



# Radiation Detection and Measurement

Lecture 4

Chapter 2: Radiation Interactions

# Interactions with fast electrons: specific energy loss

- Lose energy at a lower rate than heavy charged particles, leading to longer path lengths.
- Specific Energy Loss:
  - Losses due to collision (ionization and excitation)

$$-\left(\frac{dE}{dx}\right)_c = \frac{2\pi e^4 N z}{m_o v^2} \left( \ln \frac{m_o v^2 E}{2I^2 (1 - \beta^2)} - (\ln 2) \left( 2\sqrt{1 - \beta^2} - 1 + \beta^2 \right) + (1 - \beta^2) + \frac{1}{8} \left( 1 - \sqrt{1 - \beta^2} \right)^2 \right)$$

- where  $\beta = v/c$ , and  $I$  is the average excitation and ionization potential of the absorber.

# Interactions with fast electrons: specific energy loss

- Losses due to radiation (i.e. Bremsstrahlung)

$$-\left(\frac{dE}{dx}\right)_r = \frac{NEz(z+1)e^4}{137m_o^2c^4} \left(4 \cdot \ln \frac{2E}{m_o c^2} - \frac{4}{3}\right)$$

- Total stopping power is the sum of the two energy loss equations.

$$\frac{dE}{dx} = \left(\frac{dE}{dx}\right)_c + \left(\frac{dE}{dx}\right)_r$$

# Interactions with fast electrons: specific energy loss

- The ratio of the specific energy loss is

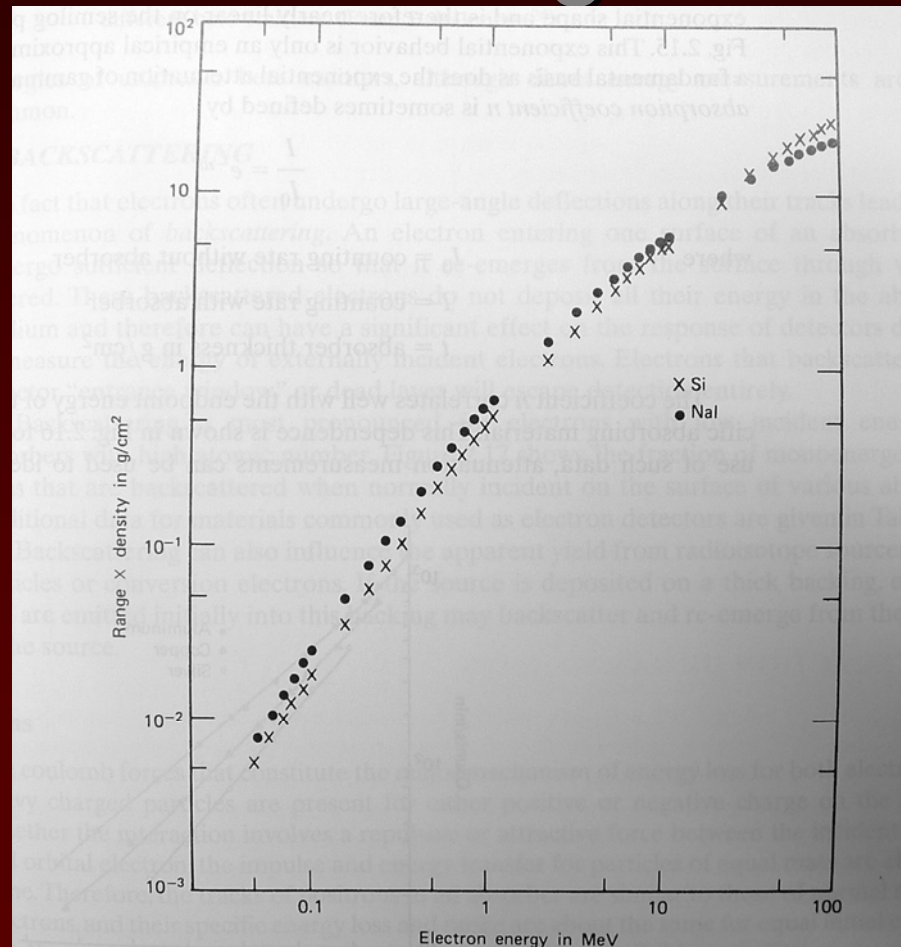
$$\frac{(dE / dx)_r}{(dE / dx)_c} = \frac{EZ}{700}$$

