Outline

## Part I Introduction



## Part II Applications



When are two events similar?

*When they have similar distributions of energy*

The Energy Mover's Distance

*Work to rearrange one event into another.*

Movie Time

*Visualize energy movement and jet formation.*

Observables

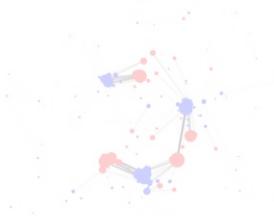
*Conceptually rich connections to EMD.*

Quantifying event modifications

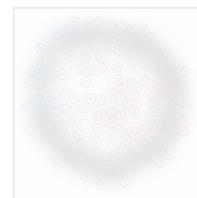
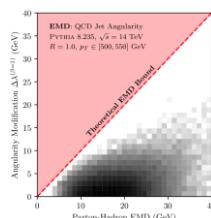
Exploring the Space of Events

# Outline

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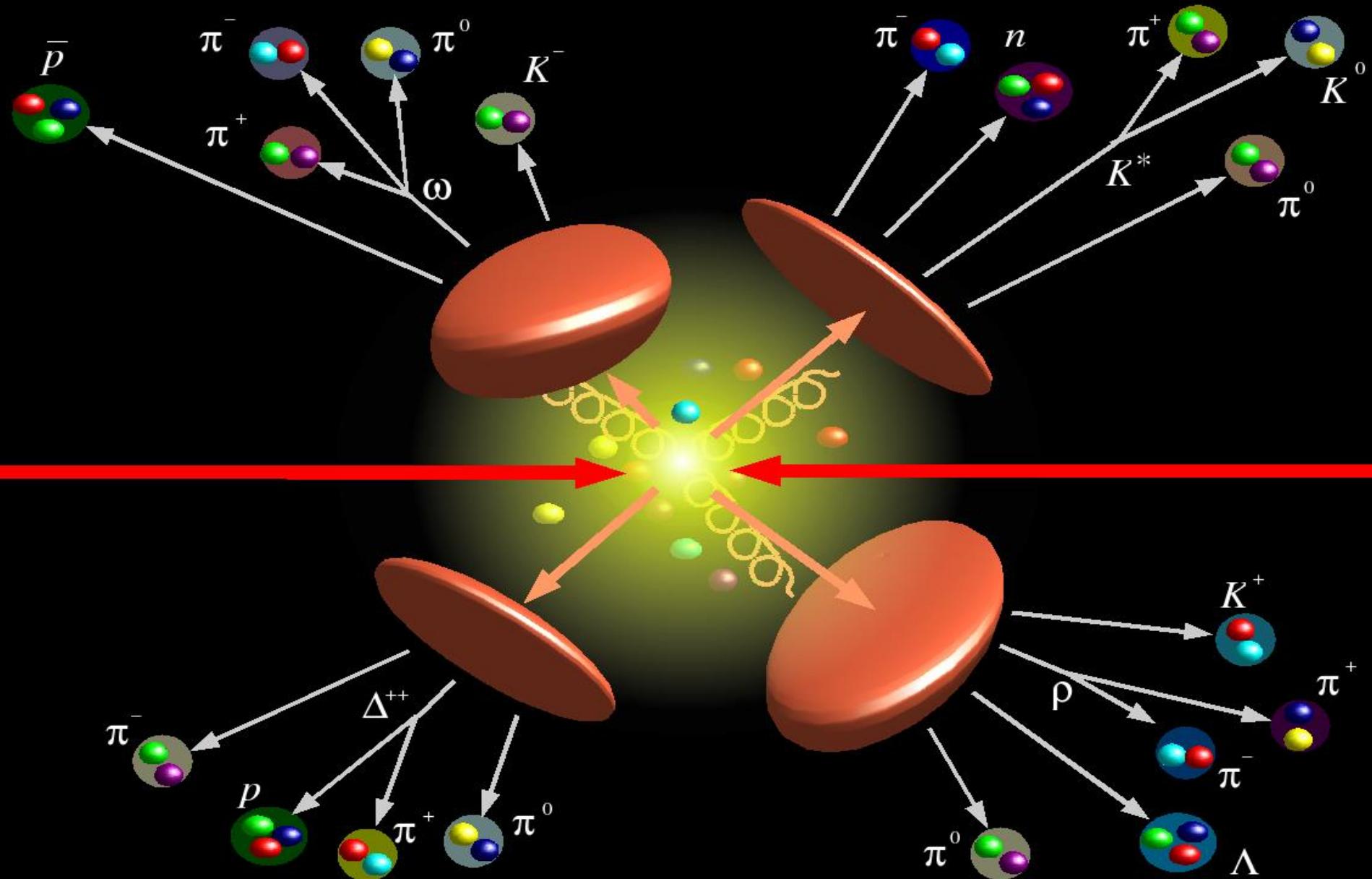
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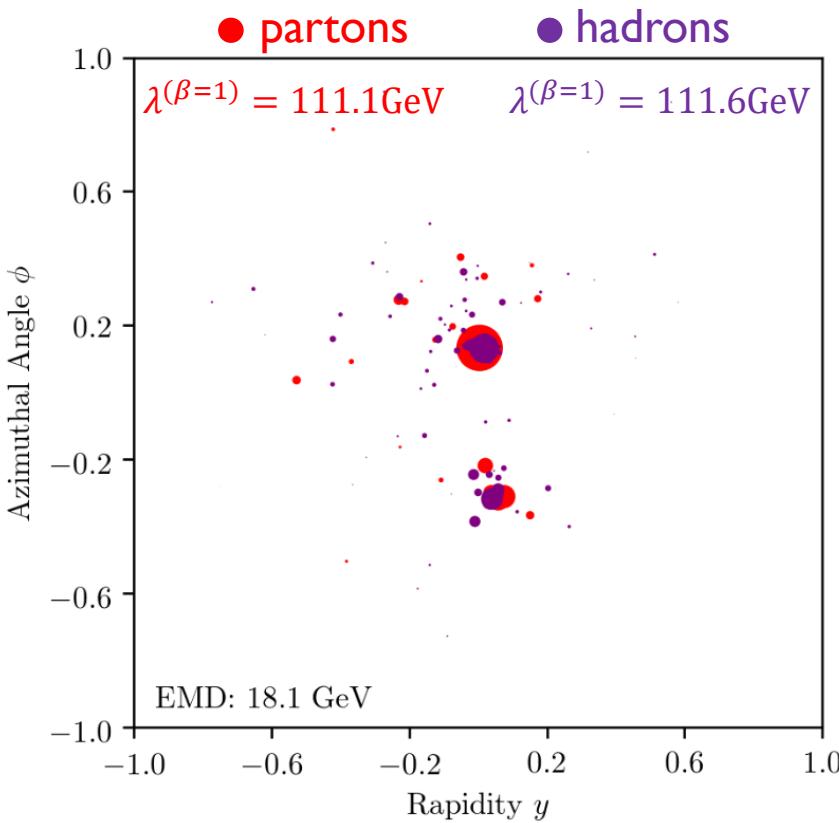
Exploring the Space of Events

# Quantifying event modifications: Hadronization



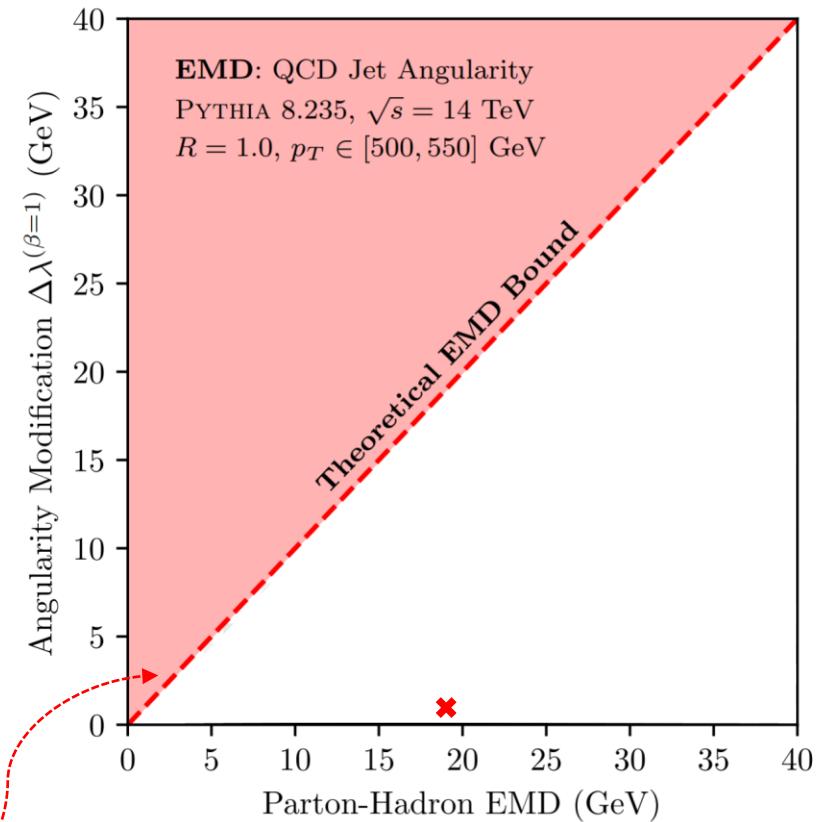
# Quantifying event modifications: Hadronization

