

ASP3262 Introduction to coding

Week 11 Lab

Numerical Solutions to Nuclear Reaction Networks

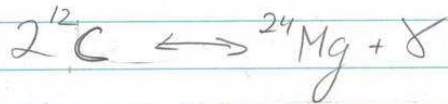
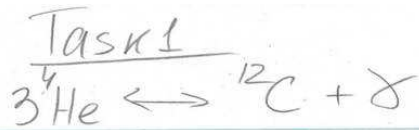
Written by Evgenii Neumerzhitskii

Oct 17, 2019

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Task 1



Forward reaction rates

$$\begin{aligned} \frac{d}{dt} Y_{^4\text{He}} &= \left(\lambda_{3^4\text{He} \rightarrow {}^{12}\text{C}} \right) \frac{(0-3)}{(3!0!)} Y_{^4\text{He}}^3 Y_{^{12}\text{C}}^0 \\ &= \left(\lambda_{3^4\text{He} \rightarrow {}^{12}\text{C}} \right) \left(-\frac{1}{2} \right) \left(Y_{^4\text{He}}^3 \right) \end{aligned}$$

$$\begin{aligned} \frac{d}{dt} Y_{^{12}\text{C}} &= \left(\lambda_{3^4\text{He} \rightarrow {}^{12}\text{C}} \right) \frac{(1-0)}{(3!0!)} Y_{^4\text{He}}^3 Y_{^{12}\text{C}}^0 \\ &\quad + \left(\lambda_{2^{12}\text{C} \rightarrow {}^{24}\text{Mg}} \right) \frac{(0-2)}{(2!0!)} Y_{^{12}\text{C}}^2 Y_{^{24}\text{Mg}}^0 \\ &= \left(\lambda_{3^4\text{He} \rightarrow {}^{12}\text{C}} \right) \left(\frac{1}{6} \right) Y_{^4\text{He}}^3 + \left(\lambda_{2^{12}\text{C} \rightarrow {}^{24}\text{Mg}} \right) \left(-1 \right) Y_{^{12}\text{C}}^2 \end{aligned}$$

$$\begin{aligned} \frac{d}{dt} Y_{^{24}\text{Mg}} &= \lambda_{2^{12}\text{C} \rightarrow {}^{24}\text{Mg}} \frac{(1-0)}{(2!0!)} Y_{^{12}\text{C}}^2 Y_{^{24}\text{Mg}}^0 \\ &= \left(\lambda_{2^{12}\text{C} \rightarrow {}^{24}\text{Mg}} \right) \left(\frac{1}{2} \right) Y_{^{12}\text{C}}^2 \end{aligned}$$

Reverse reaction rates



$$\begin{aligned}\frac{d}{dt} Y_{^3\text{He}} &= (\lambda_{^{12}\text{C} \rightarrow 3\,^3\text{He}}) \left(\frac{3-0}{1! 0!} \right) Y_{^{12}\text{C}}^1 Y_{^3\text{He}}^0 \\ &= (\lambda_{^{12}\text{C} \rightarrow 3\,^3\text{He}}) (3) Y_{^{12}\text{C}}^1\end{aligned}$$

$$\begin{aligned}\frac{d}{dt} Y_{^{12}\text{C}} &= (\lambda_{^{12}\text{C} \rightarrow 3\,^3\text{He}}) \left(\frac{0-1}{1! 0!} \right) Y_{^{12}\text{C}}^1 Y_{^3\text{He}}^0 \\ &\quad + (\lambda_{^{24}\text{Mg} \rightarrow 2\,^{12}\text{C}}) \left(\frac{2-0}{1! 0!} \right) Y_{^{24}\text{Mg}}^1 Y_{^{12}\text{C}}^0 \\ &= -(\lambda_{^{12}\text{C} \rightarrow 3\,^3\text{He}}) (1) Y_{^{12}\text{C}}^1 + (\lambda_{^{24}\text{Mg} \rightarrow 2\,^{12}\text{C}}) (2) Y_{^{24}\text{Mg}}^1\end{aligned}$$

$$\begin{aligned}\frac{d}{dt} Y_{^{24}\text{Mg}} &= (\lambda_{^{24}\text{Mg} \rightarrow 2\,^{12}\text{C}}) \left(\frac{0-1}{1} \right) Y_{^{24}\text{Mg}}^1 \\ &= -(\lambda_{^{24}\text{Mg} \rightarrow 2\,^{12}\text{C}}) (Y_{^{24}\text{Mg}}^1)\end{aligned}$$

Task 2

The rates of the nuclear reactions are shown on Figures 1 - 4. We can see that the rates for reverse reactions are identical for the two different pressures. In addition, we can see that reverse reactions only happen at temperatures higher than about 1×10^9 K, while forward reactions happen at lower temperatures.

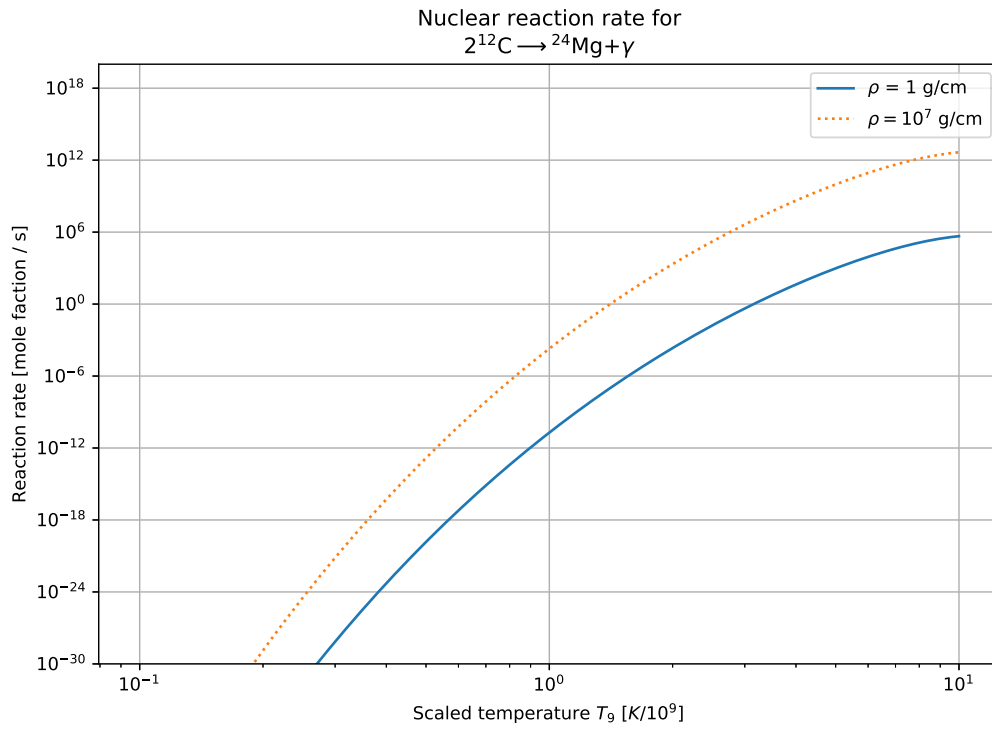


Figure 1: Nuclear reaction rates for $2^{12}\text{C} \rightarrow {}^{24}\text{Mg}$.

