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Testing production scenarios for (anti-)(hyper-)nuclei with multiplicity-dependent measurements at the LHC

F. Bellini (CERN, Geneva) † and A. P. Kalweit (CERN, Geneva)

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The production of light anti- and hyper-nuclei provides unique observables to characterise the system created in high energy proton-proton (pp), proton-nucleus (pA) and nucleus-nucleus (AA) collisions. In particular, nuclei and hyper-nuclei are special objects with respect to non-composite hadrons (such as pions, kaons, protons, etc.), because their size is comparable to a fraction or the whole system created in the collision. Their formation is typically described within the framework of coalescence and thermal-statistical production models. In order to distinguish between the two production scenarios, we propose to measure the coalescence parameter B_A for different anti- and hyper-nuclei (that differ by mass, size and internal wave-function) as a function of the size of the particle emitting source. The latter can be controlled by performing systematic measurements of light anti- hyper- nuclei in different collision systems (pp, pA, AA) and as a function of the multiplicity of particles created in the collision. While it is often argued that the coalescence and the thermal model approach give very similar predictions for the production of light nuclei in heavy-ion collisions, our study shows that large differences can be expected for hypernuclei with extended wave-functions, as the hyper-triton. We compare the model predictions with data from the ALICE experiment and we discuss perspectives for future measurements with the upgraded detectors during the High-Luminosity LHC phase in the next decade.

1. Introduction: the "anti-nuclei" puzzle

The formation of light anti- and hyper-nuclei in high energy protonproton (pp), proton-nucleus (pA) and nucleus-nucleus (AA) collisions provides unique observables for the study of the system created in these reactions, and can be used to understand both the internal structure and the formation mechanisms of loosely-bound composite objects. The production of (anti-)(hyper-)nuclei in high-energy collisions is commonly described

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[†] Presenter. For correspondence: francesca.bellini@cern.ch

by following two distinct approaches: formation by nucleon coalescence at the system (kinetic) freeze-out [1–6] or thermal-statistical production at the chemical freeze-out [7,8]. Thanks to the large data samples of pp, p-Pb and 36 Pb-Pb collisions collected during the first ten years of operations of the 37 CERN Large Hadron Collider (LHC), A Large Ion Collider Experiment 38 (ALICE) Collaboration has measured the production of light nuclei and anti-nuclei at several centre-of-mass energies [9–15], thus providing a cru-40 cial experimental input and a boost to theoretical and phenomenological 41 investigations [16–22]. In small collision systems, the experimental results 42 seem to confirm the validity of the coalescence picture, with the most recent 43 multiplicity-differential measurements pointing toward a dependence of the 44 coalescence process on the volume of the particle-emitting source ("source 45 size" hereafter). In heavy-ion collisions, coalescence approaches that do not 46 take into account the source size are not able to reproduce the data. At 47 the same time, the production of light nuclei, anti-nuclei and hypertriton as measured in Pb–Pb collisions is found to be consistent with statistical-49 thermal model predictions and a non-zero deuteron elliptic flow is observed. This is surprising as (anti-)nuclei produced at chemical freeze-out are not 51 expected to survive the hadronic phase: the deuteron is a "fragile object" 52 when surrounded by the fireball created in heavy-ion collisions, because its 53 binding energy ($B_E = 2.2 \text{ MeV}$) is much lower than the characteristic tem-54 peratures of the system ($T_{chem} \approx 153 \text{ MeV}$, $T_{kin} \approx 100 \text{ MeV}$). Moreover, the 55 cross-section for pion-induced deuteron breakup is significantly larger than 56 the typical (pseudo)-elastic cross-sections for the re-scattering of hadronic 57 resonance decay products [22–25]. These observations pose the "(anti-)nuclei 58 puzzle": how can loosely-bound composite objects survive in the dense and 59 hot fireball, freeze-out and develop collective flow like the other light-flavour 60 non-composite hadrons?

