

## Part 3 Essay Excerpt

EXCERPT BEGINS ... We now take a closer look at the possible cloud support mechanisms already alluded to. Here we assume a globally bound molecular cloud and derive a more complete version of the Virial Theorem. The theory of un-bound clouds is discussed in the following Chapter.

The generalised hydrostatic equation for a fluid which has an ambient magnetic field permeating it is:

$$\rho \frac{Du}{Dt} = -\vec{\nabla} P - \rho \vec{\nabla} \Phi_g + (\vec{J} \times \vec{B}) \quad (1)$$

where  $Du/Dt$  is the ‘convective’/‘Lagrangian’ time-derivative, defined by:

$$\frac{D}{Dt} = \frac{\partial}{\partial t} + (\vec{u} \cdot \vec{\nabla}) \quad (2)$$

Of course from Maxwell’s equations we know that:

$$\vec{\nabla} \times \vec{B} = \mu_0 \left( \vec{J} + \epsilon_0 \frac{\partial \vec{E}}{\partial t} \right) \quad (3)$$

For hereon we assume that the electric field changes slowly and we ignore this term in the equation so that we can re-cast Equation (1) as:

$$\rho \frac{Du}{Dt} = -\vec{\nabla} P - \rho \vec{\nabla} \Phi_g + \frac{1}{\mu_0} (\vec{\nabla} \times \vec{B} \times \vec{B}) \quad (4)$$

which expands to:

$$\rho \frac{Du}{Dt} = -\vec{\nabla} P - \rho \vec{\nabla} \Phi_g + \frac{1}{\mu_0} (\vec{B} \cdot \vec{\nabla}) \vec{B} + \frac{1}{2\mu_0} (\vec{\nabla} |\vec{B}|^2) \quad (5)$$

We can simply see what each term in this expression corresponds to: terms 1 and 2 on the RHS are the normal thermal pressure and gravity terms; term 4 is a magnetic pressure term; while term 3 is an extra tension due to curved magnetic field lines. To push forward to derive the Virial Theorem we now take the scalar product of Equation (5) with the position vector  $\vec{r}$  and integrate this over the volume,  $V$ :

$$\int_V \rho \frac{\partial u}{\partial t} \cdot \vec{r} dV + \int_V \rho (\vec{u} \cdot \vec{\nabla}) \vec{u} \cdot \vec{r} dV = - \int_V \vec{\nabla} P \cdot \vec{r} dV - \int_V \rho \vec{\nabla} \Phi_g \cdot \vec{r} dV + \frac{1}{\mu_0} \int_V (\vec{B} \cdot \vec{\nabla}) \vec{B} \cdot \vec{r} dV + \frac{1}{2\mu_0} \int_V (\vec{\nabla} |\vec{B}|^2) \cdot \vec{r} dV \quad (6)$$

This integration is normally performed in components. Thus the LHS of Equation (6) becomes:

$$\int_V \rho x_i \frac{\partial u_i}{\partial t} dV + \int_V \rho x_i u_j \partial_j u_i dV \quad (7)$$

Now by noting the continuity equation which (in components) is  $\partial_j (\rho u_j) = -\frac{\partial \rho}{\partial t}$  and also that  $\rho dV = dm$  the LHS becomes:

$$LHS = \frac{1}{2} \ddot{I} - 2T \quad (8)$$

where the moment of inertia,  $I$ , and the kinetic energy,  $T$ , are defined by the usual expressions:

$$I = \int_M x_i x_j dm \quad (9)$$

$$T = \frac{1}{2} \int_M u_i u_j dm \quad (10)$$

The first term on the RHS can be worked out by integration by parts

$$- \int x_i \partial_i P dV = - \int \partial_i (x_i P) dV + \int (\partial_i x_i) P dV \quad (11)$$

and adopting  $P = (5/3 - 1)u$  we get:

$$\Rightarrow - \int x_i \partial_i P dV = - \int P \vec{r} \cdot d\vec{S} + 2U \quad (12)$$

The second RHS term is:

$$- \int_V \rho \vec{\nabla} \Phi_g \cdot \vec{r} dV = - \int \vec{r} \cdot \vec{g} dm \quad (13)$$

But we can write this as:

$$\int \vec{r} \cdot \vec{g} dm = - \int \int \frac{G(\vec{r} - \vec{r}') \cdot \vec{r}}{|\vec{r} - \vec{r}'|^3} dmdm' = - \int \int \frac{G(\vec{r}' - \vec{r}) \cdot \vec{r}'}{|\vec{r} - \vec{r}'|^3} dmdm' \quad (14)$$

Adding these two ways together and dividing by two we get:

$$\int \vec{r} \cdot \vec{g} dm = - \frac{1}{2} \int \int G \frac{dmdm'}{|\vec{r} - \vec{r}'|} = \Omega \quad (15)$$

We next evaluate the third and fourth RHS terms noting the definition that:

$$M_B = \frac{1}{2\mu_0} \int_V |\vec{B}|^2 dV \quad (16)$$

The third term is given by:

$$\frac{1}{\mu_0} \int x_i B_j \partial_j B_i dV = \frac{1}{\mu_0} \int \partial_j (x_i B_j B_i) dV - \frac{1}{\mu_0} \int B_i \partial_j (x_i B_j) dV \quad (17)$$

$$\Rightarrow \frac{1}{\mu_0} \int x_i B_j \partial_j B_i dV = \frac{1}{\mu_0} \int \partial_j (x_i B_j B_i) dV - \frac{1}{\mu_0} \int B_i B_i dV \quad (18)$$

$$\Rightarrow \frac{1}{\mu_0} \int (\vec{r} \cdot \vec{B}) \vec{B} \cdot \vec{n} dV = \frac{1}{\mu_0} \int \partial_j (x_i B_j B_i) dV - 2M_B \quad (19)$$

The fourth term is:

$$\frac{1}{2\mu_0} \int x_i \partial_i (B_j B_j) dV = \frac{1}{2\mu_0} \int \partial_i (x_i B_j B_j) dV + \frac{1}{2\mu_0} \int (\partial_i x_i) B_j B_j dV \quad (20)$$

$$\Rightarrow \frac{1}{2\mu_0} \int x_i \partial_i (B_j B_j) dV = \frac{1}{2\mu_0} \int B^2 \vec{r} \cdot \vec{n} dV + 3M_B \quad (21)$$

Combining all our results we can now state the Virial Theorem.

$$\frac{1}{2} \ddot{I} = 2T + 2U + \Omega + M_B - \int_V \left( P + \frac{1}{\mu_0} B^2 \right) \vec{r} \cdot \vec{n} dV + \frac{1}{\mu_0} \int_V (\vec{r} \cdot \vec{B}) \vec{B} \cdot \vec{n} dV \quad (22)$$

It is usual also to ignore the two last surface integrals on the RHS as these involve effects of external pressure. We can safely ignore these terms for strongly self-gravitating objects. This is definitely justified for dense cores but may not be ok to do on a more global scale across the entire cloud complex. Furthermore if the clouds are stable (i.e. in a force balance) for large times we can ignore the LHS also. With these terms ignored we have our working Virial Theorem:

$$\boxed{2T + 2U + \Omega + M_B = 0} \tag{23}$$

i.e. for stability  $\Omega$  must be balanced by the three other terms. We now investigate the relative importance of these three support mechanisms ... EXCERPT ENDS