OVERVIEW

- Molecular speeds
- Maxwell distribution
- Boltzmann distribution

Molecular speeds I

Temperature is the measure of the average (translational) kinetic energy of the molecules

$$\frac{m\langle v^2\rangle}{2} = \frac{3}{2}k_BT$$

Define the so-called root mean square (rms) speed

$$v_{\rm rms} = \sqrt{\langle v^2 \rangle} = \sqrt{\frac{3k_BT}{m}}$$

$v_{\rm rms}$

- increases with temperature
- is smaller for heavier molecules

Molecular speeds II

What exactly is $\langle ... \rangle$? For *N* molecules

$$\langle v_x^2 \rangle \equiv \frac{1}{N} \underbrace{\left(v_{x;1}^2 + v_{x;2}^2 + \dots + v_{x;N}^2\right)}_{N \ \mathrm{terms}}$$

where the sum is over all the molecules.

Q: How many molecules have a given value of v_x ?

A: Can't really tell, v_x is a continuous variable

Q: How many molecules have v_x in a certain interval dv_x ?

Velocity distribution

The number of molecules with v_x between v_x and $v_x + dv_x$:

$$dN = Nf(v_x) dv_x$$

where $f(v_x)$ is the distribution function.

Two equivalent interpretations: $f(v_x) dv_x$ is

- the fraction of molecules with x-projections of velocity between v_x and $v_x + dv_x$
- the *probability* that a randomly chosen molecule has v_x in this interval

Velocity distribution II

Normalization (the probablilities sum to unity):

$$\int_{-\infty}^{\infty} f(v_x) \, dv_x = 1$$

$$\int_{v_1}^{v_2} f(v_x) dv_x = \text{fraction of molecules with } v_x \text{ between } v_1 \text{ and } v_2$$

Averages:

$$\langle v_x \rangle = \int_{-\infty}^{\infty} v_x f(v_x) dv_x$$

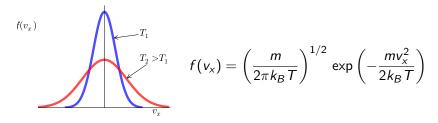
 $\langle v_x^2 \rangle = \int_{-\infty}^{\infty} v_x^2 f(v_x) dv_x$

By symmetry
$$f(-v_x) = f(v_x)$$
 \Rightarrow $\langle v_x \rangle = 0$, but $\langle v_x^2 \rangle \neq 0$

MAXWELL DISTRIBUTION

Ideal gas only

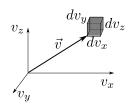
Maxwell distribution (for velocity projection):



- is symmetric around zero
- the width is given by the temperature

MAXWELL DISTRIBUTION

Q: What is the distribution of the *velocity*, rather than velocity *projection*?



$$v = |\vec{v}| = \sqrt{v_x^2 + v_y^2 + v_z^2}$$

Need to: (i) find probability that all three projections of \vec{v} are within a cube $d\Omega$, and then (ii) integrate over the possible directions of \vec{v} .

$$\begin{aligned} \textbf{Prob}(v_{x}, v_{y}, v_{z} \in d\Omega) &= f(v_{x}) \, dv_{x} \cdot f(v_{y}) \, dv_{y} \cdot f(v_{z}) \, dv_{z} \\ &\propto \exp\left(-\frac{mv^{2}}{2k_{B}T}\right) dv_{x} dv_{y} dv_{z} \end{aligned}$$

Maxwell distribution

$$\begin{aligned} \textbf{Prob}(v_{x}, v_{y}, v_{z} \in d\Omega) &= f(v_{x}) \, dv_{x} \cdot f(v_{y}) \, dv_{y} \cdot f(v_{z}) \, dv_{z} \\ &\propto \exp\left(-\frac{mv^{2}}{2k_{B}T}\right) dv_{x} dv_{y} dv_{z} \end{aligned}$$

Since this does not have any angular dependence, integrating over the directions is easy:

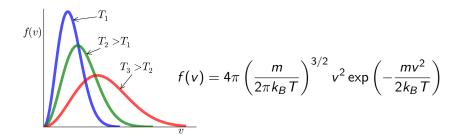
$$\iint_{\text{angles}} (\cdots) \, dv_x dv_y dv_z = 4\pi v^2 dv$$

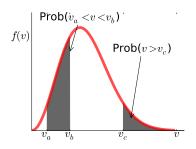
Finally,

$$f(v) = 4\pi \left(\frac{m}{2\pi k_B T}\right)^{3/2} v^2 \exp\left(-\frac{mv^2}{2k_B T}\right)$$



MAXWELL DISTRIBUTION





Normalization:

$$\int_{0}^{\infty} f(v) dv =$$

CHARACTERISTIC VELOCITIES OF THE MAXWELL DISTRIBUTION

$$f(v)$$
 $w_{
m av}$ $w_{
m av}$ $w_{
m mp}$ $w_{
m mp} = \sqrt{\frac{2k_BT}{m}}$

Average speed:
$$v_{\rm av} = \langle v \rangle = \int_0^\infty v \, f(v) \, dv = \sqrt{\frac{8k_BT}{\pi m}}$$

rms speed:
$$v_{\rm rms} = \sqrt{\langle v^2 \rangle} = \left(\int_0^\infty v^2 f(v) \, dv \right)^{1/2} = \sqrt{\frac{3k_BT}{m}}$$



MAXWELL, BOLTZMANN, GIBBS

MAXWELL DISTRIBUTION

$$\mathsf{Prob}(v_{\mathsf{x}} \in [u, u + dv_{\mathsf{x}}]) \propto \exp\left(-\frac{\mathrm{mu}^2/2}{k_B T}\right) dv_{\mathsf{x}}$$

GIBBS DISTRIBUTION

In fact, a more general result holds: if the energy of a system depends on some variable ξ , then in thermal equilibrium

Prob
$$(\xi \in [\xi, \xi + d\xi]) \propto \exp\left(-\frac{E(\xi)}{k_B T}\right) d\xi$$

For ideal gases, this was first derived by L. Boltzmann, later generalized by J. Gibbs.



Example: Maxwell-Boltzmann distribution

Isothermal atmosphere: T = const

The energy of a single molecule is $E = \frac{mv^2}{2} + mgh$, hence the *joint distribution* of v and h,

$$f(v,h) \propto v^2 \exp\left(-\frac{mv^2}{2k_BT}\right) \exp\left(-\frac{mgh}{k_BT}\right)$$

Intuitively, f(v, h) is a fraction of molecules with velocities around v at altitudes around h.

Formally:

$$f(v) = \int_0^\infty f(v,h) \, dh \qquad \text{(Maxwell distribution)}$$

$$f(h) = \int_0^\infty f(v,h) \, dv \qquad \text{gives density variation with the altitude}$$

Real atmosphere is a little more complicated.

Image credits: NOAA.