In our study [17] we have extended and combined known formalisms used to describe (anti-)(hyper-)nuclei production in order to allow, for the first time, a direct comparison of the thermal and coalescence models as well as a direct comparison to the ALICE data. Identifying the coalescence parameter (B_A) as the key observable, we present a consistent picture across different collision systems (pp, p-Pb, Pb-Pb) for light (anti-)(hyper-)nuclei with mass number A = 2, 3 and 4. We also suggest to address the open questions by looking at the production of nuclei and hyper-nuclei up to A = 4 that differ by size and properties, measured as a function of multiplicity used as a proxy for the source size. Whereas the (anti-)deuteron production has been measured multi-differentially and quite precisely with the LHC Run 1 and 2 data, the study of heavier objects with A = 3 and 4 will greatly profit from the increase in integrated luminosity foreseen at the LHC Runs 3 and 4 in all collision systems [26]. A comprehension of (anti-)nuclei production

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mechanisms is not only relevant for nuclear and hadronic physics, but has applications in astrophysics and indirect Dark Matter searches [27]. In recent years, it has been suggested that the detection of light anti-nuclei in space could provide a signature for the presence of Dark Matter in the Cosmos, see for instance [28, 29]. Anti-deuterons and ${}^{3}\overline{\text{He}}$ might indeed be produced by coalescence of antiprotons and antineutrons coming from the annihilation of Weakly Interacting Massive Particles into Standard Model particles, for which anti-nuclei created in reactions between primary cosmic ray protons and interstellar matter (pp, pA collisions) represent a source of background.

The main features of the theoretical frameworks employed for our study are briefly summarised in Sec. 2, while we address the reader to [17] for the full details. Section 3 presents the main results and conclusions follow.

2. Modelling light (anti-)(hyper-)nuclei production

For our study, we consider nuclei and hyper-nuclei with mass number 89 A=2,3 and 4, whose properties are summarised in Tab. 1. Those properties 90 are the same as for their anti-matter counterparts and we assume that the 91 same formation mechanisms are at play for matter and anti-matter¹. Nuclei 92 and hyper-nuclei are special objects with respect to non-composite hadrons 93 (pions, protons, etc.), because their size is comparable to a fraction or the whole system created in pp, p-Pb and Pb-Pb collisions. The size is typically 95 defined in two ways: a) as the rms of the (charge) distribution (λ_A) , typically 96 measured in electron scattering experiments, or b) as the size parameter of 97 the object wave-function (r_A) , typically taken as the gaussian solution of an isotropic harmonic oscillator potential in coalescence models. For light 99 nuclei, $\lambda_A \approx 2$ fm. For the hyper-triton, theoretical calculations indicate 100 a charge rms radius $\lambda_A \approx 5$ fm [31], driven by the average separation of 101 the Λ relative to the two other nucleons. Assuming a similar structure (e.g. 102 a s-wave interaction for a bound state of a n or a Λ with a deuteron), the 103 hypertriton results in a much larger object than the other non-strange nuclei 104 with A=3. A simple relation holds between λ_A and r_A , see [17]. In Tab. 1 105 the binding energy (B_E) is also reported. The most tightly bound nucleus 106 is ${}^{4}\text{He}$, whereas the most loosely bound object is ${}^{3}_{\Lambda}\text{H}$, that is also the largest 107 one. For the latter, we report the separation energy of the Λ baryon from 108 the deuteron ($B_{\Lambda} = 0.13 \text{ MeV}$). The large size and the low binding energy of the $^3_{\Lambda}{\rm H}$ with respect to the other (hyper-)nuclei has important consequences 110 on its production, as discussed in what follows.

For brevity, in the following we refer to "nuclei" and "hyper-nuclei" but we imply both matter and anti-matter.

