1

Testing production scenarios for (anti-)(hyper-)nuclei with multiplicity-dependent measurements at the LHC

F. Bellini (CERN, Geneva) †
and A. P. Kalweit (CERN, Geneva)

The production of light anti- and hyper-nuclei provides unique observables to characterise the system created in high energy proton-proton (pp), proton-nucleus (pA) and nucleus-nucleus (AA) collisions. In particular, nuclei and hyper-nuclei are special objects with respect to non-composite hadrons (such as pions, kaons, protons, etc.), because their size is comparable to a fraction or the whole system created in the collision. Their formation is typically described within the framework of coalescence and thermal-statistical production models. In order to distinguish between the two production scenarios, we propose to measure the coalescence parameter B_A for different anti- and hyper-nuclei (that differ by mass, size and internal wave-function) as a function of the size of the particle emitting source. The latter can be controlled by performing systematic measurements of light anti- hyper- nuclei in different collision systems (pp, pA, AA) and as a function of the multiplicity of particles created in the collision. While it is often argued that the coalescence and the thermal model approach give very similar predictions for the production of light nuclei in heavy-ion collisions, our study shows that large differences can be expected for hyper-nuclei with extended wave-functions, as the hyper-triton. We compare the model predictions with data from the ALICE experiment and we discuss perspectives for future measurements with the upgraded detectors during the High-Luminosity LHC phase in the next decade.

1. Introduction: the "anti-nuclei" puzzle

The formation of light anti- and hyper-nuclei in high energy proton-proton (pp), proton-nucleus (pA) and nucleus-nucleus (AA) collisions provides unique observables for the study of the system created in these reactions, and can be used to understand both the internal structure and the formation mechanisms of loosely-bound composite objects. The production of (anti-)(hyper-)nuclei in high-energy collisions is commonly described

* XXV Cracow EPIPHANY Conference on Advances in Heavy Ion Physics
† Presenter. For correspondence: francesca.bellini@cern.ch

by following two distinct approaches: formation by nucleon coalescence at the system (kinetic) freeze-out [1–6] or thermal-statistical production at the chemical freeze-out [7,8]. Thanks to the large data samples of pp, p–Pb and Pb–Pb collisions collected during the first ten years of operations of the CERN Large Hadron Collider (LHC), A Large Ion Collider Experiment (ALICE) Collaboration has measured the production of light nuclei and anti-nuclei at several centre-of-mass energies [9–15], thus providing a crucial experimental input and a boost to theoretical and phenomenological investigations [16–22]. In small collision systems, the experimental results seem to confirm the validity of the coalescence picture, with the most recent multiplicity-differential measurements pointing toward a dependence of the coalescence process on the volume of the particle-emitting source ("source size" hereafter). In heavy-ion collisions, coalescence approaches that do not take into account the source size are not able to reproduce the data. At the same time, the production of light nuclei, anti-nuclei and hypertriton as measured in Pb–Pb collisions is found to be consistent with statistical-thermal model predictions and a non-zero deuteron elliptic flow is observed. This is surprising as (anti-)nuclei produced at chemical freeze-out are not expected to survive the hadronic phase: the deuteron is a "fragile object" when surrounded by the fireball created in heavy-ion collisions, because its binding energy ($B_E = 2.2$ MeV) is much lower than the characteristic temperatures of the system ($T_{chem} \approx 153$ MeV, $T_{kin} \approx 100$ MeV). Moreover, the cross-section for pion-induced deuteron breakup is significantly larger than the typical (pseudo)-elastic cross-sections for the re-scattering of hadronic resonance decay products [22–25]. These observations pose the "(anti-)nuclei puzzle": how can loosely-bound composite objects survive in the dense and hot fireball, freeze-out and develop collective flow like the other light-flavour non-composite hadrons?

