Utili per il colcolo di zin e 9.

(asa (1):

dew one nere

2 = a + 1 b

Z= M.e 18 Ampierra: M= 121 - malulo

numero complesso

numero complessor (: 0 = arg (2) -o free

Scritto Come

scritto come

modula e fase Reisis a - parte reale

Perte immaginaria

Im(2) 5 b - s parte immeginaria

M= Va2+62

parte reale e

O = vedi formule

Cesa (2):

Z = M. e J & devo ottenere

== > t = a + j b

Scrivor e ja come:

e 10 , cos (0) + J seu (0) ,

allora:

Q = M. (8)

b = M. Jsen (0)

VECTOR DIFFERENTIAL OPERATIONS

$$\nabla \Phi = \hat{\mathbf{x}} \frac{\partial \Phi}{\partial x} + \hat{\mathbf{y}} \frac{\partial \Phi}{\partial y} + \hat{\mathbf{z}} \frac{\partial \Phi}{\partial z}$$

$$\nabla \cdot \mathbf{D} = \frac{\partial D_x}{\partial x} + \frac{\partial D_y}{\partial y} + \frac{\partial D_z}{\partial z}$$

$$\nabla \times \mathbf{H} = \hat{\mathbf{x}} \left(\frac{\partial H_z}{\partial y} - \frac{\partial H_y}{\partial z} \right) + \hat{\mathbf{y}} \left(\frac{\partial H_x}{\partial z} - \frac{\partial H_z}{\partial x} \right) + \hat{\mathbf{z}} \left(\frac{\partial H_y}{\partial x} - \frac{\partial H_x}{\partial y} \right)$$

$$\nabla^2 \Phi = \frac{\partial^2 \Phi}{\partial x^2} + \frac{\partial^2 \Phi}{\partial y^2} + \frac{\partial^2 \Phi}{\partial z^2}$$

$$\nabla^2 \mathbf{A} = \hat{\mathbf{x}} \nabla^2 A_x + \hat{\mathbf{y}} \nabla^2 A_y + \hat{\mathbf{z}} \nabla^2 A_z$$

$$\nabla \Phi = \hat{\mathbf{r}} \frac{\partial \Phi}{\partial r} + \hat{\Phi} \frac{1}{r} \frac{\partial \Phi}{\partial \phi} + \hat{\mathbf{z}} \frac{\partial \Phi}{\partial z}$$

$$\nabla \cdot \mathbf{D} = \frac{1}{r} \frac{\partial}{\partial r} (rD_r) + \frac{1}{r} \frac{\partial D_{\phi}}{\partial \phi} + \frac{\partial D_z}{\partial z}$$

$$\nabla \times \mathbf{H} = \hat{\mathbf{r}} \left[\frac{1}{r} \frac{\partial H_z}{\partial \phi} - \frac{\partial H_{\phi}}{\partial z} \right] + \hat{\Phi} \left[\frac{\partial H_r}{\partial z} - \frac{\partial H_z}{\partial r} \right] + \hat{\mathbf{z}} \left[\frac{1}{r} \frac{\partial (rH_{\phi})}{\partial r} - \frac{1}{r} \frac{\partial H_r}{\partial \phi} \right]$$

$$\nabla^2 \Phi = \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial \Phi}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2 \Phi}{\partial \phi^2} + \frac{\partial^2 \Phi}{\partial z^2}$$

$$\nabla^2 \mathbf{A} = \hat{\mathbf{r}} \left(\nabla^2 A_r - \frac{2}{r^2} \frac{\partial A_{\phi}}{\partial \phi} - \frac{A_r}{r^2} \right) + \hat{\Phi} \left(\nabla^2 A_{\phi} + \frac{2}{r^2} \frac{\partial A_r}{\partial \phi} - \frac{A_{\phi}}{r^2} \right) + \hat{\mathbf{z}} (\nabla^2 A_z)$$

$$\nabla \Phi = \hat{\mathbf{r}} \frac{\partial \Phi}{\partial r} + \hat{\mathbf{\theta}} \frac{1}{r} \frac{\partial \Phi}{\partial \theta} + \frac{\hat{\mathbf{\phi}}}{r \sin \theta} \frac{\partial \Phi}{\partial \phi}$$

$$\nabla \cdot \mathbf{D} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 D_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta D_\theta) + \frac{1}{r \sin \theta} \frac{\partial D_\phi}{\partial \phi}$$

$$\stackrel{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}}{\overset{\mathfrak{S}}}\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}}{\overset{\mathfrak{S}}}\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}}{\overset{\mathfrak{S}}}}{\overset{\mathfrak{S}}}}{\overset{\mathfrak{S}}}\overset{\mathfrak{S}}{\overset{\mathfrak{S}}}}{\overset{\mathfrak{S}}}}}{\overset{\mathfrak{S}}}\overset{\mathfrak{S}}}{\overset{\mathfrak{S}}}}{\overset{\mathfrak{S}}}}}} + \frac{1}{r^2}}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}}} \frac{1}{r^2}} \frac{1}{r^2}} \frac$$

