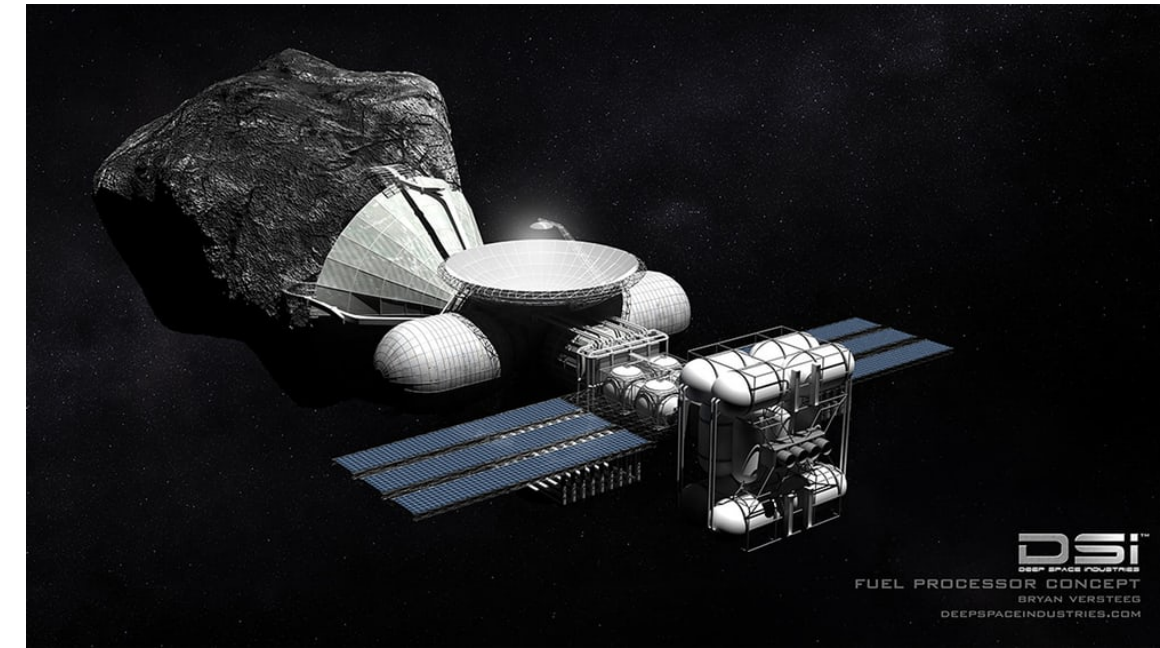


Introduction

- Asteroids and comets are of significant interest
 - Science** - Insight into early solar system formation
 - Mining** - vast quantities of useful materials
 - Impact** - high risk from hazardous Near-Earth asteroids
- Near-Earth asteroids (NEAs) are especially interesting
 - Orbit close to the Earth and are easily accessible
 - Many asteroids hold vast quantities of useful materials
 - Asteroid mining: Precious metals, propulsion fuels, semiconductors
 - Commercialization is feasible with huge amounts of possible profit
- High probability of future asteroid impacts



Asteroid Mining



Asteroid Impact

Technical Challenges

- Low-thrust propulsion systems offer innovative options
 - Electric propulsion offers much greater efficiency
 - Allows for greater velocity change with a reduced mass cost
 - Key component for long duration missions with frequent thrusting
 - Requires new methods of design
- Optimal trajectory design is complicated
 - Highly nonlinear and chaotic dynamics requires intuition by designer
 - Using low-thrust propulsion adds additional difficulties in accurately capturing the small perturbations
- Astrodynamic trajectory design typically uses direct optimal control
 - Large nonlinear programming problem inherently approximates the true optimal solution
 - High dimensionality of the solution makes it extremely computationally intensive