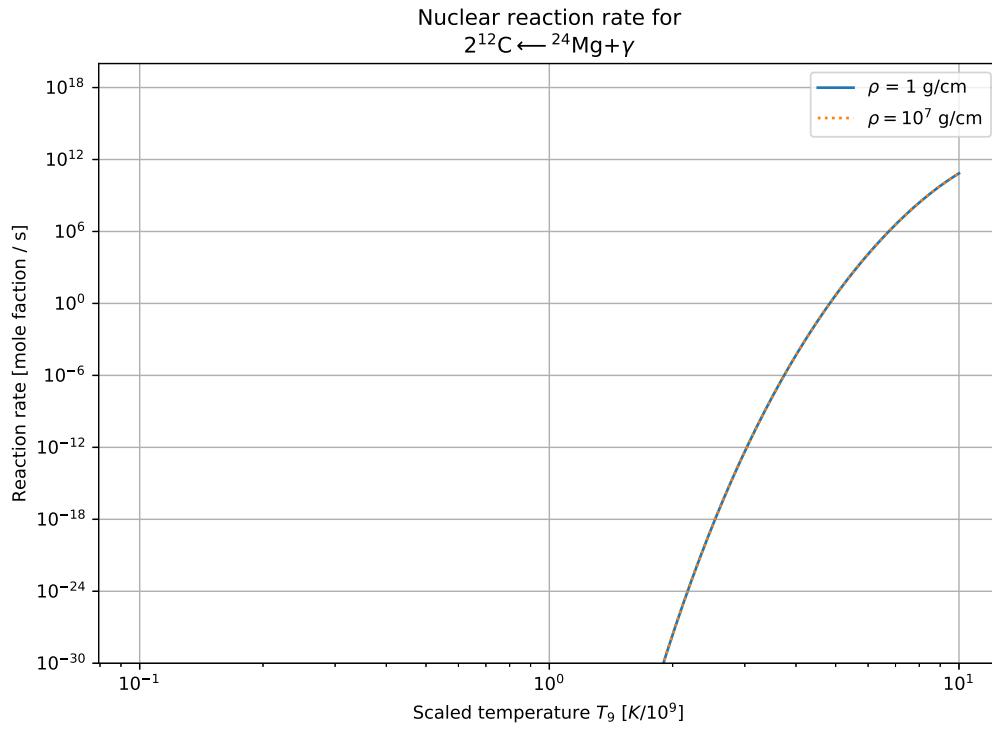


Figure 2: Nuclear reaction rates for $2^{12}\text{C} \leftarrow {}^{24}\text{Mg}$.

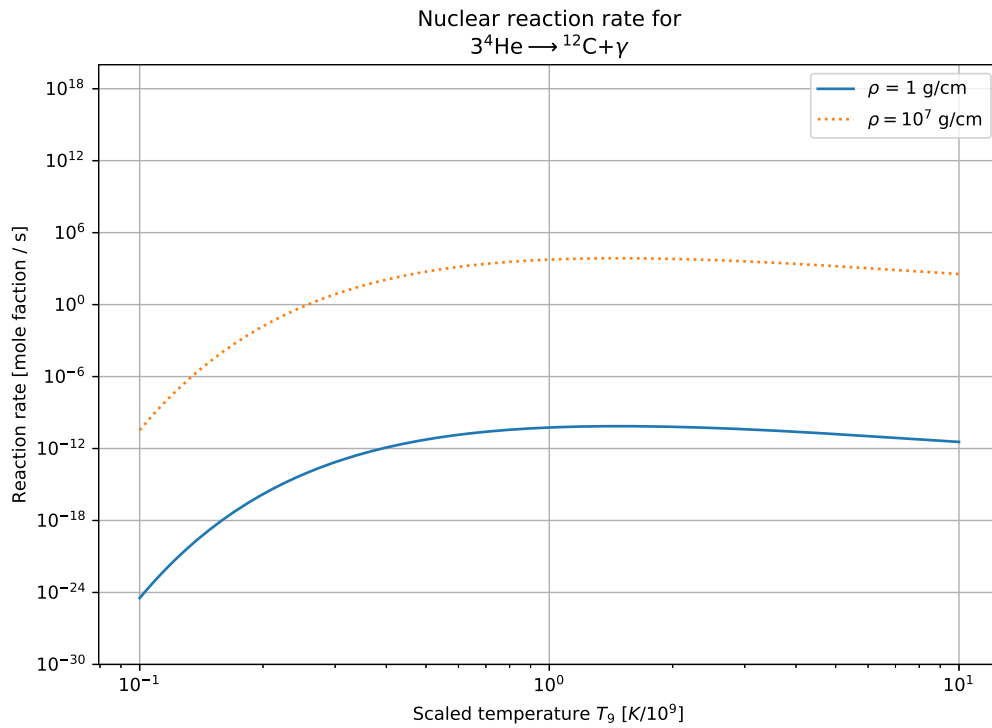


Figure 3: Nuclear reaction rates for $3^4\text{He} \rightarrow {}^{12}\text{C}$.

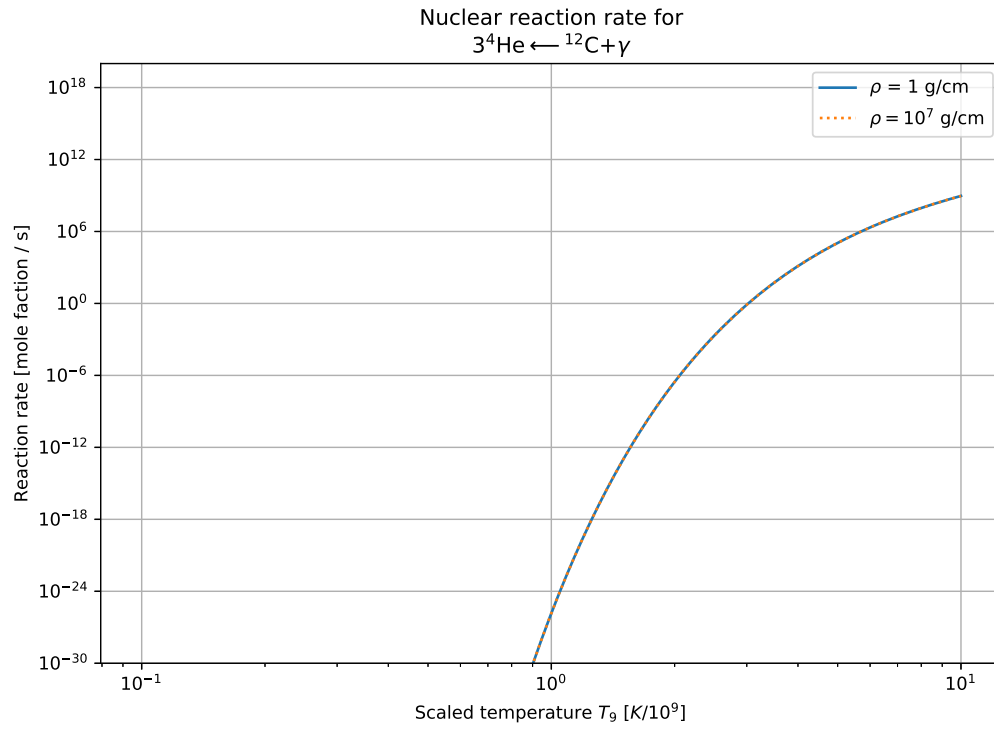


Figure 4: Nuclear reaction rates for ${}^3\text{He} \leftarrow {}^{12}\text{C}$.

Task 3

The changes of mass fraction over time for a range of temperatures are shown on Figures 5 - 14.