$$\lambda^{(\beta=1)} = \sum_{i=1}^M E_i \theta_i$$



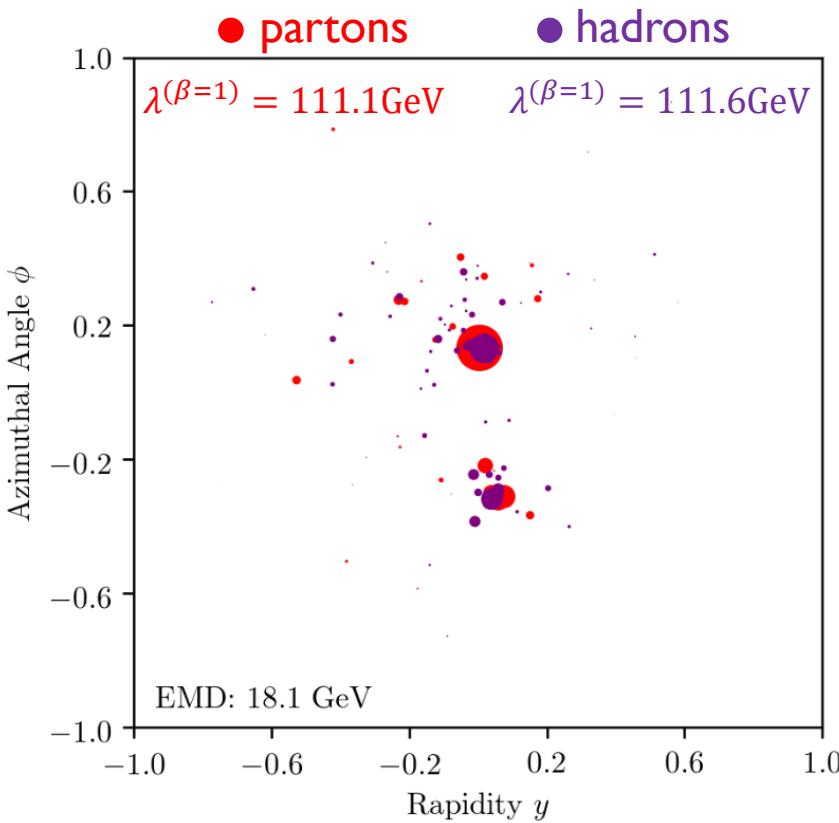
$$\begin{aligned}\mathcal{E} &= \mathcal{E}_{\text{partons}} \\ \mathcal{E}' &= \mathcal{E}_{\text{hadrons}}\end{aligned}$$

$$|\lambda^{(\beta=1)}(\mathcal{E}) - \lambda^{(\beta=1)}(\mathcal{E}')| \leq \text{EMD}(\mathcal{E}, \mathcal{E}')$$



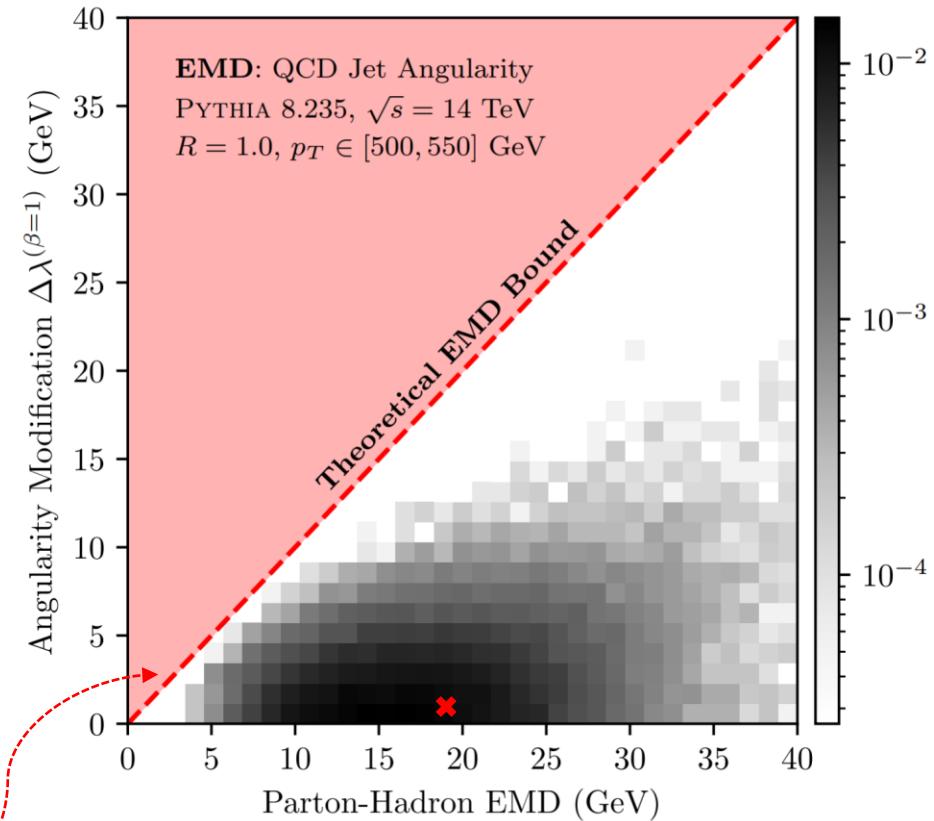
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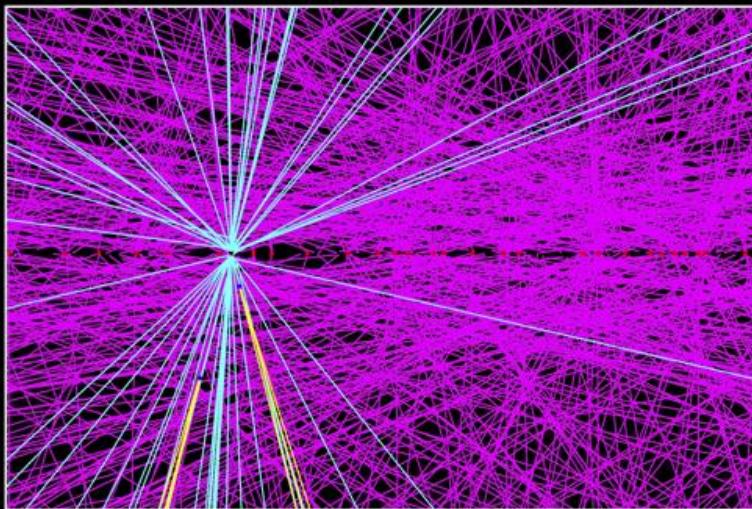


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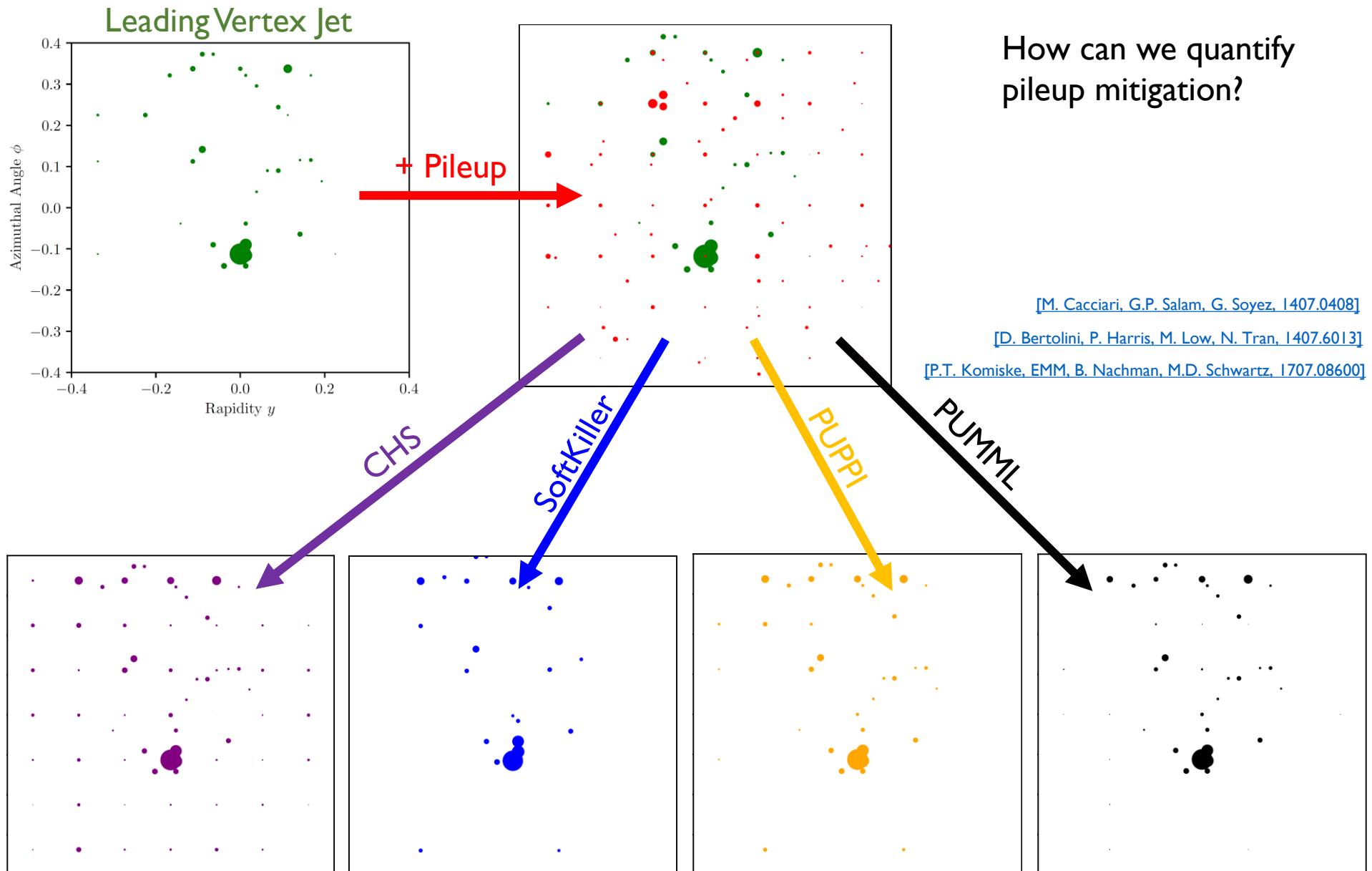
$$|\lambda^{(\beta=1)}(\mathcal{E}) - \lambda^{(\beta=1)}(\mathcal{E}')| \leq \text{EMD}(\mathcal{E}, \mathcal{E}')$$



