

Periodic solutions for a Sitnikov restricted $n + 1$ -body problem with primaries in rigid motion

Gastón Beltritti *

Dpto. de Matemática, Facultad de Ciencias Exactas Físico-Químicas y Naturales
Universidad Nacional de Río Cuarto (5800) Río Cuarto, Córdoba, Argentina,
`gbeltritti@exa.unrc.edu.ar`

Fernando D. Mazzone †

Dpto. de Matemática, Facultad de Ciencias Exactas, Físico-Químicas y Naturales
Universidad Nacional de Río Cuarto
(5800) Río Cuarto, Córdoba, Argentina,
`fmazzone@exa.unrc.edu.ar`

Martina G. Oviedo ‡

Dpto. de Matemática, Facultad de Ciencias Exactas, Físico-Químicas y Naturales
Universidad Nacional de Río Cuarto
(5800) Río Cuarto, Córdoba, Argentina,
`martinagoviedo@gmail.com`

Abstract

1 Introduction

In this paper we study the following restricted Newtonian $n + 1$ -body problem P (see figure 1):

- P_1 We have n primary bodies of masses m_1, \dots, m_n and an additional massless body.
- P_2 The primary bodies are in a central configuration rigid motion (see [13, Section 2.9]). This motion is periodic and it is carried out in a plane Π .
- P_3 The massless particle is moving on the perpendicular line to Π passing through the center of masses.

*SECyT-UNRC and CONICET

†SECyT-UNRC, FCEyN-UNLPam

‡SECyT-UNRC, CIN

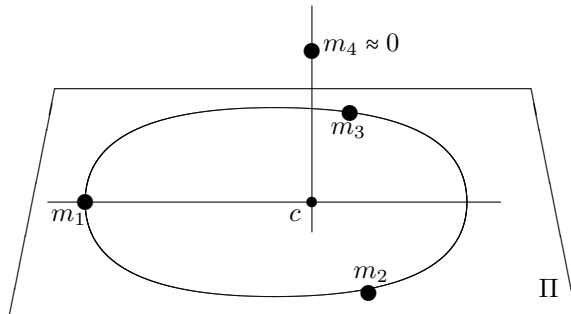


Figure 1: Four-body problem with three primaries

Problems like the one presented have been extensively discussed in the literature. In [23] K. Sitnikov considered the problem of two body in a Keplerian motion and a massless particle moving in the perpendicular line to the orbital plane passing through the center of masses. Sitnikov obtained deep results about existence of solutions, some of them periodic (see [17, III(5)]). Since then many other authors have studied Sitnikov problem, for instance Liu, Zhou, and Sun [11], Hagel and Trenkler [6], Dvorak [5], Dankowicz and Holmes [4], Llibre, Meyer and Soler [12], Chesley [3], Jiménez-Lara and Escalona-Buendía [8], Llibre and Ortega [14], Pérez, Jiménez and Lacomba [21].

Problems like the Sitnikov problem for four bodies were addressed more recently. In [24] Soulis, Papadakis and Bountis studied existence, linear stability and bifurcations for a problem similar to P . They considered a Lagrangian equilateral triangle configuration for the primaries bodies, which were supposed to have the same mass $m_1 = m_2 = m_3$. In [2] Papadakis and Bountis extend of results of [24] to $N-1$ primaries ($N \geq 4$) in a polygonal equal mass configuration. Later in [20], Pandey and Ahmad extend the analysis started in [24] to the case when the primaries are oblate (not mass points). In [27], Zhao and Zhang proved existence, by means of a variational approach, of periodic solutions for a problem similar to the one dealt in [24]. In [10] Li, Zhang and Zhao studied a special type of restricted circular $N+1$ -body problem with equal masses for the primaries in a regular polygon configuration.

In the present paper we extend the analysis in [27] to the case of a collinear central configuration for the primaries.

2 Preliminaries and Main Results

We start considering n mass points, $n > 2$, of masses m_1, \dots, m_n moving in a Euclidean 3-dimensional space according to Newton's laws of motion. We assume that $q_1(t), \dots, q_n(t)$ are the coordinates (column vectors) of the bodies in some inertial Cartesian coordinate system. We denote by $r_{ij} = |q_i - q_j|$ the Euclidean distance between q_i and q_j . We can suppose, without any loss of generality, that

the center of mass $c := \sum_j m_j q_j / M$ ($M := \sum_j m_j$) is fixed at the origin ($c = 0$).

We assume that these bodies are in a *rigid motion*. We recall that a *rigid motion*, is a solution of motions equations with r_{ij} constant. It is known (see [26]) that a rigid motion is performed in a plane Π . We assume that Π is the plane determined by the first two coordinates axes. Then a rigid motion has the form (see [13])

$$q_j(t) = Q(\nu t) q_j^0,$$

where

$$Q(\nu t) = \begin{pmatrix} \cos(\nu t) & -\sin(\nu t) & 0 \\ \sin(\nu t) & \cos(\nu t) & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

and $q_j^0 \in \Pi$, $j = 1, \dots, n$ are vectors in a planar *central configuration* (CC) in \mathbb{R}^3 , i.e. there exists $\lambda \in \mathbb{R}$ such that

$$\nabla_j U(q_1^0, \dots, q_n^0) + \lambda m_j q_j^0 = 0, \quad j = 1, \dots, n.$$

where the *potential function* U is defined by:

$$U(x) = \sum_{i < j} \frac{m_i m_j}{r_{ij}}, \quad (1)$$

and ∇_j denotes the 3-dimensional partial gradient with respect to q_j . According to [13, Eq. (2.16)] we have $\nu^2 = \lambda$. Then the primaries bodies perform a periodic motion with period $T := 2\pi/\nu$.