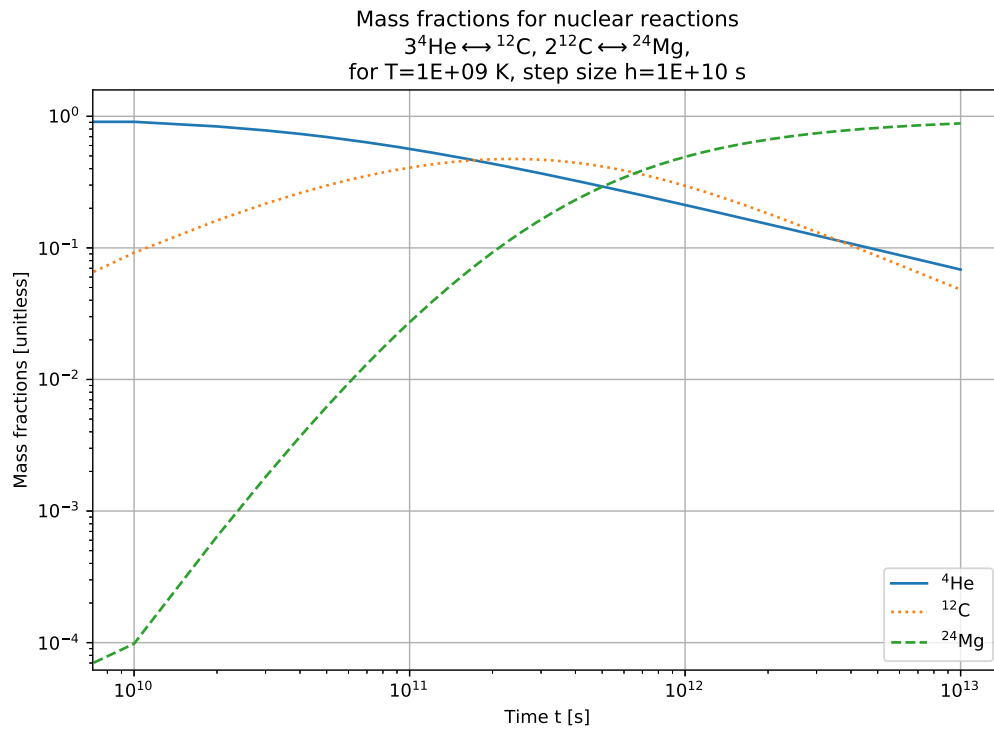


Figure 5: Changes of mass fraction over time for temperature $T = 1 \times 10^9$ K.

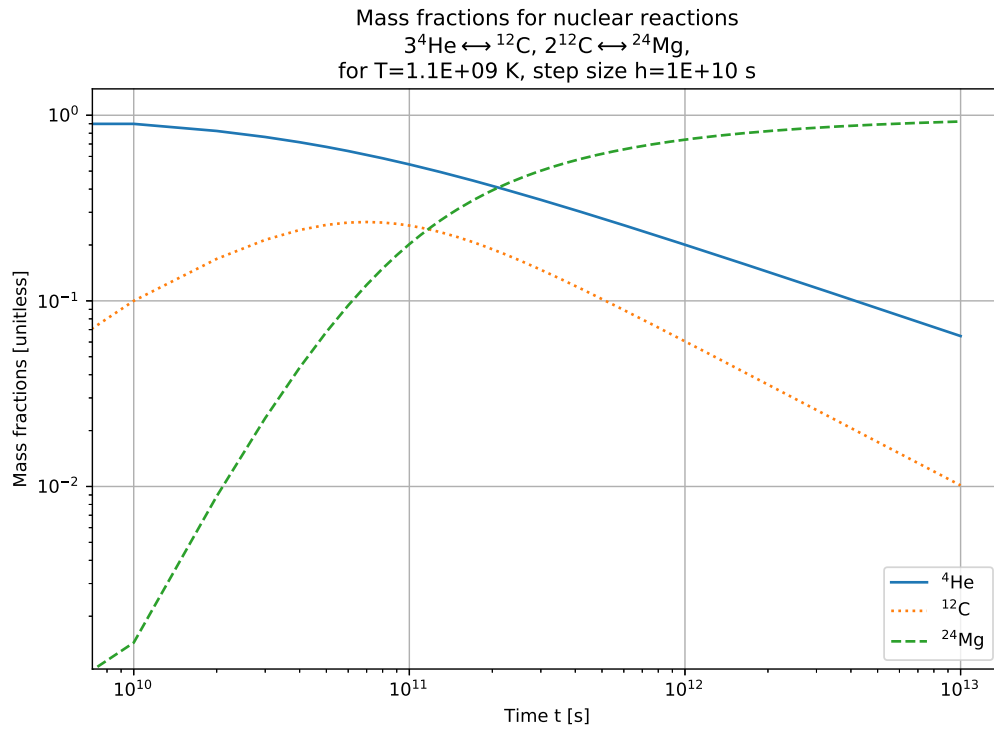


Figure 6: Changes of mass fraction over time for temperature $T = 1.1 \times 10^9$ K.

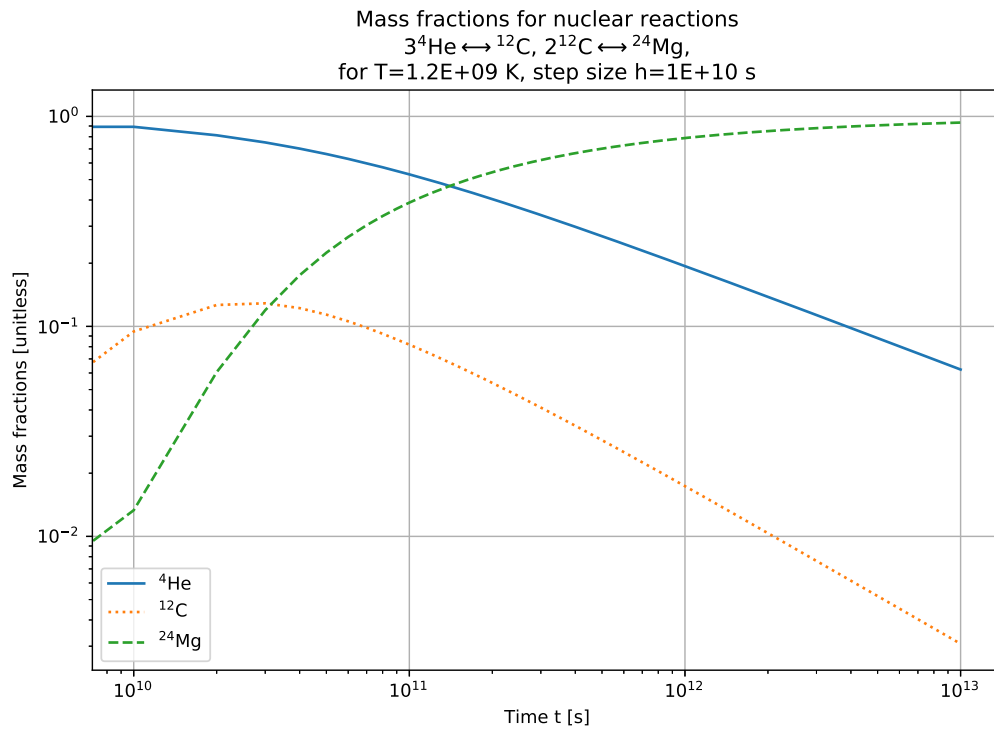


Figure 7: Changes of mass fraction over time for temperature $T = 1.2 \times 10^9$ K.

