# dwk ++ User Manual

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## 1 Introduction

dwk++ is a small code to calculate  $\delta W_k$  and solver fishbone dispersion relation for toakmak plasmas. The  $\delta W_k$  is calculate by 3D integration in phase space  $(r, \Lambda, E)$ . A slowing down distribution function and a kink like mode structure is used. With a small growth rate (imag part of  $\Omega$ ) and tokamak parameters as a input, this code find out the fishbone mode frequency and fast ion  $\beta_{h,0}$ .

# 2 The defintion of $\delta W_k$

# 2.1 The normalized $\delta W_k$ in the code:

$$\delta W_k = \sum_{p} \int_0^1 \frac{J}{q} dr \int d\Lambda \int E^3 dE \tau_b(\omega - \omega_\star) \frac{\partial F}{\partial E} \frac{|Y_p|^2}{n\omega_\phi + p\omega_b - \omega}$$
 (1)

In current version, we only keep signle n.  $\sigma = \pm 1$  is the direction of  $v_{\parallel}$ , m and n is poloidal and toroidal mode number,  $\Lambda = \frac{\mu}{E}$ , E is the fast ion energy,  $\tau_b$  is the particle bounce time. The slowing down distribution function of fast ions is:

$$F = \frac{2^{3/2}}{C_f} \hat{F}(r, \epsilon, \Lambda) = \frac{2^{3/2}}{C_f} \frac{1}{E^{3/2} + E_o^{3/2}} erfc\left(\frac{E - E_0}{\Delta E}\right) e^{-\left(\frac{r - r_0}{\Delta r}\right)^2} e^{-\left(\frac{\Lambda - \Lambda_0}{\Delta \Lambda}\right)^2}$$
(2)

$$C_f = \int d^3 \mathbf{v} \frac{1}{E^{3/2} + E_c^{3/2}} erfc\left(\frac{E - E_0}{\Delta E}\right) e^{-\left(\frac{\Lambda - \Lambda_0}{\Delta \Lambda}\right)^2}$$
(3)

And  $d^3\mathbf{v} = \frac{\sqrt{2}\pi}{b\sqrt{1-\frac{\Lambda}{b}}}d\Lambda E^{1/2}dE$ .

$$C_f = \int \frac{\sqrt{2\pi}}{b\sqrt{1 - \frac{\Lambda}{b}}} \frac{1}{E^{3/2} + E_c^{3/2}} erfc\left(\frac{E - E_0}{\Delta E}\right) e^{-\left(\frac{\Lambda - \Lambda_0}{\Delta \Lambda}\right)^2} d\Lambda E^{1/2} dE \tag{4}$$

$$\frac{\partial F}{\partial E} = \frac{-2^{3/2}}{C_f} \left[ \frac{2 \exp(-(\frac{E-E_0}{\Delta E})^2)}{\sqrt{\pi} \Delta E(E^{3/2} + E_c^{3/2})} + \frac{3\sqrt{E} erfc(\frac{E-E_0}{\Delta E})}{2(E^{3/2} + E_c^{3/2})^2} - \frac{2\Lambda(\Lambda - \Lambda_0) erfc(\frac{E-E_0}{\Delta E})}{E\Delta \Lambda^2(E^{3/2} + E_c^{3/2})} \right] e^{-(\frac{r-r_0}{\Delta r})^2} e^{-(\frac{\Lambda - \Lambda_0}{\Delta r})^2}$$
(5)

$$\frac{\partial F}{\partial r} = \frac{2^{3/2}}{C_f} \frac{erfc\left(\frac{E - E_0}{\Delta E}\right)}{E^{3/2} + E_c^{3/2}} \frac{2(r_0 - r)}{\Delta r^2} e^{-\left(\frac{r - r_0}{\Delta r}\right)^2} e^{-\left(\frac{\Lambda - \Lambda_0}{\Delta \Lambda}\right)^2}$$
(6)

The diamagnetic frequency:

$$\omega_* = \frac{nq}{2r} \frac{\rho_0}{\varepsilon_0} \frac{\partial F/\partial r}{\partial F/\partial E} \tag{7}$$

where  $\rho_0$  is the gyro radius with injection energy.  $\varepsilon_0$  is the inverse aspect-ratio, and  $\varepsilon = \frac{r}{R_0}$ . The transit frequency for passing particle is given below:

$$\omega_b = \frac{\pi\sqrt{\kappa}}{K(\kappa^{-1})} \frac{\sqrt{\varepsilon\Lambda/2}}{q} \sqrt{E} \tag{8}$$

where  $\kappa = \frac{1-\Lambda(1-\varepsilon)}{2\varepsilon\Lambda}$ , and K denotes the complete elliptic integral of the first kind. The transit frequency in toroidal: $\omega_{\phi} = q\omega_{b}$ . Particle bounce time:  $\tau_{b} = \frac{2\pi}{\omega_{b}}$ . The integral along the particle orbits:

$$Y(r, \Lambda, E) = \frac{1}{2\pi} \int_0^{2\pi} \chi d\theta B_{\Lambda} \left( \Lambda_b + 2 \left( 1 - \Lambda_b \right) \right) G(r, \theta, E) e^{-i\chi p\Theta} \tag{9}$$

where  $\chi(r,\Lambda) = \frac{\sigma\pi\sqrt{\kappa}\sqrt{\varepsilon\Lambda/2}}{K(\kappa^{-1})}$ ,  $\kappa = \frac{1-\Lambda(1-\varepsilon)}{2\varepsilon\Lambda}$ . K denotes the complete elliptic integral of the first kind.  $\Lambda_b = \frac{\Lambda}{b}$ ,  $B_{\Lambda}(r,\Lambda,\theta) = \frac{1}{b\sqrt{(1-\Lambda_b)}}$ ,  $b = B/B_0 = 1 + (r/R_0)cos\theta$ .

$$G(\mathbf{r}, \Lambda, \mathbf{E}, \theta) = (g^{\theta\theta} \kappa_{\theta} + g^{r\theta} \kappa_{r}) \xi_{\theta}(\hat{r}(\bar{r}, \rho_{d}, \theta), \theta) + (g^{rr} \kappa_{r} + g^{r\theta} \kappa_{\theta}) \xi_{r}(\hat{r}(\bar{r}, \rho_{d}, \theta), \theta)$$

$$\tag{10}$$

where  $\hat{r} = \bar{r} + \rho_d cos\theta$ .

$$\rho_d(r, \Lambda, E) = \frac{q}{2} \rho_0 \sqrt{\frac{E}{1 - \Lambda/b}} \left[ \frac{\Lambda}{b} + 2(1 - \frac{\Lambda}{b}) \right] = \frac{q}{2} \rho_0 \sqrt{\frac{E}{1 - \Lambda_b}} \left[ \Lambda_b + 2(1 - \Lambda_b) \right]$$
(11)

To simplify the code, we use  $\Lambda_0$  instead of  $\Lambda$  in  $\rho_d$ .

$$\rho_d(r, E) = \frac{q}{2} \rho_0 \sqrt{\frac{E}{1 - \Lambda_{0,b}}} \left[ \Lambda_{0b} + 2(1 - \Lambda_{0b}) \right]$$
 (12)

For  $\Lambda_0 = 0$ 

$$\rho_d(r, E) = q\rho_0\sqrt{E}$$

$$G(\mathbf{r}, \mathbf{E}, \theta) = (g^{\theta\theta} \kappa_{\theta} + g^{r\theta} \kappa_{r}) \xi_{\theta}(\hat{r}(\bar{r}, \rho_{d}, \theta), \theta) + (g^{rr} \kappa_{r} + g^{r\theta} \kappa_{\theta}) \xi_{r}(\hat{r}(\bar{r}, \rho_{d}, \theta), \theta)$$

$$\tag{13}$$

 $g^{rr}=1+\tfrac{\varepsilon cos\theta}{2},\,g^{\theta\theta}=\tfrac{1}{r^2}\left(1-\tfrac{5}{2}\varepsilon cos\theta\right),\,g^{r\theta}=-\tfrac{3}{2r}\varepsilon sin\theta.$ 

$$\Theta(\theta, r, \Lambda) = \int_0^\theta d\theta' \frac{1}{b\sqrt{(1 - \Lambda_b)}} = \int_0^\theta B_\Lambda \tag{14}$$

#### 2.2 The mode structure

In the current version, the mode structure is source code, and it is a kink structure with a fixed boundary at r = 1, and with a finite resonance layer width  $\Delta r_s$ .

$$\xi_r(r,\theta) = \xi_{r0}(r) exp(i(\phi - \theta - \omega t)), \ \xi_{\theta}(r,\theta) = -i\xi_{\theta 0}(r) rexp(i(\phi - \theta - \omega t))$$

$$\xi_{r0}(r) = \begin{cases} \xi_0 & r \le r_s - \Delta r_s/2 \\ \xi_0 \frac{\Delta r_s - r + r_s - \Delta r_s/2}{\Delta r_s} & r_s - \frac{\Delta r_s}{2} < r < r_s + \frac{\Delta r_s}{2} \\ 0 & r \ge r_s + \frac{\Delta r_s}{2} \end{cases}$$

$$\xi_{\theta0}(r) = \begin{cases} \xi_0 & r \leq r_s - \Delta r/2 \\ \xi_0 \frac{\Delta r_s - 2r + r_s - \Delta r_s/2}{\Delta r_s} & r_s - \frac{\Delta r_s}{2} < r < r_s + \frac{\Delta r_s}{2} \\ 0 & r \geq r_s + \frac{\Delta r_s}{2} \end{cases}$$

If the flag in dwk.cfg input file input\_i=1, then the code will read the structure in file with name defined in input filename.