Gravitational Modeling

- Asteroids are extended bodies - not point masses
 - Gravity is the key force in orbital mechanics
 - An accurate representation of gravity is critical to accurate and realistic analysis
- Spherical Harmonic approach is popular but not ideal
 - Model is only valid outside of circumscribing sphere
 - Composed of an infinite series - always results in an approximation
 - Model will diverge when close to the surface and is not ideal for landing missions
- Polyhedron Gravitational model** used to represent the asteroid
 - Gravity is a function of the shape model
 - Globally valid and closed-form analytical solution for gravity
 - Exact potential assumes a constant density assumption
 - Accuracy is only dependent on the shape

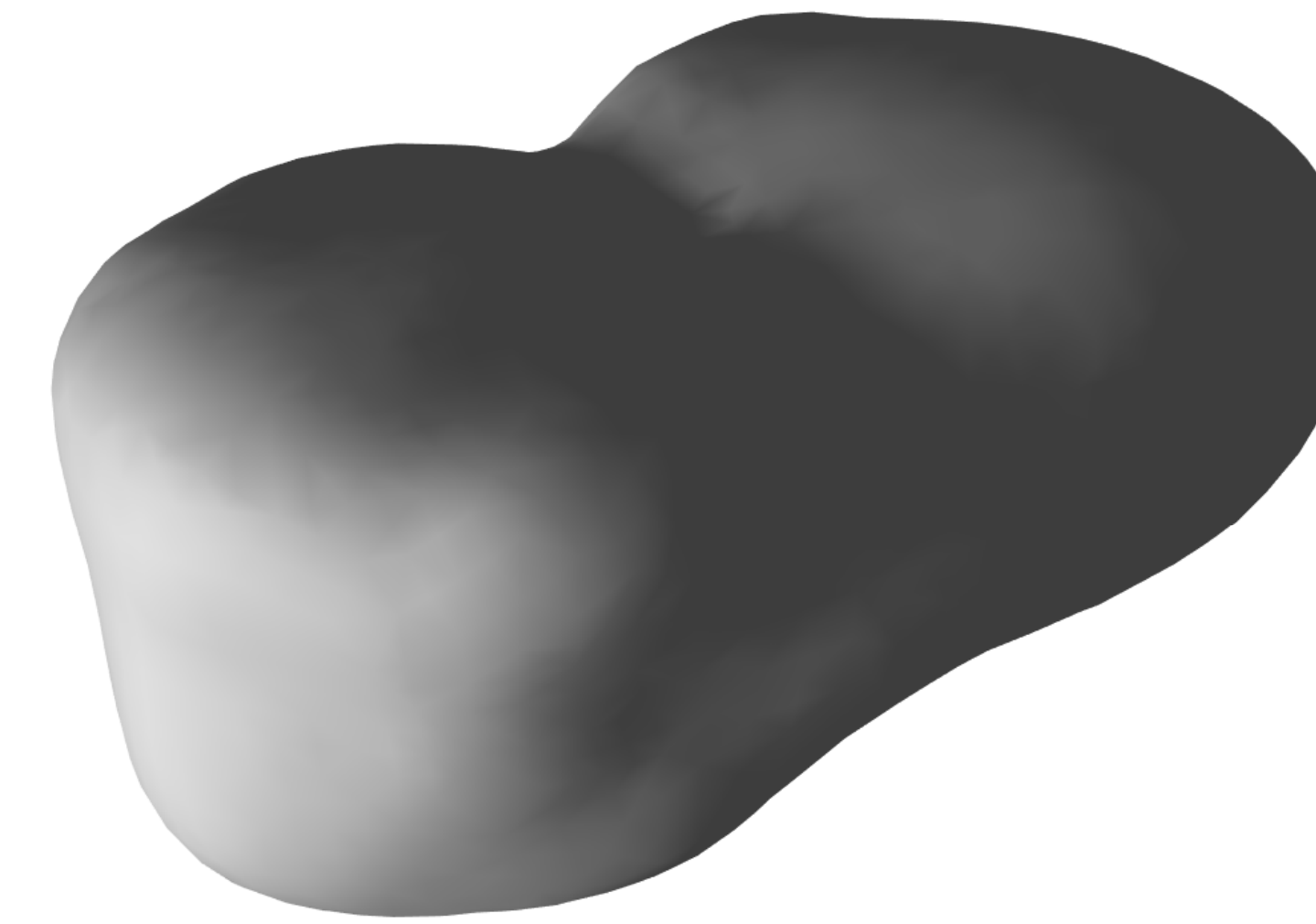
$$U(\mathbf{r}) = \frac{1}{2}G\sigma \sum_{e \in \text{edges}} \mathbf{r}_e \cdot \mathbf{E}_e \cdot \mathbf{r}_e \cdot L_e - \frac{1}{2}G\sigma \sum_{f \in \text{faces}} \mathbf{r}_f \cdot \mathbf{F}_f \cdot \mathbf{r}_f \cdot \omega_f$$

