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Testing production scenarios for (anti-)(hyper-)nuclei with multiplicity-dependent measurements at the LHC

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The production of light anti- and hyper-nuclei provides unique observables to characterise the system created in high energy proton-proton (pp), proton-nucleus (pA) and nucleus-nucleus (AA) collisions. In particular, nuclei and hyper-nuclei are special objects with respect to non-composite hadrons (such as pions, kaons, protons, etc.), because their size is comparable to a fraction or the whole system created in the collision. Their formation is typically described within the framework of coalescence and thermal-statistical production models. In order to distinguish between the two production scenarios, we propose to measure the coalescence parameter  $B_A$  for different anti- and hyper-nuclei (that differ by mass, size and internal wave-function) as a function of the size of the particle emitting source. The latter can be controlled by performing systematic measurements of light anti- hyper- nuclei in different collision systems (pp, pA, AA) and as a function of the multiplicity of particles created in the collision. While it is often argued that the coalescence and the thermal model approach give very similar predictions for the production of light nuclei in heavy-ion collisions, our study shows that large differences can be expected for hypernuclei with extended wave-functions, as the hyper-triton. We compare the model predictions with data from the ALICE experiment and we discuss perspectives for future measurements with the upgraded detectors during the High-Luminosity LHC phase in the next decade.

# 1. Introduction: the "anti-nuclei" puzzle

The formation of light anti- and hyper-nuclei in high energy protonproton (pp), proton-nucleus (pA) and nucleus-nucleus (AA) collisions provides unique observables for the study of the system created in these reactions, and can be used to understand both the internal structure and the formation mechanisms of loosely-bound composite objects. The production of (anti-)(hyper-)nuclei in high-energy collisions is commonly described

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by following two distinct approaches: formation by nucleon coalescence at the system (kinetic) freeze-out [1-4] or thermal-statistical production at the chemical freeze-out [5,6]. Thanks to the large data samples of pp, p-Pb and 36 Pb-Pb collisions collected during the first ten years of operations of the 37 CERN Large Hadron Collider (LHC), A Large Ion Collider Experiment 38 (ALICE) Collaboration has measured the production of light nuclei and anti-nuclei at several centre-of-mass energies [7–12], thus providing a cru-40 cial experimental input and a boost to theoretical and phenomenological 41 investigations [13–19]. In small collision systems, the experimental results 42 seem to confirm the validity of the coalescence picture, with the most recent 43 multiplicity-differential measurements pointing toward a dependence of the 44 coalescence process on the volume of the particle-emitting source ("source 45 size" hereafter). In heavy-ion collisions, coalescence approaches that do not 46 take into account the source size are not able to reproduce the data. At 47 the same time, the production of light nuclei, anti-nuclei and hypertriton as measured in Pb–Pb collisions is found to be consistent with statistical-49 thermal model predictions and a non-zero deuteron elliptic flow is observed. This is surprising as (anti-)nuclei produced at chemical freeze-out are not 51 expected to survive the hadronic phase: the deuteron is a "fragile object" 52 when surrounded by the fireball created in heavy-ion collisions, because its 53 binding energy ( $B_E = 2.2 \text{ MeV}$ ) is much lower than the characteristic tem-54 peratures of the system ( $T_{chem} \approx 153 \text{ MeV}$ ,  $T_{kin} \approx 100 \text{ MeV}$ ). Moreover, the 55 cross-section for pion-induced deuteron breakup is significantly larger than 56 the typical (pseudo)-elastic cross-sections for the re-scattering of hadronic 57 resonance decay products [19–22]. These observations pose the "(anti-)nuclei 58 puzzle": how can loosely-bound composite objects survive in the dense and 59 hot fireball, freeze-out and develop collective flow like the other light-flavour 60 non-composite hadrons?

In our study [14] we have extended and combined known formalisms used to describe (anti-)(hyper-)nuclei production in order to allow, for the first time, a direct comparison of the thermal and coalescence models as well as a direct comparison to the ALICE data. Identifying the coalescence parameter  $(B_A)$  as the key observable, we present a consistent picture across different collision systems (pp, p-Pb, Pb-Pb) for light (anti-)(hyper-)nuclei with mass number A = 2, 3 and 4. We also suggest to address the open questions by looking at the production of nuclei and hyper-nuclei up to A = 4 that differ by size and properties, measured as a function of multiplicity used as a proxy for the source size. Whereas the (anti-)deuteron production has been measured multi-differentially and quite precisely with the LHC Run 1 and 2 data, the study of heavier objects with A = 3 and 4 will greatly profit from the increase in integrated luminosity foreseen at the LHC Runs 3 and 4 in all collision systems [23]. A comprehension of (anti-)nuclei production