- where E [MeV]

# Electron range and transmission curves: absorption of monoenergetic $e^-$

- Similar to the discussion of  $\alpha$ 's, one can acquire an extrapolated range by plotting  $\Delta$ Intensity vs. thickness of an absorber.
- The  $e^-$ 's that penetrate the furthest are those with the straightest path (least loss of E).
- Range is not well defined since paths can vary greatly and are tortuous.
- One can make a crude estimate for range:
  - $\sim 2\text{mm/MeV}$  for low density materials.
  - $\sim 1\text{mm/MeV}$  for moderate density materials.
- For a reasonable approximation the product of the range and the density of the absorber is constant for differing materials with the same initial velocity  $e^-$ .
- Fig. 2.14 shows this relation for 2 materials (Si & NaI) which have very differing properties and atomic #'s.

# Electron range and transmission curves: absorption of monoenergetic $e^-$



**Figure 2.14** Range-energy plots for electrons in silicon and sodium iodide. If units of mass thickness (distance  $\times$  density) are used for the range as shown, values at the same electron energy are similar even for materials with widely different physical properties or atomic number. (Data from Mukoyama.<sup>24</sup>)

# Electron range and transmission curves: absorption of $\beta$ Particles

- “Soft” or low energy  $\beta$  particles are readily absorbed even in small thickness of an absorber.
- Empirically (without fundamental derivation) they behave with exponential losses:

$$\frac{I}{I_o} = e^{-nt}$$

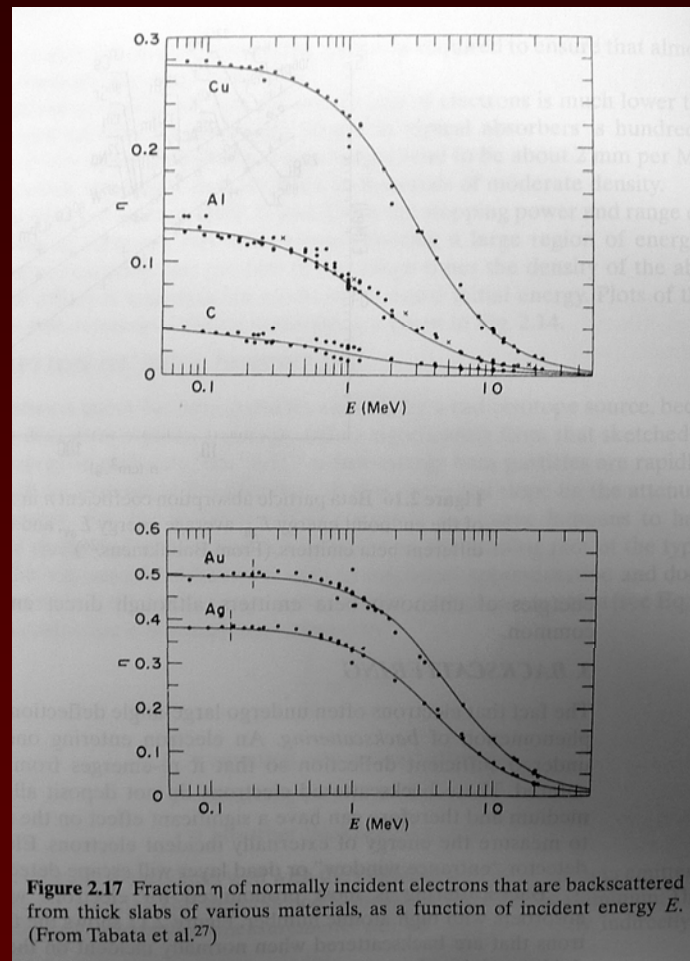
- Where  $I_o$  – counting rate without an absorber,
- $I$  – counting rate with absorber,
- $t$  – thickness of absorber in  $[\text{g}/\text{cm}^2]$ ,
- $n$  – an absorption coefficient (similar to  $\mu$  in  $\gamma$  attenuation).