Dynamics about the asteroid 4769 Castalia

- Dynamics are very similar to the famous three-body problem

$$\begin{bmatrix} \dot{\mathbf{r}} \\ \dot{\mathbf{v}} \end{bmatrix} = \begin{bmatrix} \mathbf{g}(\mathbf{r}) + \mathbf{h}(\mathbf{v}) + \mathbf{u} \end{bmatrix}$$
- Huge history of analytical tools allow for great insight into the dynamics
- Analytical insight is critical to understanding the free motion around an asteroid
 - We require an accurate understanding of the motion under the influence of gravity alone
 - Efficient use of the limited onboard fuel is dependent on exploiting the natural dynamics of the asteroid environment
- Jacobi Integral - single constant of motion which bounds the feasible regions in terms of "energy"

$$J(\mathbf{r}, \mathbf{v}) = \frac{1}{2}\omega^2(x^2 + y^2) + U(\mathbf{r}) - \frac{1}{2}(\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$$



Asteroid 4769 Castalia

Reachability on the Poincaré section

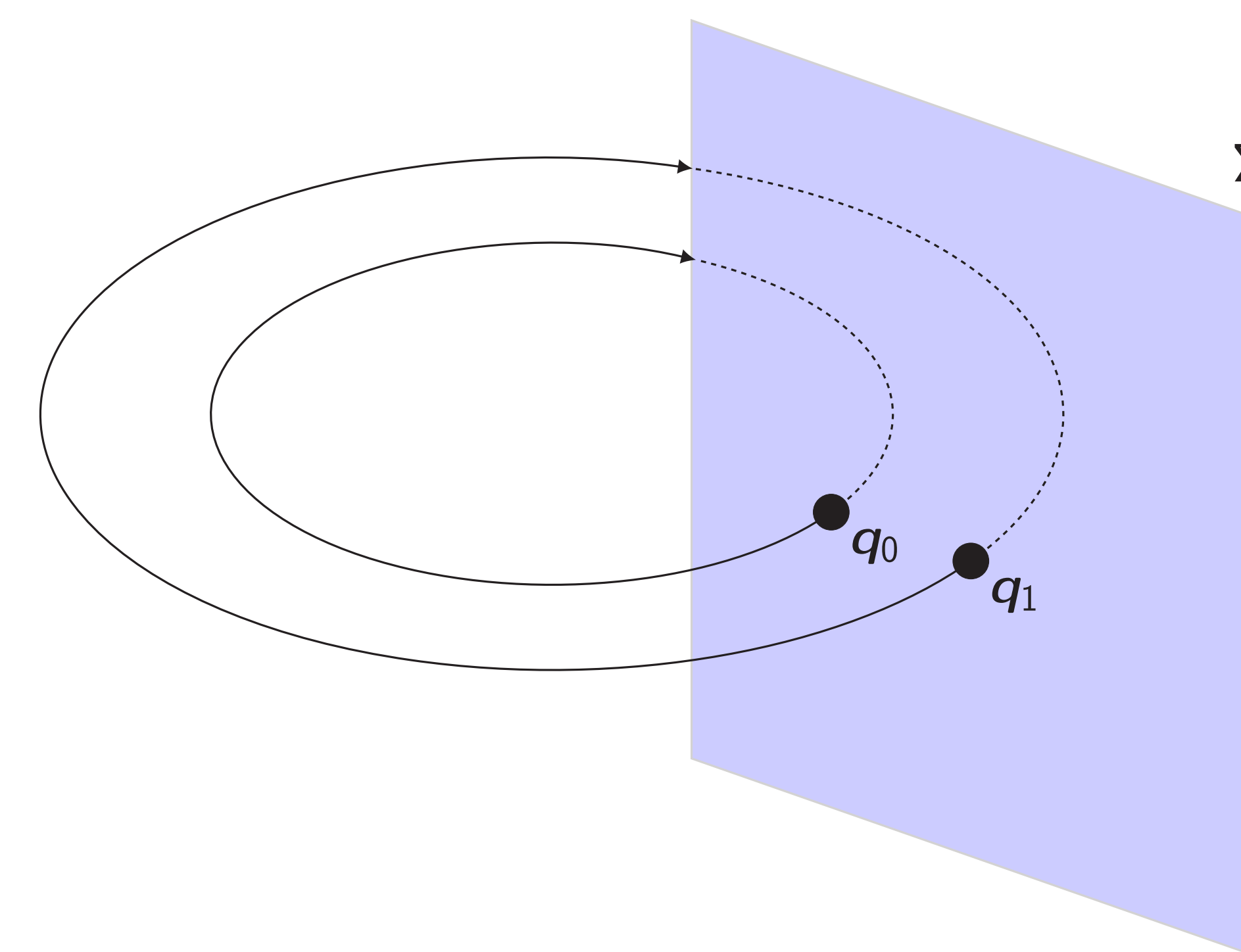
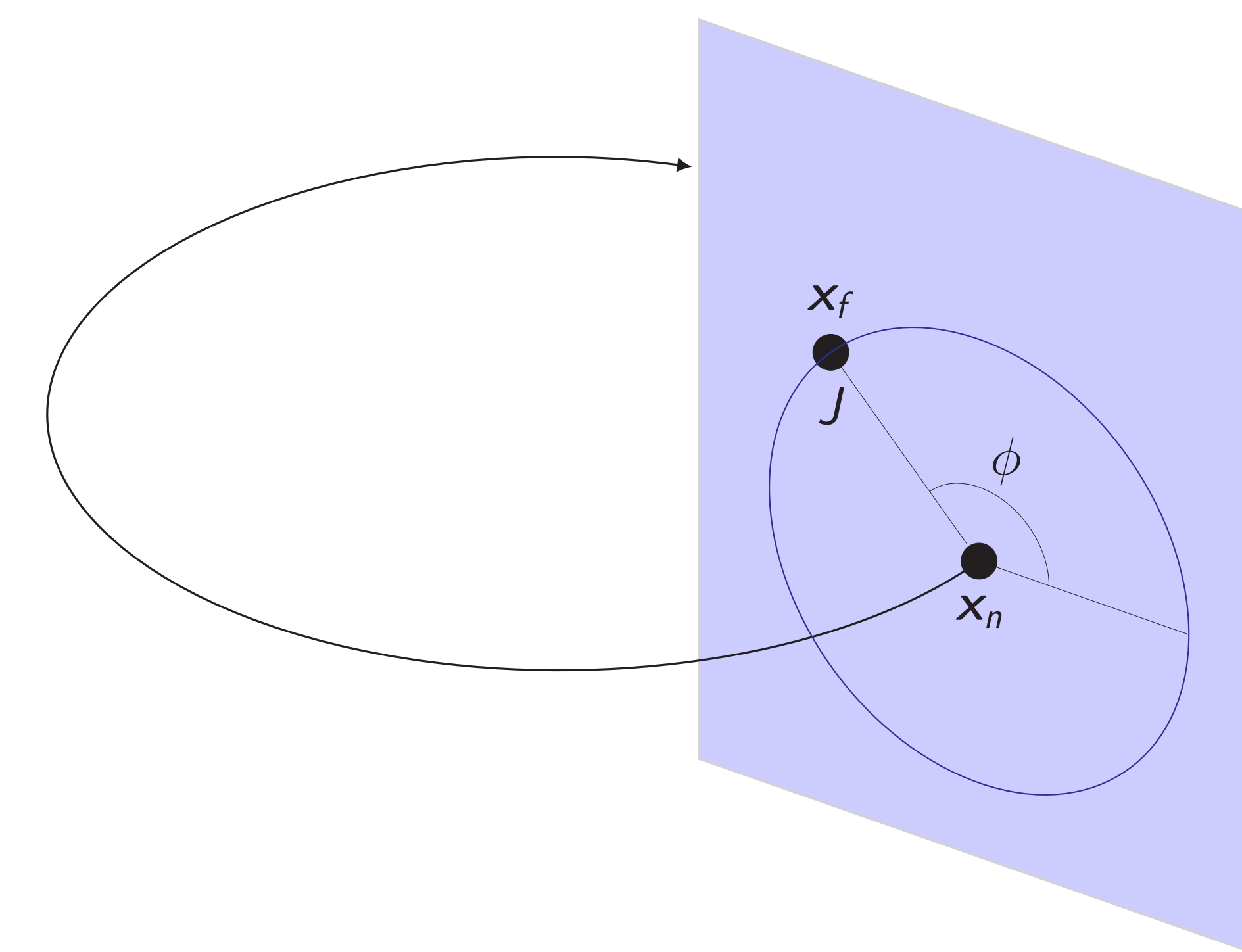


