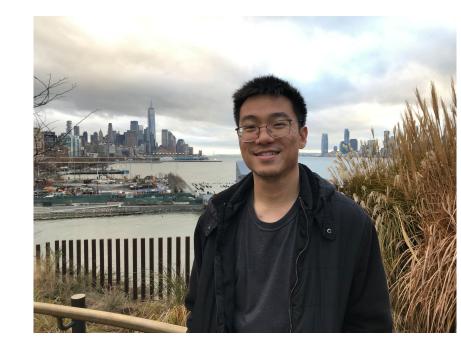
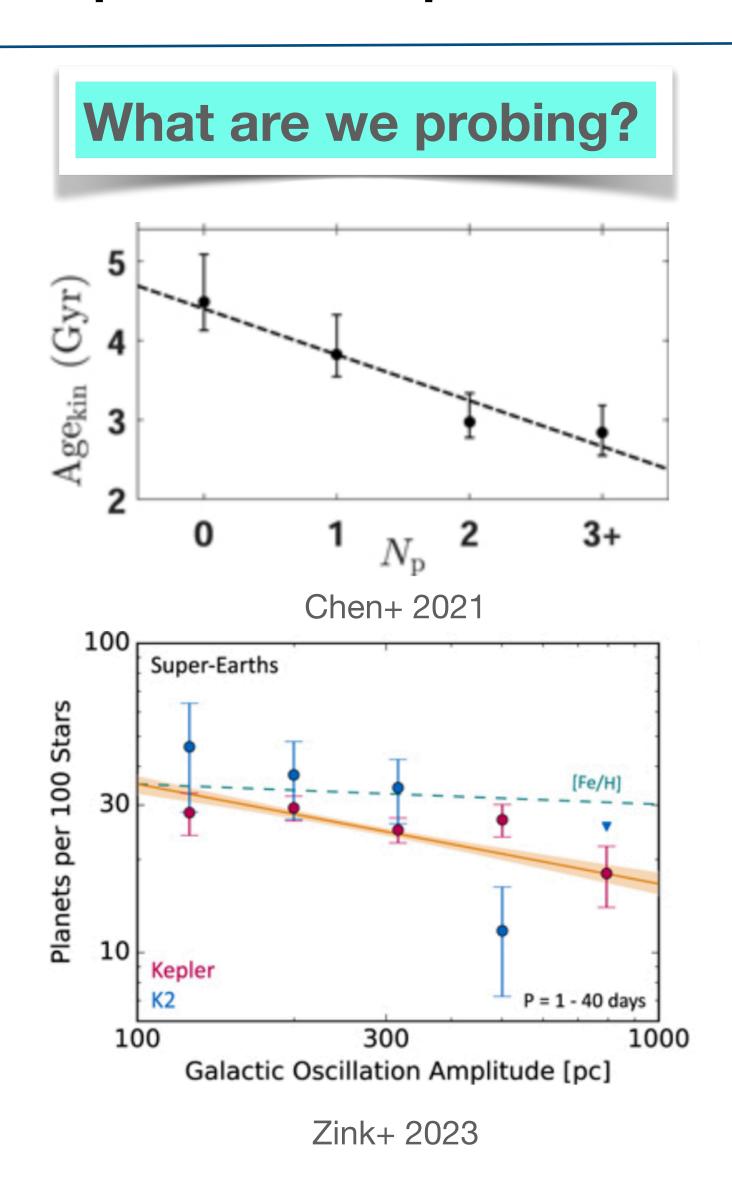
Ages of "Singles" vs "Multis": Predictions for Dynamical Sculpting over Gyr in the Kepler Sample