# Electron range and transmission curves: backscattering

- Due to large angle scattering of  $e^-$ 's, the impending  $e^-$  may re-emerge from the surface it entered. This is backscatter.
- This can be a problem in a detectors “entrance window”, where the  $e^-$ 's that backscatter are not detected.
- Effect is largest for low incident energy  $e^-$  and high atomic number absorbers.



# Electron range and transmission curves: backscattering



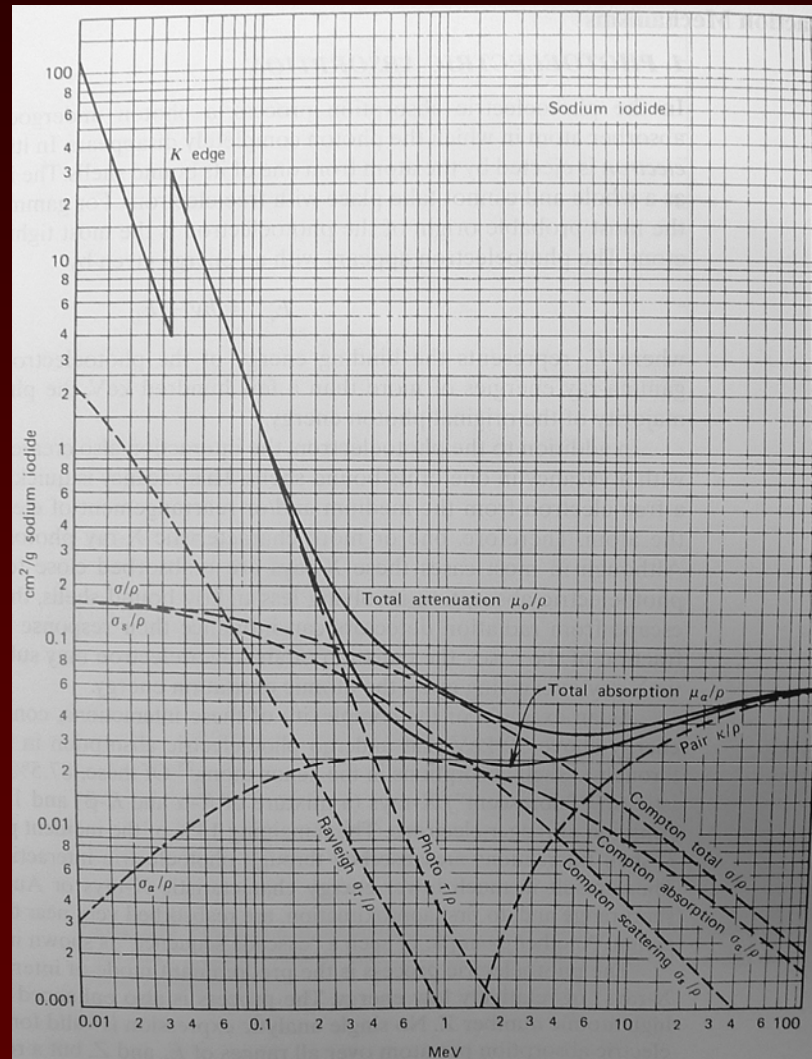
# Electron range and transmission curves: positron ( $\beta^+$ ) interactions

- Still driven by coulombic interactions, so similar to  $\beta^-$ , only interaction is annihilation.

# Interaction of $\gamma$ -rays (photons): mechanisms

- $\gamma$ 's will interact in one of four ways:
  - Compton Scattering
  - Raleigh (Coherent) Scattering
  - Photoelectric Absorption
  - Pair Production.

# Interaction of $\gamma$ -rays (photons): mechanisms



**Figure 2.18** Energy dependence of the various gamma-ray interaction processes in sodium iodide. (From *The Atomic Nucleus* by R. D. Evans. Copyright 1955 by the McGraw-Hill Book Company. Used with permission.)