We suppose that we have a massless particle with coordinates $q(t) = (x(t), y(t), z(t)) \in \mathbb{R}^3$. This particle does not disturb the rigid motion of primaries. We want to find conditions under which this particle perform a T -periodic motion on the z axis of coordinates.

The particle q satisfies the Newtonian equations of motion

$$\ddot{q} = \sum_{i=1}^n \frac{m_i (q_i - q)}{|q_i - q|^3}, \quad (2)$$

A necessary and sufficient condition for the particle has a non-trivial motion along the z axis at all times is that the horizontal component of the gravitational force originated by the primaries is null. The following theorem characterize the initial configuration (q_1^0, \dots, q_n^0) with this property.

Theorem 2.1. *There exists a non-stationary solution of (2) with $x(t) = y(t) = 0$ if and only if q_1^0, \dots, q_n^0 satisfy that for any $r > 0$, such that the set*

$$F_r := \{i : |q_i^0| = r\}$$

is non empty, that

$$\sum_{i \in F_r} m_i q_i^0 = 0. \quad (3)$$

i.e. every maximal set of bodies which are equidistant from origin has center of mass equal to 0.

We will show by means of an example that non-stationary assumption is necessary.

If condition (3) holds then equation (2) is reduced to

$$\ddot{z} = - \sum_{i=1}^n \frac{m_i z}{(s_i^2 + z^2)^{\frac{3}{2}}}, \quad (4)$$

with $s_i = |q_i^0|$.

We note that in order to get collisionless solutions of problem P we need that no primary body is located in the center of mass. We say that a CC is *admissible* if it is non-collisional and satisfies (3). In the following theorem we characterize all admissible configurations with 3 or 4 bodies.

Theorem 2.2. *The only 3-body admissible CC is the equilateral triangle with three equal masses. In the case of 4-body, an admissible CC has two pairs of equal masses and satisfies some of the following properties: it is collinear and symmetric around the center of mass or it is a rhombus with the equal masses in opposite vertices, being the minor masses near from origin. In the particular case that the four masses lie in a common circle with center of mass at the origin the CC is a equal mass square.*

If the equation (4) has a kT -periodic solution, where T is the period for the primaries and k is a positive integer, we say that the solution is kT -synchronous.

The following theorem characterize all the possible period for the massless particle.

Theorem 2.3. *We assume that q_1^0, \dots, q_n^0 is a admissible CC.*

1. *A non-trivial solution of the equation (4) is either periodic or its norm tends to infinity when t goes to infinity. The escape velocity for initial condition $z(0)$ is*

$$V_{Esc} = \left(\sum_{i=1}^n \frac{2m_i}{\sqrt{s_i^2 + z(0)^2}} \right)^{\frac{1}{2}}. \quad (5)$$

2. *A necessary and sufficient condition for that equation (4) has non trivial T_0 -periodic solutions is that*

$$\frac{4\pi^2}{T_0^2} < \sum_{i=1}^n \frac{m_i}{s_i^3} \quad (6)$$

3. *In particular, there is a T -synchronous solution if and only if*

$$\sum_{i < j} \frac{m_i m_j}{r_{ij}} < \left(\sum_{i=1}^n \frac{m_i}{s_i^3} \right) \left(\sum_{i=1}^n m_i s_i^2 \right). \quad (7)$$

We observe that for all admissible CC there exists kT synchronous solution (and therefore a periodic solution of the complete $n + 1$ -problem) when $k \in \mathbb{N}$ is large enough.

The sufficiency of the condition $n \leq 472$ in the following corollary was proved in [10].

Corollary 2.4. *We suppose that q_1^0, \dots, q_n^0 is the equal masses regular polygon configuration (this is an admissible CC). Then there exists a synchronous solution if and only if $n \leq 472$.*

Our next objective is to verify that condition (7) is satisfied for all admissible CC of 3-body or 4-body. In virtue of Corollary 2.4 and Theorem 2.2 we rest prove that condition (7) is satisfied for the symmetric collinear 4-body CC, and for the CC forming a rhombus with equal masses in opposite vertices. Let's call these central configurations CCcl and CCr respectively.