The file format is define below:

the first line is a integer for the grid number in r.

from the second line to the last is formated:

r xi r xi theta

## 2.3 The normalized quantities used for $\delta W_k$ :

 $v_0 = \sqrt{2T_0/M}$ ,  $T_0$  is the fast ions injection energy, M is the fast ion's mass.  $\omega_0 = \frac{v_0}{R_0}$ .  $F_0 = \frac{n_0}{v_0^2}$ ,  $n_0$  is the fast ion density at axis.  $r_0 = a$  is the minor radius,  $\varepsilon = a/R_0$ ,  $E_0 = T_0/M$ ,  $B_0$  is the torodial magnetic field at magnetic axis.  $\delta W_{k,0} = \pi^2 a^2 R_0 n_0 T_0$ .

## 3 Fishbone dispersion relation

Assume  $\delta W_{mhd} = 0$ , the normalized dispersion relation is:

$$\frac{4}{\pi} \left(\frac{r_s}{R_0}\right)^2 \left|\frac{\xi_0}{a}\right|^2 \left(-i\frac{\omega}{\omega_A}\right) + \beta_h C_p \delta W_k = 0 \tag{15}$$

or:

$$i\omega = C\beta_{h,0}C_{p}\delta W_{k} \tag{16}$$

where:

$$C = \frac{\omega_A}{\omega_0} \frac{1}{4(\frac{r_s}{R_0})^2 |\frac{\xi_0}{R_0}|^2} \tag{17}$$

and  $C_p = \frac{p_0}{n_0 T_0}$ . Here  $\xi_s/\xi_0 = 1$ ,  $\omega_A = \frac{v_A}{3^{1/2} R_0 s}$ ,  $v_A = \frac{B}{\sqrt{\mu_0 \rho_m}}$ ,  $s = r_s \frac{dq}{dr(r=r_s)}$ , and  $\beta_{h,0} = 8\pi n_0 T_h/B_t^2$ . Considering MHD contribution from m=1, n=1:

$$i\omega = C\beta_{h,0}C_P\delta W_k + \frac{\omega_A}{\omega_0}\delta W_T \tag{18}$$

where  $\delta W_T = 3\pi \left(\frac{r_s}{R}\right)^2 (1 - q_0) \left(\frac{13}{144} - \beta_{ps}^2\right)$ .

#### 4 How to run dwk++

#### 4.1 Compile the code

dwk++ code is using c++ language as the main language, so it need a c++ compiler to compile the code. Gnu/g++ on max os and Linux and intel compiler on Linux was tested. To compile the code, libconfig with version >1.5 is needed, and set environment variable LIBCONFIG\_DIR to the path where libconfig located. Set CXX to the c++ compiler (g++, icpc), and run 'make' at 'dwk-/' directory. If everything correct, a executable file 'dwk++' should be generated in 'dwk-/src/' directory and be copied to current directory. Then you can run dwk++ with dwk.cfg input file. A example of dwk.cfg can be find in 'dwk++/examples'.

#### 4.2 Arguments

- -h print help information.
- -i input file name, dwk.cfg by default.
- -o outfile name, omega dwk.out by default.
- -s scan dwk(omega), only find  $\Omega_0$  and  $\beta_{h,0}$ .
- -y write Yps 3D to Yps.nc.
- -x find solution for dispersion relation using iteration.
- -g find solution for dispersion relation using newton iteration with initial guess in gomegar and gomegai.
- -B fint the solution by newton for beta h to beta hb.

