

# Entrainment of low Mach number thermals in stratified atmospheres

EVAN H. ANDERS,<sup>1,2</sup> DANIEL LECOANET,<sup>3</sup> AND BENJAMIN P. BROWN<sup>1,2</sup>

<sup>1</sup>*Dept. Astrophysical & Planetary Sciences, University of Colorado – Boulder, Boulder, CO 80309, USA*

<sup>2</sup>*Laboratory for Atmospheric and Space Physics, Boulder, CO 80303, USA*

<sup>3</sup>*Stuff*

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## ABSTRACT

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### 1. INTRODUCTION

#### 2. EXPERIMENT

We perform direct numerical simulations of the evolution of dry thermals in an ideal gas whose pressure ( $P$ ), temperature ( $T$ ), and density ( $\rho$ ) follow a nondimensional ideal gas law of  $P = \rho T$ . We evolving the temperature, log density ( $\ln \rho$ ), and velocity vector ( $\mathbf{u} = u\hat{x} + v\hat{y} + z\hat{z}$ ) according to the fully compressible Navier-Stokes equations,

$$\frac{D \ln \rho}{Dt} = -\nabla \cdot (\mathbf{u}) \quad (1)$$

$$\frac{D \mathbf{u}}{Dt} = -\nabla T - T \nabla \ln \rho + \mathbf{g} + \nabla \cdot (\bar{\bar{\Pi}}), \quad (2)$$

$$\frac{DT}{Dt} + (\gamma - 1)T \nabla \cdot (\mathbf{u}) + \frac{1}{\rho c_V} \nabla \cdot (-\rho \chi \nabla T) = \quad (3)$$

$$\frac{1}{\rho c_V} (\bar{\bar{\Pi}} \cdot \nabla) \cdot \mathbf{u},$$

where  $\chi$  is the thermal diffusivity and with the viscous stress tensor given by

$$\Pi_{ij} \equiv \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} - \frac{2}{3} \delta_{ij} \nabla \cdot \mathbf{u} \right), \quad (4)$$

where  $\delta_{ij}$  is the Kronecker delta and with  $\mu = \rho \nu$  is the dynamic viscosity and  $\nu$  is the kinematic viscosity. We split our thermodynamics into a time-stationary adiabatic background component and a time-evolving fluctuation away from that background as  $T = T_0(z) + T_1(x, y, z, t)$  and  $\ln \rho = \ln \rho_0(z) + \ln \rho_1(x, y, z, t)$ . We

specify the background atmosphere as an adiabatic polytrope, as in ?, so that

$$T_0(z) = (1 + L_z - z), \quad \ln \rho_0(z) = m_{ad} \ln(T_0), \quad (5)$$

where  $L_z$  is the domain depth, and  $m_{ad} = 1/(\gamma - 1) = 1.5$  is the adiabatic polytropic index. In the construction of these atmospheres, we assume that gravity  $\mathbf{g} = -g\hat{z}$  is everywhere constant and that the background atmosphere is in hydrostatic equilibrium such that  $\partial_z T_0 + T_0 \partial_z \ln \rho_0 - g = 0$ .

We assume a Prandtl number of unity,  $\text{Pr} = \nu/\chi = 1$ , everywhere in the domain. We furthermore assume that  $\nu$  and  $\chi$  are uniform everywhere in the domain, and that a negligibly small diffusivity,  $\chi_B = 0$  acts on the pure background terms in the thermal flux such that

$$\frac{1}{\rho_0 c_V} \nabla \cdot (-\rho \chi_B \nabla T) = 0.$$

Under all of these assumptions, the equations we solve become

$$\frac{D \ln \rho_1}{Dt} = -\nabla \cdot (\mathbf{u}) - w \partial_z \ln \rho_0 \quad (6)$$

$$\frac{D \mathbf{u}}{Dt} = -\nabla T_1 - T_0 \nabla \ln \rho_1 - T_1 \partial_z \rho_0 \hat{z} - T_1 \nabla \ln \rho_1 + \nabla \cdot (\bar{\bar{\Pi}}), \quad (7)$$

$$\frac{DT_1}{Dt} + w \partial_z T_0 + (\gamma - 1)(T_0 + T_1) \nabla \cdot (\mathbf{u}) \quad (8)$$

$$- \frac{1}{c_V} \left\{ \chi (\nabla^2 T_1 + \partial_z \ln \rho_0 \partial_z T_1 + \partial_z \ln \rho_1 \partial_z T_0 + \nabla \ln \rho_1 \cdot \nabla T_1) \right\} =$$

$$\frac{1}{\rho c_V} (\bar{\bar{\Pi}} \cdot \nabla) \cdot \mathbf{u},$$

### 3. RESULTS

#### 4. DISCUSSION

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#### APPENDIX

##### A. TABLE OF SIMULATIONS