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mechanisms is not only relevant for nuclear and hadronic physics, but has applications in astrophysics and indirect Dark Matter searches [24]. In recent years, it has been suggested that the detection of light anti-nuclei in space could provide a signature for the presence of Dark Matter in the Cosmos, see for instance [25, 26]. Anti-deuterons and  ${}^{3}\overline{\text{He}}$  might indeed be produced by coalescence of antiprotons and antineutrons coming from the annihilation of Weakly Interacting Massive Particles into Standard Model particles, for which anti-nuclei created in reactions between primary cosmic ray protons and interstellar matter (pp, pA collisions) represent a source of background.

The main features of the theoretical frameworks employed for our study are briefly summarised in Sec. 2, while we address the reader to [14] for the full details. Section 3 presents the main results and conclusions follow.

## 2. Modelling light (anti-)(hyper-)nuclei production

For our study, we consider nuclei and hyper-nuclei with mass number 89 A=2,3 and 4, whose properties are summarised in Tab. 1. Those properties 90 are the same as for their anti-matter counterparts and we assume that the 91 same formation mechanisms are at play for matter and anti-matter<sup>1</sup>. Nuclei 92 and hyper-nuclei are special objects with respect to non-composite hadrons 93 (pions, protons, etc.), because their size is comparable to a fraction or the whole system created in pp, p-Pb and Pb-Pb collisions. The size is typically 95 defined in two ways: a) as the rms of the (charge) distribution  $(\lambda_A)$ , typically 96 measured in electron scattering experiments, or b) as the size parameter of 97 the object wave-function  $(r_A)$ , typically taken as the gaussian solution of an isotropic harmonic oscillator potential in coalescence calculations. For light 99 nuclei,  $\lambda_A \approx 2$  fm. For the hyper-triton, theoretical calculations indicate 100 a charge rms radius  $\lambda_A \approx 5$  fm [27], driven by the average separation of 101 the  $\Lambda$  relative to the two other nucleons. Assuming a similar structure (e.g. 102 a s-wave interaction for a bound state of a n or a  $\Lambda$  with a deuteron), the 103 hypertriton results in a much larger object than the other non-strange nuclei 104 with A=3. A simple relation holds between  $\lambda_A$  and  $r_A$ , see [14]. In Tab. 1 105 the binding energy  $(B_E)$  is also reported. The most tightly bound nucleus 106 is  ${}^{4}\text{He}$ , whereas the most loosely bound object is  ${}^{3}_{\Lambda}\text{H}$ , that is also the largest 107 one. For the latter, we report the separation energy of the  $\Lambda$  baryon from 108 the deuteron ( $B_{\Lambda} = 0.13$  MeV). The large size and the low binding energy of the  $^3_{\Lambda}{\rm H}$  with respect to the other (hyper-)nuclei has important consequences 110 on its production, as discussed in what follows.

For brevity, in the following we refer to "nuclei" and "hyper-nuclei" but we imply both matter and anti-matter.

Mass number	Nucleus	Composition	$B_E \; ({ m MeV})$	$\operatorname*{Spin}_{J_{A}}$	(Charge) rms radius $\lambda_A^{meas}$ (fm)	Harmonic oscillator size parameter $r_A$ (fm)	Refs.
A = 2	d	pn	2.224575(9)	1	$2.1413\pm0.0025$	3.2	[28, 29]
A = 3	$^3\mathrm{H}$ $^3\mathrm{He}$ $^3_\Lambda\mathrm{H}$	pnn ppn pΛn	$\begin{array}{c} 8.4817986 \ (20) \\ 7.7180428 \ (23) \\ 0.13 \pm 0.05 \end{array}$	$1/2 \\ 1/2 \\ 1/2$	$1.755 \pm 0.086$ $1.959 \pm 0.030$ $4.9 - 10.0$	2.15 $2.48$ $6.8 - 14.1$	[30] [30] [27,31]
A=4	$^{4}_{\stackrel{4}{\Lambda} H}$ $^{4}_{\stackrel{\Lambda}{\Lambda} \Lambda} H$ $^{4}_{\stackrel{\Lambda}{\Lambda} H}$	$ppnn$ $p\Lambda nn$ $p\Lambda\Lambda n$ $pp\Lambda n$	$28.29566 (20)$ $2.04 \pm 0.04$ $0.39 - 0.51$ $2.39 \pm 0.03$	0 0 1 0	$1.6755 \pm 0.0028$ $2.0 - 3.8$ $4.2 - 7.1$ $2.0 - 3.8$	$1.9 \\ 2.4 - 4.9 \\ 5.5 - 9.4 \\ 2.4 - 4.9$	[32, 33] [27, 31] [27] [27, 31]