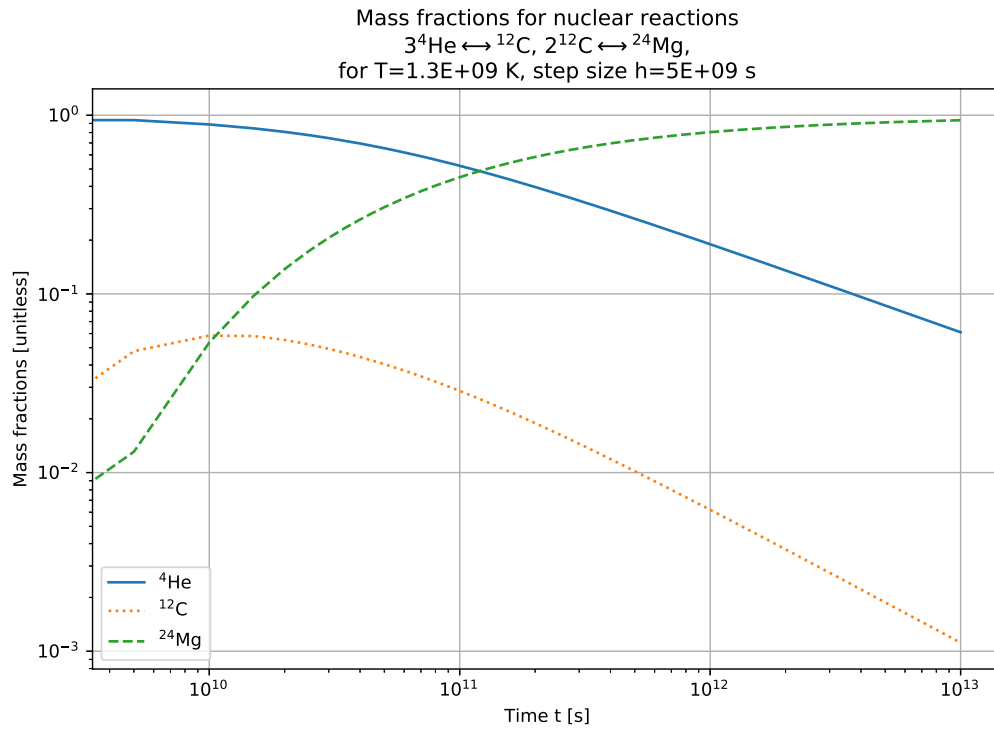


Figure 8: Changes of mass fraction over time for temperature $T = 1.3 \times 10^9$ K.

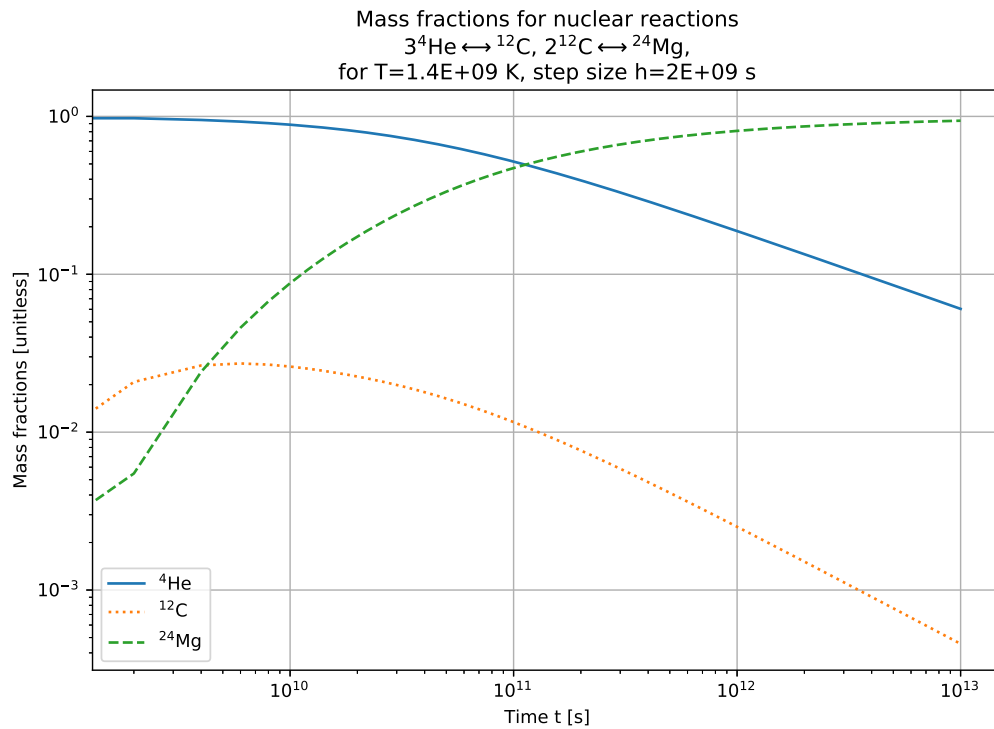


Figure 9: Changes of mass fraction over time for temperature $T = 1.4 \times 10^9$ K.

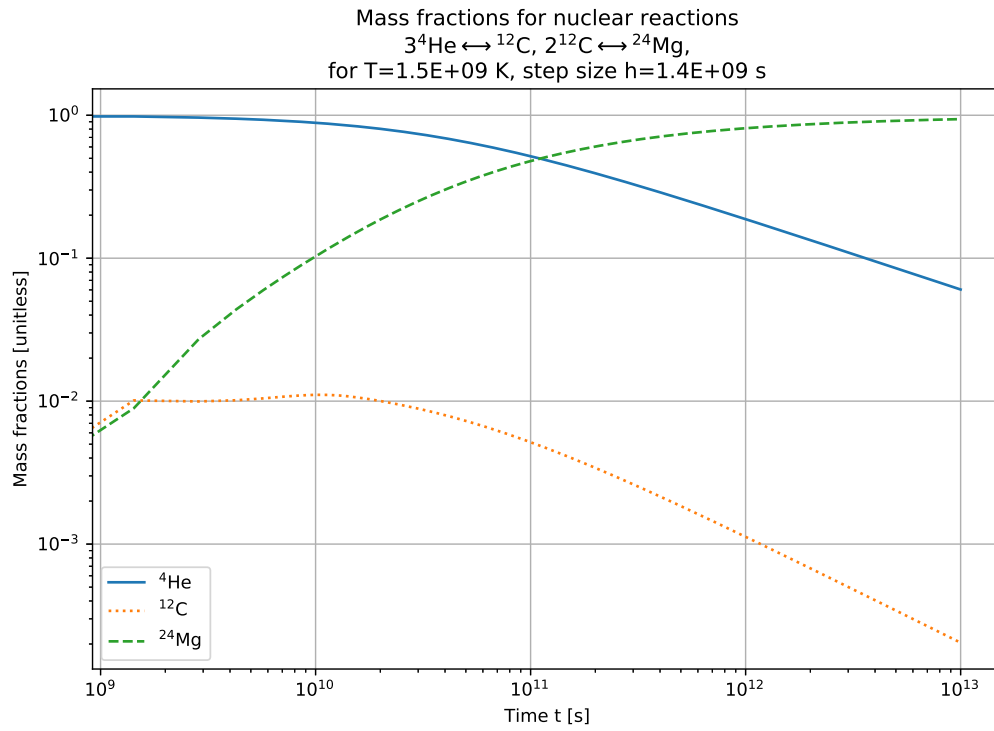


Figure 10: Changes of mass fraction over time for temperature $T = 1.5 \times 10^9$ K.

