

## Problem Set 6

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### Problem 1

For an object orbiting in the Lagrange point  $L1$  between the Earth and the Moon, the following must be true

$$F_{\text{Earth}} + F_{\text{Moon}} = F_C, \quad (1)$$

where  $F_{\text{Earth}}$  is the gravitational force on the object due to the Earth,  $F_{\text{Moon}}$  is the gravitational force on the object due to the Moon, and  $F_C$  is the centripetal force on the object.  $F_{\text{Earth}}$  and  $F_{\text{Moon}}$  are given by Newton's gravitational law and  $F_C = \omega^2 r$ . Equation 1 thus becomes

$$\frac{GM}{r^2} - \frac{Gm}{(R-r)^2} = \omega^2 r, \quad (2)$$

where  $M$  is the mass of the Earth,  $m$  is the mass of the Moon,  $r$  and  $\omega$  are the radius and angular frequency of the orbit at  $L1$ ,  $R$  is the radius of the Moon's orbit, and  $G$  is the gravitational constant. The angular frequency  $\omega$  can be expressed in terms of  $G$ ,  $R$ , and  $M$ ,

$$\omega^2 = \frac{GM}{R^3}. \quad (3)$$

We can replace our variables with the dimensionless quantities  $r' = \frac{r}{R}$  and  $m' = \frac{m}{M}$ . This gives

$$\frac{m'}{(1-r')^2} + r' - \frac{1}{r'^2} = 0. \quad (4)$$

Using Newton's method and jax for the analytic derivative, we can compute the distance to  $L1$ ,  $r$ , for two objects of any masses  $m$ ,  $M$ , orbiting at a distance  $R$  by finding the roots of equation 4. From a plot of equation 4, (fig 1) we can see that the root is somewhere near the middle of the interval  $(0, 1)$ . Thus, we will choose a starting point for Newton's method near 0.

Using following constants

Earth Mass	$5.974 \times 10^{24} \text{kg}$
Moon Mass	$7.384 \times 10^{22} \text{kg}$
Solar Mass	$1.988 \times 10^{30} \text{kg}$
Jupiter Mass	$1.898 \times 10^{27} \text{kg}$
Lunar Orbit Radius	$3.844 \times 10^8 \text{m}$
Earth Orbit Radius	$149.60 \times 10^9 \text{m}$