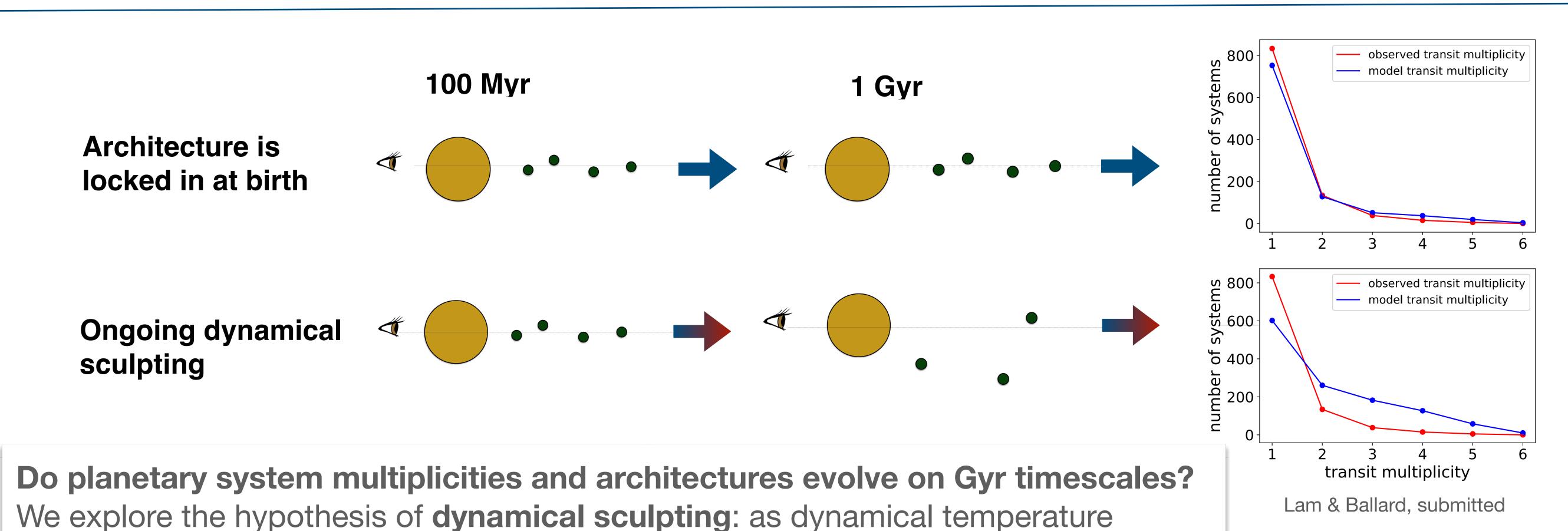


Scan for virtual poster and repo! Christopher Lam¹ & Sarah Ballard¹ ¹University of Florida

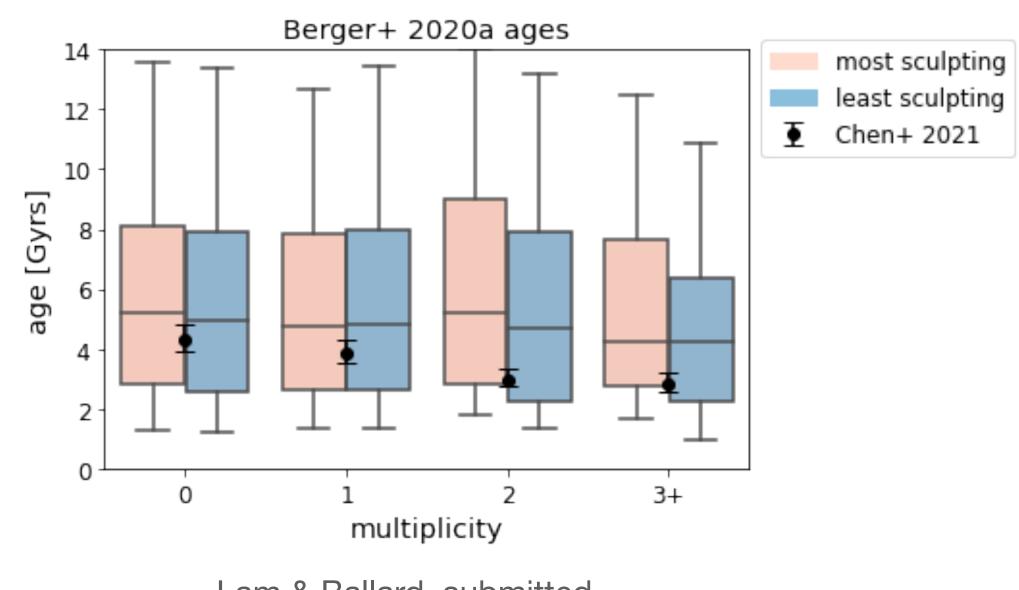
increases, planet orbits get perturbed to higher eccentricities and inclinations, or