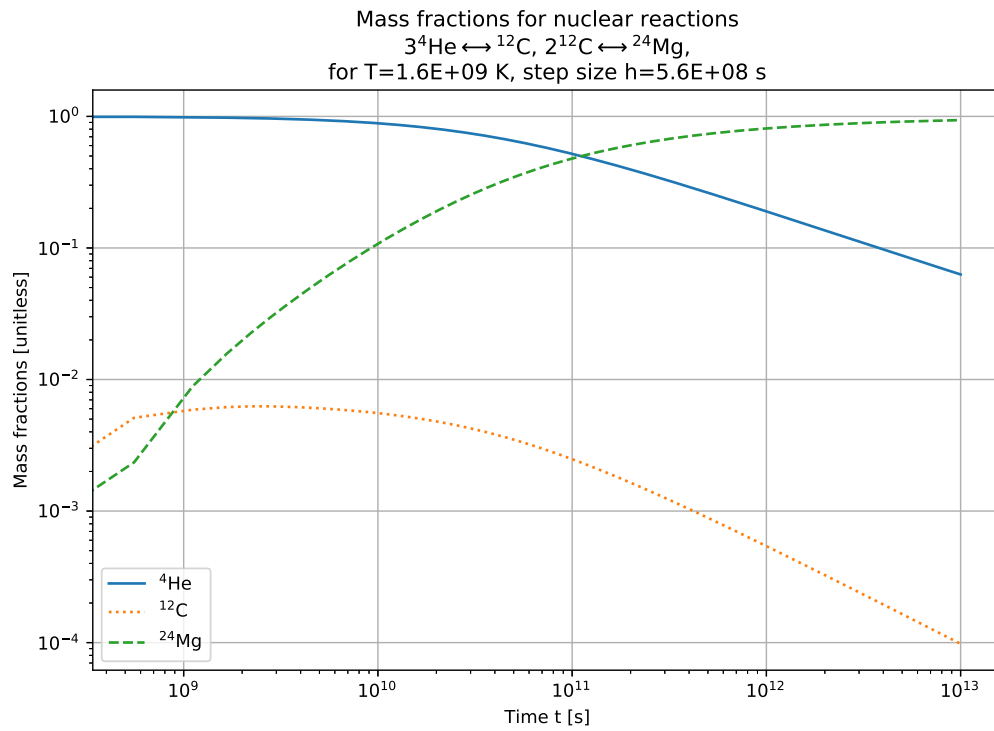


Figure 11: Changes of mass fraction over time for temperature $T = 1.6 \times 10^9$ K.

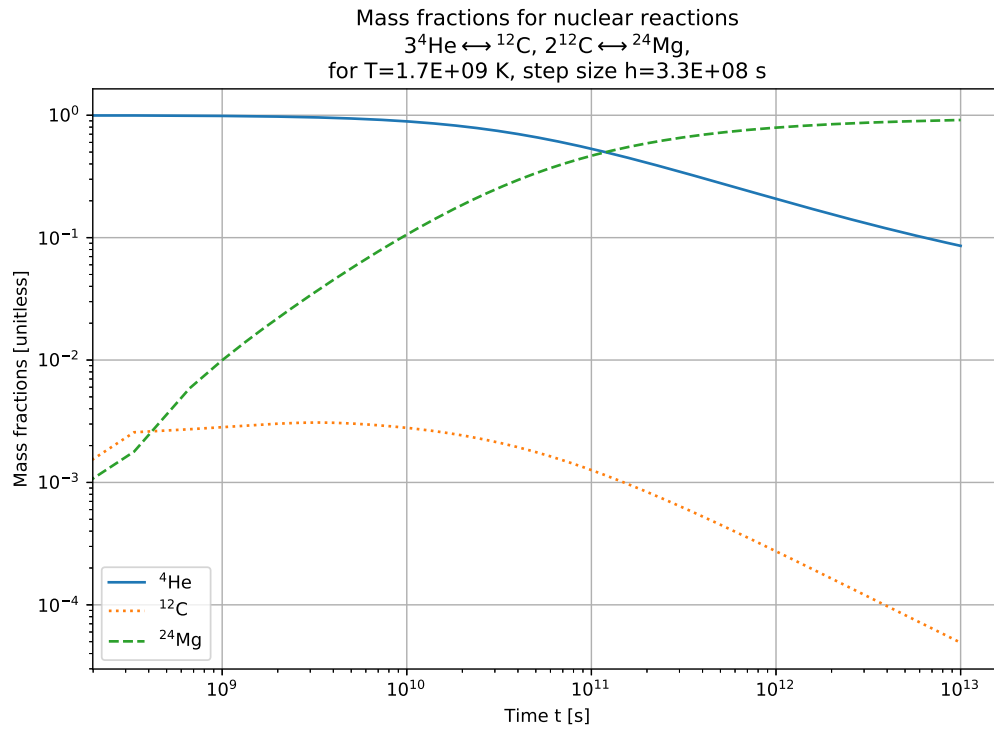


Figure 12: Changes of mass fraction over time for temperature $T = 1.7 \times 10^9$ K.

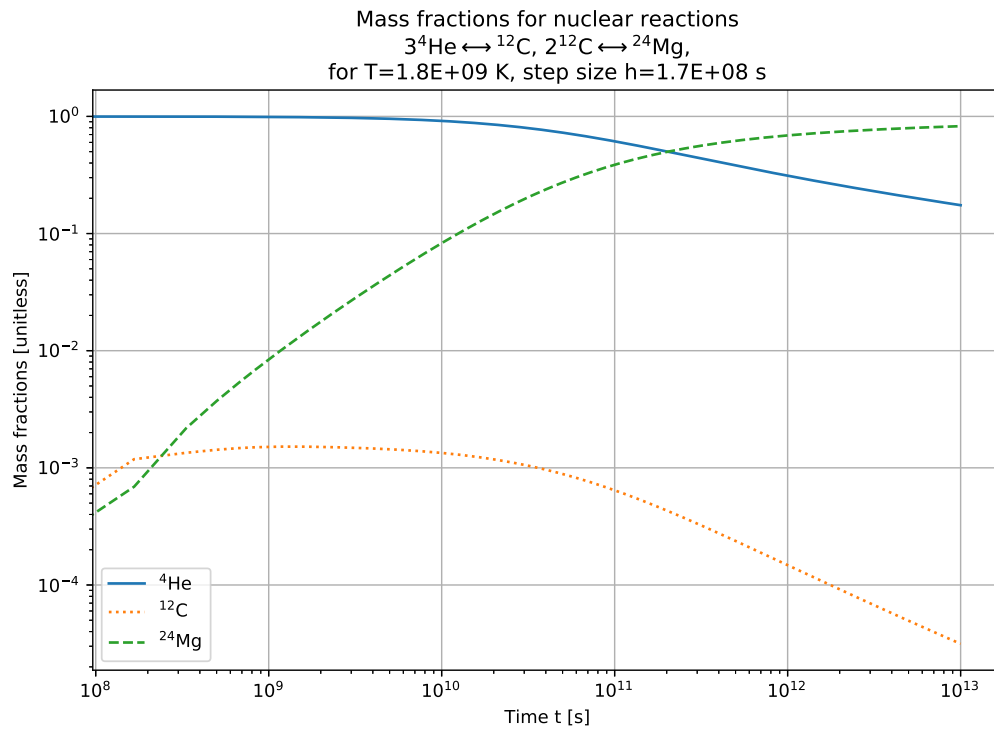


Figure 13: Changes of mass fraction over time for temperature $T = 1.8 \times 10^9$ K.