Theorem 2.5. *The central configurations CCcl and CCr satisfy condition (7).*

Corollary 2.6. *For all admissible CC of 3-body or 4-body problem P has a T -synchronous solution.*

3 Proofs

Lemma 3.1. *For $c > 0$ we define the function $y_c(t) := (c + t)^{-3/2}$. If $0 < t_1 < t_2 < \dots < t_k$ then the functions $y_j(t) := y_{t_j}(t)$ are linearly independent on each open interval $\mathcal{I} \subset \mathbb{R}^+$.*

Proof. It is sufficient to prove that Wronskian

$$W := W(y_1, \dots, y_k)(t) = \det \begin{pmatrix} y_1 & \dots & y_k \\ \frac{dy_1}{dt} & \dots & \frac{dy_k}{dt} \\ \vdots & \ddots & \vdots \\ \frac{d^{k-1}y_1}{dt^{k-1}} & \dots & \frac{d^{k-1}y_k}{dt^{k-1}} \end{pmatrix}$$

is not null on \mathcal{I} .

Using induction is easy to show that

$$\frac{d^i y_c}{dt^i} = \beta_i y_c^{\frac{2i+3}{3}}, \quad \text{for some } \beta_i \neq 0, \text{ and for all } i = 1, \dots \quad (8)$$

Fix any $t \in I$. Then, according to (8) and writing $\lambda_j := (t + t_j)^{-1}$, we have

$$\begin{aligned}
W(t) &= \det \begin{pmatrix} \lambda_1^{3/2} & \lambda_2^{3/2} & \dots & \lambda_k^{3/2} \\ \beta_1 \lambda_1^{5/2} & \beta_1 \lambda_2^{5/2} & \dots & \beta_1 \lambda_k^{5/2} \\ \vdots & \vdots & \ddots & \vdots \\ \beta_{k-1} \lambda_1^{k+1/2} & \beta_{k-1} \lambda_2^{k+1/2} & \dots & \beta_{k-1} \lambda_k^{k+1/2} \end{pmatrix} \\
&= \beta_1 \beta_2 \dots \beta_{k-1} \lambda_1^{3/2} \lambda_2^{3/2} \dots \lambda_k^{3/2} \det \begin{pmatrix} 1 & 1 & \dots & 1 \\ \lambda_1 & \lambda_2 & \dots & \lambda_k \\ \vdots & \vdots & \ddots & \vdots \\ \lambda_1^{k-1} & \lambda_2^{k-1} & \dots & \lambda_k^{k-1} \end{pmatrix} \\
&= \beta_1 \beta_2 \dots \beta_{k-1} \lambda_1^{3/2} \lambda_2^{3/2} \dots \lambda_k^{3/2} \prod_{1 \leq i < j \leq n} (\lambda_j - \lambda_i),
\end{aligned}$$

where the last equality follows of the well known Vandermonde determinant identity. Therefore $W \neq 0$ if and only if $\lambda_i \neq \lambda_j$, $i \neq j$, which in turn is equivalent to $t_i \neq t_j$, $i \neq j$. \square

Proof Theorem 2.1. We use a rotating coordinate system where the primaries are fixed. Concretely we put

$$\xi = Q(-\nu t)q.$$

In this system the motion equations are

$$\ddot{\xi} + 2\nu B \dot{\xi} + \nu^2 C \xi = \sum_{i=1}^n \frac{m_i (q_i^0 - \xi)}{|q_i^0 - \xi|^3}, \quad (9)$$

where

$$B := \begin{pmatrix} J & 0_{2 \times 1} \\ 0_{1 \times 2} & 0 \end{pmatrix}, \quad J := \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} \quad \text{and} \quad C = \begin{pmatrix} -I_2 & 0_{2 \times 1} \\ 0_{1 \times 2} & 0 \end{pmatrix},$$

where $0_{n \times m}$ and I_n denote the null $n \times m$ matrix and the identity $n \times n$ matrix respectively. Assuming that the masless particle is moving on the z -axis then $\xi = q = (0, 0, z)$ and the Coriolis and centrifugal forces, $2\nu B \dot{\xi}$ and $\nu^2 C \xi$ respectively, are null. Therefore, taking account in the first two equation in (9) and identifying the vectors q_i^0 , $i = 1, \dots, n$ with vectors in \mathbb{R}^2 , we have

$$\sum_{i=1}^n \frac{m_i q_i^0}{|q_i^0 - \xi|^3} = 0.$$

Let $D = \{|q_i^0| : i = 1, \dots, n\}$. Suppose that $D = \{r_1, \dots, r_k\}$, with $r_i \neq r_j$ for $i \neq j$, and $\{1, \dots, n\} = F_1 \cup \dots \cup F_k$, where if $i \in F_j$ then $|q_i^0| = r_j$. Then

$$\sum_{j=1}^k \left\{ \frac{1}{(r_j^2 + z^2)^{3/2}} \sum_{i \in F_j} m_i q_i^0 \right\} = 0.$$

Since we are considering a non-stationary solution, we have that $z(t)$ is not constant. Therefore there exists an interval $\mathcal{I} \subset \mathbb{R}^+$ where

$$\sum_{j=1}^k \left\{ \frac{1}{(r_j^2 + s)^{3/2}} \sum_{i \in F_j} m_i q_i^0 \right\} = 0, \quad s \in \mathcal{I}.$$

Then, according to Lemma 3.1, we obtain (3).