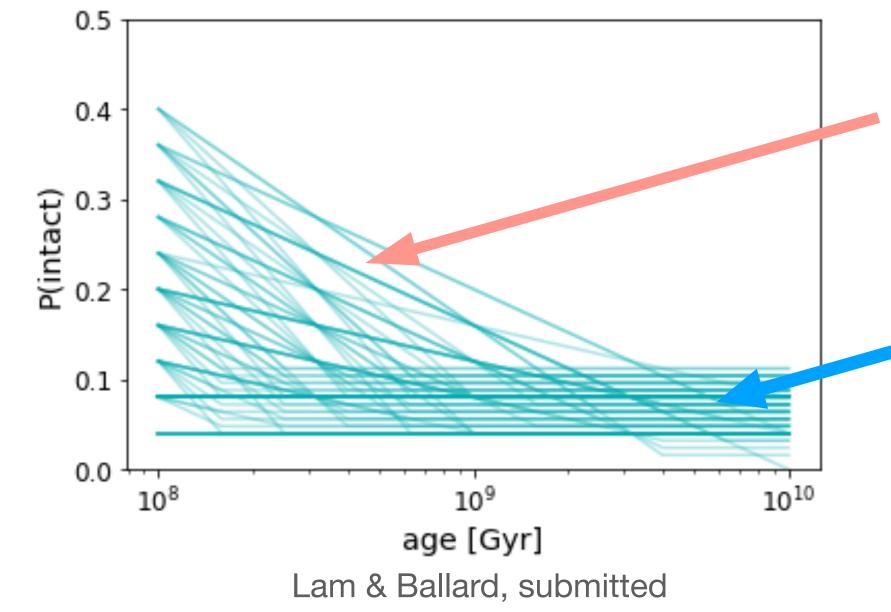


What did we find?



Lam & Ballard, submitted

Is there actually a difference in multiplicity between young and old systems? Not using the isochrone ages and their errors from the Berger+ 2020a Gaia-Kepler crossmatch.

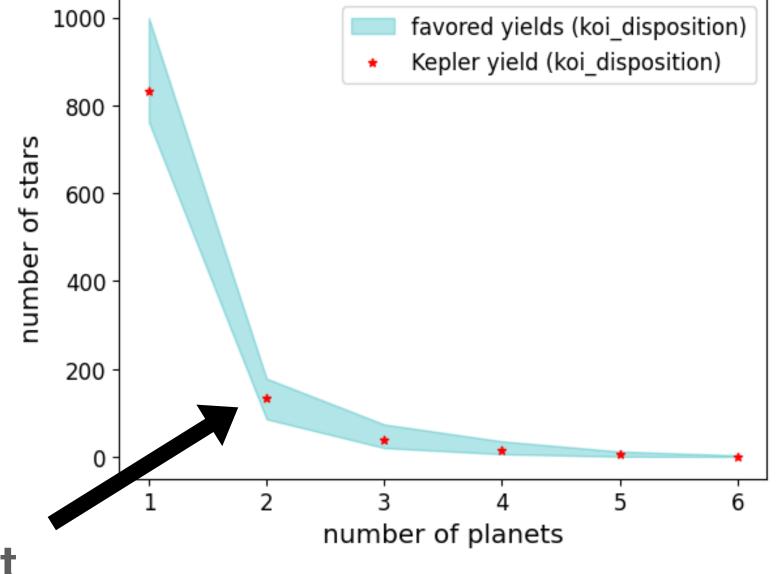


they may experience orbit crossings that lead to ejections.

Favored: the fraction of dynamically cool "intact") systems changes over Gyr

Also favored: the intact fraction starts lower but doesn't change after 100 Myr

Both models have transit multiplicity yields resembling Kepler's!