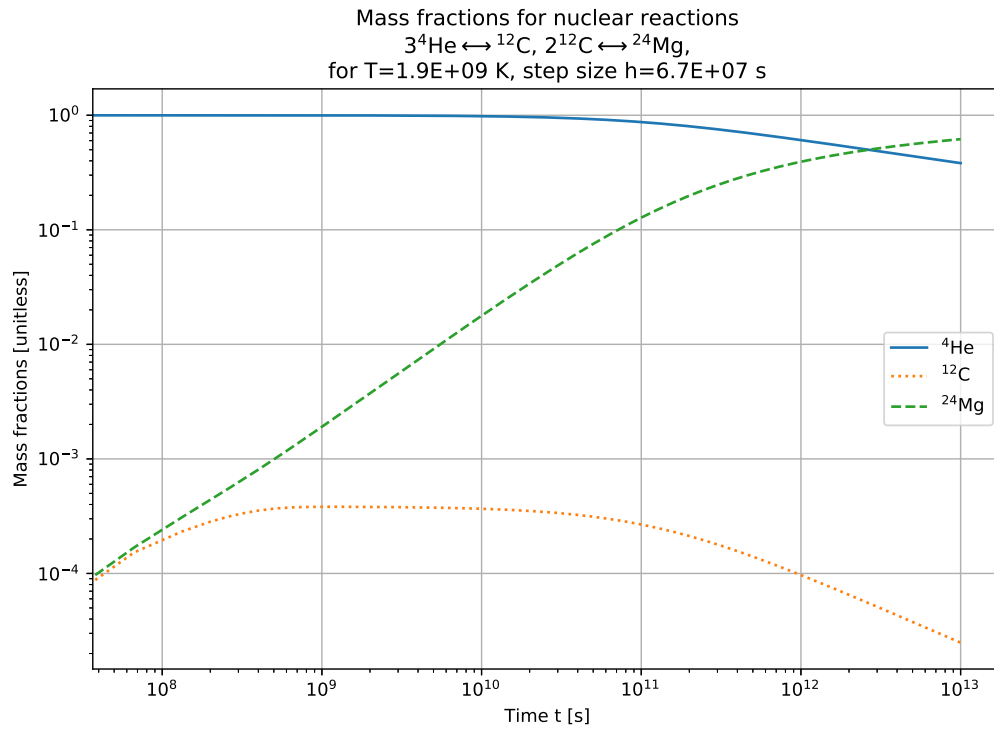


Figure 14: Changes of mass fraction over time for temperature $T = 1.9 \times 10^9$ K.

Task 4

Here we use an integrator with adaptive time step. The changes of mass fraction over time are shown on [Figure 15](#).

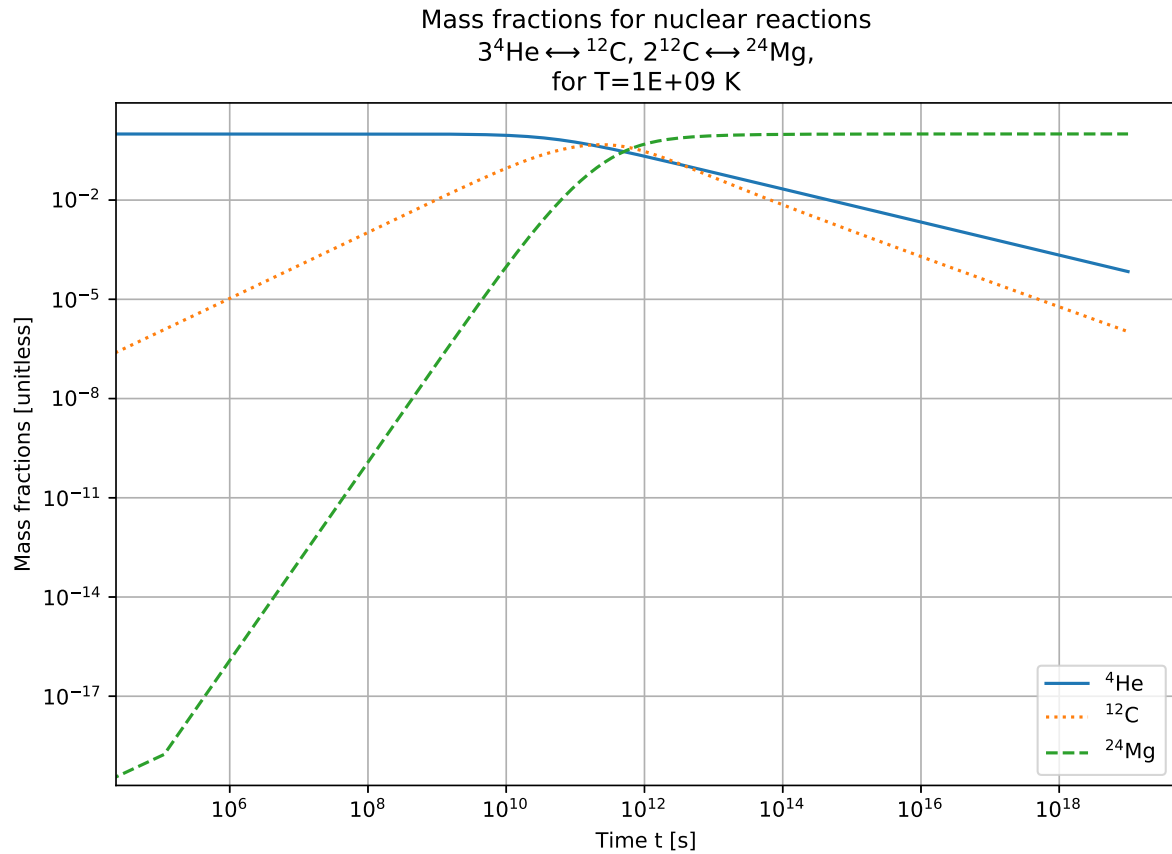


Figure 15: Changes of mass fraction over time for temperature $T = 1 \times 10^9$ K.

Task 5

Solutions obtained with an implicit solver for higher temperatures are shown on [Figure 16](#) and [Figure 17](#).