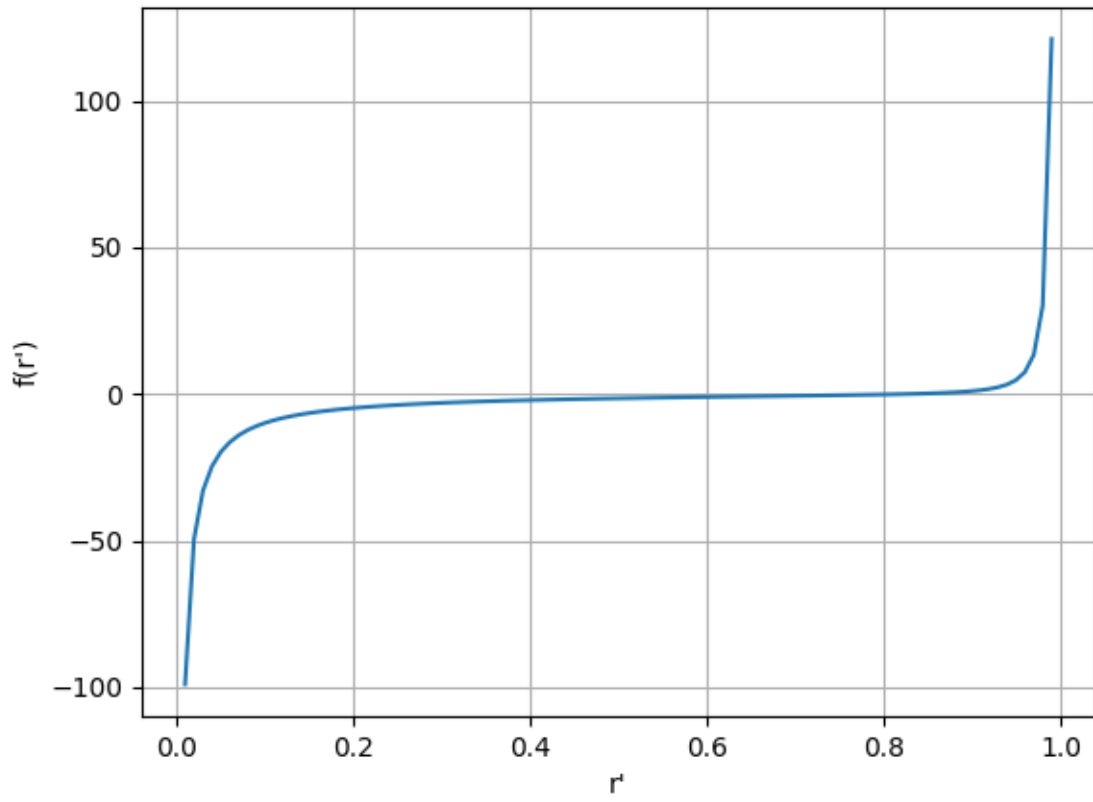


Figure 1: Plot of equation 4 as a function of  $r'$  with  $m' = \frac{m}{M}$

we calculate  $L1$  for the Earth-Moon system to be at a distance of  $3.262 \times 10^8$  meters. For the Earth-Sun system,  $L1$  is at  $1.481 \times 10^{11}$ . For a Jupiter sized planet orbiting the sun in place of Earth,  $L1$  is at a distance of  $1.396 \times 10^{11}$ .

## Problem 2

We want to minimize

$$f(x) = (x - 0.3)^2 e^x \tag{5}$$

using Brent's method. A plot of equation 5 is shown in figure 2. Brent's method uses a combination of parabolic minimization and golden section minimization. My implementation of Brent's method in python gives a minimum of 0.3000259015536602 with a tolerance of 0.0001. Scipy's built in Brent's minimization gives a minimum of 0.2999996417851 with a tolerance of 0.0001. The difference between these two methods is  $2.6259768560188412 \times 10^5$ .

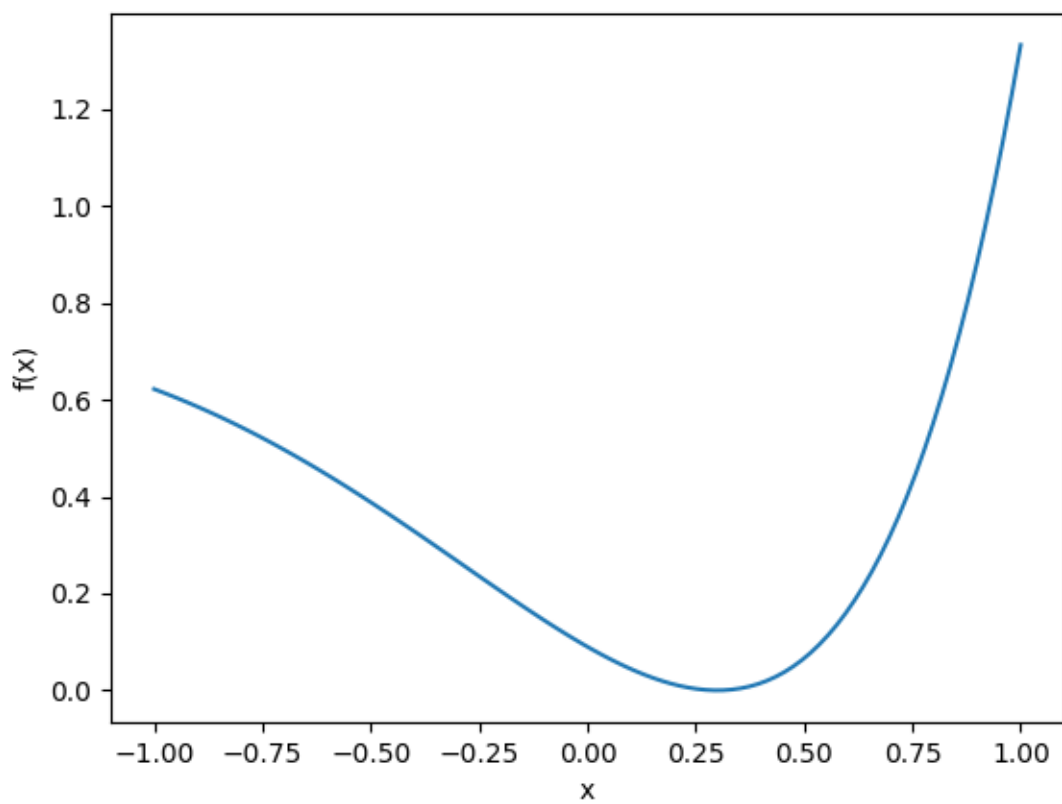


Figure 2: Plot of equation 5 over the interval -1, 1.