Mass number	Nucleus	Composition	$B_E \; ({ m MeV})$	$_{J_{A}}^{\mathrm{Spin}}$	(Charge) rms radius λ_A^{meas} (fm)	Harmonic oscillator size parameter r_A (fm)	Refs.
A = 2	d	pn	2.224575(9)	1	2.1413 ± 0.0025	3.2	[32, 33]
A = 3	³ H ³ He ³ ΛH	pnn ppn pΛn	$\begin{array}{c} 8.4817986 \ (20) \\ 7.7180428 \ (23) \\ 0.13 \pm 0.05 \end{array}$	$1/2 \\ 1/2 \\ 1/2$	1.755 ± 0.086 1.959 ± 0.030 $4.9 - 10.0$	2.15 2.48 $6.8 - 14.1$	[34] [34] [31,35]
A = 4	$^{4}_{\stackrel{4}{\Lambda} H}$ $^{4}_{\stackrel{\Lambda}{\Lambda} \Lambda} H$ $^{4}_{\stackrel{\Lambda}{\Lambda} H}$	$ppnn$ $p\Lambda nn$ $p\Lambda\Lambda n$ $pp\Lambda n$	$28.29566 (20)$ 2.04 ± 0.04 $0.39 - 0.51$ 2.39 ± 0.03	0 0 1 0	1.6755 ± 0.0028 $2.0 - 3.8$ $4.2 - 7.1$ $2.0 - 3.8$	$1.9 \\ 2.4 - 4.9 \\ 5.5 - 9.4 \\ 2.4 - 4.9$	[36, 37] [31, 35] [31] [31, 35]

Table 1. Properties of nuclei and hyper-nuclei with mass number $A \leq 4$. B_E is the binding energy in MeV. The size parameter r_A , is chosen to approximately reproduce the measured/expected rms, $\lambda_A{}^{meas}$ (fm). The proton rms charge radius $\lambda_p = 0.879(8)$ fm [38] is subtracted quadratically from the measured rms charge radius $\lambda_A{}^{meas}$ of the nucleus $\lambda_A = \sqrt{(\lambda_A{}^{meas})^2 - \lambda_p^2}$ to account for the finite extension of the constituents. Implicitly we assume here that $\lambda_\Lambda \approx \lambda_n \approx \lambda_p$. The spin of ${}^4_{\Lambda}\Lambda$ H is discussed in the text of [31].

2.1. The coalescence approach

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Starting from the model described in [5,6], we have obtained in [17] a generalised expression for the coalescence parameter B_A

$$B_A = \frac{2J_A + 1}{2^A} \frac{1}{\sqrt{A}} \frac{1}{m_T^{A-1}} \left(\frac{2\pi}{R^2 + (\frac{r_A}{2})^2} \right)^{\frac{3}{2}(A-1)} , \qquad (1)$$

which is a function of the spin of the particle J_A , its transverse mass $m_{\rm T}$, its size parameter r_A and the source radius R. Figure 1 shows the source radius dependence of B_A for different composite objects, including the nuclei and hyper-nuclei with A=2,3 and 4 whose properties are reported in Tab. 1.

2.2. Thermal-statistical models

Thermal-statistical models [7,8] have been successful in describing the production of light (anti-)(hyper-)nuclei across a wide range of energies in AA collisions, including production at the LHC. In this approach, particles are produced from a fireball in thermal equilibrium with temperatures of $T_{chem} \approx 156$ MeV.

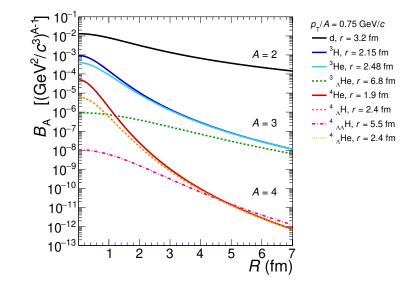


Fig. 1. (Color online) Coalescence parameter B_A as a function of the source radius R as predicted from the coalescence model (Eq. 1) for various composite objects with $p_{\rm T}/A=0.75~{\rm GeV}/c$. For each (hyper-)nucleus, the radius r used for the calculation is reported in the legend.