In our study [17] we have extended and combined known formalisms used to describe (anti-)(hyper-)nuclei production in order to allow, for the first time, a direct comparison of the thermal and coalescence models as well as a direct comparison to the ALICE data. Identifying the coalescence parameter (B_A) as the key observable, we present a consistent picture across different collision systems (pp, p–Pb, Pb–Pb) for light (anti-)(hyper-)nuclei with mass number $A = 2, 3$ and 4. We also suggest to address the open questions by looking at the production of nuclei and hyper-nuclei up to $A = 4$ that differ by size and properties, measured as a function of multiplicity used as a proxy for the source size. Whereas the (anti-)deuteron production has been measured multi-differentially and quite precisely with the LHC Run 1 and 2 data, the study of heavier objects with $A = 3$ and 4 will greatly profit from the increase in integrated luminosity foreseen at the LHC Runs 3 and 4 in all collision systems [26]. A comprehension of (anti-)nuclei production mechanisms is not only relevant for nuclear and hadronic physics, but has applications in astrophysics and indirect Dark Matter searches [27]. In recent years, it has been suggested that the detection of light anti-nuclei in space could provide a signature for the presence of Dark Matter in the Cosmos, see for instance [28, 29]. Anti-deuterons and $^{3}\overline{\mathrm{He}}$ might indeed be produced by coalescence of antiprotons and antineutrons coming from the annihilation of Weakly Interacting Massive Particles into Standard Model particles, for which anti-nuclei created in reactions between primary cosmic ray protons and interstellar matter (pp, pA collisions) represent a source of background.

The main features of the theoretical frameworks employed for our study are briefly summarised in Sec. 2, whereas we address the reader to [17] for the full details. Section 3 presents the main results and conclusions follow.

2. Modelling Light (anti-)(hyper-)nuclei production

Nuclei and hyper-nuclei are special objects with respect to non-composite hadrons (pions, protons, etc.), because their size is comparable to a fraction or the whole system created in high-energy proton-proton (pp), proton-nucleus (pA) and nucleus-nucleus (AA) collisions [34]. Their size is typically defined as the rms of their (charge) wave-function, corresponding to about 2 fm for light (anti-)nuclei as obtained from electron scattering experiments. For the hyper-triton, theoretical calculations indicate a rms of the wave-function of about 5 fm [35], significantly larger than that of non-strange nuclei with mass number $A = 3$ and driven by the average separation of the Λ relative to the two other nucleons. This difference in the wave-functions results in dramatic consequences for the production scenarios, as discussed in the following. The properties of the objects under study here are summarised in Tab. 1.

2.1. Thermal-statistical models

Thermal-statistical models [7, 8] have been successful in describing the production of light (anti-)(hyper-)nuclei across a wide range of energies in AA collisions, including production at the LHC. In this approach, particles are produced from a fireball in thermal equilibrium with temperatures of $T_{chem} \approx 156$ MeV.

2.2. The coalescence approach

Starting from the model described in [5, 6], we have obtained in [17] a generalised expression for the coalescence parameter B_A

Mass number	Nucleus	Compo-sition	B_E (MeV)	Spin J_A	(Charge) rms radius ${\lambda_A}^{meas}$ (fm)	Harmonic oscillator size parameter r_A (fm)	Refs.
A = 2	d	pn	2.224575 (9)	1	2.1413 ± 0.0025	3.2	[36, 37]
A = 3	^{3}H	pnn	8.4817986 (20)	1/2	1.755 ± 0.086	2.15	[38]
	^{3}He	ppn	7.7180428 (23)	1/2	1.959 ± 0.030	2.48	[38]
	$^3_\Lambda$H	pΛn	0.13 ± 0.05	1/2	4.9 – 10.0	6.8 – 14.1	[35, 39]
A = 4	^{4}He	ppnn	28.29566 (20)	0	1.6755 ± 0.0028	1.9	[40, 41]
	$^4_\Lambda$H	pΛnn	2.04 ±0.04	0	2.0 – 3.8	2.4 – 4.9	[35, 39]
	$^4_{\Lambda\Lambda}$H	pΛΛn	0.39 – 0.51	1	4.2 – 7.1	5.5 – 9.4	[35]
	$^4_\Lambda$He	ppΛn	2.39 ± 0.03	0	2.0 – 3.8	2.4 – 4.9	[35, 39]