If (3) is satisfied then the force field F acting on the masless particle carries the z axis in itself. Therefore, from the existence and uniqueness theorem and other elementary properties of system of ODEs we obtain a solution of (2) with $x(t) = y(t) = 0$. \square

Proof Proposition ??. Some of the following calculations were made with symbolic math software. Let's consider a particular Euler's linear central configuration formed by three collinear bodies of mass $m_1 = 4 - \mu$, $m_2 = 2 + \mu$, $m_3 = 1$, where $0 < \mu < 1$, and $q_1 = 0$, $q_2 = 1$ and $q_3 = 1 + r$ their respective positions, with r such that

$$p(r, \mu) = 0,$$

where $p(r, \mu) = 6r^5 + (-\mu + 16)r^4 + (-2\mu + 14)r^3 + (-\mu - 5)r^2 + (-2\mu - 7)r - \mu - 3$. Note that as $p(0, \mu) = -\mu - 3$ and $P(1, \mu) = -7\mu + 21$ then, for all $0 < \mu < 1$, $r \in (0, 1)$. In this case the center of mass C is equal to $\frac{\mu}{7} + \frac{r}{7} + \frac{3}{7}$, so $C \in (0, 1)$.

We denote with x the point between 0 and 1 where the sum of the forces that the primary bodies make on a massless particle located in that position is equal to zero. For this value of x we have to $f(x) = 0$, where $f(x) = -\frac{4-\mu}{x^2} + \frac{\mu+2}{(-x+1)^2} + \frac{1}{(r-x+1)^2}$. Note that the left side of the previous equation is an increasing function that tends to $-\infty$ when x goes to 0, and tends to $+\infty$ when x goes to 1, so there is a unique point $x \in (0, 1)$, such that the equality holds.

Then, if we want to have a trivial solution of the problem (2) then necessarily C has to be equal to x . Let's see that there exists $\mu \in (0, 1)$ such that $C = x$, i.e. $f(C) = 0$. For this purpose, since C is a continuous function with respect to μ , we need to see that there exists a value $\mu_1 \in (0, 1)$ such that $f(C) < 0$ and $\mu_2 \in (0, 1)$ such that $f(C) > 0$. The function $f(x)$ can be factorized as

$$f(x) = \frac{Nf(x)}{Df(x)},$$

where $Nf(x) = 2\mu r^2 x^2 - 2\mu r^2 x + \mu r^2 - 4\mu r x^3 + 8\mu r x^2 - 6\mu r x + 2\mu r + 2\mu x^4 - 6\mu x^3 + 7\mu x^2 - 4\mu x + \mu - 2r^2 x^2 + 8r^2 x - 4r^2 + 4r x^3 - 20r x^2 + 24r x - 8r - x^4 + 10x^3 - 21x^2 + 16x - 4$ and $Df(x) = x^2 (x - 1)^2 (r - x + 1)^2$. Note that $Df(x) > 0$ for all $x \in (0, 1)$. If we consider $\mu = 0$ and compute $Nf(C)$ we have that

$$Nf(C) = \frac{r^4}{2401} + \frac{1514r^3}{2401} + \frac{2245r^2}{2401} + \frac{1110r}{2401} + \frac{333}{2401} > 0,$$

on the other, hand if $\mu = 1$ then

$$Nf(C) = -\frac{71r^4}{2401} + \frac{1486r^3}{2401} + \frac{401r^2}{2401} - \frac{1480r}{2401} - \frac{592}{2401} < 0,$$

because $0 < r < 1$. □

Proof Theorem 2.2. For the case of 3-bodies, we note that the Theorem 2.1 and the fact that the center of masses is an excluded position imply that if F_r is not empty then $\#F_r \geq 2$. Hence an admissible 3-body CC consists of three equidistant bodies from the origin. Therefore, it must be the Lagrangian equilateral triangle configuration. Now, equation (3) implies that every bodies has the same mass.

In the case of 4-bodies, we have again that $\#F_r \geq 2$. We consider two cases, the first one $|q_1| \neq |q_2|$. Therefore we can suppose that $|q_1| = |q_3|$ and $|q_2| = |q_4|$. Now (3) implies that $m_1 = m_3$ and $m_2 = m_4$. We divide the plane in two cones by means of the line L joining q_1 and $q_3 = -q_1$ together with its perpendicular bisector M . From the Perpendicular Bisector Theorem (see [16]), we have that if q_2 is in a open cone, then q_4 is in the other one. But on the other hand (3) implies $q_2 = -q_4$, which is a contradiction. Then $q_2, q_4 \in M$ or $q_2, q_4 \in L$ (since $q_2 = -q_4$ the case $q_2 \in L$ and $q_4 \in M$ is impossible). In the first case the CC is a rhombus with the larger masses closer to the origin (see [22]). The second case we have a collinear CC which is also symmetric by (3). It remains to discuss the case of $|q_1| = |q_2| = |q_3| = |q_4|$. In this situation, in [7] was proved that the configuration is the equal mass square. □