# Photoelectric absorption

- $\gamma$  is absorbed into the atom exciting an  $e^-$  and ejecting it.
- Generally the  $e^-$  is from an inner shell (K).
- Energy of the (photo)  $e^-$   $E_e = h\nu - E_b$ , where  $h\nu$  is the energy of the  $\gamma$ -ray and  $E_b$  is the binding energy of the  $e^-$
- A vacancy is left and subsequent characteristic x-ray ( $\gamma$ ) or Auger  $e^-$  are produced.
- This is a primary mode of interaction for low energy  $\gamma$ 's or high  $z$  absorbers.

# Photoelectric absorption

- The probability of interaction is approximated by:

$$\tau \cong \text{const} \times \frac{Z^n}{E_\gamma^{3.5}}$$

- Where  $n$  varies from 4-5 over the  $\gamma$  ray energy region of interest.
- This shows why high  $z$  materials are used for shields (Pb) & detectors.
- Note: When looking at absorption curves, the discontinuities are commiserate with the binding energies of the  $e^-$ 's. Generally seen for the K-shell (K-edge) it is energetically favorable to emit a photoelectron just above the K-shell energy. Just below the K-edge there is not enough energy to excite the  $e^-$ , thus there is much less absorption below the K-edge.

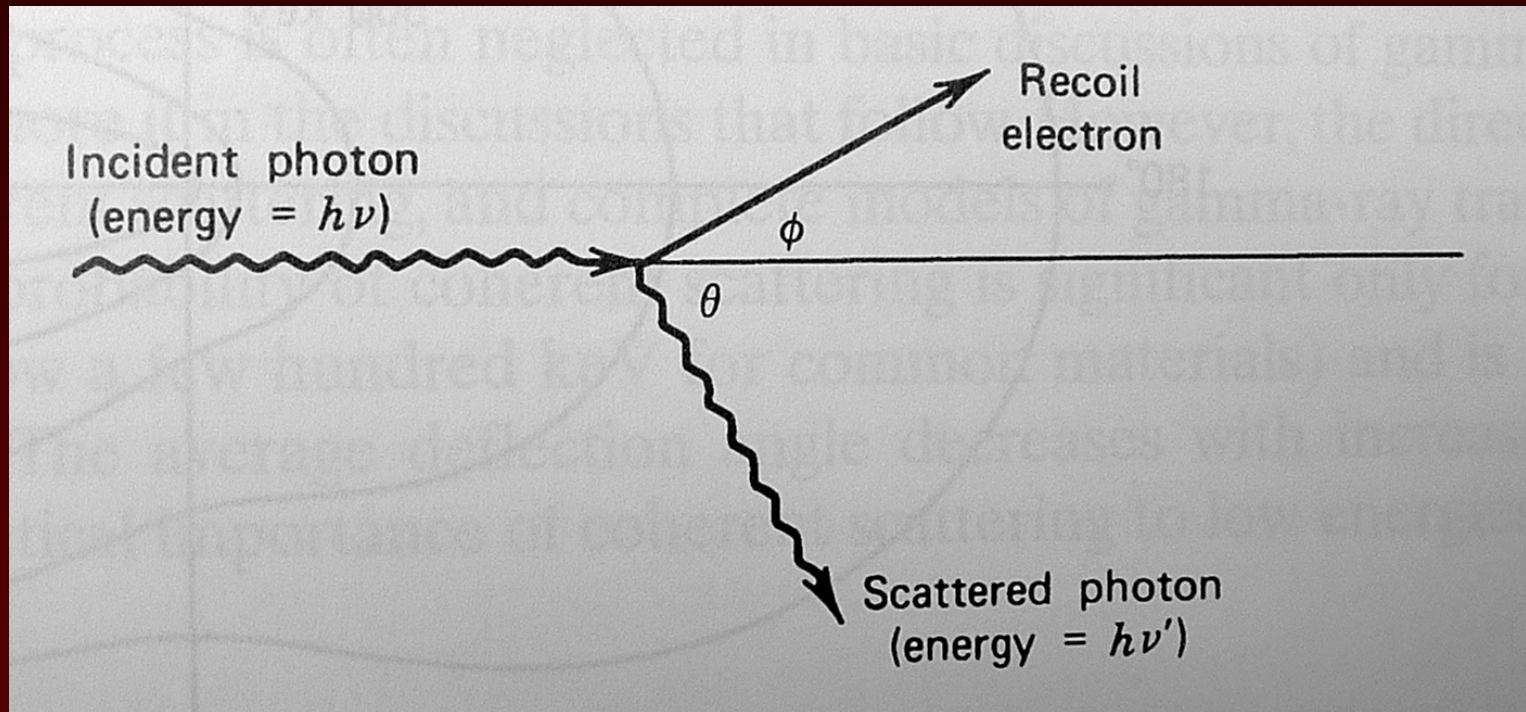
# Compton Scattering

- This is an energy & momentum conserved “billiard ball” reaction (assume the binding energy is negligible).
- It can be shown that:

$$h\nu' = \frac{h\nu}{1 + \frac{h\nu}{m_0c^2}(1 - \cos \theta)}$$

- Where  $h\nu'$  is the energy of the scattered  $\gamma$ ,  $h\nu$  is the energy of the incident  $\gamma$ ,  $m_0c^2$  is the rest mass energy of the recoil  $e^-$ , and  $\theta$  is the scattered  $\gamma$  angle.

# Compton Scattering





# Compton Scattering

- This equation shows direction but not intensity.
- Some energy is always retained by the incident  $\gamma$ , even when  $\theta=\pi$ .
- The probability of Compton scattering per atom of the absorber depends on the number of electrons available as scattering targets and thus increases with  $z$  and gradually falls off with increasing energy.