3. Comparison with data

REFERENCES

- 126 [1] S. T. Butler and C. A. Pearson Phys. Rev. 129 (1963) 836–842.
- [2] J. I. Kapusta *Phys. Rev.* C21 (1980) 1301–1310.
- 128 [3] H. Sato and K. Yazaki *Phys. Lett.* **B98** (1981) 153–157.
- [4] J. L. Nagle, B. S. Kumar, D. Kusnezov, H. Sorge, and R. Mattiello *Phys. Rev.* C53 (1996) 367–376.
- [5] R. Scheibl and U. W. Heinz Phys. Rev. C59 (1999) 1585-1602,
 arXiv:nucl-th/9809092 [nucl-th].
- [6] K. Blum, K. C. Y. Ng, R. Sato, and M. Takimoto *Phys. Rev.* **D96** no. 10, (2017) 103021, arXiv:1704.05431 [astro-ph.HE].
- 135 [7] A. Andronic, P. Braun-Munzinger, J. Stachel, and H. Stocker *Physics Letters* B **697** no. 3, (2011) 203 207.
- [8] A. Andronic, P. Braun-Munzinger, K. Redlich, and J. Stachel *Nature* 561
 no. 7723, (2018) 321–330, arXiv:1710.09425 [nucl-th].
- [9] ALICE Collaboration, J. Adam et al. Phys. Rev. C93 no. 2, (2016) 024917,
 arXiv:1506.08951 [nucl-ex].
- [10] ALICE Collaboration, J. Adam et al. Phys. Lett. B754 (2016) 360-372,
 arXiv:1506.08453 [nucl-ex].
- 143 [11] Anielski, Jonas J. Phys. Conf. Ser. 612 no. 1, (2015) 012014.
- 144 [12] ALICE Collaboration, M. Puccio Nucl. Phys. A982 (2019) 447–450.
- [13] ALICE Collaboration, S. Acharya *et al. Phys. Rev.* C97 no. 2, (2018)
 024615, arXiv:1709.08522 [nucl-ex].
- [14] ALICE Collaboration, S. Acharya et al. Eur. Phys. J. C77 no. 10, (2017)
 658, arXiv:1707.07304 [nucl-ex].
- 149 [15] ALICE Collaboration, S. Acharya et al. arXiv:1902.09290 [nucl-ex].
- [16] S. Mrowczynski Acta Phys. Polon. B48 (2017) 707, arXiv:1607.02267
 [nucl-th].
- 152 [17] F. Bellini and A. P. Kalweit arXiv:1807.05894 [hep-ph].
- [18] S. Bazak and S. Mrowczynski Mod. Phys. Lett. A33 no. 25, (2018) 1850142,
 arXiv:1802.08212 [nucl-th].
- [19] W. Zhao, L. Zhu, H. Zheng, C. M. Ko, and H. Song *Phys. Rev.* C98 no. 5,
 (2018) 054905, arXiv:1807.02813 [nucl-th].
- 157 [20] K.-J. Sun, C. M. Ko, and B. Doenigus arXiv:1812.05175 [nucl-th].
- 158 [21] X. Xu and R. Rapp arXiv:1809.04024 [nucl-th].
- [22] D. Oliinychenko, L.-G. Pang, H. Elfner, and V. Koch arXiv:1809.03071
 [hep-ph].
- 161 [23] H. Garcilazo Phys. Rev. Lett. 48 (1982) 577–580.
- 162 [24] S. A. Bass et al. Prog. Part. Nucl. Phys. 41 (1998) 255-369,
 163 arXiv:nucl-th/9803035 [nucl-th].
- 164 [25] J. Schukraft Nucl. Phys. A967 (2017) 1-10, arXiv:1705.02646 [hep-ex].

