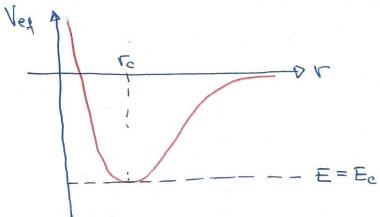
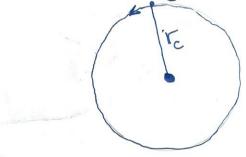
Clase 3 Lu 300821



El período en la órbita àroular: Tenemos que

En un periodo t = T, el werpo m completa un singulo $\phi = 2\pi$



$$\frac{2\pi}{2eriodo} = \frac{2\pi}{T} = \frac{L}{mr_c^2}$$

$$\frac{\Delta d}{\Delta t} = \frac{2\pi}{T} = 2\pi \cdot \frac{L}{mr_c^2}$$

$$T = 2\pi \cdot \frac{M}{L} \cdot \left(\frac{L^2}{GMm^2}\right)^2 = 2\pi \cdot \frac{M}{L} \cdot \frac{L^4}{G^2M^2m^4}$$

$$\Rightarrow T = 2\pi \cdot \frac{L^3}{G^2M^2m^3}$$

si Ec < E < 0, sun podemos user là constante de movimiento, pero

 $L = Mr^2 \phi$ Y supone que conocemos $r = r(\phi)$ ó r = r(t):

$$\frac{d\phi}{dt} = \frac{L}{mr^2}$$

$$\frac{L}{m} \int_0^{2\pi} dt = \int_0^{r^2} (\phi) d\phi$$

$$\frac{L}{m} \int_0^{2\pi} r^2(\phi) d\phi$$

o bien

$$\frac{M}{L} \cdot 2\pi = \int \frac{dt}{r^2(t)}$$

¿ como determinamos v?

Padriamos partir desole

resolver esto implica encontrar r=r(t)

$$\Rightarrow \pm \sqrt{\frac{2}{m}} dt = \int \frac{dr}{\sqrt{E-U_{ef}(r)}} + cte$$

Sin embergo, para obtener $r=r(\phi)$ debemos ocupar la otra constante de movimi ento:

$$F = \frac{dr}{dt} = \frac{d\phi}{dt} \frac{dr}{d\phi} = \frac{\phi}{dt} \frac{dr}{d\phi} = \pm \sqrt{\frac{2}{m}} \sqrt{E-U_{eff}}$$

$$\frac{L}{m} \frac{dr/r^2}{d\phi} = \pm \sqrt{\frac{2}{m}} \sqrt{E-U_{eff}}$$

$$\frac{dr/r^2}{\sqrt{E-U_{eff}}} = \pm \frac{\sqrt{2m}}{L} \sqrt{\frac{d\phi}{d\phi}} = \pm \sqrt{\frac{2m}{L}} \cdot \Delta\phi$$

$$\frac{E-U_{eff}}{\sqrt{E-U_{eff}}} = \frac{L^2}{2m} + \frac{GMm}{L}$$

$$Con \qquad M = \frac{1}{r} = \frac{dM}{2m} - \frac{dM}{2m} = \frac{dM}{L} \cdot \Delta\phi$$

$$E - \frac{L^2}{2m} \left(M^2 - 2 \frac{GMm^2}{2m} M \right)$$

$$E - \frac{L^2}{2m} \left(M^2 - 2 \frac{GMm^2}{2m} M \right)$$

$$E + \frac{G^2 L^2 M^2 M^3}{2m} - \frac{L^2}{2m} \left[M^2 - 2 \frac{GMm^2}{2m} M + \frac{GMm^2}{2m} M \right]$$

$$E + \frac{G^2 L^2 M^2 M^3}{2m} - \frac{L^2}{2m} \left[M^2 - 2 \frac{GMm^2}{2m} M + \frac{GMm^2}{2m} \right]$$

$$(E+\frac{G^{2}r^{2}w^{3}L^{2}}{2}) - \frac{L^{2}}{2mm} \left(\mu - \alpha\right)^{2}; \quad \alpha = GMm^{2}$$

$$Lwago \quad \chi = \mu - \alpha \quad \text{queolendo}$$

$$\int \frac{dx}{\sqrt{a_{1} - a_{2}x^{2}}} = \frac{1}{\sqrt{a_{2}}} \int \frac{dx}{\sqrt{\frac{a_{1}}{a_{2}} - x^{2}}}$$

$$\chi = \frac{a_{1}}{a_{2}}\cos S \quad \omega \quad dx = -\frac{a_{1}}{a_{2}}\sin \omega \quad d\omega$$

$$= -\frac{1}{\sqrt{a_{2}}} \int \frac{a_{1}}{a_{2}} \sin \omega \quad d\omega = -\sqrt{\frac{a_{1}}{a_{2}}} \int d\omega$$

$$= -\sqrt{\frac{a_{1}}{a_{2}}} \omega = -\sqrt{\frac{a_{1}}{a_{2}}} \cdot Arc \cos \left(\frac{a_{2}}{a_{1}} \left(\mu - \alpha\right)\right)$$

$$\therefore -\sqrt{\frac{a_{1}}{a_{2}}} Arc \cos \left[\frac{a_{2}}{a_{1}} \left(\frac{1 - \alpha}{r}\right)\right] = \pm \sqrt{\frac{a_{2}}{a_{1}}} \cdot \sqrt{2m} \cdot \left(\frac{d - d_{1}}{r}\right)$$

$$\Rightarrow \quad T(d) = \frac{b_{1}}{1 + b_{2}} \cos \left(\frac{b_{3}(d - d_{1})}{r}\right)$$

Entonces, para determinar of período

$$T = \frac{M}{L} \cdot b_1^2 \int \frac{d\phi}{\left(1 + b_2 \cos\left(b_3(\phi - \phi_i)\right)\right)^2}$$

TAKEA: Determinar, con sus constantes, el período de unz orbita elíptica.