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Studies of phase transitions in magnetic systems

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Contents

1	Introduction	3
2	Theory	4
2.1	the Ising model	4
2.1.1	Periodic boundary conditions	4
2.2	Statistical physics	4
2.2.1	the partition function	4
2.2.2	Calculation of values	4
2.3	Phase transition	5
2.4	Randomness	5
3	Method	6
4	Result & Discussion	7
4.1	Analytic 2x2	7
4.1.1	Microstates 2x2	7
4.1.2	Quantities	8
4.2	Simulation 2x2	10
4.3	Phase transition	11
4.3.1	Numerical studies of phase transition	11
4.3.2	Extracting the critical temperature	12
5	Conclusion	13
6	References	14
7	Appendix	15

Abstract

1 Introduction

These laws are not enough to solve the motion of the planets. From the laws one can derive differential equations for the motion, which are not trivial or even possible to solve analytically. This is where computational methods are useful. With the tools developed in computational physics we can make a prediction to the motion of the planets in our solar system.¹ And because of our assignment we kind of have to do this to pass the course.[2]

¹[Semester page for FYS3150 - Autumn 2017.](#)

2 Theory

2.1 the Ising model

The Ising model describes a coupled system. Where only the nearest neighbor affect each other. In this report the Ising model will be applied to a two dimensional magnetic system. This will be a grid of spins, where each spin s_i can either have 1 or 0 as value. The total energy is expressed as:

$$E = - \sum_{\langle i,j \rangle} J_{i,j} s_i s_j$$

Where the symbol $\langle kl \rangle$ indicates that we sum over nearest neighbors only. If we assume that each coupling has the same magnitude J , then the energy is expressed as:

$$E = -J \sum_{\langle i,j \rangle} s_i s_j \quad (1)$$

2.1.1 Periodic boundary conditions

When working with a finite matrix we run into a problem with the boundaries. They are missing neighbours. We solve this by introducing periodic boundary conditions. This means that the right neighbour for S_n is assumed to take the value of S_1 .

2.2 Statistical physics

2.2.1 the partition function

Boltzmann distribution is used as the probability distribution. Boltzmann distribution states the probability for E_i is proportional to $e^{-\beta E_i}$, where β is $\frac{1}{k_B T}$. k is the boltzmann constant. For this to be a probability distribution, it needs to be normalized. To normalize the distribution divide the sum of probabilities by a constant Z :

$$1 = \frac{\sum_i e^{-\beta E_i}}{Z}$$

$$Z = \sum_i e^{-\beta E_i}$$

Z is called the partition function.

2.2.2 Calculation of values

The partition function is very useful. In combination with the boltzmann distribution we get a expression for the probability.

$$P(E_i) = \frac{e^{-\beta E_i}}{Z}$$

For finding a mean value, one can simply make a sum over $P(E_i)$ multiplied by the value of interest. For instance the mean energy is given by:

$$\langle E \rangle = \sum_i E_i P(E_i)$$

Expressions for important expectation values can be derived such for the energy E , magnetic moment $|M|$, specific heat capacity C_v and the susceptibility χ . The expressions used in this report are listed below^[1]:²

$$\langle E \rangle = \sum_i E_i P(E_i) \quad (2)$$

$$\langle |M| \rangle = \sum_i M_i P(E_i) \quad (3)$$

$$\langle C_V \rangle = \frac{1}{kT^2} (\langle E^2 \rangle - \langle E \rangle^2) \quad (4)$$

$$\langle \chi \rangle = \frac{1}{kT} (\langle M^2 \rangle - \langle |M| \rangle^2) \quad (5)$$

2.3 Phase transition

2.4 Randomness

²lecture note page 427

3 Method

4 Result & Discussion

4.1 Analytic 2x2

4.1.1 Microstates 2x2

Table 1: This shows the different microstates that is possible for a 2x2 spinmatrix. It also states the energy and magnetic moment for each microstate.

State	Energy	Magnetic moment	State	Energy	Magnetic moment
$\uparrow\uparrow$ $\uparrow\uparrow$	-8J	4	$\downarrow\downarrow$ $\downarrow\downarrow$	-8J	-4
$\downarrow\uparrow$ $\uparrow\uparrow$	0J	2	$\uparrow\downarrow$ $\downarrow\downarrow$	0J	-2
$\uparrow\downarrow$ $\uparrow\uparrow$	0J	2	$\downarrow\uparrow$ $\downarrow\downarrow$	0J	-2
$\uparrow\uparrow$ $\downarrow\uparrow$	0J	2	$\downarrow\downarrow$ $\uparrow\downarrow$	0J	-2
$\uparrow\uparrow$ $\uparrow\downarrow$	0J	2	$\downarrow\downarrow$ $\downarrow\uparrow$	0J	-2
$\downarrow\downarrow$ $\uparrow\uparrow$	0J	0	$\uparrow\uparrow$ $\downarrow\downarrow$	0J	0
$\downarrow\uparrow$ $\downarrow\uparrow$	0J	0	$\uparrow\downarrow$ $\uparrow\downarrow$	0J	0
$\uparrow\downarrow$ $\downarrow\uparrow$	8J	0	$\downarrow\uparrow$ $\uparrow\downarrow$	8J	0

Table 2: The table shows a summary from table 4.1.1.