Proof Theorem 2.3. The second order equation (4) is consevative, therefore solutions conserve the energy

$$E := \frac{|v|^2}{2} - \sum_{i=1}^n \frac{m_i}{(s_i^2 + z^2)^{\frac{1}{2}}}, \quad (10)$$

i.e. $E(z(t), \dot{z}(t))$ is constant. An elementary analysis shows that the energy leves curves are non bounded when $E \geq 0$. Moreover for a fix energy E we have two branch for velocity

$$v = v(E, z) = \pm \left(2E + 2 \sum_{i=1}^n \frac{m_i}{\sqrt{s_i^2 + z^2}} \right)^{\frac{1}{2}}.$$

We note that, for $z, v \geq 0$, v is decreasing with respect to z . And energy curve cut the v -axis at $v_E = (2E + 2 \sum_{i=1}^n m_i s_i^{-1})^{\frac{1}{2}}$. An energy curve cut the z -axis, only in the case that $E < 0$, at z_E which satisfies $-\sum_{i=1}^n m_i (s_i^2 + z_E^2)^{-\frac{1}{2}} = E$. Therefore solutions with $E \geq 0$ correspond to unbounded motions. Since that energy curves does not contain stationary points, solutions with $-\sum_{i=1}^n m_i s_i^{-1} \leq E < 0$ correspond to periodic motions (see Figure 1). The case $E = 0$ corresponds to the critical value that separates periodic of unbounded motions. From this we deduce the formula (5).

Now we will prove the second item of the theorem. We consider a T_0 -periodic solution ($E < 0$). As a consequence of the symmetries of the equation we have that the curve $t \mapsto (z(t), v(t))$, for $0 \leq t \leq T_0/4$, joins the points $(0, v_E)$ and

3 Proofs

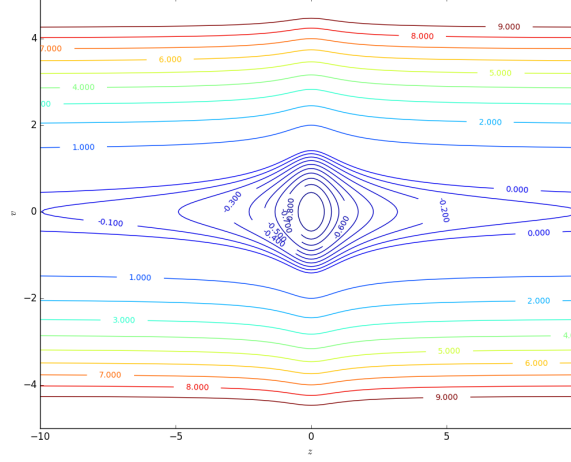


Figure 2: Energy level for two equal mass primaries

$(z_E, 0)$. Then, taking account (10) we have that

$$\begin{aligned}
 \frac{T_0}{4} &= \frac{1}{\sqrt{2}} \int_0^{T_0/4} \left(E + \sum_{i=1}^n m_i (s_i^2 + z^2)^{-\frac{1}{2}} \right)^{-\frac{1}{2}} z'(t) dt \\
 &= \frac{1}{\sqrt{2}} \int_0^{z_E} \left(\sum_{i=1}^n m_i \left((s_i^2 + z^2)^{-\frac{1}{2}} - (s_i^2 + z_E^2)^{-\frac{1}{2}} \right) \right)^{-\frac{1}{2}} dz \\
 &= \frac{1}{\sqrt{2}} \int_0^{z_E} (z_E^2 - z^2)^{-\frac{1}{2}} f(z, z_E) dz \\
 &= \frac{1}{\sqrt{2}} \int_0^1 (1 - u^2)^{-\frac{1}{2}} f(z_E u, z_E) du,
 \end{aligned} \tag{11}$$

where

$$f(z, z_E) = \left(\sum_{i=1}^n m_i \{ (s_i^2 + z^2)(s_i^2 + z_E^2) \}^{-\frac{1}{2}} \{ (s_i^2 + z^2)^{-\frac{1}{2}} + (s_i^2 + z_E^2)^{-\frac{1}{2}} \} \right)^{-\frac{1}{2}}.$$

We note that $f(z_E u, z_E)$ is an increasing function with respect to z_E for $u \in [0, 1]$ fix. Additionally

$$\lim_{z_E \rightarrow 0} f(z_E u, z_E) = \left(\sum_{i=1}^n \frac{m_i}{2s_i^3} \right)^{-\frac{1}{2}} \quad \text{and} \quad \lim_{z_E \rightarrow \infty} f(z_E u, z_E) = \infty.$$

Therefore, from the dominated convergence theorem and monotone convergence

theorem we have that

$$\lim_{z_E \rightarrow 0} T_0 = 2\pi \left(\sum_{i=1}^n \frac{m_i}{s_i^3} \right)^{-\frac{1}{2}} \quad \text{and} \quad \lim_{z_E \rightarrow \infty} T_0 = \infty.$$

Finally, since $T_0 = T_0(z_E)$ is continuous and increasing respect to z_E we conclude the affirmation in the item 2.

The item 3 is consequence that $T^2 = 4\pi^2 \sum_{i=1}^n m_i |q_i|^2 / U$ (see [13, p. 109]). \square

Let us to show a second proof of item 2 of Theorem 2.3. For the sufficiency we follow arguments of [27] and [10], based on variational principles. For the necessity of condition (6) we use Sturm's comparison theorem.