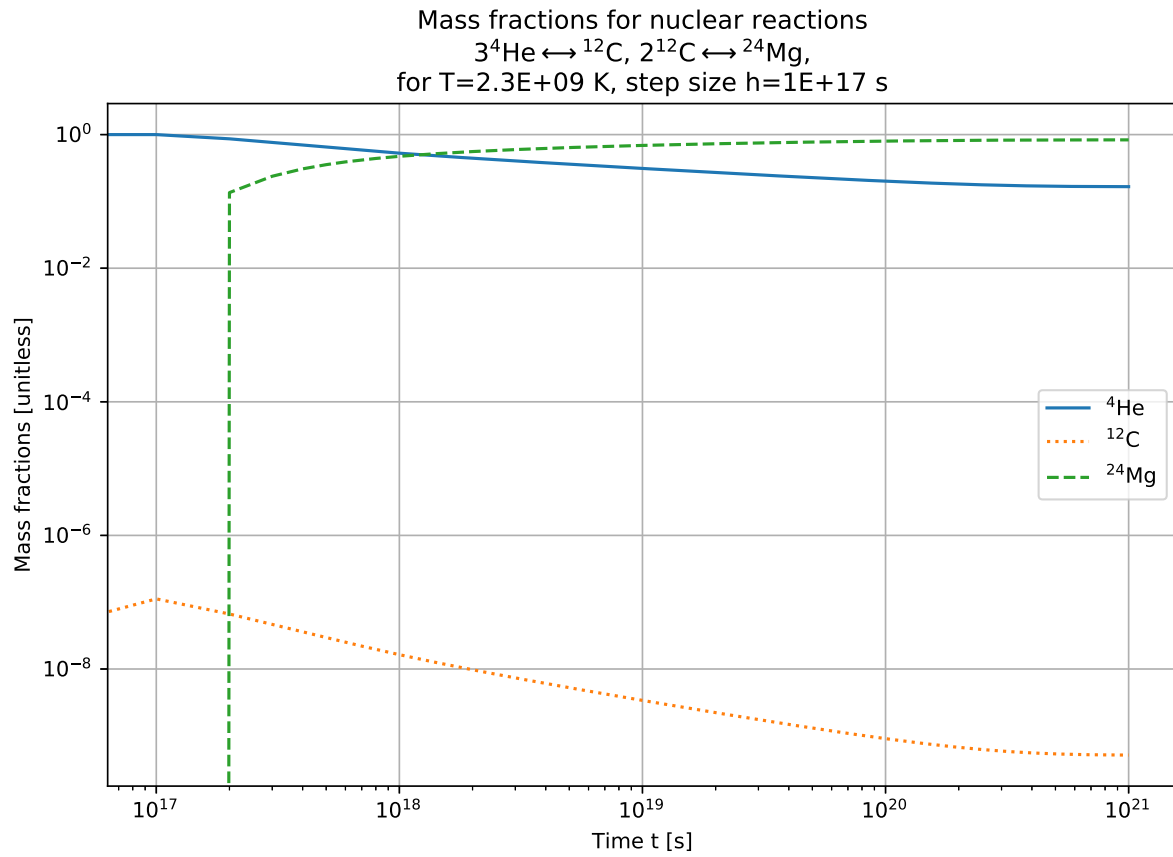


Figure 16: Changes of mass fraction over time for temperature $T = 2.3 \times 10^9$ K.

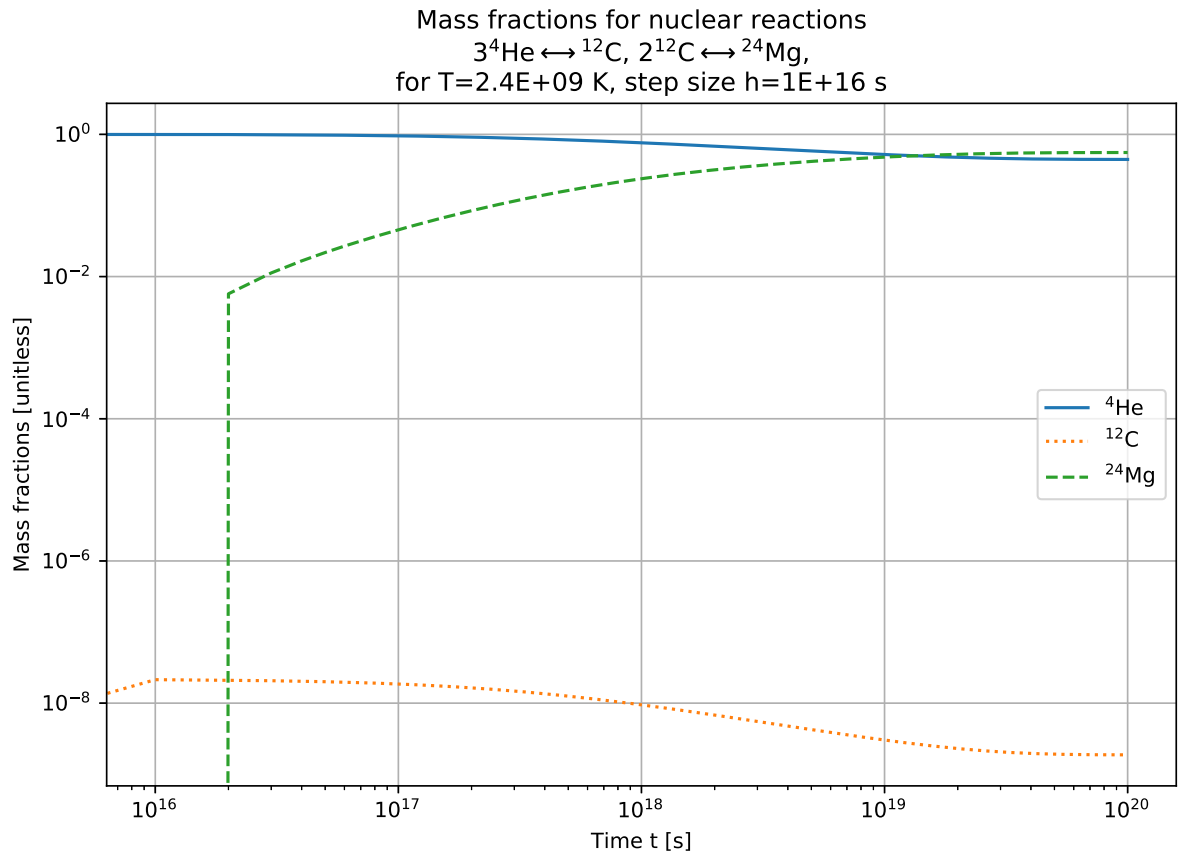


Figure 17: Changes of mass fraction over time for temperature $T = 2.4 \times 10^9$ K.

Task 6

Solutions obtained with an adaptive implicit solver are shown on [Figure 18](#). The final mass fractions for Helium, Carbon and Magnesium are 0.0156, 0 and 0.984 respectively.

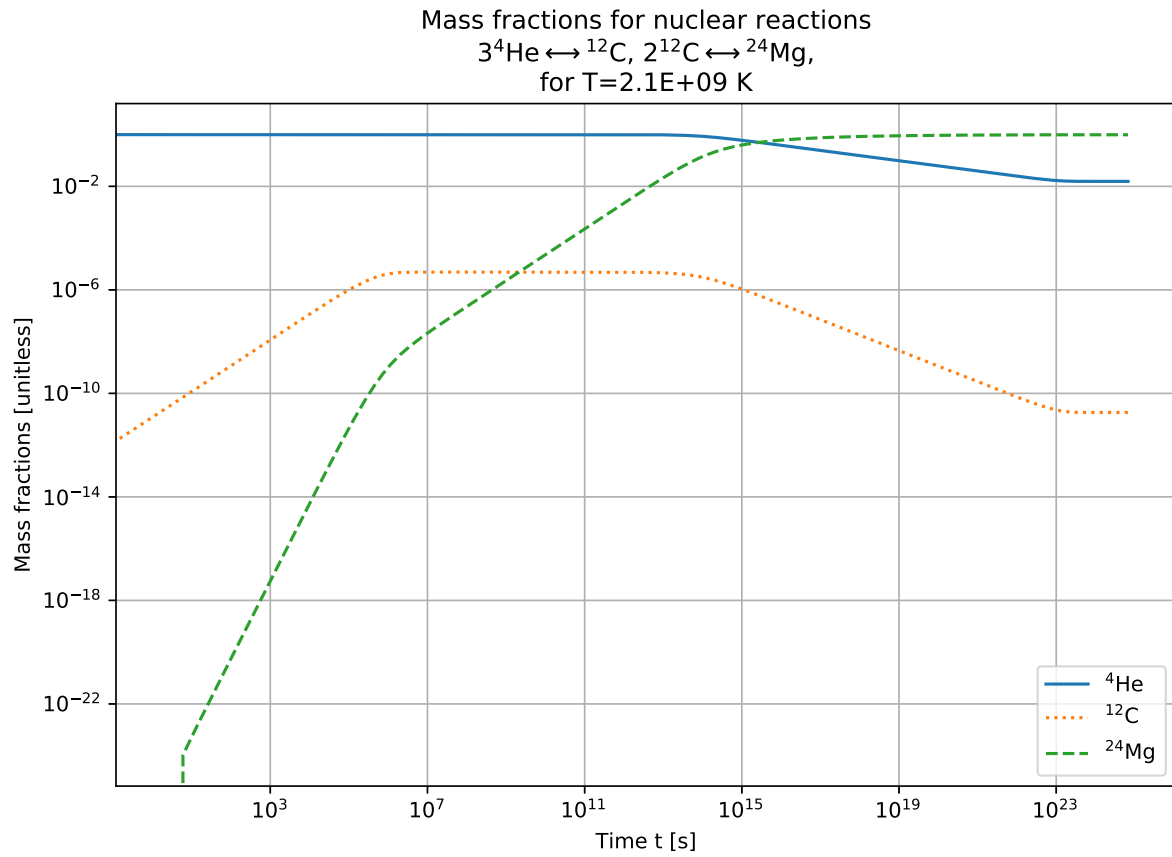


Figure 18: Changes of mass fraction over time for temperature $T = 2.1 \times 10^9$ K calculated with adaptive implicit solver.