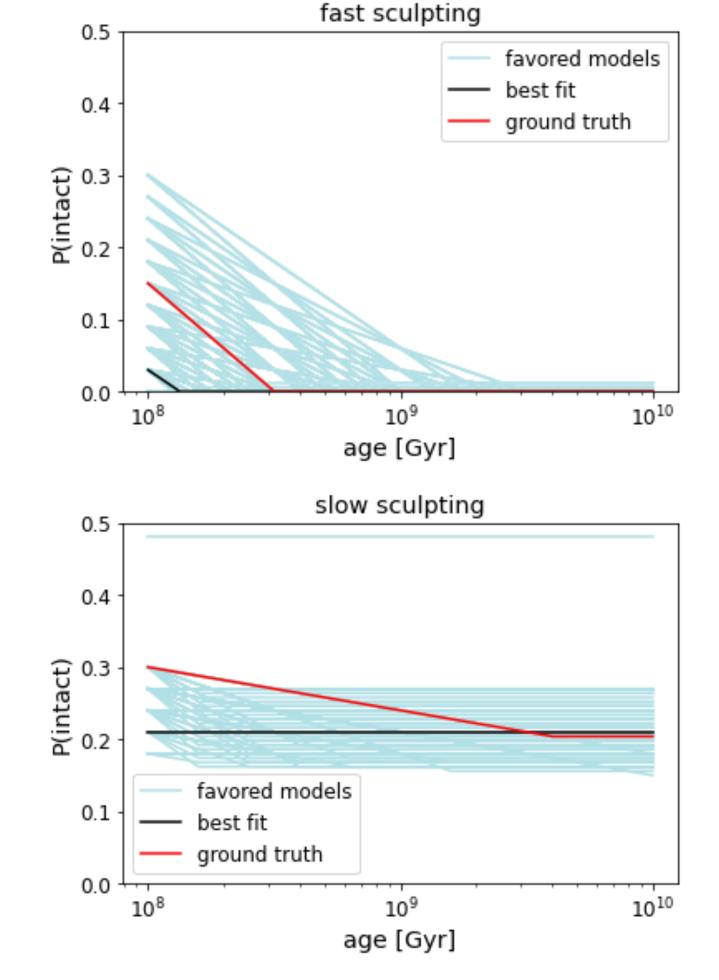
Yet, there are many different dynamical sculpting laws that match the Kepler transit multiplicity. Thus, transit multiplicity alone is insufficient for constraining Gyr sculpting. We may need some law that turns on later, or it may imprint through some other observables.

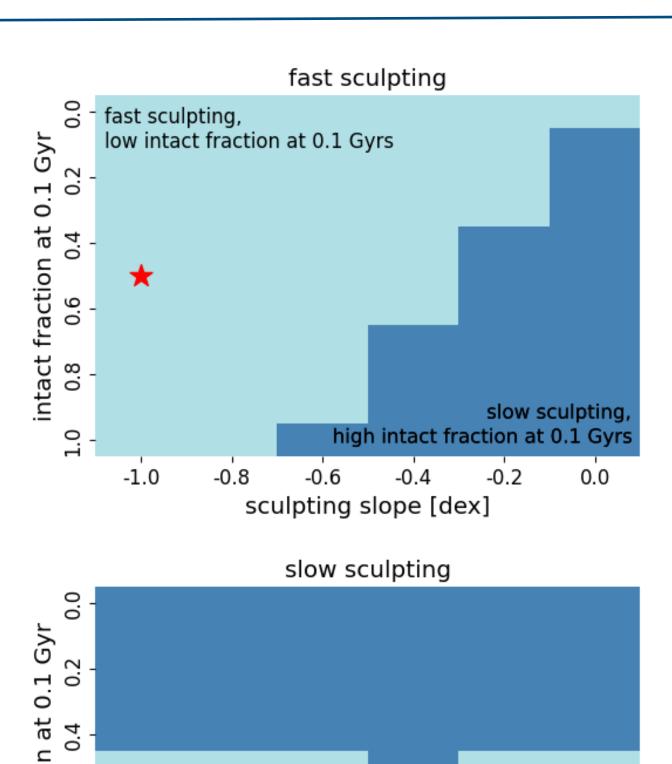
How did we do it?