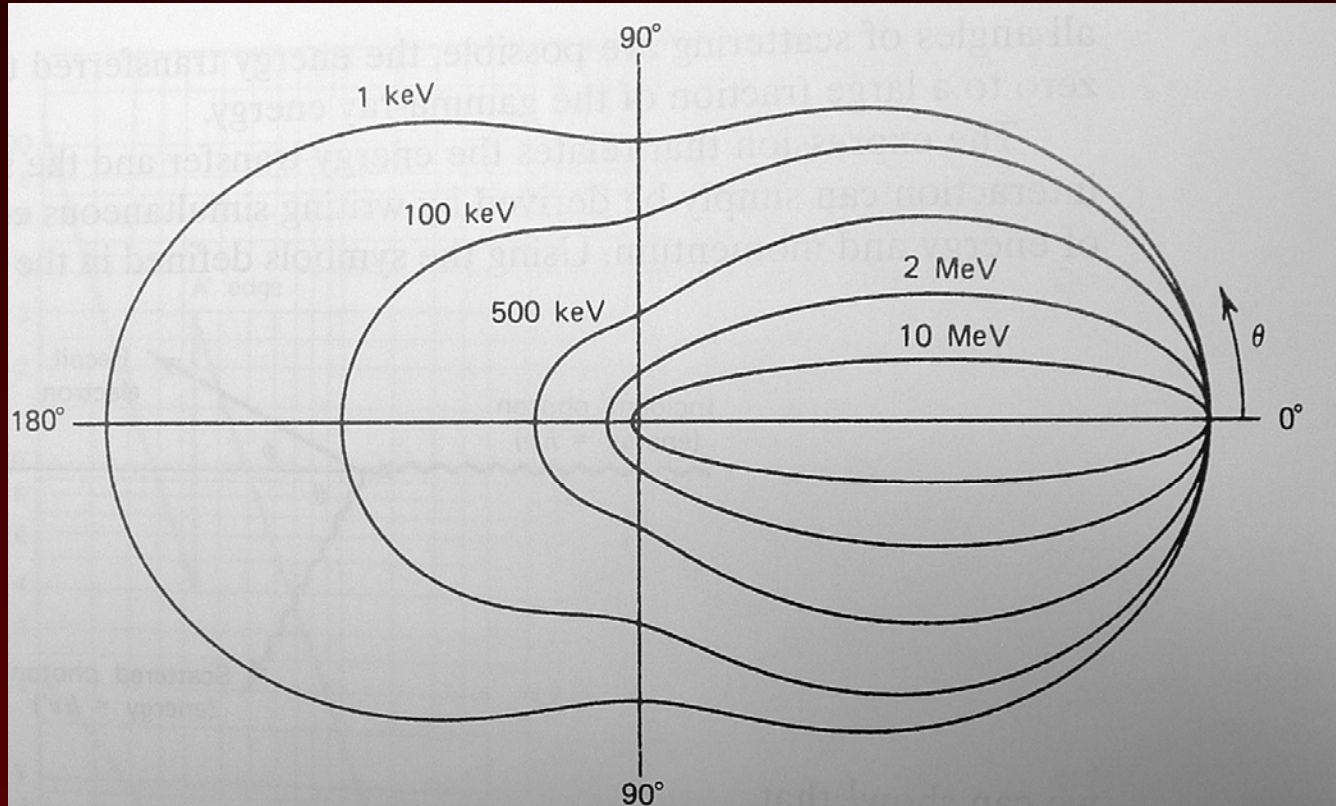
# Compton Scattering

- The angular scattering of Compton scattering is predicted by the Klein-Nashina formula for differential scattering cross section ( $d\sigma/d\Omega$ ):

$$\frac{d\sigma}{d\Omega} = z r_o^2 \left( \frac{1}{1 + \alpha(1 - \cos \theta)} \right)^2 \left( \frac{1 + \cos^2 \theta}{2} \right) \left( 1 + \frac{\alpha^2 (1 - \cos \theta)^2}{(1 + \cos^2 \theta)[1 + \alpha(1 - \cos \theta)]} \right)$$

- Where  $\alpha \equiv \frac{h\nu}{m_o c^2}$ ,  $r_o$  is the classical electron radius, and  $d\sigma/d\Omega$  has azimuthal symmetry. A graph of the distribution is seen in Fig. 2.19.
- Fig. 2.19 is a polar plot of the number of photons Compton scattered into a unit solid angle at the scattering angle  $\theta$ .
- Note that the distribution moves forward with increasing energy.