Alternitave Proof Theorem 2.3 (2). First we prove that (6) is a necessary condition for the existence of a T_0 -periodic solution. We assume that z is a T_0 -periodic solution of (4). Using Sturm's Comparison Theorem (see [1]) with equations $z'' + q_1(z)z = 0$, where $q_1(z) = \sum_{i=1}^n m_i (s_i^2 + z^2)^{-3/2}$, and $z'' + (\sum_{i=1}^n m_i s_i^{-3})z = 0$ we deduce (6).

Let $T_0 > 0$ satisfying (6). We consider the action integral

$$\mathcal{I}(z) = \int_0^{T_0} \frac{1}{2} |z'|^2 + \sum_{i=1}^n \frac{m_i}{\sqrt{s_i^2 + z^2}} dt,$$

Then T_0 -periodic solutions of (4) are critical points of \mathcal{I} in the space $H^1(\mathbb{T}, \mathbb{R})$, where $\mathbb{T} = \mathbb{R}/T_0\mathbb{Z}$, of the functions absolutely continuous, T_0 -periodic with $z' \in L^2(\mathbb{T}, \mathbb{R})$ (see [15, Cor. 1.1]). We prove the existence of critical points by means of the direct method of calculus of variations, i.e. we will prove that \mathcal{I} has minimum. The functional \mathcal{I} is not coercive in $H^1(\mathbb{T}, \mathbb{R})$, this deficiency is drawn with symmetry techniques (see [25]). The group \mathbb{Z}_2 acts on $H^1(\mathbb{T}, \mathbb{R})$ according to the following assignments $(\bar{0} \cdot z)(t) = z(t)$ and $(\bar{1} \cdot z)(t) = -z(t + \frac{T_0}{2})$. The functional \mathcal{I} is \mathbb{Z}_2 -invariant, i.e. $\mathcal{I}(g \cdot z) = \mathcal{I}(z)$. We define the space of all \mathbb{Z}_2 -symmetric (this simmetry is called the italian simmetry) funcions

$$\Lambda(\mathbb{T}, \mathbb{R}) := \{z \in H^1(\mathbb{T}, \mathbb{R}) | \forall g \in \mathbb{Z}_2 : z = g \cdot z\}.$$

The funcional \mathcal{I} restricted to Λ is coercive. This fact follows from an obvious adaptation of proposition 4.1 of [25]. We note that $F(z) := \sum_{i=1}^n m_i (s_i^2 + z^2)^{-\frac{1}{2}}$ satisfies the condition (A) in [15, p. 12], then \mathcal{I} is continuously differentiable and weakly lower semicontinuous on $H^1(\mathbb{T}, \mathbb{R})$ (see [15, p. 13]). Therefore \mathcal{I} has a minimum z_0 in $\Lambda(\mathbb{T}, \mathbb{R})$. Then by the Palais' principle symmetric criticality, z_0 is a critical point of \mathcal{I} in $H^1(\mathbb{T}, \mathbb{R})$ (see [25] and [19]).

We use the second variation $\delta^2 \mathcal{I}$ in order to show that $z_0 \neq 0$. It is well known (see [9, Th. 1.3.1]) that if z_0 is a minimum of \mathcal{I} on $H^1(\mathbb{T}, \mathbb{R})$ then $\delta^2 \mathcal{I}(z_0, \varphi) \geq 0$ for all $\varphi \in H^1(\mathbb{T}, \mathbb{R})$. In our case

$$\delta^2 \mathcal{I}(0, \varphi) = \int_0^{T_0} |\varphi'|^2 - \sum_{i=1}^n \frac{m_i}{r_i^3} \varphi^2 dt,$$

(see [9, Eq. 1.3.6]). In particular for $\varphi(t) = \sin(2\pi t/T_0)$ it follows from (6) that

$$\delta^2 \mathcal{I}(0, \varphi) = \left(\frac{4\pi^2}{T_0^2} - \sum_{i=1}^n \frac{m_i}{r_i^3} \right) \frac{T_0}{2} < 0. \quad (12)$$

It is sufficient to guarantee that $z_0 \equiv 0$ is not a minimum. \square

Remark 1. We note that this second proof of Theorem 2.3 (2), unlike the first one, does not guarantee that T_0 is the minimum period for z_0 . It could happen that z_0 had period T_0/m , with natural $m \in \mathbb{N}$. By the italian symmetry would have to occur that m is odd.

Proof Corollary 2.4. Our condition (7) is reduced to the inequality (41) in [10], where was proved that this inequality holds if and only if $n \leq 472$. \square

Proof Theorem 2.5. Let's start by analyzing the central configuration CCr. We can suppose without loss of generality that $q_1 = -q_3 = (0, y)$ for $y > 0$, $q_2 = -q_4 = (1, 0)$, $m_1 = m_3 = M$, $m_2 = m_4 = m$ and $M > m$. Then, necessarily $y < 1$ (see [22]). For this CC the condition (7) becomes

$$\frac{M^2}{2y} + \frac{4Mm}{\sqrt{1+y^2}} + \frac{m^2}{2} < \left(\frac{2M}{y^3} + 2m \right) (2My^2 + 2m).$$

As $M^2/(2y) < 4M^2/y$, $m^2/2 < 4m^2$ and $4Mm/\sqrt{1+y^2} < 4Mm/y^3$ (since $y < 1$), we have that the inequality holds.