# Quantifying event modifications: Pileup

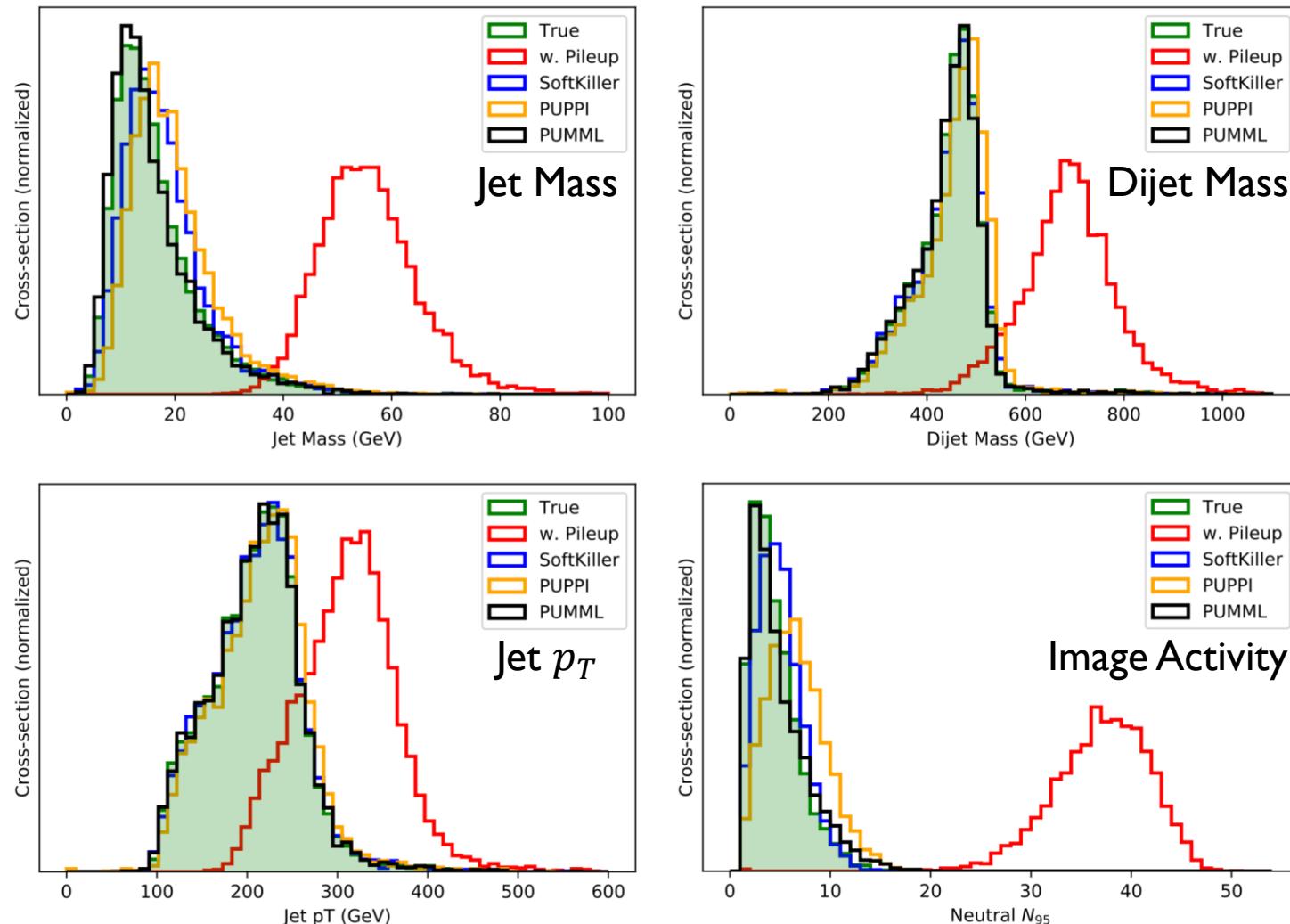


# Quantifying event modifications: Pileup



# Quantifying event modifications: Pileup

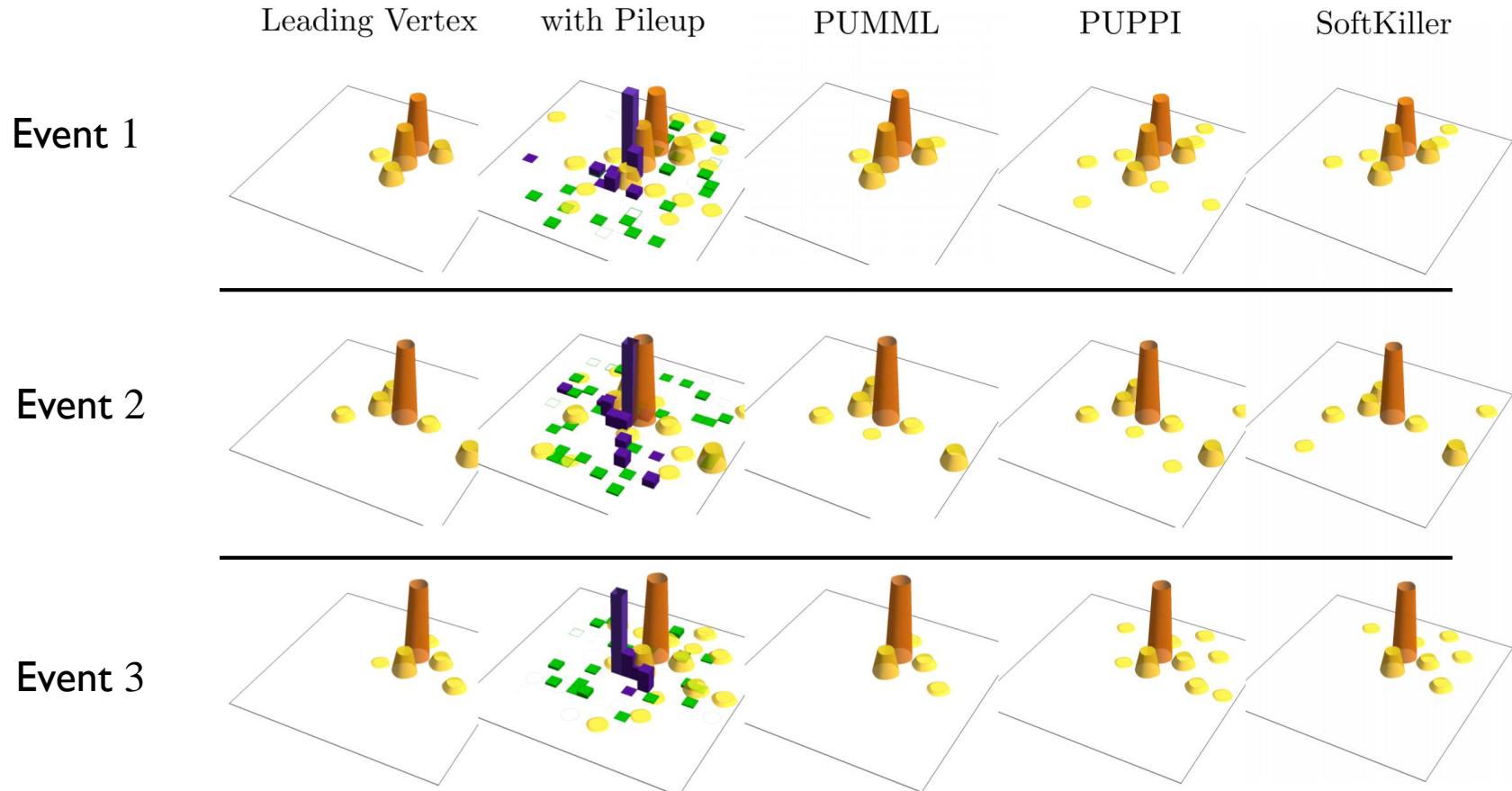
Compare on a collection of observables?



Requires ad hoc choices of observables.