# Compton Scattering

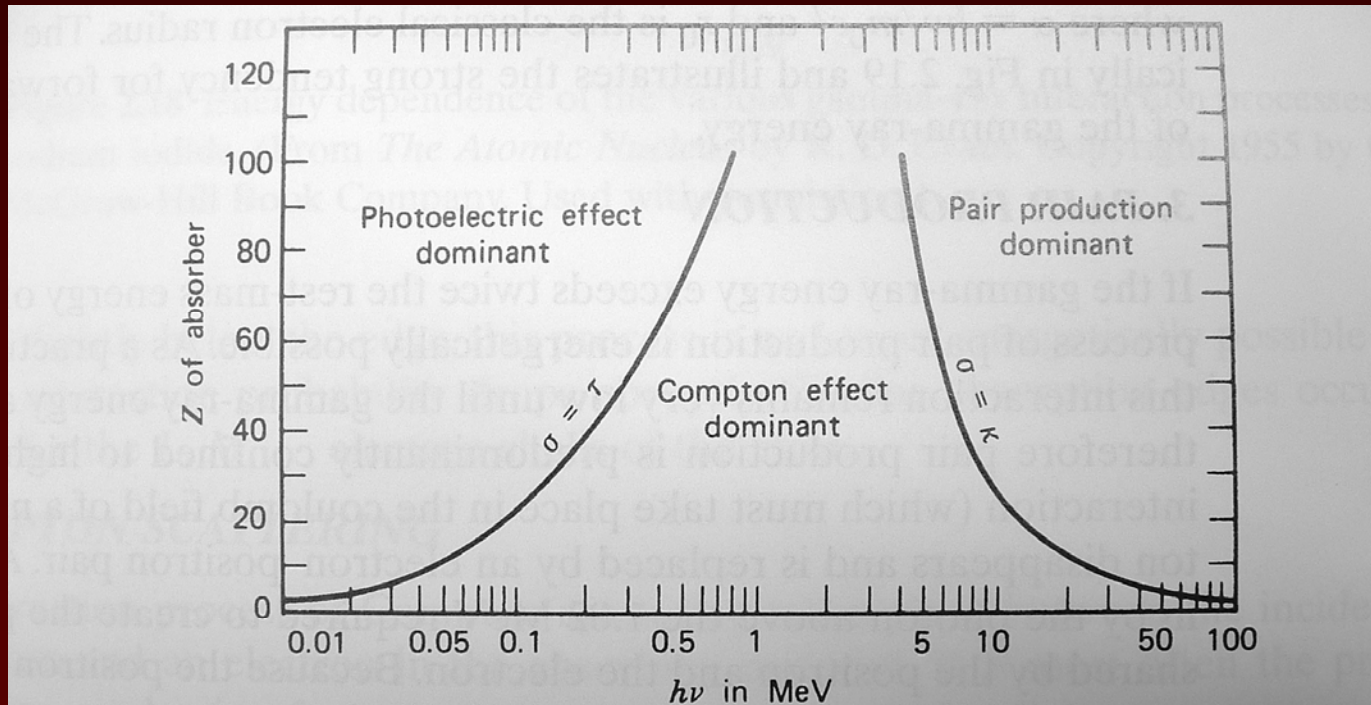


**Figure 2.19** A polar plot of the number of photons (incident from the left) Compton scattered into a unit solid angle at the scattering angle  $\theta$ . The curves are shown for the indicated initial energies.

# Pair Production

- $\gamma$  energy must exceed the level to produce 2 electrons (1.02 MeV).
- Probability of pair production is low until  $E_\gamma$  reaches several MeV, and rises sharply with energy.
- Must take place in the coulomb field of the nucleus.
- Fig. 2.20 shows the relative importance of each of the interactions with energy.

# Relative importance of the three major types of gamma-ray interaction



**Figure 2.20** The relative importance of the three major types of gamma-ray interaction. The lines show the values of  $Z$  and  $h\nu$  for which the two neighboring effects are just equal. (From *The Atomic Nucleus* by R. D. Evans. Copyright 1955 by McGraw-Hill Book Company. Used with permission.)

# Coherent Scattering (Raleigh)

- No excitation or ionization, but involves all of the electrons in the absorber atom.
- Incident and scattered photon directions are different (can contribute to noise in imaging).
- Probability is high for low energy  $\gamma$ 's (< a few 100 keV for common materials).
- Most prominent in high z absorbers.
- Gives the sky its overall blue color.
- Average deflection angle increases with decreasing energy.

# Gamma-ray attenuation

- Describe the statistical removal of  $\gamma$ 's from the mono-energetic beam.
- Defined by the probabilities of each interaction previously described:

$$\mu = \tau(\textit{photoelectric}) + \sigma(\textit{Compton}) + \kappa(\textit{pair})$$

–  $\mu$  is the linear attenuation coefficient.

$$\frac{I}{I_o} = e^{-\mu t}$$

- (loss of relative intensity for thickness (t) of the absorber).