Table 1. Properties of nuclei and hyper-nuclei with mass number  $A \leq 4$ .  $B_E$  is the binding energy in MeV. The size parameter  $r_A$ , is chosen to approximately reproduce the measured/expected rms,  $\lambda_A{}^{meas}$  (fm). The proton rms charge radius  $\lambda_p = 0.879(8)$  fm is subtracted quadratically from the measured rms charge radius  $\lambda_A{}^{meas}$  of the nucleus  $\lambda_A = \sqrt{(\lambda_A{}^{meas})^2 - \lambda_p^2}$  to account for the finite extension of the constituents. Implicitly we assume here that  $\lambda_A \approx \lambda_p$ .

## 2.1. The coalescence approach

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In the coalescence picture, nucleons produced in the collision coalesce into nuclei if they are close in space and have similar velocities [1–3]. The coalescence probability is encoded in the coalescence parameter,  $B_A$ . Considering that at LHC energies the number of produced protons and neutrons at midrapidity as well as their momentum distributions are expected to be equal,  $B_A$  is defined as

$$E_A \frac{\mathrm{d}^3 N_A}{\mathrm{d}p_A^3} = B_A \left( E_{p,n} \frac{\mathrm{d}^3 N_{p,n}}{\mathrm{d}p_{p,n}^3} \right)^A \Big|_{\vec{p}_p = \vec{p}_n = \frac{\vec{p}_A}{A}}, \tag{1}$$

where  $p_{p,n}$  are the proton and neutron momenta and  $E_{p,n}$  their energies. Equation 1 represents also the operative definition of  $B_A$  that is used by experiments like ALICE to extract the coalescence probability starting from the measured nucleus and nucleon (proton) distributions. Starting from the model described in [3,4], we have obtained in [14] a generalised expression for  $B_A$ 

$$B_A = \frac{2J_A + 1}{2^A} \frac{1}{\sqrt{A}} \frac{1}{m_T^{A-1}} \left( \frac{2\pi}{R^2 + (\frac{r_A}{2})^2} \right)^{\frac{3}{2}(A-1)} , \qquad (2)$$

which is a function of the spin of the particle  $J_A$ , its transverse mass  $m_{\rm T}$ , its size parameter  $r_A$  and the source radius R. Very importantly, Eq. 2

takes explicitly into account the source size (R), as the coalescence probability naturally decreases for nucleons with similar momenta that are produced far apart in configuration space. Moreover, the source size is identified with the effective sub-volume of the whole system that is governed by the (momentum-dependent) homogeneity length of the interacting nucleons and experimentally accessible with Hanbury-Brown-Twiss (HBT) interferometry [3,4]. Figure 1 shows the source radius dependence of  $B_A$  for nuclei and hyper-nuclei with A=2,3 and 4 whose properties are reported in Tab. 1. For the cases in which more than one estimate for  $r_A$  is available, we have adopted the lowest value for the calculations in Fig. 1. We observe that the

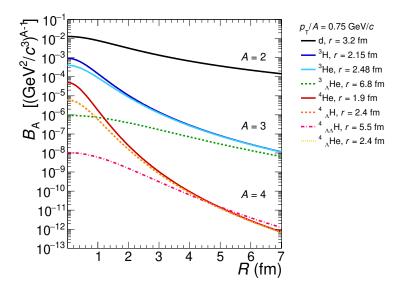


Fig. 1. (Color online) Coalescence parameter  $B_A$  as a function of the source radius R as predicted from the coalescence model (Eq. 2) for various composite objects with  $p_{\rm T}/A=0.75~{\rm GeV}/c$ . For each (hyper-)nucleus, the radius r used for the calculation is reported in the legend.

coalescence probability decreases with increasing mass number and  $B_A$  decreases with increasing volume. For a given A, the larger the object radius, the lower is  $B_A$ , as clearly visible by comparing <sup>3</sup>He and <sup>3</sup><sub>A</sub>H. For objects with same A, mass and spin (e.g. the isobars <sup>3</sup>H and <sup>3</sup>He),  $B_A$  differs only due to the different radius  $r_A$ . This difference is more relevant in small systems, because in large systems the difference between nucleus radii is much smaller than the size of the source. Incidentally, this could be experimentally verified with high precision measurements of the production of <sup>3</sup>H relative

to <sup>3</sup>He in pp collisions. The production of objects with radius larger than the source is strongly suppressed, indicating that the process is driven by the length scale defined by the object radius relative to the source radius.