9) Costanti di uso frequente

Costante dielettrica del vuoto : $\epsilon_o=8.85~10^{-12}~F/m$ Permeabilitá magnetica del vuoto : $\mu_o=4\pi~10^{-7}~H/m$

Carica dell'elettrone : $e = 1.60 \ 10^{-19} \ C$ Massa dell'elettrone : $m_e = 9.1 \ 10^{-31} \ kg$

Rapporto e/m dell'elettrone : $e/m = 1.76 \ 10^{11} \ C/kg$

Massa del protone : $m_p = 1.67 \ 10^{-27} \ kg$

Velocitá delle onde e.m. nel vuoto : $c = 3.0 \ 10^8 \ m/s$

Impedenza del vuoto : $Z_o = 376.7 \Omega$

Costante di Planck : $h = 6.626 \ 10^{-34} \ J \cdot s$

Magnetone di Bohr : $\mu_B = 9.42~10^{-24}~A~m^2$

Costante gravitazionale : $G = 6.672 \ 10^{-11} m^3 \ kg - 1 \ s^{-2}$

Numero di Avogadro : $N_A = 6.02252 \ 10^{23} \ mol^{-1}$

Costante di Boltzmann : $k=1.38054\ 10^{-23}\ J\ K^{-1}$

Costante dei gas : $R = 8.314 \ J/(mol \ K)$ = 1.986 $cal/(mol \ K)$

Volume di una mole(STP gas ideale) : $k = 22.414 \ 10^{-3} \ m^3 mol^{-1}$

Unitá astronomica : $AU = 1.49598 \ 10^{11} \ m$

Raggio(equatoriale) della terra : $R_{\bigoplus} = 6.378 \ 10^6 \ m$

Massa della terra : $M_{\bigoplus} = 5.973 \ 10^{24} \ kg$

Massa del sole : $M_{\odot} = 1.989 \ 10^{30} \ kg$

Formule di Teoria dei Segnali

L. Verdoliva

Formule di trigonometria

$$\cos(\alpha + \beta) = \cos \alpha \cos \beta - \sin \alpha \sin \beta$$

$$\sin(\alpha + \beta) = \sin \alpha \cos \beta + \sin \beta \cos \alpha$$

$$\cos^{2} \alpha = \frac{1 + \cos 2\alpha}{2}$$

$$\sin^{2} \alpha = \frac{1 - \cos 2\alpha}{2}$$

$$\sin 2\alpha = 2\sin \alpha \cos \alpha$$

$$\cos 2\alpha = \cos^{2} \alpha - \sin^{2} \alpha = 2\cos^{2} \alpha - 1 = 1 - 2\sin^{2} \alpha$$

$$\cos \alpha \cos \beta = \frac{1}{2} [\cos(\alpha - \beta) + \cos(\alpha + \beta)]$$

$$\sin \alpha \sin \beta = \frac{1}{2} [\cos(\alpha - \beta) - \cos(\alpha + \beta)]$$

$$\sin \alpha \cos \beta = \frac{1}{2} [\sin(\alpha + \beta) + \sin(\alpha - \beta)]$$

Formule di Eulero

$$\cos\alpha \ = \ \frac{e^{j\alpha}+e^{-j\alpha}}{2} \qquad \qquad \sin\alpha = \frac{e^{j\alpha}-e^{-j\alpha}}{2j} \qquad \qquad e^{j\alpha} = \cos\alpha + j\sin\alpha$$

Proprietà $\delta(t)$ e $\delta(n)$

$$\int_{t_1}^{t_2} x(t) \, \delta(t) \, dt = \begin{cases} x(0) & 0 \in (t_1, t_2) \\ 0 & \text{altrimenti} \end{cases} \qquad \delta(n) = \begin{cases} 1 & n = 0 \\ 0 & \text{altrimenti} \end{cases}$$

$$\int_{-\infty}^{+\infty} \delta(t) \, dt = 1 \qquad \qquad \sum_{k=-\infty}^{+\infty} \delta(n-k) = 1$$

$$\sum_{n=-\infty}^{+\infty} x(t) \delta(t-t_0) \, dt = x(t_0) \qquad \qquad \sum_{n=-\infty}^{+\infty} x(n) \delta(n-n_0) = x(n_0)$$