# Quantifying event modifications: Pileup

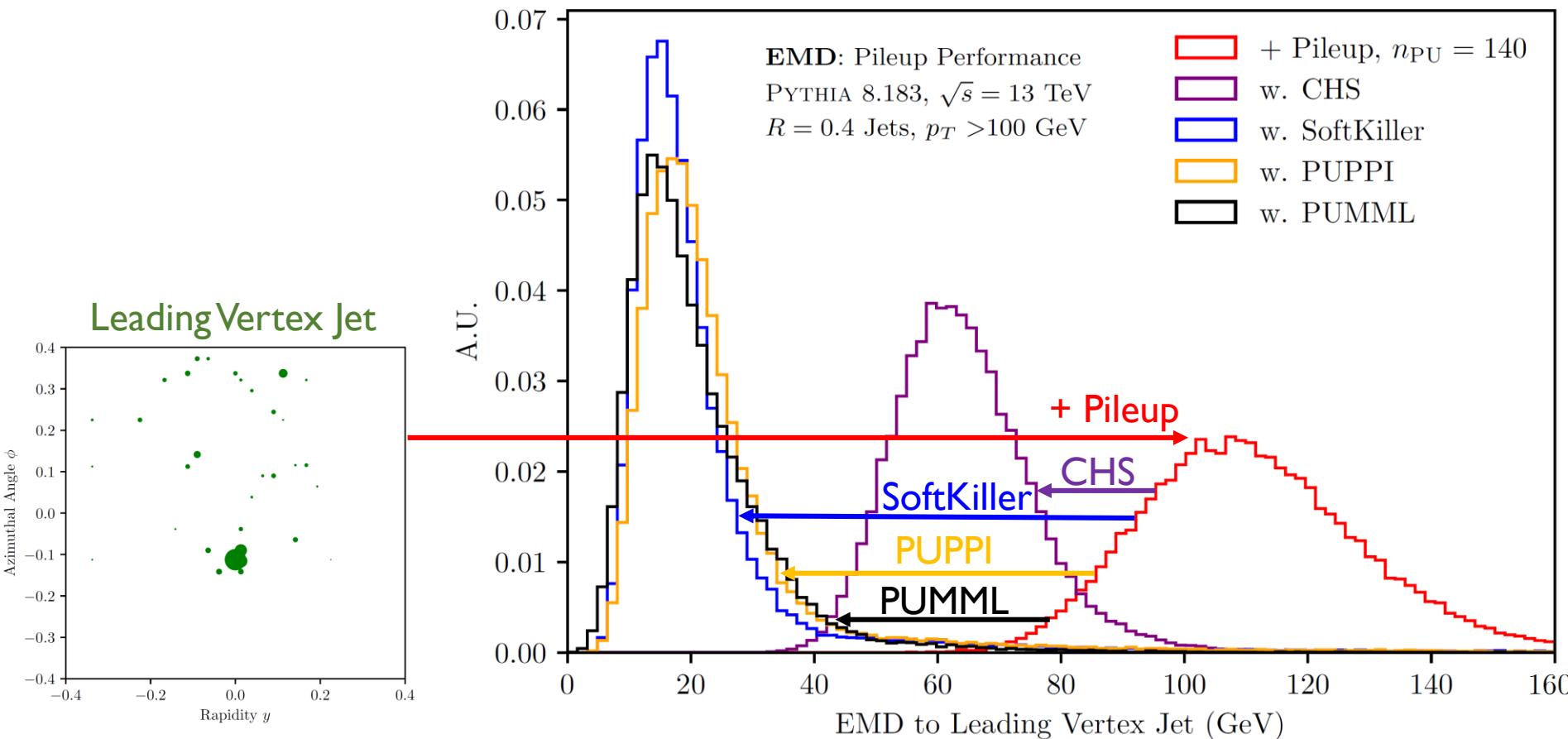
Compare calorimeter images pixel by pixel?



Discontinuous under physically-sensible single-pixel perturbations.  
Undesirable behavior with increasing resolution.

# Quantifying event modifications: Pileup

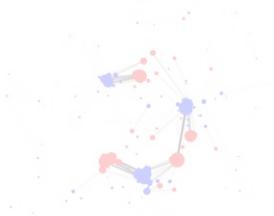
Measure pileup mitigation performance with EMD!



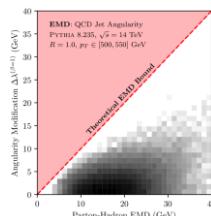
Guarantees performance on IRC safe observables.  
Stable under physically-sensible perturbations.  
Train to optimize EMD with machine learning?

# Outline

## Part I Introduction



## Part II Applications



When are two events similar?

*When they have similar distributions of energy*

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*Work to rearrange one event into another.*

Movie Time

*Visualize energy movement and jet formation.*

Observables

*Conceptually rich connections to EMD.*

Quantifying event modifications

*Hadronization, pileup, detector effects*

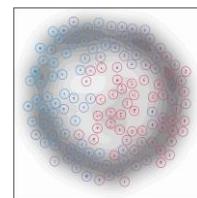
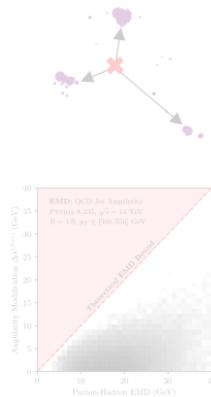
Exploring the Space of Events

# Outline

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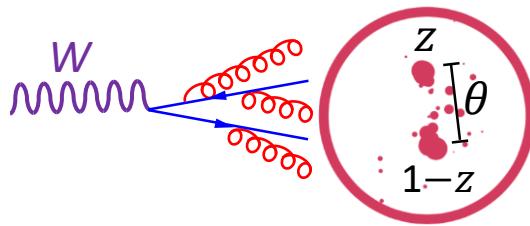
*Conceptually rich connections to EMD.*

Quantifying event modifications

*Hadronization, pileup, detector effects*

Exploring the Space of Events

# Exploring the Space of Events: $W$ jets



$W$  jets are 2-pronged:

$z$ : Energy Sharing of Prongs

$\theta$ : Angle between Prongs

$\varphi$ : Azimuthal orientation

Constrained by  $W$  mass:

$$z(1 - z)\theta^2 = \frac{p_{\mu J}^2}{p_T^2} = \frac{m_W^2}{p_T^2}$$

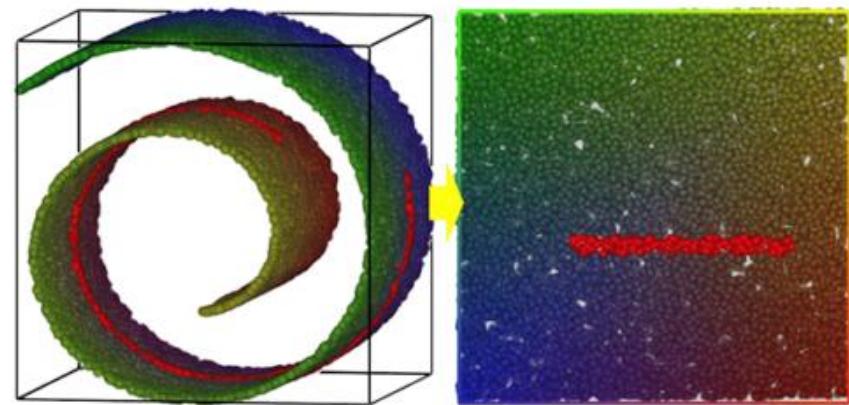
Hence we expect a **two-dimensional** space of  $W$  jets.

After  $\varphi$  rotation: **one-dimensional**

Visualize the space of events with t-Distributed Stochastic Neighbor Embedding (t-SNE).

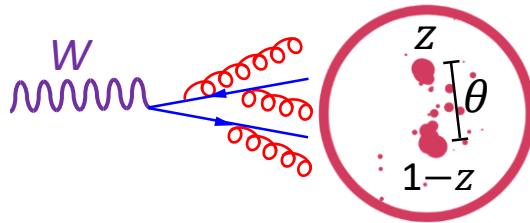
[\[L. van der Maaten, G. Hinton\]](#)

Finds an embedding into a low-dimensional manifold that respects distances.



Src: <http://web-ext.u-aizu.ac.jp/~shigeo/home.html>

# Exploring the Space of Events: $W$ jets



$W$  jets are 2-pronged:

$z$ : Energy Sharing of Prongs

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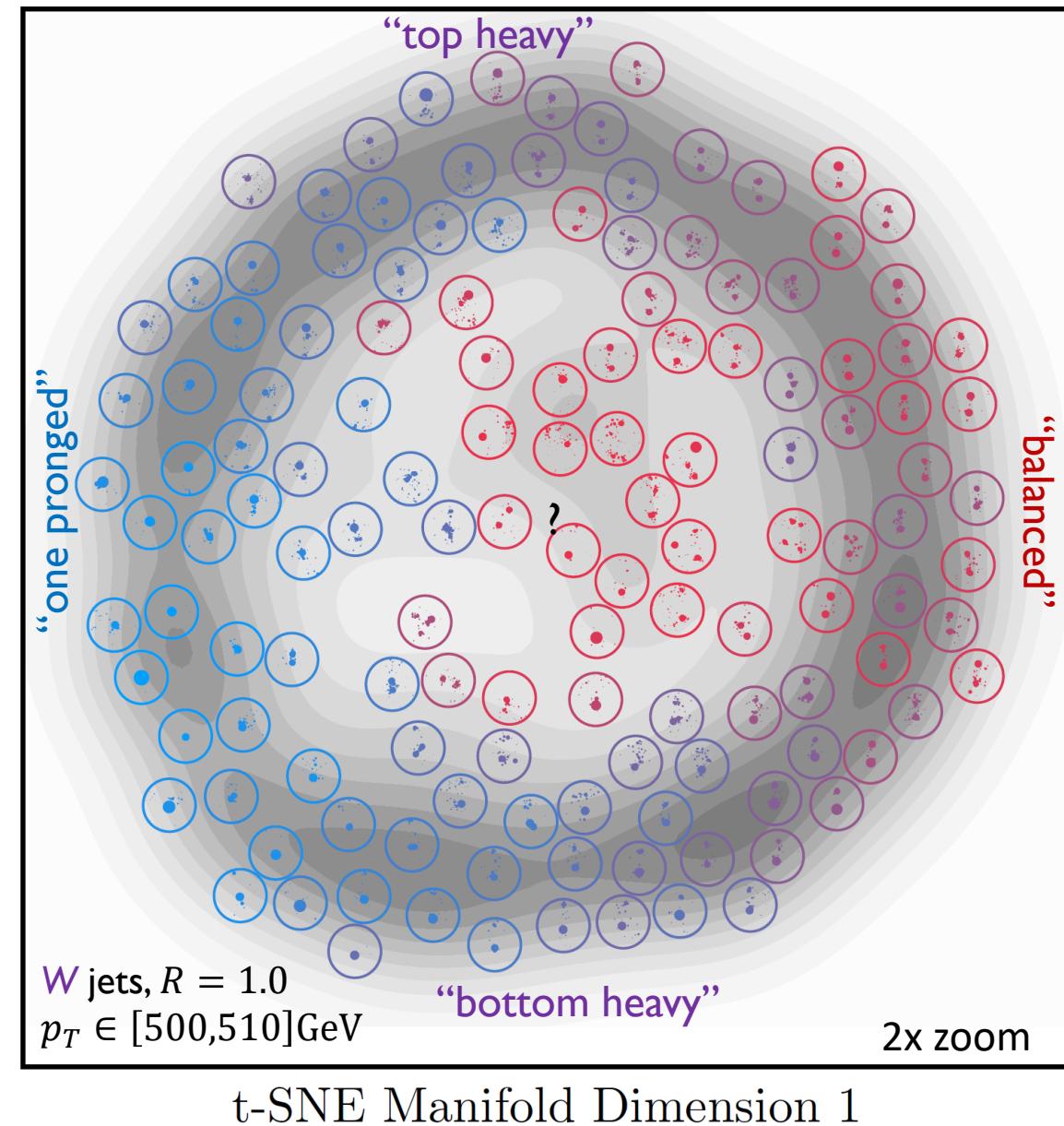
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# Exploring the Space of Jets: Correlation Dimension