- [26] Z. Citron et al. in HL/HE-LHC Workshop: Workshop on the Physics of
 HL-LHC, and Perspectives at HE-LHC Geneva, Switzerland, June 18-20,
 2018. 2018. arXiv:1812.06772 [hep-ph].
- 168 [27] T. Aramaki et al. Phys. Rept. 618 (2016) 1-37, arXiv:1505.07785
 169 [hep-ph].
- [28] M. Cirelli, N. Fornengo, M. Taoso, and A. Vittino JHEP 08 (2014) 009,
 arXiv:1401.4017 [hep-ph].
- [29] M. Korsmeier, F. Donato, and N. Fornengo *Phys. Rev.* **D97** no. 10, (2018)
 103011, arXiv:1711.08465 [astro-ph.HE].
- [30] **ALICE** Collaboration, J. Adam *et al. Phys. Rev.* **C93** no. 2, (2016) 024905, arXiv:1507.06842 [nucl-ex].
- 176 [31] H. Nemura, Y. Suzuki, Y. Fujiwara, and C. Nakamoto *Prog. Theor. Phys.* 103 (2000) 929–958, arXiv:nucl-th/9912065 [nucl-th].
- 178 [32] C. Van Der Leun and C. Alderliesten Nucl. Phys. A380 (1982) 261–269.
- 179 [33] P. J. Mohr, D. B. Newell, and B. N. Taylor *Rev. Mod. Phys.* **88** no. 3, (2016) 035009, arXiv:1507.07956 [physics.atom-ph].
- 181 [34] J. E. Purcell and C. G. Sheu Nucl. Data Sheets 130 (2015) 1–20.
- 182 [35] D. H. Davis Nucl. Phys. A754 (2005) 3–13.
- [36] M. Wang, G. Audi, F. Kondev, W. Huang, S. Naimi, and X. Xu *Chinese Physics C* 41 no. 3, (2017) 030003.
- [37] I. Angeli and K. P. Marinova Atom. Data Nucl. Data Tabl. 99 no. 1, (2013)
 69–95.
- [38] A1 Collaboration, J. C. Bernauer et al. Phys. Rev. Lett. 105 (Dec, 2010)
 242001.
- [39] V. Vovchenko, B. Doenigus, and H. Stoecker *Phys. Lett.* B785 (2018)
 171–174, arXiv:1808.05245 [hep-ph].

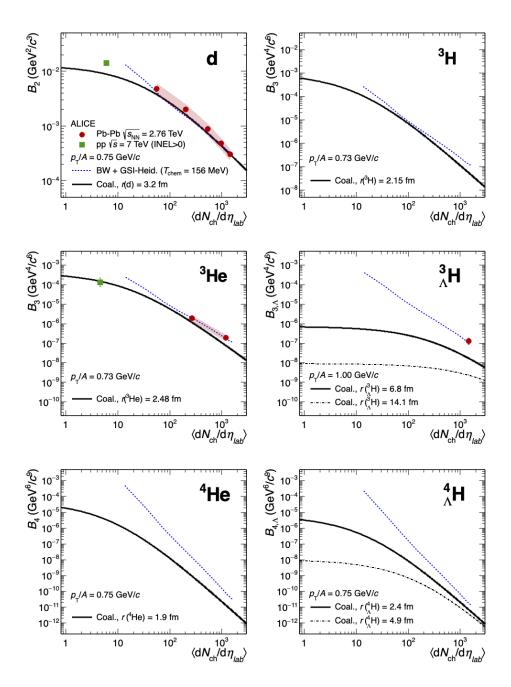


Fig. 2. (Color online) Coalescence parameter B_A as a function of the average charged particle multiplicity density for various (hyper-)nuclei, up to A=4. The coalescence calculations (continuous or dashed-dotted black lines) are compared to the thermal+blast-wave predictions (dashed blue line), as well as to pp (green square) and Pb-Pb (red circles) collision data from ALICE [9,10,13].

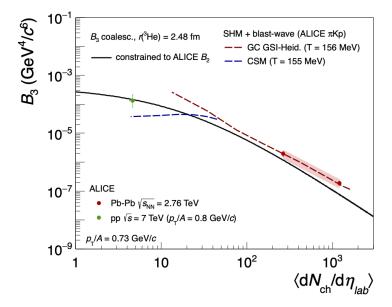


Fig. 3. (Color online) Coalescence parameter B_3 for ${}^3\mathrm{He}$ as a function of the average charged particle multiplicity density. The coalescence calculation (continuous black line) is compared to two thermal+blast-wave predictions (dashed lines), obtained by using the Grand Canonical (GC, red) [8] and Canonical Statistical Model (CSM, blue) [39] expectations for the ${}^3\mathrm{He}$ yield, respectively. ALICE data from pp (green circles) and Pb–Pb (red circles) collisions [9,13] are reported.