Table 1. Properties of nuclei and hyper-nuclei with mass number $A \leq 4$. B_E is the binding energy in MeV. The size of the nucleus is given in terms of the (charge) rms radius of the wave-function, λ_A. The size parameter of the wave-function of the harmonic oscillator potential, r_A, is chosen such that the measured/expected rms is approximately reproduced. Please note that the proton rms charge radius $\lambda_p = 0.879(8)$ fm [42] is subtracted quadratically from the measured rms charge radius λ_A^{meas} of the nucleus $\lambda_A = \sqrt{(\lambda_A^{meas})^2 - \lambda_p^2}$ to account for the finite extension of the constituents. Implicitly we assume here that $\lambda_\Lambda \approx \lambda_n \approx \lambda_p$. References are given in the last column. The spin of $^4_{\Lambda\Lambda}$H is discussed in the text of [35].

$$\boxed{B_A = \frac{2J_A + 1}{2^A} \frac{1}{\sqrt{A}} \frac{1}{m_T^{A-1}} \left(\frac{2\pi}{R^2 + (\frac{r_A}{2})^2} \right)^{\frac{3}{2}(A-1)}} \quad , \tag{1}$$

which is a function of the spin of the particle J_A, its transverse mass m_{T}, its size parameter r_A and the source radius R. Figure 1 shows the source radius dependence of B_A for different composite objects, including the nuclei and hyper-nuclei with $A = 2, 3$ and 4 whose properties are reported in Tab. 1.

3. Comparison with data

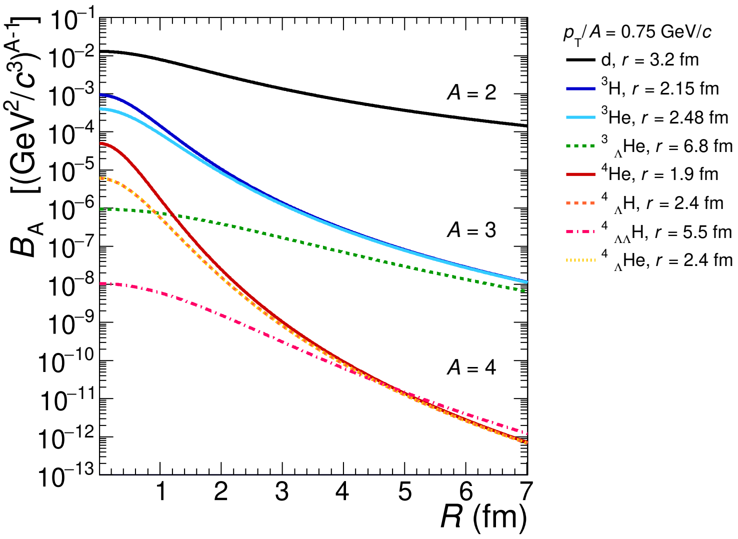

Fig. 1. (Color online) Coalescence parameter B_A as a function of the source radius R as predicted from the coalescence model (Eq. 1) for various composite objects with $p_T/A = 0.75$ GeV/c. For each (hyper-)nucleus, the radius r used for the calculation is reported in the legend.

REFERENCES

[1] S. T. Butler and C. A. Pearson *Phys. Rev.* **129** (1963) 836–842.

[2] J. I. Kapusta *Phys. Rev.* **C21** (1980) 1301–1310.

[3] H. Sato and K. Yazaki *Phys. Lett.* **B98** (1981) 153–157.

[4] J. L. Nagle, B. S. Kumar, D. Kusnezov, H. Sorge, and R. Mattiello *Phys. Rev.* **C53** (1996) 367–376.

[5] R. Scheibl and U. W. Heinz *Phys. Rev.* **C59** (1999) 1585–1602, `arXiv:nucl-th/9809092 [nucl-th]`.

[6] K. Blum, K. C. Y. Ng, R. Sato, and M. Takimoto *Phys. Rev.* **D96** no. 10, (2017) 103021, `arXiv:1704.05431 [astro-ph.HE]`.