VOLUME 50, NUMBER 5

PHYSICAL REVIEW LETTERS

31 JANUARY 1983

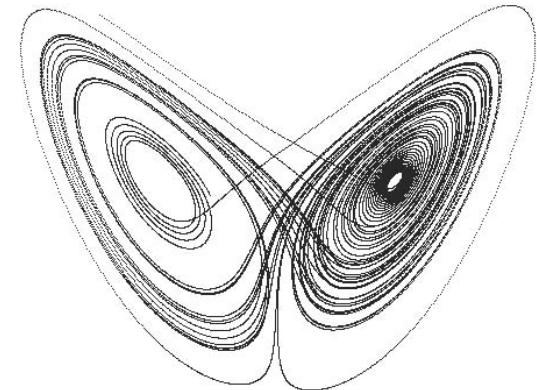
## Characterization of Strange Attractors

Peter Grassberger<sup>(a)</sup> and Itamar Procaccia

*Chemical Physics Department, Weizmann Institute of Science, Rehovot 76100, Israel*

(Received 7 September 1982)

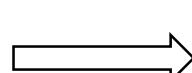
A new measure of strange attractors is introduced which offers a practical algorithm to determine their character from the time series of a single observable. The relation of this new measure to fractal dimension and information-theoretic entropy is discussed.



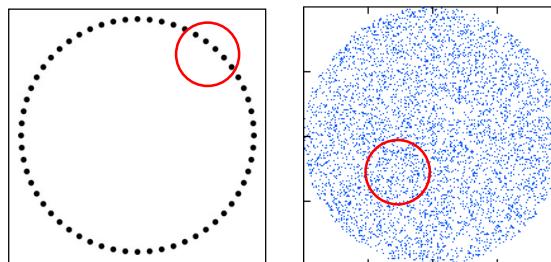
Intuition:

$$N_{\text{neighboring}}(r) \propto r^{\dim}$$

points



$$\dim(r) = r \frac{\partial}{\partial r} \ln N_{\text{neighbors}}(r)$$



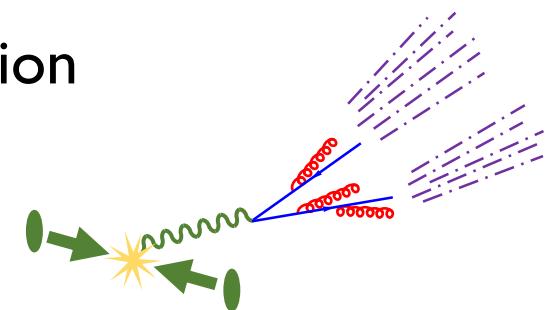
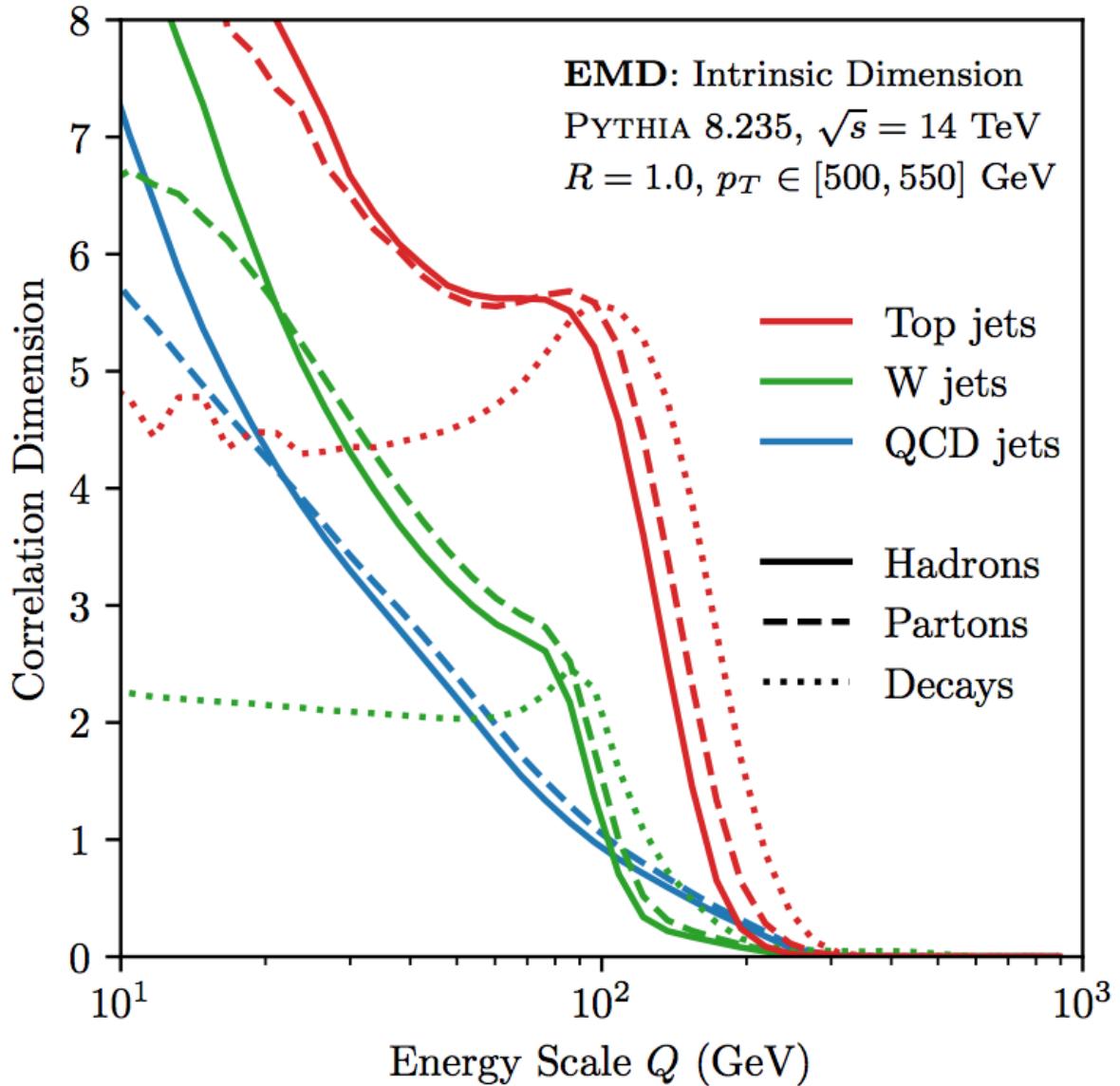
Correlation dimension:

$$\dim(Q) = Q \frac{\partial}{\partial Q} \ln \sum_{i=1}^N \sum_{j=1}^N \Theta[\text{EMD}(\varepsilon_i, \varepsilon_j) < Q]$$

Energy scale  $Q$   
dependence

Count neighbors in  
ball of radius  $Q$

# Exploring the Space of Jets: Correlation Dimension



QCD jets are simplest.

W jets are more complicated.

Top jets are most complex.

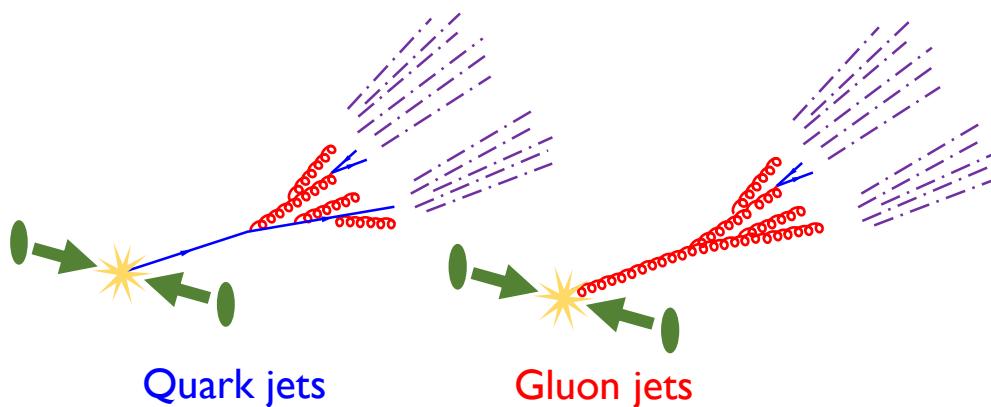
“Decays” have ~constant dimension.