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#### 2.2. Thermal + blast-wave model

Thermal-statistical models [5,6] have been successful in describing the 149 production of light (anti-)(hyper-)nuclei across a wide range of energies in 150 AA collisions, including production at the LHC. In this approach, par-151 ticles are produced from a fireball in thermal equilibrium with tempera-152 tures of  $T_{chem} \approx 156$  MeV. Particle yields are derived from the partition function assuming a Grand Canonical ensemble<sup>2</sup> and they depend only 153 154 on the mass of the particle and the temperature of chemical freeze-out, 155  $\mathrm{d}N/\mathrm{d}y \propto \exp(-m/T_{\mathrm{chem}})$ . The thermal model cannot – alone – be com-156 pared to the  $p_{\rm T}$ -dependent coalescence description because it provides only 157 predictions for  $p_{\rm T}$ -integrated yields. A thermal particle production implies 158 a Boltzmann distribution of the momenta only for a static source, which 159 is not the case of the rapidly expanding system produced in heavy-ion col-160 lisions. The thermal model (i.e. the GSI-Heidelberg implementation we 161 have considered here) needs to be complemented by a hydrodynamic description of the rapidly expanding source. To that end, we use the Blast-163 Wave model [35], which has been proven to describe reasonably well the 164 measured momentum distributions of protons [36]. Our "thermal+blast-165 wave" approach results from the combination of the two models: the  $p_{T}$ -166 dependence is modelled by the blast-wave whereas the normalisation (i.e. 167 the  $p_{\rm T}$ -integrated yield) is taken from the thermal model predictions. In 168 particular, to shape the (hyper-)nucleus transverse momentum distributions 169 we use the blast-wave parameters obtained from the simultaneous fit to 170 pion, kaon and proton spectra measured in Pb-Pb collisions by ALICE 171 for several centralities/multiplicities and reported in [36]. We therefore in-172 herit the multiplicity-dependence of the radial expansion of the system from the measurements. The object size does not enter in the formulation of 174 the blast-wave model. In our thermal+blast-wave model, for each centrality/multiplicity class, the coalescence parameter is extracted according to 176 Eq. 1 (as done for data) from the predicted (hyper-)nucleus spectrum and the proton spectrum measured by ALICE [36].

<sup>&</sup>lt;sup>2</sup> Extensions to small systems employ a canonical ensemble partition function to account for the exact conservation of quantum numbers in a finite size system, see for instance [34].

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### 3. Results and discussion

As discussed in the previous sections, the coalescence model provides an analytical expression for  $B_A$  as a function of the source size R. The thermal+blast-wave approach allows us to extract the  $B_A$  as a function of the average charged particle multiplicity density,  $\langle dN_{ch}/d\eta_{lab}\rangle$ . Data are also available as a function of multiplicity. In order to compare models and data, we therefore need to map multiplicity into source radius or viceversa. As discussed more extensively in [14], we perform this mapping based on the parameterisation  $R=0.473\,\langle {\rm d}N_{\rm ch}/{\rm d}\eta_{lab}\rangle^{1/3}$ . The value of the empirical slope parameter is obtained by tuning the parameterisation such that the measured (anti-)deuteron  $B_2$  in the most central Pb-Pb class falls onto the coalescence prediction. In this way, we constrain the coalescence volume with the more differential (anti-)deuteron data and assume that it is the same for all anti- and hyper-nuclei. In [14], we used the parameterisation to map the  $\langle dN_{ch}/d\eta_{lab}\rangle$  into R. Here, we chose to apply the inverse relation and show the comparison between  $B_A$  from models and data as a function of the experimentally accessible observable  $\langle dN_{ch}/d\eta_{lab}\rangle$ . The two choices are equivalent.

We compare  $B_A$  as obtained from the coalescence (Eq. 2) and the thermal+blast-wave model approach in Fig. 2 as a function of  $\langle dN_{\rm ch}/d\eta_{lab}\rangle$ . With respect to [14], we have extended our study up to A=4 (hyper-)nuclei. The available data from ALICE are also reported.