[7] A. Andronic, P. Braun-Munzinger, J. Stachel, and H. Stocker *Physics Letters B* **697** no. 3, (2011) 203 – 207.

[8] A. Andronic, P. Braun-Munzinger, K. Redlich, and J. Stachel *Nature* **561** no. 7723, (2018) 321–330, `arXiv:1710.09425 [nucl-th]`.

[9] **ALICE** Collaboration, J. Adam *et al. Phys. Rev.* **C93** no. 2, (2016) 024917, `arXiv:1506.08951 [nucl-ex]`.

[10] **ALICE** Collaboration, J. Adam *et al. Phys. Lett.* **B754** (2016) 360–372, `arXiv:1506.08453 [nucl-ex]`.

[11] Anielski, Jonas *J. Phys. Conf. Ser.* **612** no. 1, (2015) 012014.

[12] **ALICE** Collaboration, M. Puccio *Nucl. Phys.* **A982** (2019) 447–450.

[13] **ALICE** Collaboration, S. Acharya *et al. Phys. Rev.* **C97** no. 2, (2018) 024615, `arXiv:1709.08522 [nucl-ex]`.

[14] **ALICE** Collaboration, S. Acharya *et al. Eur. Phys. J.* **C77** no. 10, (2017) 658, `arXiv:1707.07304 [nucl-ex]`.

[15] **ALICE** Collaboration, S. Acharya *et al.* `arXiv:1902.09290 [nucl-ex]`.

[16] S. Mrowczynski *Acta Phys. Polon.* **B48** (2017) 707, `arXiv:1607.02267 [nucl-th]`.

[17] F. Bellini and A. P. Kalweit `arXiv:1807.05894 [hep-ph]`.

[18] S. Bazak and S. Mrowczynski *Mod. Phys. Lett.* **A33** no. 25, (2018) 1850142, `arXiv:1802.08212 [nucl-th]`.

[19] W. Zhao, L. Zhu, H. Zheng, C. M. Ko, and H. Song *Phys. Rev.* **C98** no. 5, (2018) 054905, `arXiv:1807.02813 [nucl-th]`.

[20] K.-J. Sun, C. M. Ko, and B. Doenigus `arXiv:1812.05175 [nucl-th]`.

[21] X. Xu and R. Rapp `arXiv:1809.04024 [nucl-th]`.

[22] D. Oliinychenko, L.-G. Pang, H. Elfner, and V. Koch `arXiv:1809.03071 [hep-ph]`.

[23] H. Garcilazo *Phys. Rev. Lett.* **48** (1982) 577–580.

[24] S. A. Bass *et al. Prog. Part. Nucl. Phys.* **41** (1998) 255–369, `arXiv:nucl-th/9803035 [nucl-th]`.

[25] J. Schukraft *Nucl. Phys.* **A967** (2017) 1–10, `arXiv:1705.02646 [hep-ex]`.

[26] Z. Citron *et al.*, "Future physics opportunities for high-density QCD at the LHC with heavy-ion and proton beams," in *HL/HE-LHC Workshop: Workshop on the Physics of HL-LHC, and Perspectives at HE-LHC Geneva, Switzerland, June 18-20, 2018.* 2018. `arXiv:1812.06772 [hep-ph]`.

[27] T. Aramaki *et al. Phys. Rept.* **618** (2016) 1–37, `arXiv:1505.07785 [hep-ph]`.

[28] M. Cirelli, N. Fornengo, M. Taoso, and A. Vittino *JHEP* **08** (2014) 009, `arXiv:1401.4017 [hep-ph]`.

[29] M. Korsmeier, F. Donato, and N. Fornengo *Phys. Rev.* **D97** no. 10, (2018) 103011, `arXiv:1711.08465 [astro-ph.HE]`.

[30] **ALICE** Collaboration, B. Abelev *et al. Phys. Rev.* **C88** (2013) 044910, `arXiv:1303.0737 [hep-ex]`.

[31] P. Castorina and H. Satz `arXiv:1901.10407 [hep-ph]`.