Fragmentation becomes more complex at lower energy scales.

Hadronization becomes relevant at scales around 20 GeV.

Can we understand this analytically?

# Exploring the Space of Jets: Correlation Dimension

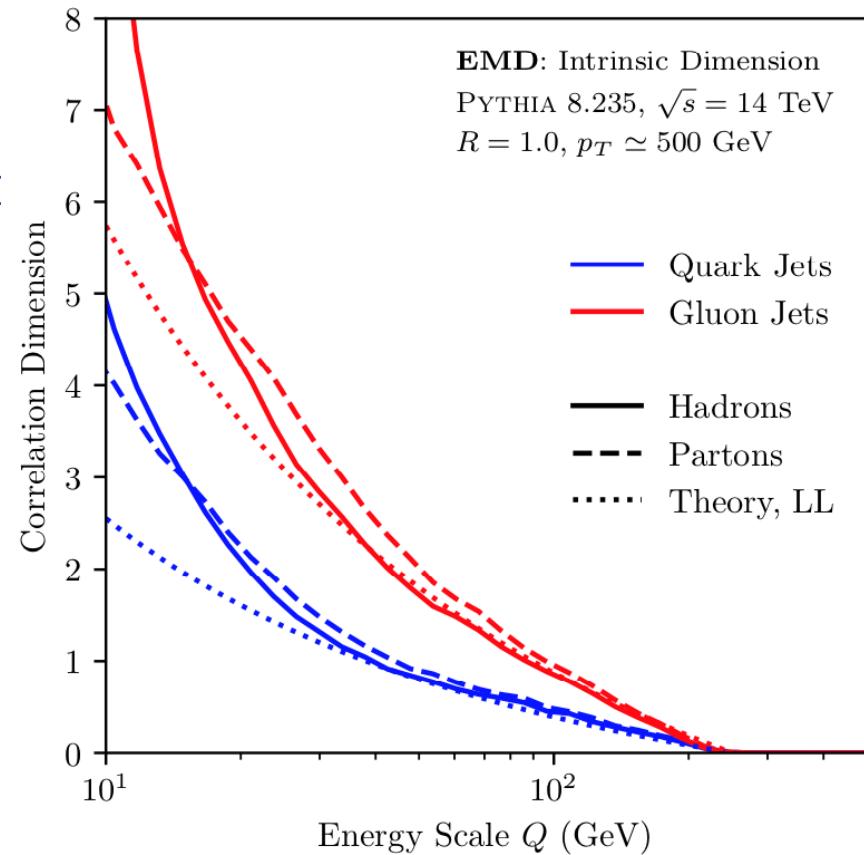


$$\text{At LL: } \dim_{q/g}(Q) = -\frac{8\alpha_s C_{q/g}}{\pi} \ln \frac{Q}{p_T/2}$$

$$+ \text{1-loop running of } \alpha_s$$

$$C_q = C_F = \frac{4}{3}$$

$$C_g = C_A = 3$$

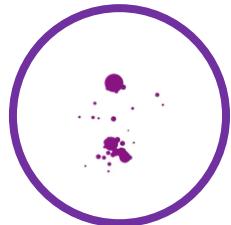


Dimension blows up at low energies.

Jets are “more than fractal”?

# Exploring the Space of Events: Jet Classification

Classify  $W$  jets vs. QCD jets



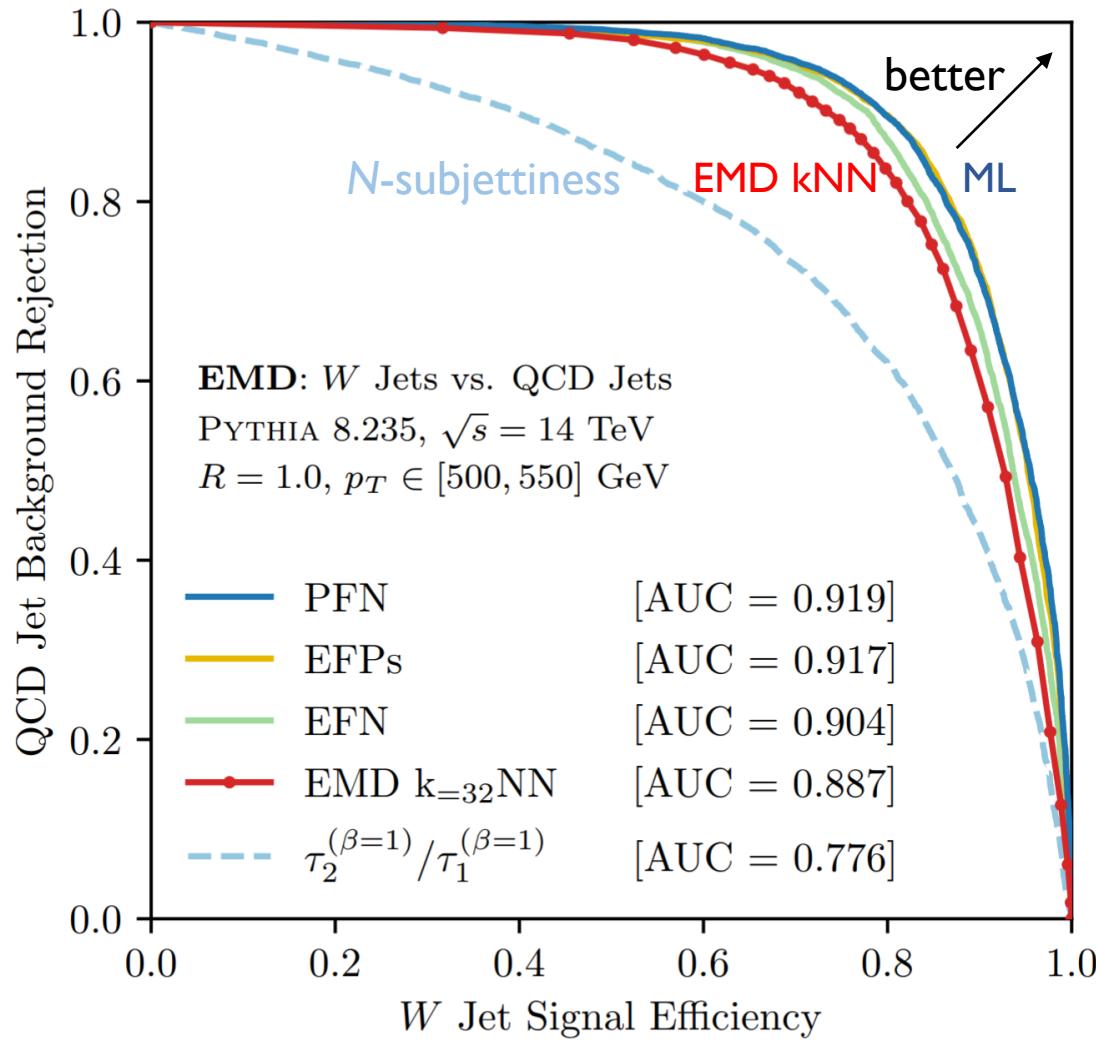
vs.



Look at a jet's nearest neighbors (kNN) to predict its class.

Optimal IRC-safe classifier with enough data.

Nearing performance of ML.



# Exploring the Space of Events

*Clustering events*

Use EMD as a measure of event similarity

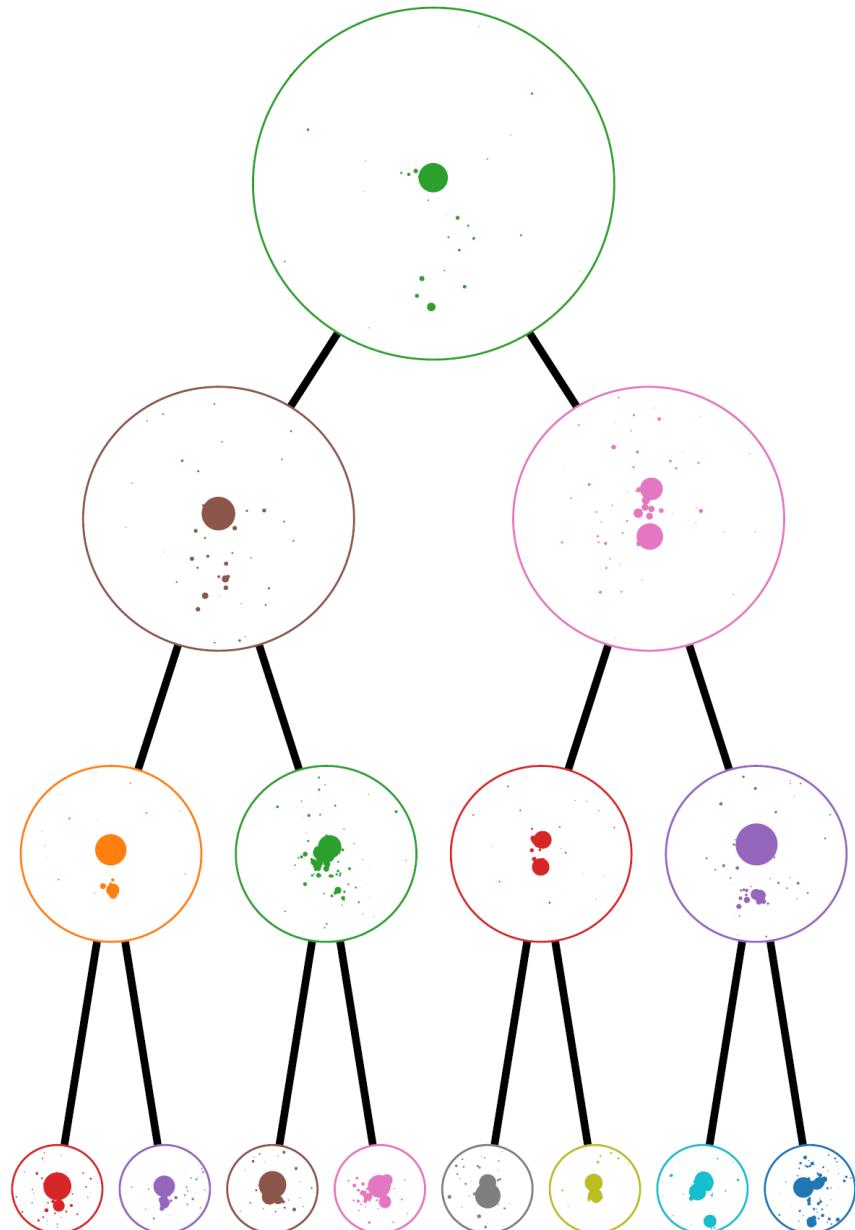
Unsupervised clustering algorithms can  
be used to cluster events

