C11 Images - SBIG ST-7xme camera

Bias: 13 March 2018, 22 March 2018

Dark: 14-20 March 2018
B Flats: 13 March 2018
V Flats: 13 March 2018
R Flats: 13 March 2018
I Flats: 13 March 2018

NGC 4755 (open cluster): 07 March 2018, 18 March 2018, 29 March 2018

Wasp 87b (exoplanet): 08 March 2018, 29 March 2018

NGC 3201 (globular cluster): 09 March 2018, 29 March 2018, 30 April 2018

NGC 2301 (open cluster): 13 March 2018

NGC 2362 (open cluster): 21 March 2018, 22 March 2018, 29 March 2018

The photometric calibration from 29 March 2018 for the C11 is:

B = $19.237 - 2.5\log(f/t) - 0.330*A + 0.143*(B-V)$, RMS= 0.025, n=33, r=12 pix V = $19.696 - 2.5\log(f/t) - 0.210*A - 0.056*(B-V)$, RMS= 0.019, n=33, r=12 pix R = $19.679 - 2.5\log(f/t) - 0.119*A - 0.072*(V-R)$, RMS= 0.011, n=33, r=12 pix I = $19.079 - 2.5\log(f/t) - 0.134*A + 0.000*(R-I)$, RMS= 0.013, n=33, r=12 pix

The photometric calibration from 30 April 2018 for the C11 is:

 $B = 19.213 - 2.5log(f/t) - 0.286*A + 0.156*(B-V), RMS= 0.015, n=18, r=12 pix \\ V = 19.700 - 2.5log(f/t) - 0.173*A - 0.062*(B-V), RMS= 0.007, n=18, r=12 pix \\ R = 19.715 - 2.5log(f/t) - 0.102*A - 0.082*(V-R), RMS= 0.006, n=18, r=12 pix \\ I = 19.099 - 2.5log(f/t) - 0.100*A + 0.009*(R-I), RMS= 0.009, n=18, r=12 pix \\ I = 19.099 - 2.5log(f/t) - 0.100*A + 0.009*(R-I), RMS= 0.009, n=18, r=12 pix \\ I = 19.099 - 10.098*(R-I) + 0.0098*(R-I) + 0$

C14 Images - SBIG STT-8300m camera

Bias: 13 March 2018

Dark: 14-17 March 2018, 19-20 March 2018

B Flats: 13 March 2018 V Flats: 13 March 2018 R Flats: 13 March 2018 I Flats: 13 March 2018

Photometric Calibration: 29 March 2018, 30 April 2018
NGC 2997 (galaxy): 07 March 2018, 29 March 2018
NGC 1672 (galaxy): 08 March 2018, 09 March 2018
NGC 2323 (open cluster): 13 March 2018, 29 March 2018
NGC 2362 (open cluster): 18 March 2018, 29 March 2018

Corot T-1 (exoplanet):21 March 2018 * * "Entered_Coordinates"

NGC 2477 (open cluster): 22 March 2018

The photometric calibration from 29 March 2018 for the C14 is:

 $B = 21.735 - 2.5\log(f/t) - 0.237*A + 0.205*(B-V)$, RMS= 0.030, n=39, r=30 pix

$$\label{eq:V} \begin{split} V &= 21.935 - 2.5log(f/t) - 0.175*A - 0.079*(B-V), \ RMS= 0.017, \ n=39, \ r=30 \ pix \\ R &= 21.823 - 2.5log(f/t) - 0.133*A - 0.157*(V-R), \ RMS= 0.010, \ n=39, \ r=30 \ pix \\ I &= 20.806 - 2.5log(f/t) - 0.094*A - 0.034*(R-I), \ RMS= 0.013, \ n=39, \ r=30 \ pix \end{split}$$

The photometric calibration from 30 April 2018 for the C14 is:

 $B = 21.792 - 2.5log(f/t) - 0.248*A + 0.199*(B-V), RMS= 0.016, n=18, r=30 pix \\ V = 21.973 - 2.5log(f/t) - 0.166*A - 0.075*(B-V), RMS= 0.007, n=18, r=30 pix \\ R = 21.841 - 2.5log(f/t) - 0.116*A - 0.161*(V-R), RMS= 0.006, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.038*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.038*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.038*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.038*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.038*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.038*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.008*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.008*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.008*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.008*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.008*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.008*(R-I), RMS= 0.013, n=18, r=30 pix \\ I = 20.820 - 2.5log(f/t) - 0.069*A - 0.008*(R-I), RMS= 0.008*(R-I),$