

## Nside = 4: Covariance and likelihood.

### Data

We use only CAMB generated scalar  $C_l$  values and CAMB generated maps. See Lewis, Challinor, Lasenby: astro-ph/9911177

Maps noiseless, seeded randomly, Gaussian.

If FITS files located in local "Downloads" directory after download/unzipped .tar file, everything works.

$N_{\text{side}} = 4$  means  $N_{\text{pix}} = 12 \times \text{side}^2 = 192$  total pixels.

$C_l$  scalar values were set to  $l_{\text{max}} = 2N_{\text{pix}}$ .

### Covariance results

The pixel-pixel temperature covariance matrix from CAMB maps is defined as

$$C_{ij} = \langle \Delta T_i \Delta T_j \rangle = \frac{1}{N_{\text{pix}}} \sum_{p=1}^{\text{pix}} (T^i(p) - \bar{T}^i)(T^j(p) - \bar{T}^j)$$

The real-space (temperature space) covariance matrix is defined as

$$C_{ij} = \sum_{l=0}^l \frac{2l+1}{4\pi} C_l^{\text{theor}} P_l(\cos \alpha_{ij})$$

where  $C_{ij}$  is the covariance between pixel  $i$  and pixel  $j$ ,  $C_l^{\text{theor}}$  is the theoretical value of  $C_l$ ,  $P_l$  are the Legendre polynomial, and  $\alpha_{ij}$  is the angle between. The dot product is defined by unit vectors,  $\cos \alpha_{ij} = \hat{n}_i \cdot \hat{n}_j$ .

In spherical harmonic space, we have mean zero and covariance

$$\sum a_{lm} a_{l'm'}^\dagger = (2l+1) C_l \quad (1)$$

### Likelihoods

Likelihood outputs will not be exact, because of some constants. They are approximate.

In spherical harmonic space, the expression is

$$-2 \ln \mathcal{L} = \sum_l (2l+1) \left[ \ln \left( \frac{C_l^{\text{th}}}{\hat{C}_l} \right) + \left( \hat{C}_l / C_l^{\text{th}} \right) - 1 \right] \quad (2)$$

where

$$\hat{C}_l = \frac{1}{2l+1} \sum_m |\hat{a}_{lm}|^2 \quad (3)$$

from HEALPix anafast.

In temperature space (real space), the expression is

$$-2 \ln \mathcal{L} \propto \mathbf{T} \mathbf{S}^{-1} \mathbf{T} + \ln \det \mathbf{S} + N \ln 2\pi \quad (4)$$