Jets are clusters of particles  
???? are clusters of jets

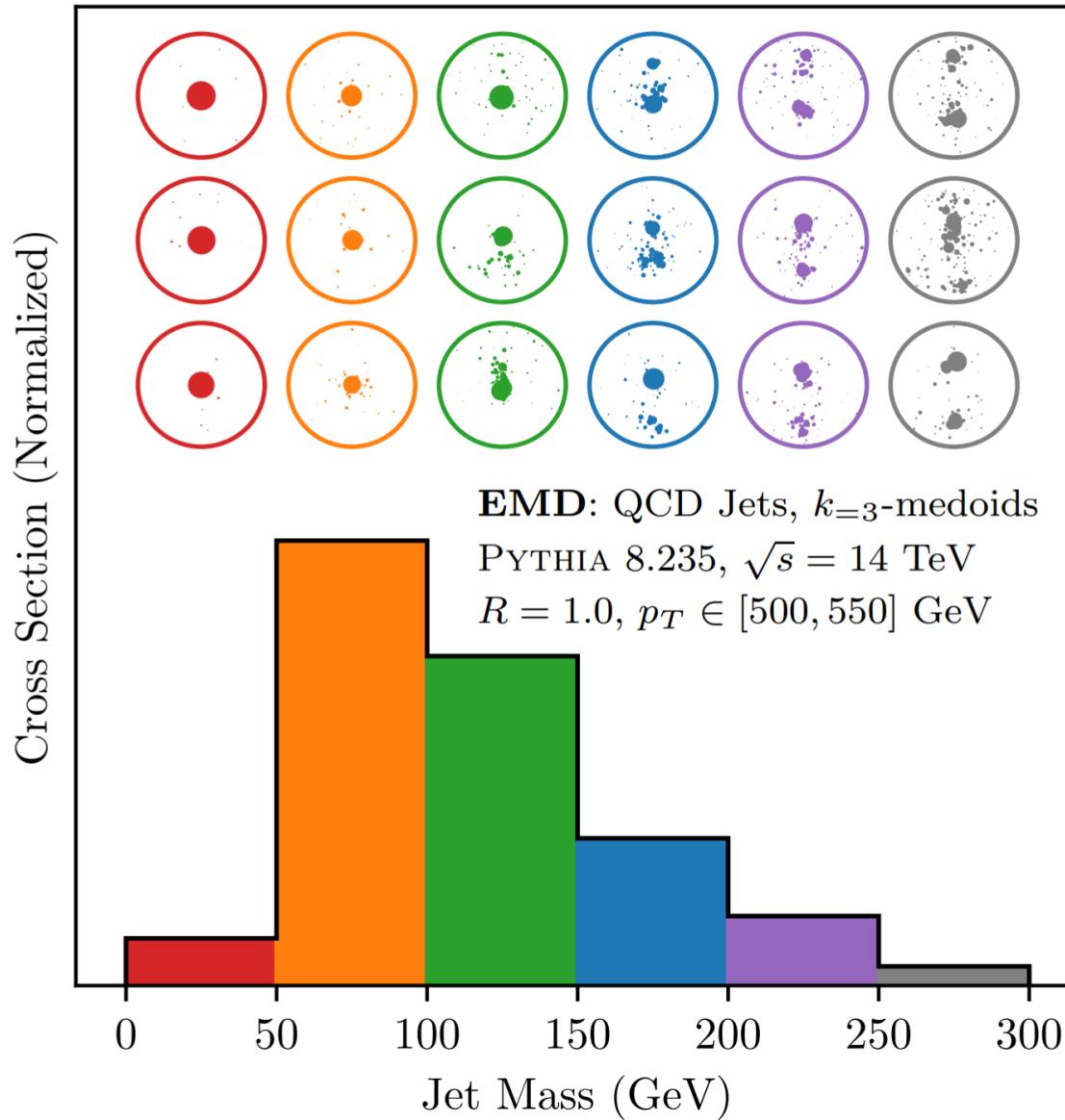
VP Tree:  $O(\log(N))$  neighbor query time

Much more to explore.

Vantage Point (VP) Tree

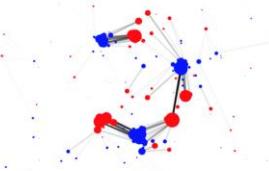
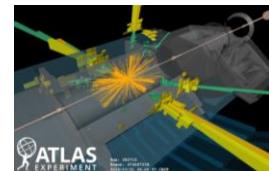


# Exploring the Space of Events: $k$ -medoids

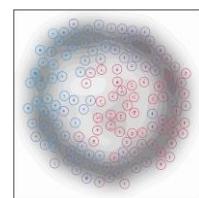
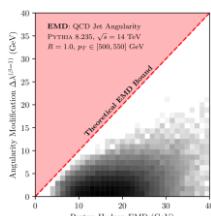
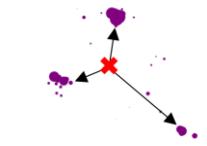


# Summary

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Quantifying event modifications

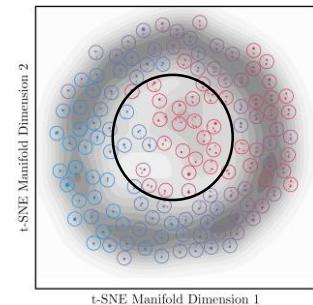
*Hadronization, pileup, detector effects*

Exploring the Space of Events

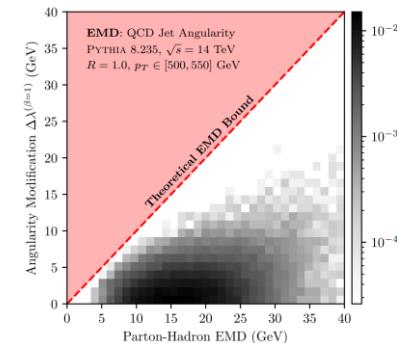
*Unlock new ideas and techniques with EMD*

# Going Beyond

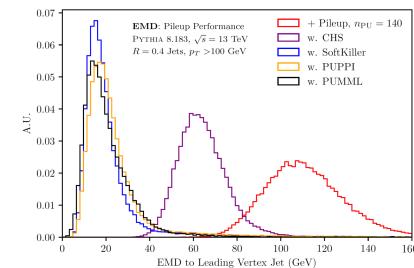
- Model (in)dependent anomaly detection?



- Sharpen the parton-hadron duality of energy flow?



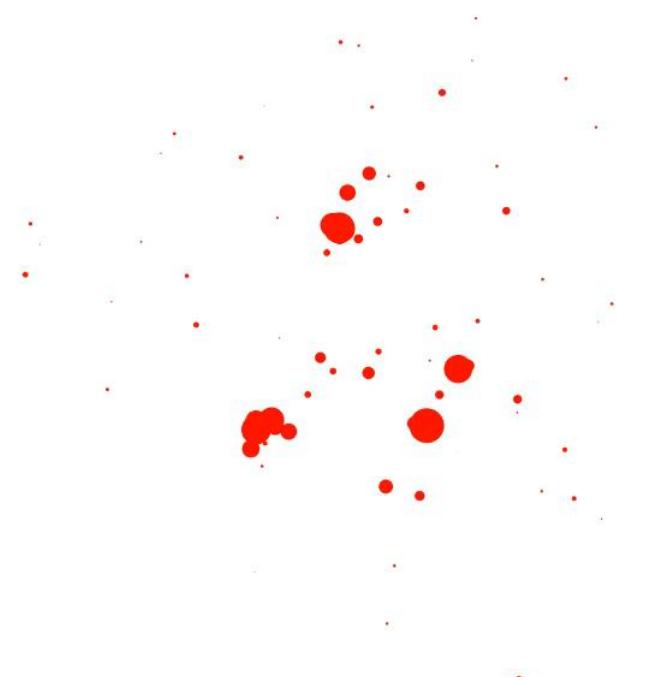
- Train ML models to optimize EMD directly?