Number of \uparrow	Multiplicity	Energy	Magnetic moment
4	1	-8J	4
3	4	0J	2
2	2	8J	0
2	4	0J	0
1	4	0J	-2
0	1	-8J	-4

4.1.2 Quantities

We will use the equations from section 2.2.2.

For energy the eq. 2 will result in:

$$Z = \sum_i e^{-\beta E_i}$$

$$T = kT/J = 1$$

$$Z = \sum_i e^{-\beta E_i} = 2e^8 + 2e^{-8} + 12$$

For energy the eq. 2 will give the result:

$$\langle E \rangle = \sum_i E_i P(E_i)$$

$$T = kT/J = 1$$

$$\langle E \rangle = \frac{1}{Z} \sum_i E_i e^{-E_i}$$

$$\langle E \rangle = \frac{1}{Z} (16e^8 - 16e^{-8}) = 7.9839$$

$$\langle E \rangle / N = \frac{\langle E \rangle}{4} = 1.9959$$

For energy the eq. 3 will give the result:

$$\langle |M| \rangle = \sum_i M_i P(E_i)$$

$$T = kT/J = 1$$

$$\langle |M| \rangle = \frac{1}{Z} \sum_i M_i e^{-E_i}$$

$$\langle |M| \rangle = \frac{1}{Z} (4 \cdot 1e^8 + 2 \cdot 4e^0 + 0 \cdot 2e^{-8} + 0 \cdot 4e^0 + 2 \cdot 4e^0 + 4 \cdot 1e^8)$$

$$\langle |M| \rangle = \frac{1}{Z} (16 + 8e^8) = 3.9946$$

$$\langle |M| \rangle / N = \frac{\langle |M| \rangle}{4} = 0.9986$$

For C_V we need to calculate $\langle E^2 \rangle$:

$$\langle E^2 \rangle = \sum_i E_i^2 P(E_i)$$

$$T = kT/J = 1$$

$$\langle E^2 \rangle = \frac{1}{Z} \sum_i E_i^2 e^{-E_i}$$

$$\langle E^2 \rangle = \frac{1}{Z} (128e^8 + 128e^{-8})$$

$$C_V = \langle E^2 \rangle - \langle E \rangle^2 = 0.12832$$

$$C_V/N = 0.03208$$

For χ we need to calculate $\langle M^2 \rangle$:

$$\langle M^2 \rangle = \sum_i M_i^2 P(E_i)$$

$$T = kT/J = 1$$

$$\langle M^2 \rangle = \frac{1}{Z} \sum_i M_i^2 e^{-E_i}$$

$$\langle M^2 \rangle = \frac{1}{Z} (16 \cdot 1e^8 + 4 \cdot 4e^0 + 0 \cdot 2e^{-8} + 0 \cdot 4e^0 + 4 \cdot 4e^0 + 16 \cdot 1e^8)$$

$$\langle M^2 \rangle = \frac{1}{Z} (32 + 32e^8) = 15.9732$$

$$\langle \chi \rangle = 0.01604$$

$$\langle \chi \rangle/N = 0.004010$$

Below you can see a summary for the quantities:

$$\langle E \rangle/N = 1.9959$$

$$C_V/N = 0.03208$$

$$\langle |M| \rangle/N = 0.9986$$

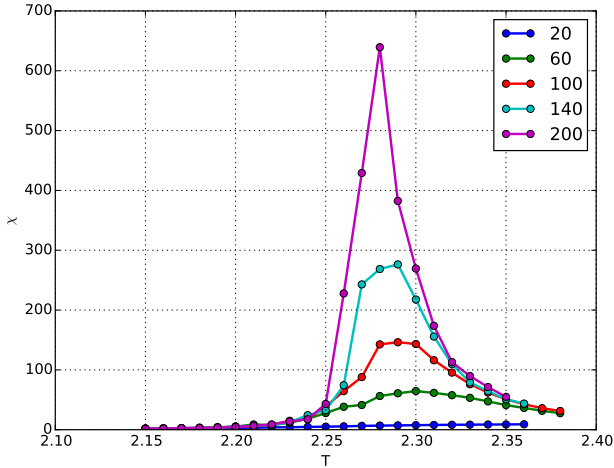
$$\langle \chi \rangle/N = 0.004010$$

4.2 Simulation 2x2

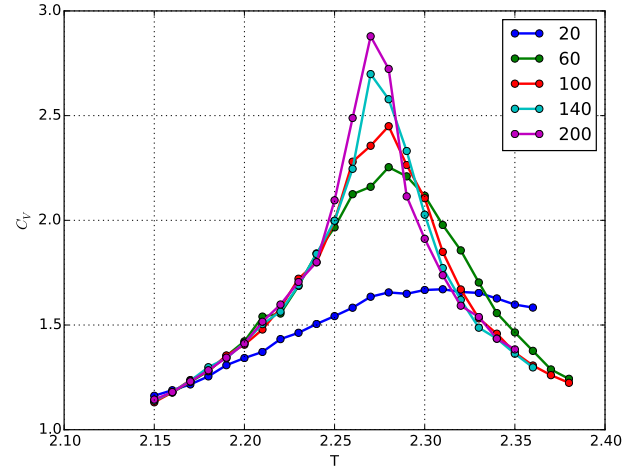
4.3 Phase transition

4.3.1 Numerical studies of phase transition

Several simulations were made with different L . Below a selected few of these simulations are shown. All the results are available at my [github](#)

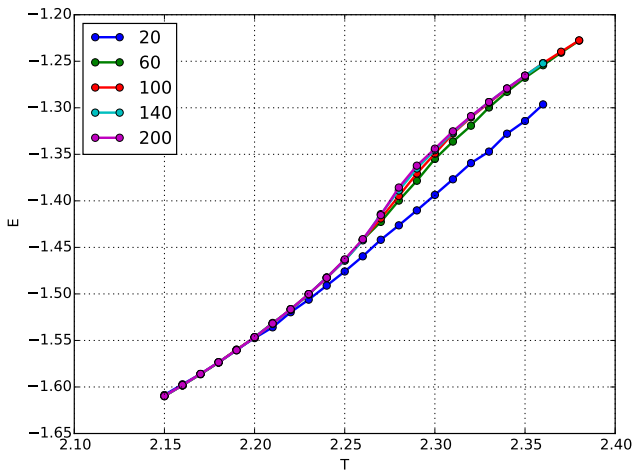


(a)

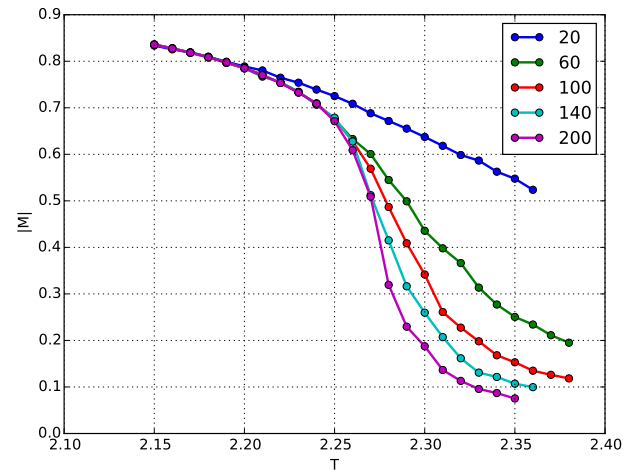


(b)

Figure 1: a) shows the orbit of earth around the sun. The initial velocity is set to 2π in y direction and the start position to 1 au in x direction. b) shows how the error develops. The initial values should give a perfect circular motion. So the error is calculated by $r_i - r_0$. It is apparent that the Verlet-Velocity method is a better approximation. This simulation was with 1000 points with the end time of 1 year. Both simulations was produced by [plot_earth_sun.py](#)



(a)



(b)

Figure 2: a) shows the orbit of earth around the sun. The initial velocity is set to 2π in y direction and the start position to 1 au in x direction. b) shows how the error develops. The initial values should give a perfect circular motion. So the error is calculated by $r_i - r_0$. It is apparent that the Verlet-Velocity method is a better approximation. This simulation was with 1000 points with the end time of 1 year. Both simulations was produced by [plot_earth_sun.py](#)

4.3.2 Extracting the critical temperature

$$T_C(L) - T_C(L = \infty) = aL^{-\frac{1}{v}}$$

To find a we set the second L to a known value

$$T_C(L = 100) - T_C(L = 140) = aL^{-\frac{1}{v}}$$

$$a = (T_C(L = 100) - T_C(L = 140)) L^{\frac{1}{v}}$$

$$T_C(L = \infty) = T_C(L = 100) - aL^{-\frac{1}{v}}$$

$$T_C(L = \infty) = T_C(L = 100) - (T_C(L = 100) - T_C(L = 140))$$

$$T_C(L = \infty) = 2.26$$

5 Conclusion

6 References

References

- [1] Morten Hjorth-Jensen. *Computational Physics*. Lecture notes. 2015. URL: <https://github.com/CompPhysics/ComputationalPhysics/blob/master/doc/Lectures/lectures2015.pdf>.
- [2] Morten Hjorth-Jensen. *Computational Physics*. Project-4. 2017. URL: <https://github.com/CompPhysics/ComputationalPhysics/blob/master/doc/Projects/2017/Project4/pdf/Project4.pdf>.

7 Appendix

```
//FLOPs FOR POSITION :: EULER
// 2 FLOPs * 3 directions
x = x + t_step*Vx
//TOTAL FLOPs = 6 FLOPs
```