Task 7

Solutions obtained using variable temperatures are shown on [Figure 19](#) and [Figure 20](#). For the last solution we had to decrease the temperature to 7×10^9 K in order for the plot to show up.

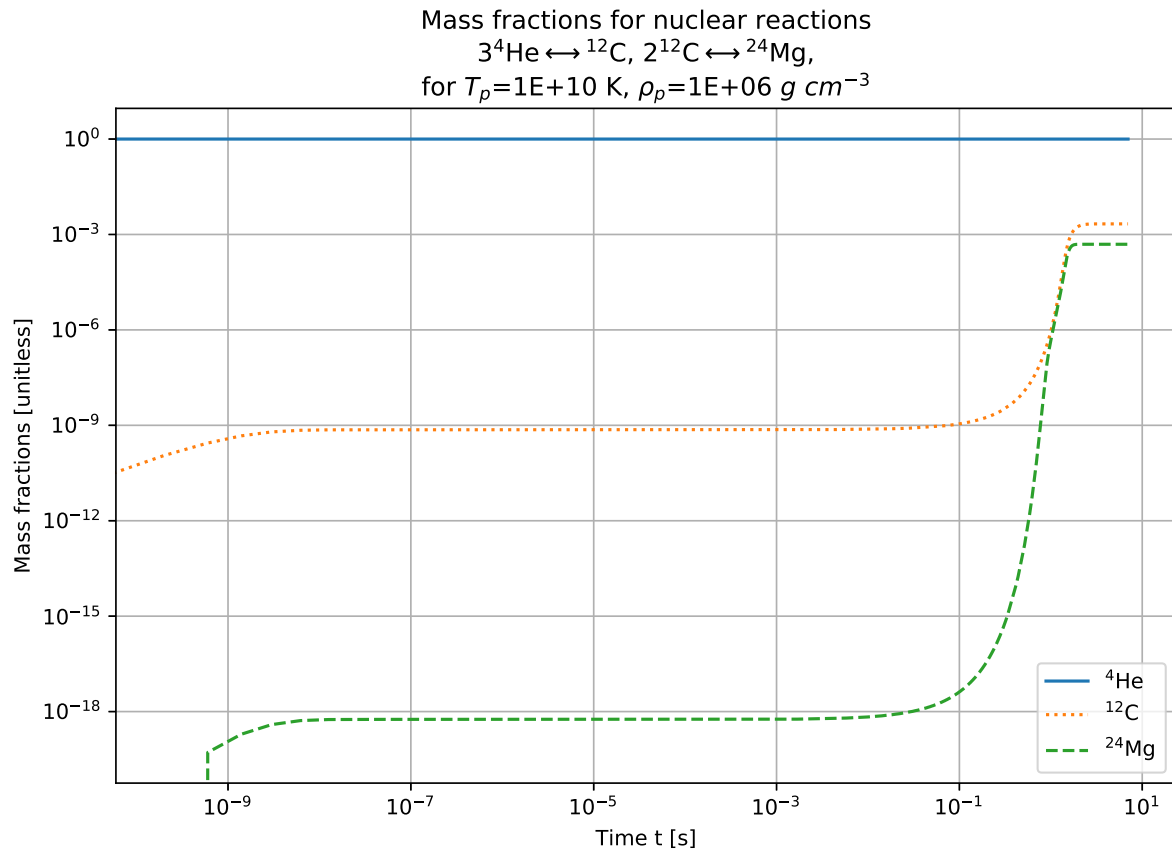


Figure 19: Changes of mass fraction over time for variable temperature.

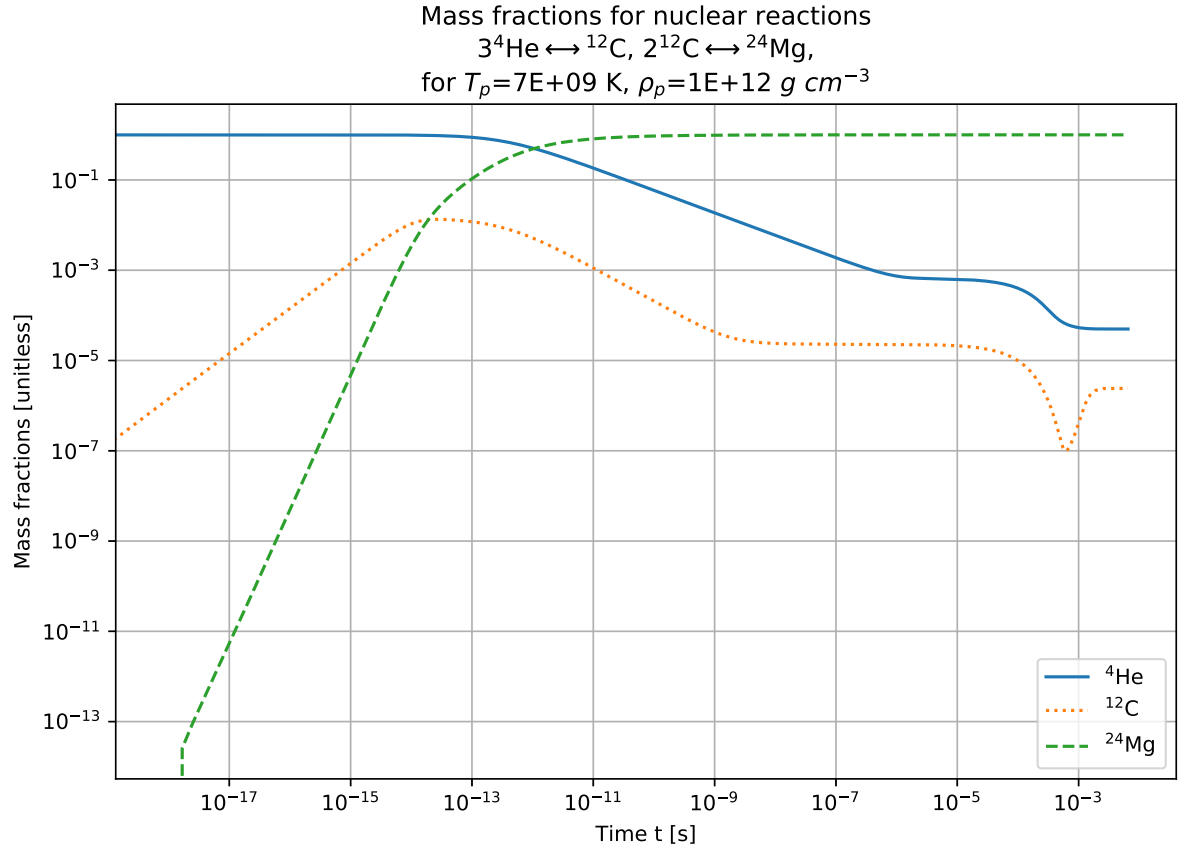


Figure 20: Changes of mass fraction over time for variable temperature.