Diagram of the Poincaré map



Reachability set on a Poincaré section

- Poincaré** map is a useful tool for the analysis of dynamical systems
 - Rather than considering the entire state (6D position and velocity) we simply investigate the intersections with a lower dimensional space
 - This reduces the complexity of analysing the dynamics and allows for visualization of highly complex dynamic interactions
- A periodic orbit on the Poincaré map is identified by fixed points q_0 and q_1
- Using the low-thrust propulsion system of the spacecraft we can enlarge the space that is achievable
 - Reachability Set** - the set of states which are attainable subject to the constraints of the system
 - The thruster of the spacecraft is used to design a transfer trajectory by repeatedly maximizing the reachability set
- Optimal Control** is used to calculate the reachability set
 - A series of constraints ensures that we will intersect the Poincaré section

$$m_1 = y = 0$$

$$m_2 = (\sin \phi_{1_d}) (x_1^2 + x_2^2 + x_3^2 + x_4^2) - x_1^2 = 0$$

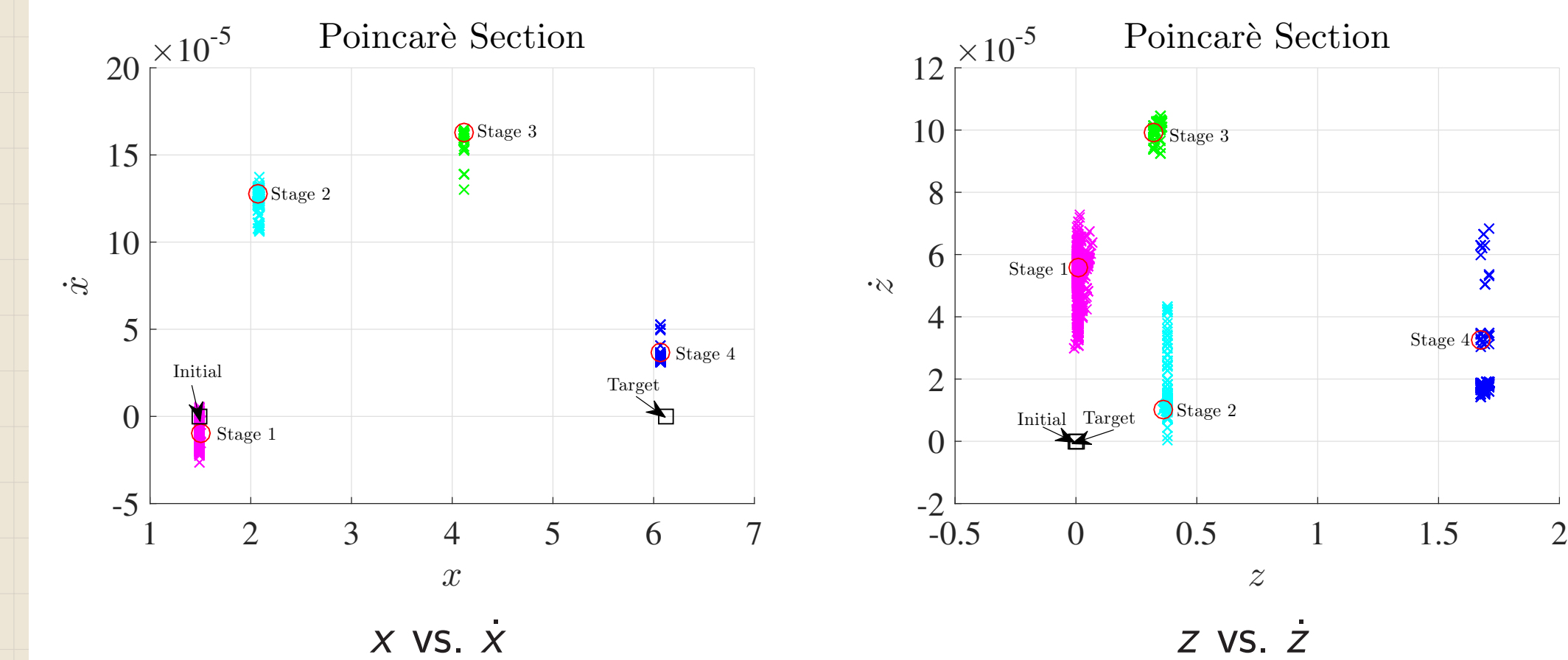
$$m_3 = (\sin \phi_{2_d}) (x_2^2 + x_3^2 + x_4^2) - x_2^2 = 0$$

$$m_4 = (\sin \phi_{3_d}) \left(2x_3^2 + 2x_3 \sqrt{x_4^2 + 2x_4^2} \right) - x_3 - \sqrt{x_4^2 + x_3^2} = 0$$
- The goal is to maximize the distance on the Poincaré section using the effect of the low thrust propulsion
- The thruster is limited to ensure a realistic scenario