We ran injection-recovery tests on different sculpting laws.

Depending on the ground truth, we can rule out different amounts of parameter space.

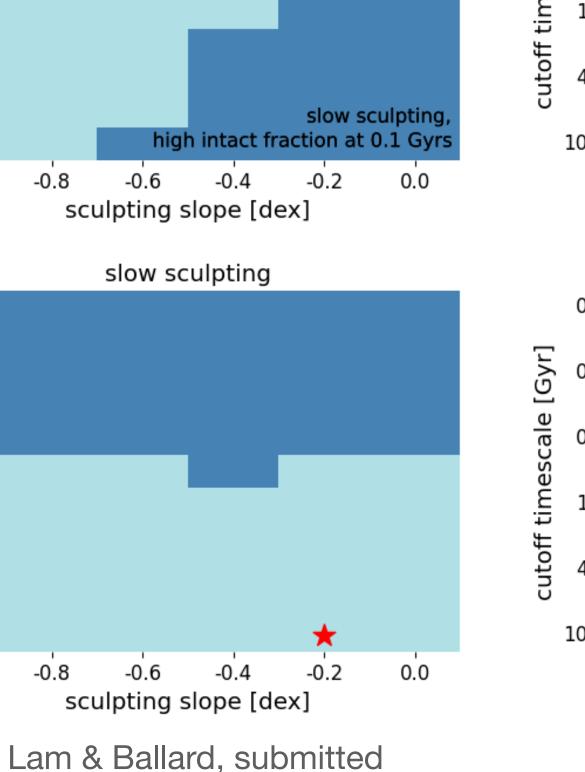
Across the board, we found that the primary driver of a model's likelihood is the presentday intact fraction it yields.

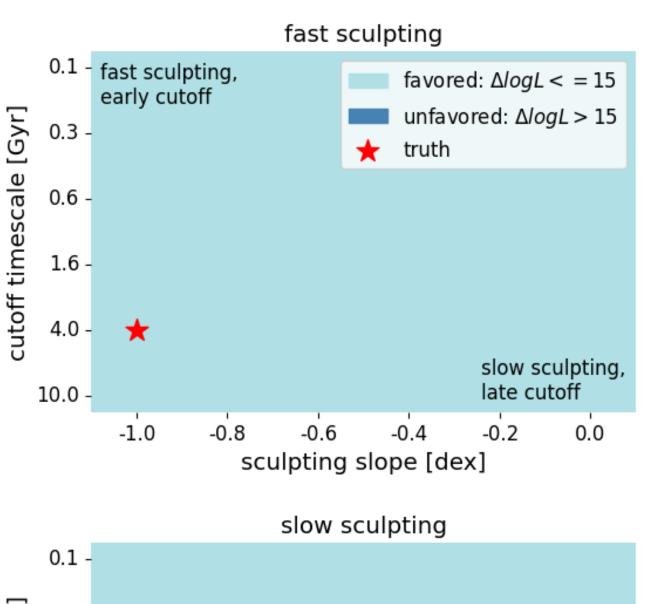


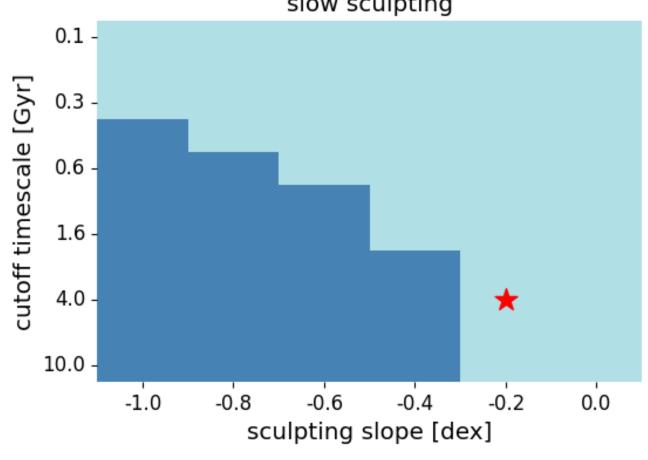


-0.6

sculpting slope [dex]







intact 0.8

-1.0