[32] U. Heinz, "Coalescence model involving HBT and flow." Presentation at EMMI Workshop in Torino, November 2017.

[33] **ALICE** Collaboration, B. B. Abelev *et al. Phys. Rev.* **C91** (2015) 024609, `arXiv:1404.0495 [nucl-ex]`.

[34] **ALICE** Collaboration, J. Adam *et al. Phys. Rev.* **C93** no. 2, (2016) 024905, arXiv:1507.06842 [nucl-ex].

[35] H. Nemura, Y. Suzuki, Y. Fujiwara, and C. Nakamoto *Prog. Theor. Phys.* **103** (2000) 929–958, arXiv:nucl-th/9912065 [nucl-th].

[36] C. Van Der Leun and C. Alderliesten *Nucl. Phys.* **A380** (1982) 261–269.

[37] P. J. Mohr, D. B. Newell, and B. N. Taylor *Rev. Mod. Phys.* **88** no. 3, (2016) 035009, arXiv:1507.07956 [physics.atom-ph].

[38] J. E. Purcell and C. G. Sheu *Nucl. Data Sheets* **130** (2015) 1–20.

[39] D. H. Davis *Nucl. Phys.* **A754** (2005) 3–13.

[40] M. Wang, G. Audi, F. Kondev, W. Huang, S. Naimi, and X. Xu *Chinese Physics C* **41** no. 3, (2017) 030003.

[41] I. Angeli and K. P. Marinova *Atom. Data Nucl. Data Tabl.* **99** no. 1, (2013) 69–95.

[42] **A1** Collaboration, J. C. Bernauer *et al. Phys. Rev. Lett.* **105** (Dec, 2010) 242001.

[43] V. Vovchenko, B. Doenigus, and H. Stoecker *Phys. Lett.* **B785** (2018) 171–174, arXiv:1808.05245 [hep-ph].

