

1 Nondimensional equations

The nondimensional system of equations we solve is

$$\nabla \cdot \mathbf{u} = 0 \quad (1)$$

$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} = -\nabla \varpi + T \hat{z} + \text{Re}_0^{-1} \nabla^2 \mathbf{u} \quad (2)$$

$$\partial_t T + \mathbf{u} \cdot \nabla T + w \nabla_{\text{ad}} = \text{Pe}_0^{-1} \nabla^2 T' + \nabla \cdot [k \nabla \overline{T}] + Q \quad (3)$$

These are just the Boussinesq equations of motion, with some slight tweaks:

1. There is a nonzero adiabatic temperature gradient, $\partial_z T_{\text{ad}} = -\nabla_{\text{ad}}$. We choose this sign convention (∇_{ad} and similarly ∇_{rad} will be positive) to align better with the stellar structure community's intuition. $|\nabla T_0|$ and ∇_{ad} may be much larger in magnitude than the convective fluctuations. The background temperature field T_0 does not enter into the momentum equation, as it is automatically canceled by a background $\nabla \varpi$.
2. Convection is driven by internal heating, Q . I've confirmed with some simple tests that $|u|^2 \approx Q$, and therefore the convective frequency is $f_{\text{conv}}^2 \approx Q$.
3. The control freefall Péclet number, Pe_0 , and the diffusivity associated with it, only acts on fluctuations, not on the $k_x = 0$ mode. For the $k_x = 0$ mode, we model the radiative flux as $F_{\text{rad}} = -k \nabla \overline{T}$, where k is a nonconstant coefficient, $k(z)$. k may be a larger or smaller diffusivity than Pe_0^{-1} , and its magnitude is set by $\int Q dz \approx -k \nabla \overline{T}$.

2 Model Setup

We will study a two-layer model in $z = [0, L_z]$ in which the conductivity, $k(z)$ is a step function such that

$$k(z) = \begin{cases} k_b & z < 1 \\ k_t & z \geq 1 \end{cases} \quad (4)$$

where subscript “t” refers to the top layer and “b” refers to the bottom. (These won’t be true discontinuities, but will be smooth step-like functions achieved with erfs.)

There will be two important input parameters in this system:

1. The stiffness,

$$\mathcal{S} = \frac{N_t^2}{f_{\text{conv}}^2}, \quad (5)$$

which is the ratio of the buoyancy frequency (in the stable layer) to the square convective frequency.

2. \mathcal{P} , a parameter which determines what the magnitude of the *negative* convective flux would be in an adiabatic penetrative layer with $k = k_t$. We will define $\mathcal{P} = -F_{\text{conv,rz}}/F_{\text{conv}}$, and our expectation is that we get a lot of convective penetration when $\mathcal{P} \ll 1$ and none when $\mathcal{P} \gg 1$.

In addition to these parameters, the magnitude of the heating, Q and the vertical extent of the heating layer, δ_H , must be specified as they set $F_{\text{conv}} = Q\delta_H$. We will choose $Q = 1$ and $\delta_H = 0.2$ so that $F_{\text{conv}} = 0.2$; we will furthermore choose and $F_{\text{bot}} = \zeta F_{\text{conv}}$, where ζ is some arbitrary scalar that sets how much smaller k is in the cz than the RZ; we’ll set $\zeta = 10^{-3}$ for now.

Under these choices, the system is determined by five equations,

1. At some point, the adiabatic gradient is equal to the radiative gradient,

$$k_{\text{ad}} \nabla_{\text{ad}} = F_{\text{tot}}. \quad (6)$$

2. at $z = 0$, the base of the CZ, the adiabatic temperature gradient is the radiative gradient, $\nabla_{\text{ad}} = \nabla_{\text{rad}}$, and carries F_{bot} ,

$$k_b \nabla_{\text{ad}} = F_{\text{bot}} \quad (7)$$

3. The radiative gradient can carry the flux in the RZ,

$$k_t \nabla_{\text{rad,t}} = F_{\text{tot}} \quad (8)$$

4. Assuming $\nabla T = -\nabla_{\text{ad}}$ everywhere, the convective flux in the top layer is $F_{\text{conv},t} = F_{\text{conv}} - \Delta k \nabla_{\text{ad}}$, where $F_{\text{conv}} = Q\delta_H$ and $\Delta k = k_t - k_b$. Thus, definitionally,

$$\mathcal{P} = -\frac{F_{\text{conv},t}}{F_{\text{conv}}} = \left(\frac{\Delta k \nabla_{\text{ad}}}{Q\delta_H} - 1 \right). \quad (9)$$

5. Defining $f_{\text{conv}}^2 \approx Q$ and $N^2 \approx -(\nabla_{\text{rad},t} - \nabla_{\text{ad}})$, the stiffness gives

$$\mathcal{S} = \frac{N^2}{f_{\text{conv}}^2} = \frac{-(\nabla_{\text{rad},t} - \nabla_{\text{ad}})}{Q} = \frac{F_{\text{tot}}}{Q}(k_{\text{ad}}^{-1} - k_t^{-1}) \quad (10)$$

Solving this system of equations gives

$$k_t = \frac{\delta_H \mathcal{P}}{\mathcal{S}}, \quad k_b = k_t \frac{\zeta}{1 + \zeta + \mathcal{P}}, \quad k_{\text{ad}} = k_t \frac{1 + \zeta}{1 + \zeta + \mathcal{P}}, \quad (11)$$

$$\nabla_{\text{ad}} = \frac{Q\mathcal{S}}{\mathcal{P}}(1 + \zeta + \mathcal{P}), \quad \nabla_{\text{rad}} = \frac{Q\mathcal{S}}{\mathcal{P}}(1 + \zeta). \quad (12)$$

Just to be safe, I'm going to offset the internal heating from the bottom boundary. We wouldn't need to do this if we didn't have a bottom, impenetrable boundary condition with a viscous boundary layer. But, since $w \rightarrow 0$ at the bottom boundary, I don't want to inject heating into the domain somewhere where it must be carried by a superadiabatic temperature gradient (I explicitly don't want to study boundary-driven convection). So

$$Q(z) = \begin{cases} 0 & z < 0.1 \\ Q & 0.1 \leq z < 0.1 + \delta_H, \\ 0 & \text{elsewhere} \end{cases} \quad (13)$$

We then choose T_0 to satisfy

$$\partial_z T_0 = - \begin{cases} \nabla_{\text{ad}} & z \leq 1 \\ \nabla_{\text{rad}} & z > 1 \end{cases} \quad (14)$$

(again, perhaps not perfectly this, but with smooth transitions). And we're off to the races.