As an example, we have also computed  $B_A$  for <sup>3</sup>He according to the thermal+blast-wave approach for the p-Pb system, by using yield predictions from the Canonical Statistical Model from [34] and the blast-wave parameters from [37]. The result is reported in Fig. 3.

Sigh! To be finished.

4. Conclusions

Sic! To be finished.

### REFERENCES

- <sup>208</sup> [1] S. T. Butler and C. A. Pearson *Phys. Rev.* **129** (1963) 836–842.
- <sup>209</sup> [2] J. I. Kapusta *Phys. Rev.* **C21** (1980) 1301–1310.
- 210 [3] R. Scheibl and U. W. Heinz *Phys. Rev.* **C59** (1999) 1585–1602, 211 arXiv:nucl-th/9809092 [nucl-th].
- <sup>212</sup> [4] K. Blum, K. C. Y. Ng, R. Sato, and M. Takimoto *Phys. Rev.* **D96** no. 10, (2017) 103021, arXiv:1704.05431 [astro-ph.HE].
- [5] A. Andronic, P. Braun-Munzinger, J. Stachel, and H. Stocker *Physics Letters* B 697 no. 3, (2011) 203 207.
- <sup>216</sup> [6] A. Andronic, P. Braun-Munzinger, K. Redlich, and J. Stachel *Nature* **561** no. 7723, (2018) 321–330, arXiv:1710.09425 [nucl-th].
- [7] **ALICE** Collaboration, J. Adam *et al. Phys. Rev.* **C93** no. 2, (2016) 024917, arXiv:1506.08951 [nucl-ex].
- 220 [8] **ALICE** Collaboration, J. Adam *et al. Phys. Lett.* **B754** (2016) 360-372, arXiv:1506.08453 [nucl-ex].
- [9] **ALICE** Collaboration, S. Acharya *et al. Phys. Rev.* **C97** no. 2, (2018) 024615, arXiv:1709.08522 [nucl-ex].
- [10] **ALICE** Collaboration, S. Acharya *et al. Eur. Phys. J.* **C77** no. 10, (2017) 658, arXiv:1707.07304 [nucl-ex].
- 226 [11] ALICE Collaboration, M. Puccio Nucl. Phys. A982 (2019) 447–450.
- 227 [12] ALICE Collaboration, S. Acharya et al. arXiv:1902.09290 [nucl-ex].
- 228 [13] S. Mrowczynski *Acta Phys. Polon.* **B48** (2017) 707, arXiv:1607.02267 [nucl-th].
- 230 [14] F. Bellini and A. P. Kalweit arXiv:1807.05894 [hep-ph].
- [15] S. Bazak and S. Mrowczynski Mod. Phys. Lett. A33 no. 25, (2018) 1850142,
   arXiv:1802.08212 [nucl-th].
- 233 [16] W. Zhao, L. Zhu, H. Zheng, C. M. Ko, and H. Song *Phys. Rev.* **C98** no. 5, (2018) 054905, arXiv:1807.02813 [nucl-th].
- 235 [17] K.-J. Sun, C. M. Ko, and B. Doenigus arXiv:1812.05175 [nucl-th].
- $_{\mbox{\scriptsize 236}}$  [18] X. Xu and R. Rapp arXiv:1809.04024 [nucl-th].
- 237 [19] D. Oliinychenko, L.-G. Pang, H. Elfner, and V. Koch arXiv:1809.03071 [hep-ph].
- 239 [20] H. Garcilazo Phys. Rev. Lett. 48 (1982) 577–580.
- 240 [21] S. A. Bass *et al. Prog. Part. Nucl. Phys.* **41** (1998) 255–369, arXiv:nucl-th/9803035 [nucl-th].
- 242 [22] J. Schukraft Nucl. Phys. A967 (2017) 1-10, arXiv:1705.02646 [hep-ex].
- 243 [23] Z. Citron et al. in HL/HE-LHC Workshop: Workshop on the Physics of
  244 HL-LHC, and Perspectives at HE-LHC Geneva, Switzerland, June 18-20,
  245 2018. arXiv:1812.06772 [hep-ph].
- <sup>246</sup> [24] T. Aramaki *et al. Phys. Rept.* **618** (2016) 1–37, arXiv:1505.07785 [hep-ph].