$$x(t) \, \delta(t-t_0) = x(t_0) \, \delta(t-t_0) \qquad \qquad x(n) \, \delta(n-n_0) = x(n_0) \, \delta(n-n_0)$$

$$\delta(t) = \delta(-t) \qquad \qquad \delta(n) = \delta(-n)$$

$$\sum_{k=-\infty}^{+\infty} x(k) \delta(n-k) = x(n) * \delta(n) = x(n)$$

$$\sum_{k=-\infty}^{+\infty} x(k) \delta(n-k) = x(n) * \delta(n) = x(n)$$

$$\sum_{k=-\infty}^{n} \delta(k) = u(n) \leftrightarrow \delta(n) = u(n) - u(n-1)$$

FORMULARIO CEM

Formule utili per exercisio (4):

$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2} & \text{Latin} = \frac{450}{2 \text{cm}} = \frac{c}{2\pi} & \sqrt{\left(\frac{\pi}{a}\right)^{1}} & [H_1] \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{04} = \frac{c}{2} & \frac{c}{2\pi} & \frac{150}{6 \text{cm}} = \frac{c}{2\pi} & \sqrt{\left(\frac{\pi}{a}\right)^{1}} & [H_1] \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{04} = \frac{c}{2\pi} & \frac{c}{2\pi} & \sqrt{\left(\frac{\pi}{a}\right)^{2}} & [H_1] \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2\pi} & \frac{c}{2\pi} & \frac{c}{2\pi} \\
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$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2\pi} & \frac{c}{2\pi} \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2\pi} & \frac{c}{2\pi} \\
\end{cases}$$

$$\begin{cases}
c & \text{TE}_{10} = \frac{c}{2\pi} \\
\end{cases}$$

Formule utili per esercitio (8): CASO TE: 1 TE ZIN - ZI [Numero puro] ξ, Te : η / (ωs(δ,) [Λ] $\frac{1}{2}$ $\frac{\eta_1}{\omega_{S(\theta_1)}}$ $\frac{\eta_1/\omega_{S(\theta_1)} + f(\eta_1/\omega_{S(\theta_1)}) t_1}{\eta_1/\omega_{S(\theta_1)} + f(\eta_1/\omega_{S(\theta_1)}) t_1}$ [2] CASO TH : TIN - En [numero puro] 7, th. m, ws(01) [n] 7, 605(91) + Jm cos(01) E1 [1] Validità yenerale: $\frac{\text{Validità yenerale:}}{\text{Snele:}}$ $\frac{\text{Snele:}}{\text{N1 Slu}(\mathfrak{I}_1)} = \text{N1 seu}(\mathfrak{I}_1)$ $\frac{\text{Snele:}}{\text{N2 Slu}(\mathfrak{I}_1)} = \text{N1 seu}(\mathfrak{I}_1)$ $\frac{\text{Snele:}}{\text{N2 Slu}(\mathfrak{I}_1)} = \text{N2 seu}(\mathfrak{I}_1)$ $\frac{\text{Risultato deve}}{\text{eisere in gradie}}$ κ, - ω· √μ, ξ. [ω-1] , κ; ω √μ, ε, η; [ω-1] ni = VEri [human puro], Bis Ki cos (Si) W = 2 [red/:]

Angolo di Brewster:

Oinc = OB = tan (ninc) <=> E2=0 , ave se: K2 cos (O1) d = n =

Continue:

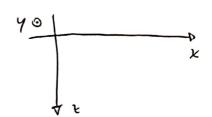
Potenta trasmessa: Pro (1 - |) · Pina [W/w.]

Porou en medie incidente: pine - { (Re) [Eil [W/w], récolcole con il veltore di Paynting.

COSO TE:

Trasvers elettres: comps & lungs q

compo H lungo ê e ê.



Situatione .

| Vioro | Metts | Voor |
|-------|----------------|------------|
| | ٤,, | |
| | (dielettics) | |
| | | , f |
| ツェニツ・ | ウ ₂ | 73 70 0 74 |