Now consider the central configuration CCl. Remark first that some of the following calculations were computed using a symbolic mathematics software. We can suppose without loss of generality that $q_1 = -q_3 = 1$, $q_2 = -q_4 = x$ with $0 < x < 1$, and $m_1 = m_3 = \mu$, $m_2 = m_4 = 1 - \mu$, with $0 < \mu < 1$. Then the inequality (7) becomes

$$\frac{2\mu(1-\mu)}{1-x} + \frac{2\mu(1-\mu)}{1+x} + \frac{\mu^2}{2} + \frac{(1-\mu)^2}{2x} < 4\mu^2 + 4\mu(1-\mu)x^2 + \frac{4\mu(1-\mu)}{x^3} + \frac{4(1-\mu)^2}{x}.$$

As $\frac{\mu^2}{2} < 4\mu^2$ and $\frac{(1-\mu)^2}{2x} < \frac{4(1-\mu)^2}{x}$ (without taking into account the term $4\mu(1-\mu)x^2$) we just have to show that

$$\frac{2\mu(1-\mu)}{1-x} + \frac{2\mu(1-\mu)}{1+x} < \frac{4\mu(1-\mu)}{x^3},$$

and this is equivalent to see that

$$\frac{x^3}{1-x^2} < 1. \quad (13)$$

The values of x involved in the above inequality are such that the configuration for the vector mass $(\mu, 1-\mu, 1-\mu, \mu)$ is central, by Moulton [18], fixed a mass μ there is only one value x satisfying this condition. So, we can define $x(\mu)$ as

such value of x . If we can see that the function $x(\mu)$ is a decreasing function, then we have that $\frac{x(\mu)^3}{1-x(\mu)^2} \leq \lim_{\mu \rightarrow 0} \frac{x(\mu)^3}{1-x(\mu)^2}$. If also we demonstrate that

$$\lim_{\mu \rightarrow 0} \frac{x(\mu)^3}{1-x(\mu)^2} < 1 \quad (14)$$

we have tested (13).

Let's first prove that $x(\mu)$ is a decreasing function. The relationship between μ and x follows from the fact that the bodies are in a central configuration. Therefore the equation

$$\frac{\mu}{4} - \frac{\mu}{x(x+1)^2} + \frac{\mu}{x(-x+1)^2} + \frac{-\mu+1}{(x+1)^2} + \frac{-\mu+1}{(-x+1)^2} - \frac{1}{x^3} \left(-\frac{\mu}{4} + \frac{1}{4} \right) = 0$$

must be satisfied, simplifying this expression we have

$$\frac{p(\mu, x)}{q(\mu, x)} = 0,$$

where $p(\mu, x) = \mu x^7 - 10\mu x^5 + \mu x^4 + 9\mu x^3 - 2\mu x^2 + \mu + 8x^5 - x^4 + 8x^3 + 2x^2 - 1$ and $q(\mu, x) = 4x^3(x^4 - 2x^2 + 1)$. So the relationship between μ and x is given that $p(\mu, x) = 0$. We can derive implicitly the last equation and we obtain

$$\frac{dx}{d\mu} = \frac{Np(x)}{Dp(x, \mu)},$$

where $Np(x) = -(x-1)(x+1)(x^5 - 9x^3 + x^2 - 1)$ and $Dp(x, \mu) = \mu x^2(7x^4 - 10x^2 + 51) + (1-\mu)(40x^4 + 20x^3 + 4x)$. The denominator $Dp(x, \mu)$ is clearly positive for $0 < x < 1$ and $0 < \mu < 1$. To prove that the numerator $Np(x)$ is negative let's see that the polynomial $x^5 - 9x^3 + x^2 - 1$ is negative for all $0 < x < 1$. In fact, calculating their real roots with a software we have that these are -3.0483999 , -0.449322 and 2.94956549 , then it is easy to see that $x^5 - 9x^3 + x^2 - 1$ is negative in the interval $(0, 1)$, hence $Np(x) < 0$ for $0 < x < 1$. This implies that $\frac{dx}{d\mu}$ is negative for all $0 < x < 1$ and $0 < \mu < 1$.

