

hasta ahora las
fuerzas son:

$$F - \mu mg = Ma,$$

$$\mu mg = ma_2$$

instante

$$t_0 = \frac{\mu g}{\alpha} (M+m)$$

$$|F = \alpha t$$

$$\text{unidad } \alpha = \left[\text{kg} \frac{\text{m}}{\text{s}^3} \right]$$

descripción de la aceleración

$$t = [0, t_0]$$

$$\bar{a}_1 = \bar{a}_2 = \bar{a} = \frac{\alpha t}{(M+m)} \hat{e}$$

se considera como
un solo cuerpo.
de masa $(M+m)$

$$t =]b, b[\quad / \quad \lim_{b \rightarrow +\infty}$$

$$\bar{a}_1 = \frac{\alpha t}{M} \quad \text{la de masa } M, \text{ va incrementando}$$

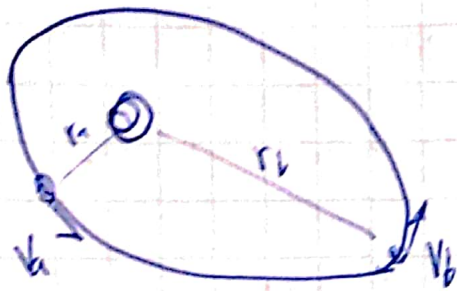
$$\bar{a}_2 = 0 \quad \text{luego del quiebre de fricción continua a velocidad}$$

$$v = \frac{\alpha t_0}{(M+m)} \cdot t_0 = (M+m)(\mu g)^2 \frac{1}{\alpha}$$

pero no acelera ya que no hay
fuerzas presentes.

Ley de inercia.

problema gravitación



$$V(r) = - \left(\frac{GM}{r} \right) \left(1 + \alpha \frac{GM}{c^2 r} \right) \quad (1)$$

encuentra que este potencial produce una órbita a precesión.

y encuentra la constante α para los datos.

la energía se conserva $E_a = E_b$

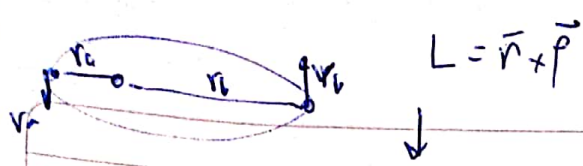
$$\frac{1}{2} m v_a^2 - \left(\frac{GM}{r_a} \right) \left(1 + \alpha \frac{GM}{c^2 r_a} \right) = \frac{1}{2} m v_b^2 - \left(\frac{GM}{r_b} \right) \left(1 + \alpha \frac{GM}{c^2 r_b} \right)$$

$$v_a^2 - v_b^2 = 2GM \left(\frac{1}{r_a} - \frac{1}{r_b} \right) + \alpha \frac{GM}{c^2} \left(\frac{1}{r_a^2} - \frac{1}{r_b^2} \right)$$

$$2\alpha \frac{GM^2}{c^2 r_a^2} - 2\alpha \frac{GM^2}{c^2 r_b^2} = 2\alpha \left(\frac{GM}{c} \right)^2 \left(\frac{1}{r_a^2} - \frac{1}{r_b^2} \right)$$

$$v_a^2 - v_b^2 = 2GM \left[\left(\frac{1}{r_a} - \frac{1}{r_b} \right) + \frac{\alpha}{c^2} GM \left(\frac{1}{r_a^2} - \frac{1}{r_b^2} \right) \right]$$

expresión de velocidades



(2)

$$V_a = L / m r_a \quad V_b = L / m r_b \quad (\text{solo en los pto de la orbita})$$

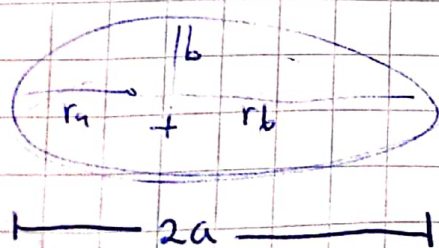
$$\frac{L^2}{m^2} \left(\frac{1}{r_a^2} - \frac{1}{r_b^2} \right) = 2GM \left[\left(\frac{1}{r_a} - \frac{1}{r_b} \right) + \frac{\alpha}{c^2} GM \left(\frac{1}{r_a^2} - \frac{1}{r_b^2} \right) \right]$$

$$\left(\frac{L^2}{m^2} - 2 \frac{\alpha}{c^2} G^2 M^2 \right) \left(\frac{1}{r_a^2} - \frac{1}{r_b^2} \right) = 2GM \left[\left(\frac{1}{r_a} - \frac{1}{r_b} \right) \right]$$

$$\therefore \left(\frac{r_b^2 - r_a^2}{(r_a r_b)^2} \right) \quad \therefore \left(\frac{r_b - r_a}{r_a r_b} \right)$$

$$\frac{L^2}{m^2} - 2 \frac{\alpha}{c^2} G^2 M^2 = 2GM \left(\frac{r_b - r_a}{r_a r_b} \right) \left(\frac{(r_a r_b)^2}{(r_b - r_a)(r_b + r_a)} \right)$$

$$\frac{L^2}{m^2} = 2GM \left(\frac{r_a \cdot r_b}{r_b + r_a} + GM \frac{\alpha}{c^2} \right)$$



$$\boxed{\frac{L^2}{m^2} = 2GM \left(\frac{b^2}{2a} + GM \frac{\alpha}{c^2} \right)}$$

$$r_a \cdot r_b = b^2$$

Via 3^{ra} ley de Kepler.

$$T = \frac{\text{area}}{\text{ratio}} \quad \text{ratio} = \frac{L}{2m} \quad \text{area elips} = \pi ab.$$

clasico $\alpha = 0$ $\frac{L^2}{m^2} = 2GM \left(\frac{b^2}{2a} \right) \quad \cdot \frac{1}{4} \rightarrow \frac{L^2}{4m^2} = \frac{GMb^2}{4a}$

$$T^2 = (\pi ab)^2 / \left(\frac{L^2}{4m^2} \right) = \frac{(\pi ab)^2 4a}{GM b^2}$$

$$T^2 = \frac{4\pi^2}{GM} a^3 \quad \text{constante de Kepler.}$$

modificado $\frac{L^2}{4m^2} = \frac{GM}{\lambda} \left(\frac{b^2}{2a} + \frac{GM\alpha}{c^2} \right)$

$$T^2 = \frac{\pi^2 a^2 b^2}{GM} \frac{1}{\left(\frac{b^2}{2a} + \frac{GM\alpha}{c^2} \right)} = \frac{2\pi^2 a^2 b^2}{GM} \frac{2ac^2}{b^2 c^2 + 2aGM\alpha}$$

$$T^2 = \frac{4\pi^2}{GM} a^3 \frac{b^2 c^2}{b^2 c^2 + 2aGM\alpha} \cdot \frac{\frac{1}{(bc)^4}}{\frac{1}{(bc)^2}} = \frac{4\pi^2}{GM} a^3 \left(\frac{1}{1 + \frac{2aGM\alpha}{(bc)^2}} \right)$$