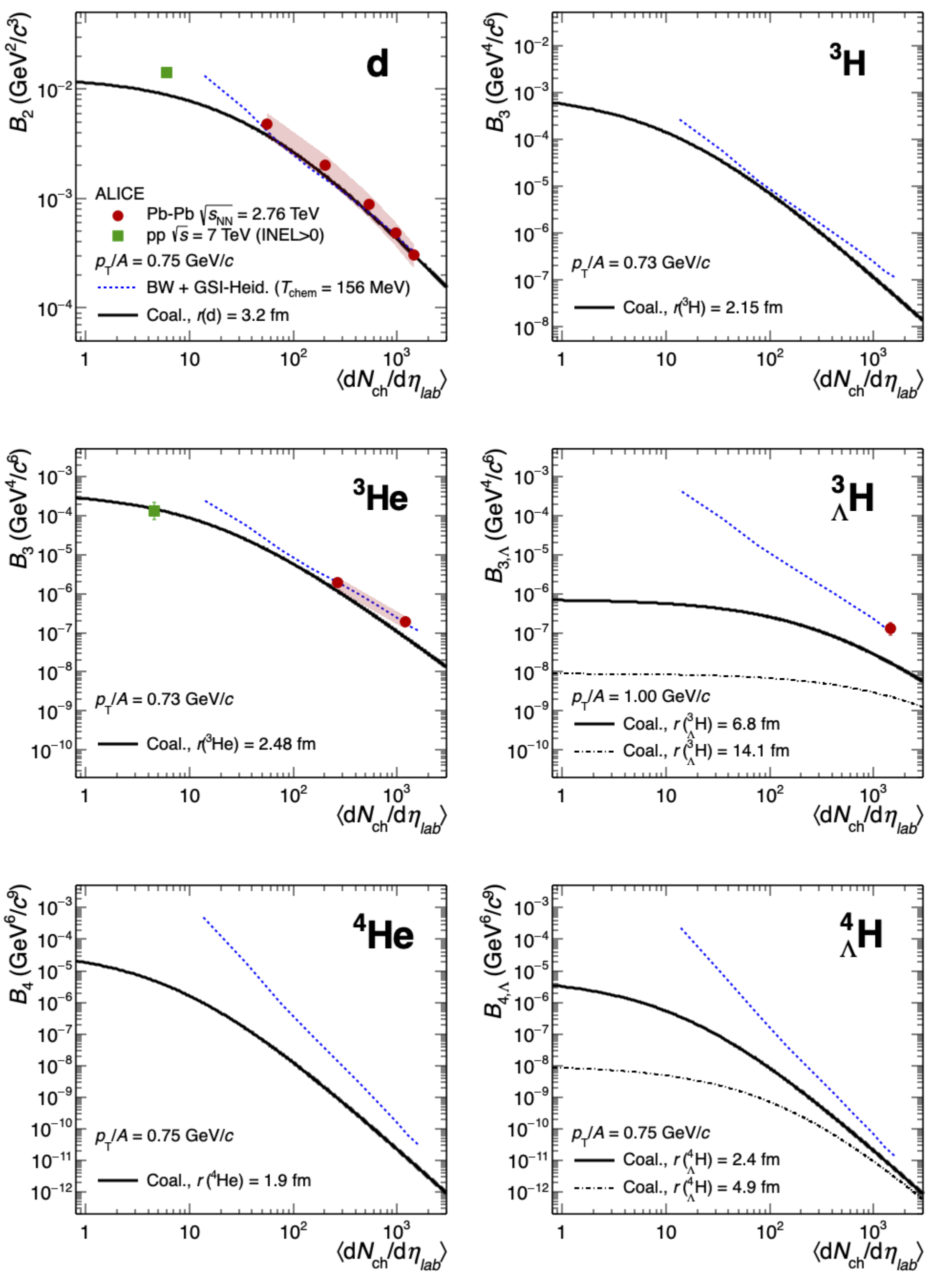

Fig. 2. (Color online) Coalescence parameter B_A as a function of the average charged particle multiplicity density for various (hyper-)nuclei, up to $A = 4$. The coalescence calculations (continuous or dashed-dotted black lines) are compared to the thermal+blast-wave predictions (dashed blue line), as well as to pp (green square) and Pb–Pb (red circles) collision data from ALICE [9, 10, 13].

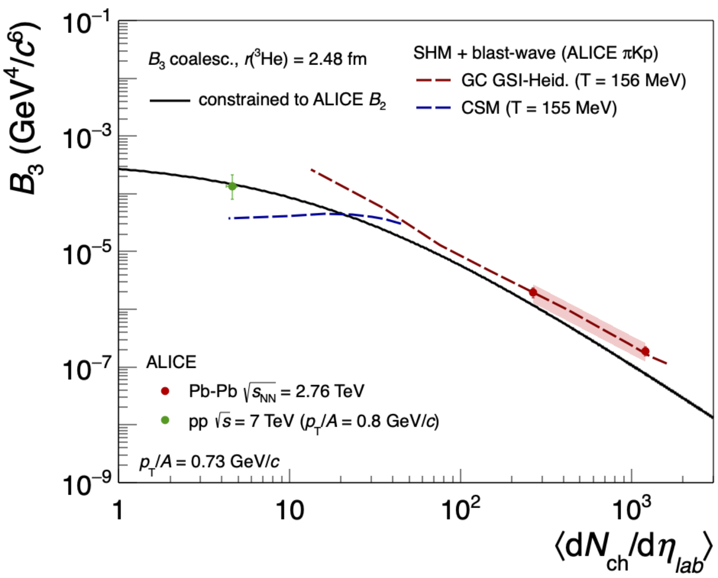

Fig. 3. (Color online) Coalescence parameter B_3 for ^{3}He as a function of the average charged particle multiplicity density. The coalescence calculation (continuous black line) is compared to two thermal+blast-wave predictions (dashed lines), obtained by using the Grand Canonical (GC, red) [8] and Canonical Statistical Model (CSM, blue) [43] expectations for the ^{3}He yield, respectively. ALICE data from pp (green circles) and Pb–Pb (red circles) collisions [9, 13] are reported.