Regione 1: anda incidente e aflessa

Regione 2: diclettino

Regione 3: onda tresmessa

(Regione 1)

$$E_{1}(x,y,z) \cdot \left[E_{inc}^{*}\cdot\left(e^{-JK_{1}(\omega)(\theta_{1})^{2}}-\int_{1}^{2}e^{-JK_{1}(\omega)(\omega_{1})^{2}}\right)e^{-JK_{1}\omega\mu(\theta_{1})x}\right]\hat{q}$$

$$H_{1}x = -\frac{1}{J\omega\mu_{1}}\nabla_{x}E_{1}\left[x\right] = \frac{-1}{J\omega\mu_{1}}\partial_{z}E_{1}y + \frac{-\beta_{1}}{\omega\mu_{2}}E_{1}y$$

$$H_{1}$$
, $\frac{-1}{Jw\mu_{0}}$ $\nabla_{x} \in \mathcal{L}_{1}$ $= \frac{1}{-Jw\mu_{0}} \partial_{x} \in \mathcal{L}_{1}$, $\frac{\beta_{1}}{w\mu_{0}} \in \mathcal{L}_{1}$

(Regione 2):

$$H_{12} = -\frac{1}{J\omega\mu_0} \nabla_x \bar{\epsilon}_1 \Big|_{\Sigma} = 0$$

$$\frac{\beta_1}{\omega\mu_0} \epsilon_{2y} \hat{x}$$

$$H_{1+} = \frac{1}{J\omega\mu} \nabla_{\lambda} \tilde{\epsilon}_{\lambda} \Big|_{\tilde{\epsilon}} = \frac{\beta_{1}}{\omega\mu} \epsilon_{1} \gamma \tilde{\epsilon}$$

In too (tru ses)

$$E_{2}^{\dagger} = E_{1}^{\dagger}$$
 ($e_{1} = E_{2}^{\dagger}$) ($e_{2} = E_{2} = E_{3} = E_$

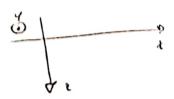
(i)
$$\frac{\beta_i}{\omega_{\mu}}$$
 ϵ_i^{\dagger} (1-7) = $\frac{\beta_1}{\omega_{\mu}}$ (ϵ_z^{\dagger} - ϵ_z^{-}) (rupre nel ce so spre)

$$\bar{E}_{2} = \left[\left(E_{3}^{+} e^{j K_{1} \cos(\theta_{2}) + e^{-j K_{1} \sin(\theta_{2}) \times e}} \right) \right] \hat{q}, \text{ point} \quad K_{1} \cos(\theta_{2}) \cdot K_{3} \cos(\theta_{3}),$$

$$= \left[\left(E_{3}^{+} e^{j K_{1} \cos(\theta_{2}) + e^{-j K_{2} \sin(\theta_{2}) \times e}} \right) \right] \hat{q}, \text{ point} \quad K_{1} \cos(\theta_{2}) \cdot K_{3} \cos(\theta_{3}),$$

$$H_{3}$$
 = $\frac{-4}{Jw\mu}$, $\nabla \times \bar{\epsilon}_{3}$ = $-\frac{1}{Jw\mu}$, $\partial_{\epsilon} E_{3}$, $\hat{\epsilon}_{3}$, $\frac{\beta_{1}}{w\mu}$, $\hat{\epsilon}_{3}$, $\hat{\epsilon}_{3}$

Tresverso megnetico: compo E lungo é e à compo H lungo j



Siquatione procedente

Per snell:

My sen (9 2) = n. sen(0), devo no essere due quantità vyudi

K 2 seu (J2) = K2 seu (D2)

Dunque, so che:

Ka=Ko.na e & Kirko.ni

Ourali aveci:

Kon, seu (ou) + Kon, seu (ou).

Amens del fettore No, che prot essere semplificatio, ritrovo la legge di Snell. Persunto le due espressioni seno uguelis.

Posso scrivere:

6 - INT ser (21)x

che:

e - j k, seu (9:1 ×

| Ricevimento Morini (01/06/2020) | |
|---|-------|
| Polanitatione TE: | |
| Unando deux trovere le espressioni dei eaux; devo | r: |
| maltiplicare une partione del campo may cretico (Hz | e h |
| × totto il compo elettrico, che so Trone lungo y (Ey) |). |
| | |
| | |
| | |
| Polavizzazione TM: | |
| Unando desto tro vere le espressioni dei camp; devo: | |
| maltiplicare una partione del compo elettrico (Exed Ez) | |
| X to HO i) comper magnetizor, che 5: trone lungo y (H | (۲) - |
| | |
| | |
| | |
| | |
| -> Se b = a che succède 9 | |
| La boude di mono mo dolità è tra la Preg. di taplia del | |
| TE 15 e quella del TE 20 | |
| | |
| | |
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| | |
| | + |
| | |
| | |
| | |
| | + |

```
Je is be le situatione:

V_{0000}

E_{r}: 2

E_{r}: 2

E_{r}: 3

E_{r}: 3

E_{r}: 3

E_{r}: 4

E_{
```