- <sup>248</sup> [25] M. Cirelli, N. Fornengo, M. Taoso, and A. Vittino *JHEP* **08** (2014) 009, arXiv:1401.4017 [hep-ph].
- [26] M. Korsmeier, F. Donato, and N. Fornengo *Phys. Rev.* **D97** no. 10, (2018)
   103011, arXiv:1711.08465 [astro-ph.HE].
- [27] H. Nemura, Y. Suzuki, Y. Fujiwara, and C. Nakamoto *Prog. Theor. Phys.* 103 (2000) 929–958, arXiv:nucl-th/9912065 [nucl-th].
- <sup>254</sup> [28] C. Van Der Leun and C. Alderliesten Nucl. Phys. **A380** (1982) 261–269.
- [29] P. J. Mohr, D. B. Newell, and B. N. Taylor Rev. Mod. Phys. 88 no. 3, (2016)
   035009, arXiv:1507.07956 [physics.atom-ph].
- <sup>257</sup> [30] J. E. Purcell and C. G. Sheu *Nucl. Data Sheets* **130** (2015) 1–20.
- 258 [31] D. H. Davis Nucl. Phys. A754 (2005) 3–13.
- 259 [32] M. Wang, G. Audi, F. Kondev, W. Huang, S. Naimi, and X. Xu Chinese Physics C 41 no. 3, (2017) 030003.
- [33] I. Angeli and K. P. Marinova Atom. Data Nucl. Data Tabl. 99 no. 1, (2013)
   69–95.
- [34] V. Vovchenko, B. Doenigus, and H. Stoecker *Phys. Lett.* B785 (2018)
   171–174, arXiv:1808.05245 [hep-ph].
- 265 [35] E. Schnedermann, J. Sollfrank, and U. W. Heinz *Phys. Rev.* C48 (1993) 2462–2475, arXiv:nucl-th/9307020 [nucl-th].
- [36] ALICE Collaboration, B. Abelev et al. Phys. Rev. C88 (2013) 044910,
   arXiv:1303.0737 [hep-ex].
- 269 [37] **ALICE** Collaboration, B. B. Abelev *et al. Phys. Lett.* **B728** (2014) 25–38, 270 arXiv:1307.6796 [nucl-ex].

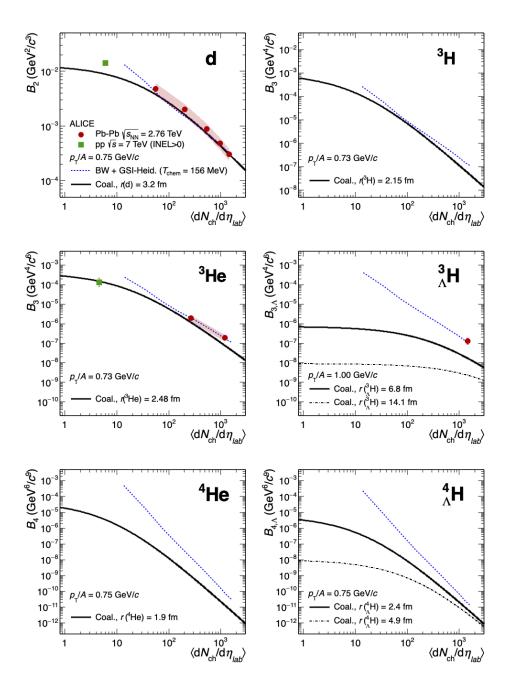


Fig. 2. (Color online) Coalescence parameter  $B_A$  as a function of the average charged particle multiplicity density for various (hyper-)nuclei, up to A=4. The coalescence calculations (continuous or dashed-dotted black lines) are compared to the thermal+blast-wave predictions (dashed blue line), as well as to pp (green square) and Pb-Pb (red circles) collision data from ALICE [7–9].

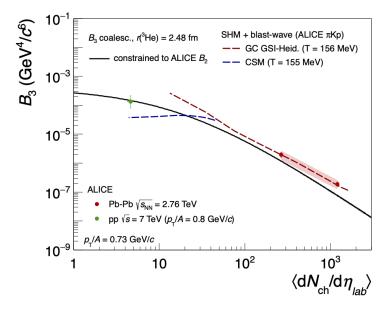


Fig. 3. (Color online) Coalescence parameter  $B_3$  for <sup>3</sup>He as a function of the average charged particle multiplicity density  $\langle dN_{\rm ch}/d\eta_{lab}\rangle$ . The coalescence calculation (continuous black line) is compared to two thermal+blast-wave predictions (dashed lines), obtained by using the Grand Canonical (GC, red) [6] and Canonical Statistical Model (CSM, blue) [34] expectations for the <sup>3</sup>He yield, respectively. ALICE data from pp (green circles) and Pb–Pb (red circles) collisions [7,9] are reported.