Let's see now that (14) holds. Since $x(\mu)$ is a continuous function we need to prove $\frac{x(0)^3}{1-x(0)^2} < 1$. For simplicity we will write x instead of $x(0)$. This value x is such that $p(0, x) = 8x^5 - x^4 + 8x^3 + 2x^2 - 1 = 0$, then $8x^5 + 8x^3 = (x^2 - 1)^2$, and this implies that $8x^5 + 8x^3 < 1$, hence $x^3 < 1/(8(x^2 + 1)) < 1/8$, $x^2 < (1/8)^{\frac{2}{3}}$ and

$$\frac{x^3}{1-x^2} < \frac{1/8}{1-1/8^{\frac{2}{3}}} = \frac{1}{6} < 1,$$

as we wanted to prove. □

Acknowledgments

References

- [1] G. Birkhoff and G.C. Rota. *Ordinary Differential Equations*. Wiley, 1989.
- [2] T Bountis and KE Papadakis. The stability of vertical motion in the n-body circular sitnikov problem. *Celestial Mechanics and Dynamical Astronomy*, 104(1):205–225, 2009.
- [3] Steven R Chesley. A global analysis of the generalized Sitnikov problem. *Celestial Mechanics and Dynamical Astronomy*, 73(1-4):291–302, 1999.
- [4] Harry Dankowicz and Philip Holmes. The existence of transverse homoclinic points in the sitnikov problem. *Journal of differential equations*, 116(2):468–483, 1995.
- [5] R Dvorak. Numerical results to the sitnikov-problem. In *Qualitative and Quantitative Behaviour of Planetary Systems*, pages 71–80. Springer, 1993.
- [6] Johannes Hagel and Thomas Trenkler. A computer aided analysis of the sitnikov problem. In *Qualitative and Quantitative Behaviour of Planetary Systems*, pages 81–98. Springer, 1993.
- [7] MARSHALL HAMPTON. Co-circular central configurations in the four-body problem. In *EQUADIFF 2003*, pages 993–998. World Scientific, 2005.
- [8] Lidia Jiménez-Lara and Adolfo Escalona-Buendía. Symmetries and bifurcations in the sitnikov problem. *Celestial Mechanics and Dynamical Astronomy*, 79(2):97–117, 2001.
- [9] J. Jost and X. Li-Jost. *Calculus of Variations*. Cambridge Studies in Advanced Mathematics. Cambridge University Press, 1998.
- [10] Fengying Li, Shiqing Zhang, and Xiaoxiao Zhao. The characterization of the variational minimizers for spatial restricted $N + 1$ -body problems. *Abstract and Applied Analysis*, 2013(Article ID 845795), 2013.
- [11] Jie Liu, Ji-lin Zhou, and Yi-sui Sun. Numerical research on the sitnikov problem. *Chinese astronomy and astrophysics*, 15(3):339–344, 1991.
- [12] Jaume Llibre, Kenneth R Meyer, and Jaume Soler. Bridges between the generalized sitnikov family and the lyapunov family of periodic orbits. *Journal of Differential Equations*, 154(1):140–156, 1999.
- [13] Jaume Llibre, Richard Moeckel, and Carles Simó. *Central Configurations, Periodic Orbits, and Hamiltonian Systems*. Advanced Courses in Mathematics - CRM Barcelona. Birkhäuser, 2015, nov 2015.

- [14] Jaume Llibre and Rafael Ortega. On the families of periodic orbits of the sitnikov problem. *SIAM Journal on Applied Dynamical Systems*, 7(2):561–576, 2008.
- [15] J. Mawhin and M. Willem. *Critical Point Theory and Hamiltonian Systems*. Applied Mathematical Sciences. Springer, 1989.
- [16] Richard Moeckel. On central configurations. *Mathematische Zeitschrift*, 205(1):499–517, 1990.
- [17] J. Moser. *Stable and Random Motions in Dynamical Systems: With Special Emphasis on Celestial Mechanics*. Annals Mathematics Studies. Princeton University Press, 1973.
- [18] Forest Ray Moulton. The straight line solutions of the problem of n bodies. *The Annals of Mathematics*, 12(1):1–17, 1910.
- [19] Richard Palais. *The principle of symmetric criticality*. Communications in Mathematical Physics, 1979.
- [20] LP Pandey and I Ahmad. Periodic orbits and bifurcations in the sitnikov four-body problem when all primaries are oblate. *Astrophysics and Space Science*, 345(1):73–83, 2013.
- [21] Hugo Jiménez Pérez and Ernesto A Lacomba. On the periodic orbits of the circular double sitnikov problem. *Comptes Rendus Mathématique*, 347(5):333–336, 2009.
- [22] Ernesto Perez-Chavela and Manuele Santoprete. Convex four-body central configurations with some equal masses. *Archive for rational mechanics and analysis*, 185(3):481–494, 2007.
- [23] K Sitnikov. The existence of oscillatory motions in the three-body problem. In *Dokl. Akad. Nauk SSSR*, volume 133, pages 303–306, 1960.
- [24] PS Soulis, KE Papadakis, and T Bountis. Periodic orbits and bifurcations in the sitnikov four-body problem. *Celestial Mechanics and Dynamical Astronomy*, 100(4):251–266, 2008.
- [25] Davide L. Ferrario; Susanna Terracini. On the existence of collisionless equivariant minimizers for the classical n-body problem. *Inventiones mathematicae*, 155, 02 2004.
- [26] Aurel Wintner. *The Analytical Foundations of Celestial Mechanics*. Courier Corporation, jun 2014.
- [27] Xiaoxiao Zhao and Shiqing Zhang. Nonplanar periodic solutions for spatial restricted 3-body and 4-body problems. *Boundary Value Problems*, 2015(1):1, 2015.