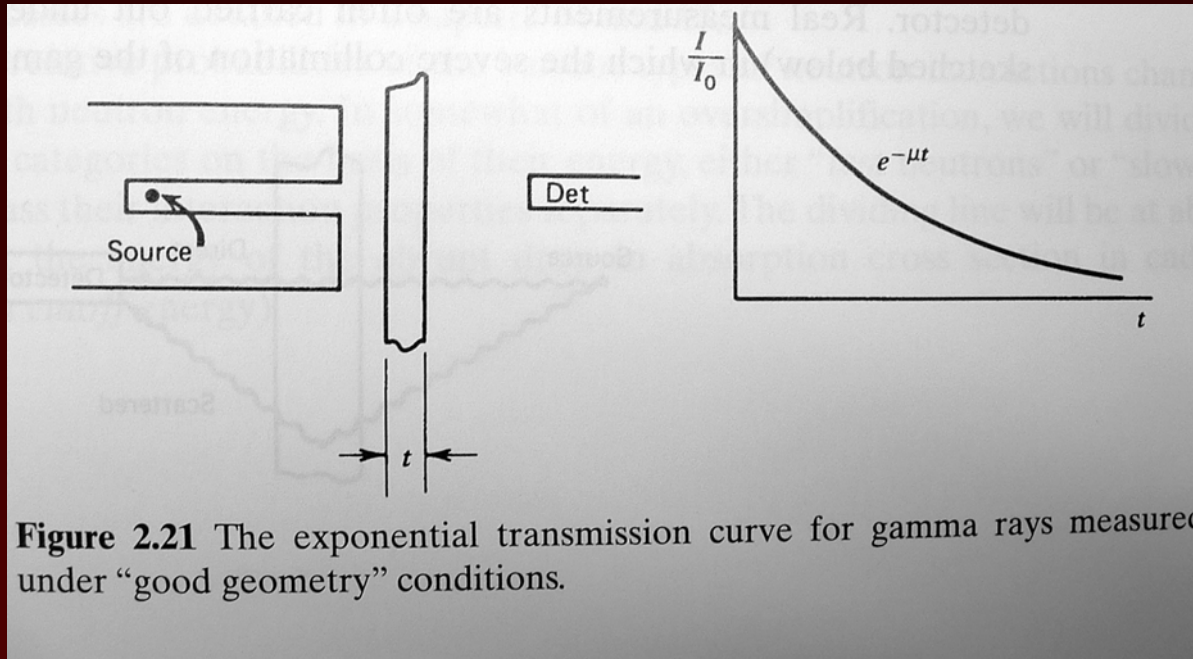
# Gamma-ray attenuation

- Can also define a mean free path,  $\lambda$ :

$$\lambda = \frac{\int_0^{\infty} x e^{-\mu x} dx}{\int_0^{\infty} e^{-\mu x} dx} = \frac{\frac{e^{-\mu x}}{-\mu} \left( x - \frac{1}{\mu} \right) \Big|_0^{\infty}}{\frac{e^{-\mu x}}{-\mu} \Big|_0^{\infty}} = \frac{0 - \frac{1}{\mu^2}}{0 - \frac{1}{\mu}} = \frac{1}{\mu}$$



# Gamma-ray attenuation



**Figure 2.21** The exponential transmission curve for gamma rays measured under “good geometry” conditions.

- Note that the curve in Fig. 2.21 is under “good” geometry conditions (i.e. point source, large distance between source and detector etc.).

# Gamma-ray attenuation

- The mass attenuation coefficient =  $\mu/\rho$ , where  $\rho$  is the density of the medium.
- $\mu/\rho$  for a compound (like scaling) is

$$\left(\frac{\mu}{\rho}\right)_c = \sum_i w_i \left(\frac{\mu}{\rho}\right)_i$$

- where  $w_i$  is the weight fraction of element  $i$  in  $C$ .

# Gamma-ray attenuation

- Using  $\mu/\rho$  we can rewrite the attenuation relation:

$$\frac{I}{I_o} = e^{-\left(\frac{\mu}{\rho}\right)\rho t}$$

- where  $\rho t$  is the mass thickness.
- This is useful for discussing the energy loss in fast electrons and charged particles since for absorber materials of similar protons and neutrons, the particle will encounter about the same number of electrons for a given (equal) mass thickness.
- Stopping power and range expressed in  $\rho t$  will be about the same for absorbers of similar  $z$ .

# Buildup

- To account for poor geometry in a source detector set up we can rewrite the attenuation as:

$$\frac{I}{I_o} = B(t, E_\gamma) e^{-\mu t}$$

- This correction for broad beam geometry keeps the exponential term for the major changes and uses a simple multiplicative correction factor.

# Buildup

- $B(t, E_\gamma)$  will vary with  $\gamma$ -energy and detector type and accounts for secondary (scattered)  $\gamma$ 's in the absorber (generally the source has little or no geometry).
- $B(t, E_\gamma)$  is greater than one and is determined experimentally.
- Rule of thumb:  $B(t, E_\gamma)$  for a thick slab tends to be equal to the thickness of the absorber in units of  $\lambda$ , provided the detector responds to a broad range of  $\gamma$ 's.

# Buildup