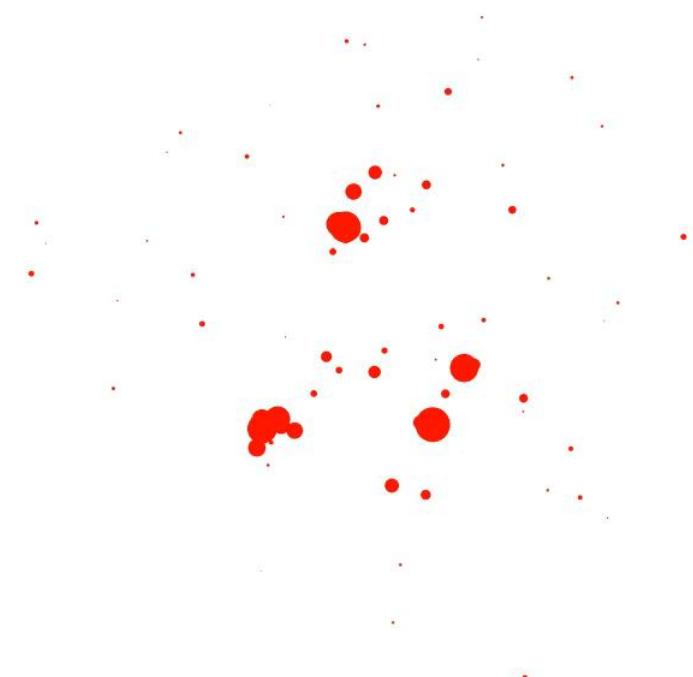
- Include flavor information?

# The End

## Thank you!



# Extra Slides



# Exploring the Space of Jets: Correlation Dimension

Sketch of leading log (one emission) calculation:

$$\dim_{q/g}(Q) = Q \frac{\partial}{\partial Q} \ln \sum_{i=1}^N \sum_{j=1}^N \Theta[\text{EMD}(\varepsilon_i, \varepsilon_j) < Q]$$

$$= Q \frac{\partial}{\partial Q} \ln \Pr [\text{EMD} < Q]$$

$$= Q \frac{\partial}{\partial Q} \ln \Pr [\lambda^{(\beta=1)} < Q; C_{q/g} \rightarrow 2 C_{q/g}]$$

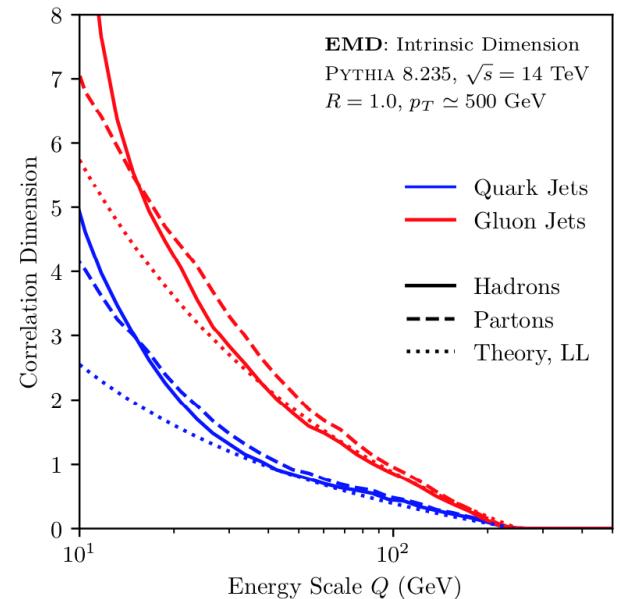
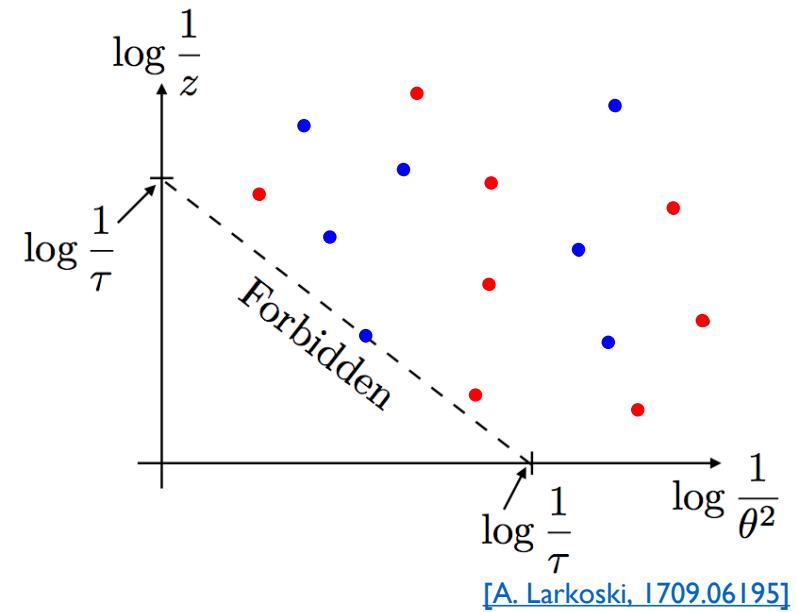
$$= Q \frac{\partial}{\partial Q} \ln \exp \left( - \frac{4\alpha_s C_{q/g}}{\pi} \ln^2 \frac{Q}{p_T/2} \right)$$

$$= - \frac{8\alpha_s C_{q/g}}{\pi} \ln \frac{Q}{p_T/2}$$

+ 1-loop running of  $\alpha_s$

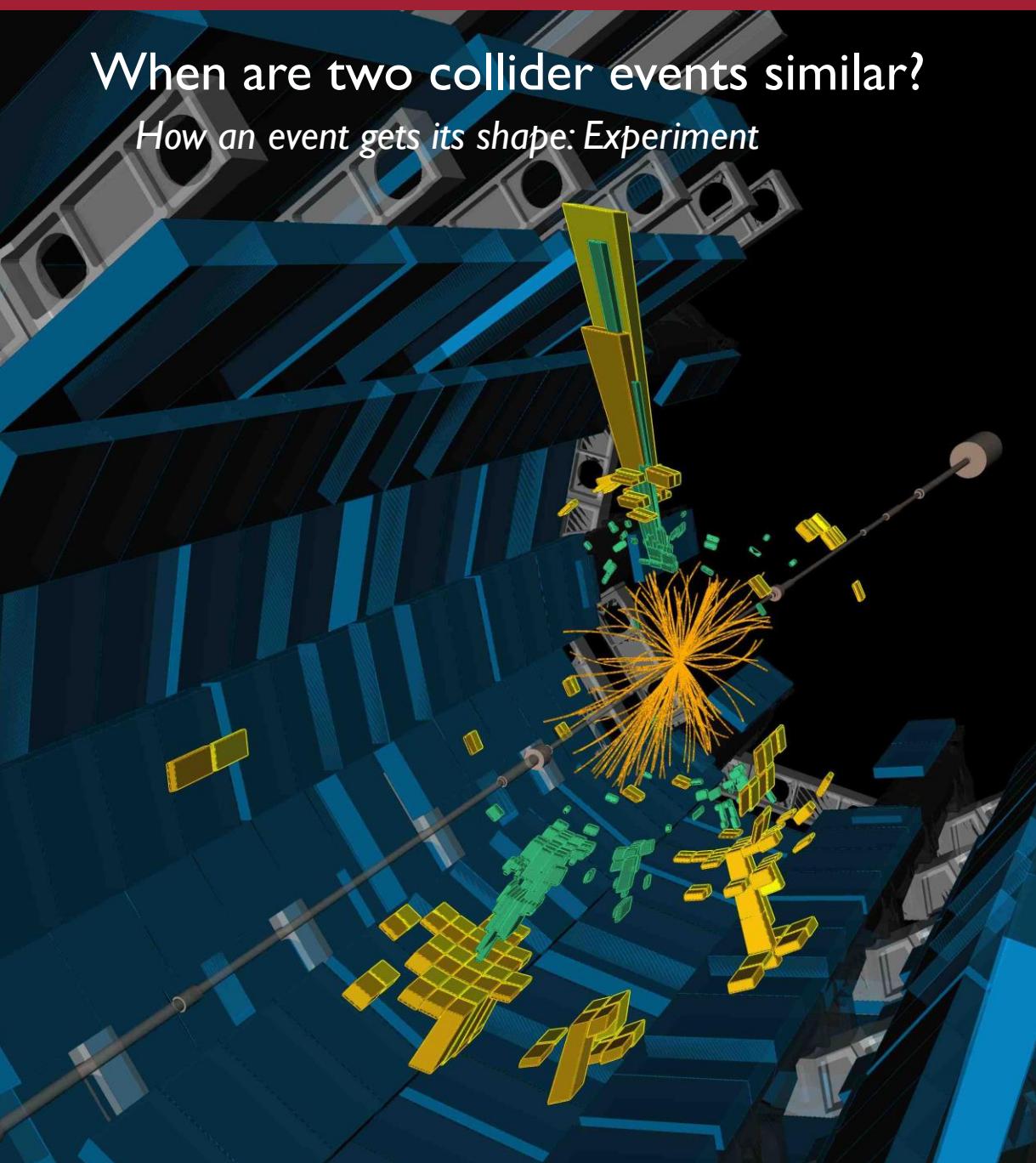
$$C_q = C_F = \frac{4}{3}$$

$$C_g = C_A = 3$$



# When are two collider events similar?

*How an event gets its shape: Experiment*



tracker	ECAL	HCal	
			$\gamma$ photon
			$e^\pm$ electron
			$\mu^\pm$ muon
			$\pi^\pm$ pion
			$K^\pm$ kaon
			$K_L^0$ K-long
			$p/\bar{p}$ proton
			$n/\bar{n}$ neutron

# Pileup Mitigation with PUMML