$$c(\mathbf{u}) = \mathbf{u}^T \mathbf{u} - u_m^2 \leq 0$$

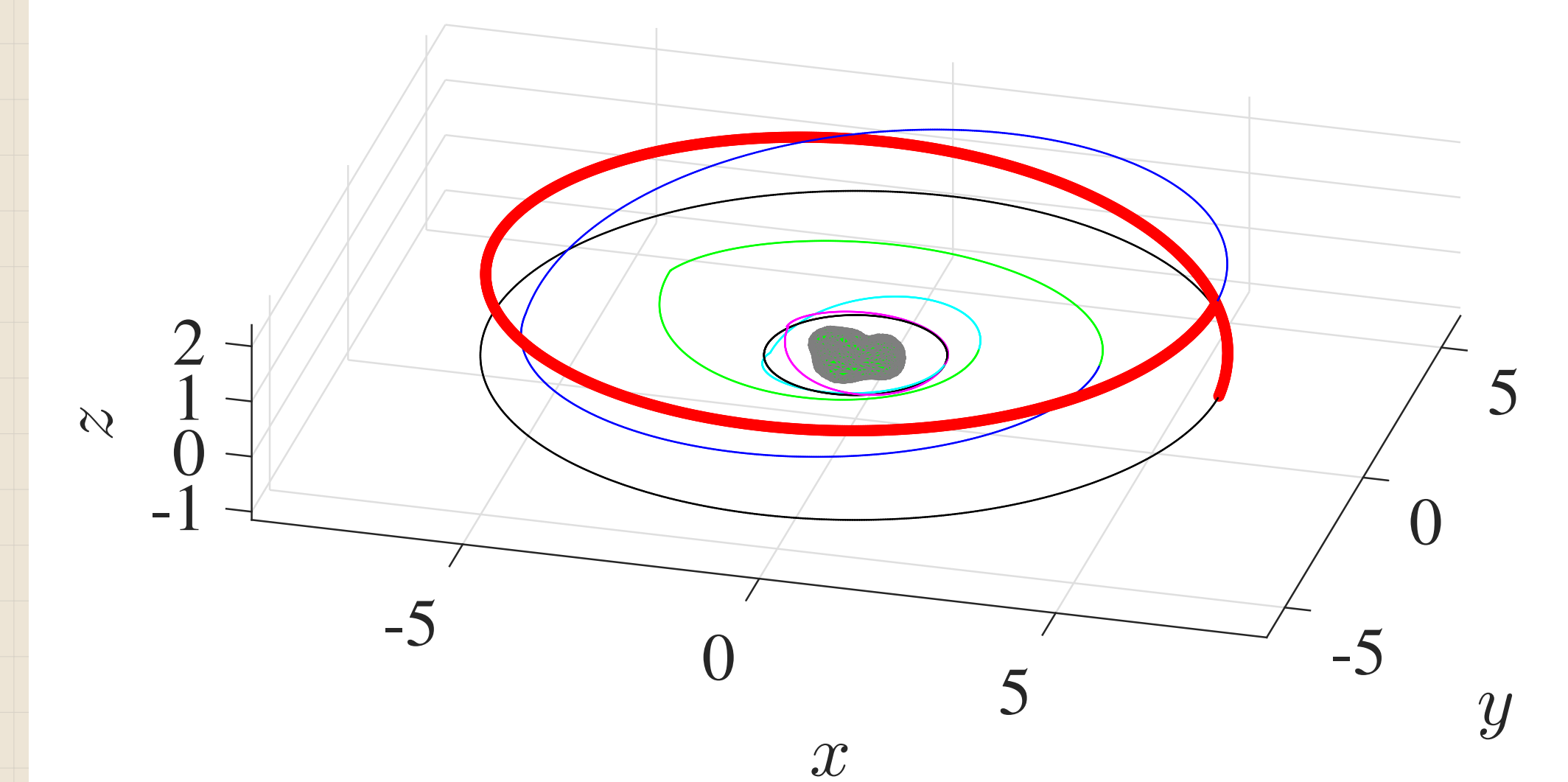
Simulation Results

- A transfer around the asteroid 4769 Castalia is completed
 - Discovered in 1989, Castalia is a potentially hazardous asteroid and passes close to the Earth
 - In 1989, Castalia passed close enough to the Earth to allow for high resolution radar imagery
- We wish to transfer a spacecraft between two periodic orbits of 4769 Castalia
 - The thruster is modeled after current state of the art low thrust propulsion systems
 - Combining multiple iterations of the **reachability** computation allows for general transfers
- Transfer is computed by combining four iterations of the reachability set
 - Each iteration of the reachability set enlarges the achievable states
 - From the reachability set we chose a direction which gets us closest to the target



Poincaré sections

Trajectory



Transfer Trajectory

Conclusions

- Orbital transfer around an asteroid using **reachability sets**
- Automatically gain insight into the feasible region of motion for the spacecraft
- Future work will extend this principle to landing trajectories
 - Irregular shape of asteroids requires innovative techniques for controlling both position and orientation