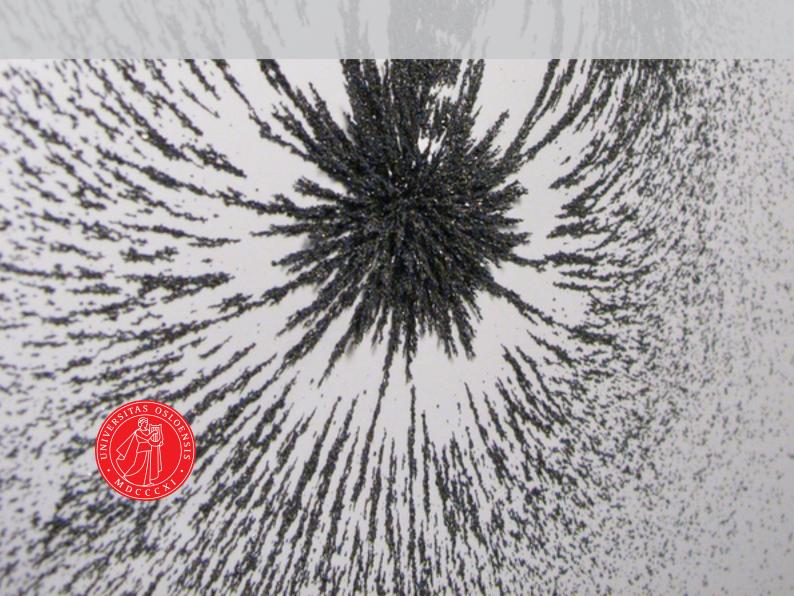


# UiO Department of Physics University of Oslo

# MAGNETIC SYSTEMS FYS4150

Erik Skaar



# Contents

1	Introduction					
<b>2</b>	Theory					
	2.1 the Ising model	3				
	2.1.1 Partition function	3				
	2.1.2 Energy	3				
	2.1.3 the mean magnetic moment $ M $	3				
	2.1.4 the specific heat $C_V$	3				
	2.1.5 the susceptibility $\chi$	3				
3	Method	3				
4	Result & Discussion	3				
	4.1 Earth-Sun system	3				
	4.1.1 Stability	3				
5	Conclusion	4				
6	References	5				
7	Appendix	6				

CONTENTS Page 1 of 7

## Abstract

### 1 Introduction

These laws are not enough to solve the motion of the planets. From the laws one can derive differential equations for the motion, which are not trivial or even possible to solve analytically. This is where computational methods are useful. With the tools developed in computational physics we can make a prediction to the motion of the planets in our solar system.<sup>1</sup> And because of our assignment we kind of have to do this to pass the course.[1]

1 INTRODUCTION Page 2 of 7

<sup>&</sup>lt;sup>1</sup>Semester page for FYS3150 - Autumn 2017.

# 2 Theory

#### 2.1 the Ising model

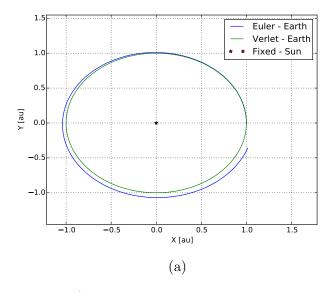
- 2.1.1 Partition function
- 2.1.2 Energy
- 2.1.3 the mean magnetic moment |M|
- 2.1.4 the specific heat  $C_V$
- 2.1.5 the susceptibility  $\chi$

#### 3 Method

#### 4 Result & Discussion

#### 4.1 Earth-Sun system

#### 4.1.1 Stability



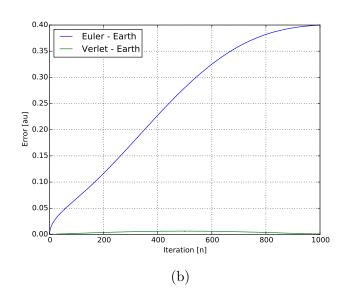


Figure 1: a) shows the orbit of earth around the sun. The intial velocity is set to  $2\pi$  in y direction and the start position to 1 au in x direction. b) shows how the error develops. The intial values should give a perfect circular motion. So the error is calculated by  $r_i - r_0$ . It is apparent that the Verlet-Velocity method is a better approximation. This simulation was with 1000 points with the end time of 1 year. Both simulations was produced by plot\_earth\_sun.py

Table 1: Time table for the different algorithms. The algorithms use nearly the same time. This is not a shocker since the number of FLOPs for the algorithms are similar, see section (??). Note: this is only the result from one test, but several was done. Both algorithms were very close and it seems to be random which is fastest.

n	Forward-Euler	Verlet-Velocity	fastest	slowest fastest
10	0.000136	0.000148	Euler	1.08823529412
100	0.000208	0.000179	Verlet	1.16201117318
1000	0.000392	0.000389	Verlet	1.00771208226
10000	0.002427	0.002426	Verlet	1.00041220115
100000	0.022931	0.022293	Verlet	1.02861884897
1000000	0.167022	0.175944	Euler	1.05341811258
10000000	1.58721	1.52666	Verlet	1.03966174525
100000000	15.1786	15.1176	Verlet	1.00403503202

# 5 Conclusion

5 CONCLUSION Page 4 of 7

# 6 References

# References

[1] Morten Hjorth-Jensen. Computational Physics. Project-4. 2017. URL: https://github.com/CompPhysics/ComputationalPhysics/blob/master/doc/Projects/2017/Project4/pdf/Project4.pdf.

REFERENCES Page 5 of 7

# 7 Appendix

```
//FLOPs FOR ACCELERATION
// 1 FLOP * 3 directions
dx = x1 - x2
// 3 FLOPs
r = sqrt(dx*dx + dy*dy + dz*dz)
// 7 FLOPs
a = - (Gconst*m*M/(r*r)) / (m*r)
// 2 FLOPs * 3 directions
a = a + a*(x1-x2);
//TOTAL FLOPs = 19 FLOPs
//FLOPs FOR POSITION :: EULER
// 2 FLOPs * 3 directions
x = x + t_step*Vx
//TOTAL FLOPs = 6 FLOPs
//FLOPs FOR VELOCITY :: EULER
// 2 FLOPs * 3 directions
Vx = Vx + t_step*ax
//TOTAL FLOPs = 6 FLOPs
//FLOPs FOR POSITION :: Verlet
// 6 FLOPs * 3 directions
x = x + t_step*Vx + (0.5*t_step*t_step*a);
//TOTAL FLOPs = 21 FLOPs
//FLOPs FOR VELOCITY :: Verlet
// 4 FLOPs * 3 directions
Vx = Vx + (0.5*t_step*(Ax+Ax_old));
//TOTAL FLOPs = 12 FLOPs
```

7 APPENDIX Page 6 of 7

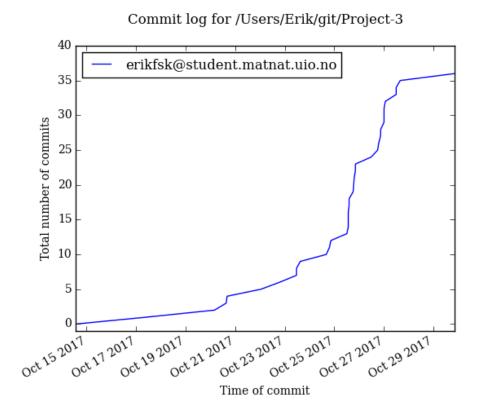


Figure 2: Our retarded workflow... Next time maybe it will be better?

7 APPENDIX Page 7 of 7