#### 4.3 Input file

Here is an example of input file:

```
Listing 1: a input file example
/*the input file for dwk++, parameters units is in ( ) */
//tokamak parameters
tokamak=
{
        a = 0.38;
                        //minor radius (m).
        R0 = 1.30:
                        //major radius (m).
        Bt = 0.84;
                        //Toroidal magnetic field at axis without plasma, (Tesla).
        n0 = 1.7e19;
                        //thermal plasma density, (m^-3).
                        //ion mass, (protom mass, m p).
        mi = 2.0;
        E_i0 = 25.142;
                        //Fost ion injection energy, (KeV).
                        //fat ion mass (protom mass,m_p).
        m ep = 2.0;
        //qc[0:7] q profile, q=qc[0] + qc[1]*r + qc[2] *r^2 ..... qc[7]*r^7.
        qc = [0.8, 0.0, 1.384189];
                        //q at resonance surface
        q s = 1.0;
        beta h = 0.0055;
                       //Fast ion beta
        beta hb=0.02;
        nbeta = 10;
//grid parameters
grid =
        nx=100; //grid size should be 3n + 1, n is a positive integer.
        nL = 100;
        nE = 100:
        ntheta = 100;
//fast ion distribution
slowing=
{
        rflag = 1;
                        //0 for exp profile, 1 for polynomial profile
                        //(a)
        r0 = 0.1;
        rd = 0.2;
                        //(a)
        //rc[0:8], F(r) = rc[0] + rc[1] * r + rc[2] * r^2 ... rc[8] * r^8;
        L0 = 0.01:
                        //Lambda 0
        Ld = 0.02;
                        //Delta Lambda
        E0 = 1.0;
                        //(E i0)
        Ed = 0.01;
                        //(E i0)
        Ec = 0.01;
                        //(E_i0)
                        //\cos(1) ? count (-1)
        sigma = 1;
//perturbation and omega range for the soultion
mode =
{
                //toroidal mode number.
        n=1;
                //poloidal mode number.
        pa=0:
                //resources: sum Yps/(n*omega_phi +p*omega_theta -\omega).
                //sum from pa to pb.
        pb=0;
        delta_r = 0.001; //step function width for kink. (a)
                        //1: using mode structure from input file
        input i=1;
        \verb|input_filename="mode.dat"; // mode structure data|\\
        omega_0=0.1; //find mode frequency and scan dwk between omega_0 and omega_1.
        omega_1=0.95; //omega unit is (v_0/R_0), v_0 is the fast ion injection speed.
        omega\_i=0.005; //image part of omega, the growth rate.
        omega n=100;
                          //scan steps.
        omega err=1.0e-5; //residual of omega 0.
                          //maximum iteration number to find the omega 0.
        \max iter = 100;
        \max iterg = 2;
```

```
dw f = 0.00;
                            //dw mhd.
                            //1: with drift orbit width effect, 0 without.
        zero rhod=0;
                            //1: with iner layer, 0 without.
        zero_iner=1;
        xi 0 = 0.01;
                            //(a) displacement at r=0.
        gomegar = 0.04;
        gomegai = 0.001;
}
dwkopt =
{
                                  //1: omega_star term on, 0: omega_star term off
        omega\_star\_off=1;
                                   //1: omega term on, 0; omega term off
        omega off=1;
}
```

#### 4.4 Output file

- omega\_dwk.out.
- dwk.log.
- Yps.nc.

#### 4.5 Utilities to plot results

There are some matlab & python scripts in 'dwk++/utilities' directory.

## 5 Benchmark

5.1 Compare with (WANG, Destabilization of internal kink modes at high frequency by energetic circulating ions. Physical Review Letters, 2001, 86) case ('dwk-/runp0 wsj').

Considering deep passing particles ( $\Lambda = 0$ ), and using distribution function:

$$F = \frac{p_h}{\pi n_0 T_0 C_p} \frac{1}{E^{3/2}} \delta(\Lambda) H(E_0 - E)$$
(19)

The  $\delta W_k$  analytical results:

$$\delta W_k = (\delta W_{k,d} + \delta W_{k,s})$$

$$\delta W_{k,d} = -\frac{8}{\varepsilon_0} \frac{\rho_h}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 [\Omega^3 ln(1 - \frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}] \int_0^{r_s} dr \frac{dp_h}{dr} q$$
 (20)

$$\delta W_{k,s} = \frac{8}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 \left[ -\frac{\Omega}{\Omega - 1} + \Omega + \Omega^2 ln(1 - \frac{1}{\Omega}) \right] \int_0^{r_s} r p_h dr$$
 (21)

Note that the second term  $\delta W_{k,s}$  in Eq.21 is different with Eq(13) in Wang's PRL paper. Using PBX parameters:  $B=0.84\mathrm{T},~\omega_{\zeta,0}/2\pi=190kHz,~R_0=1.3m,~a=0.38m,$  the injection energy  $T_0=25.142keV.~n_i=1.7\times10^{19}m^{-3},$   $\varepsilon_s=1/9,~r_s=0.1444~\mathrm{m},~s=0.4.$ 

Assume  $p_h = p_0 exp(-(\frac{r}{\Delta r})^2), q = c_0 + c_2 r^2$ .

$$\delta W_{k,d} = -\frac{8}{\varepsilon_0} \frac{\rho_h}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 [\Omega^3 ln(1 - \frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}] \int_0^{r_s} dr \frac{dp_h}{dr} q$$

$$= -\frac{8}{\varepsilon_0} \frac{\rho_h}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 [\Omega^3 ln(1 - \frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}] p_0 \int_0^{r_s} dr \frac{de^{-(\frac{r}{\Delta r})^2}}{dr} (c_0 + c_2 r^2)$$

$$= \frac{16}{\varepsilon_0} \frac{\rho_h}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 [\Omega^3 ln(1 - \frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}] \frac{p_0}{\Delta r^2} \int_0^{r_s} re^{-(\frac{r}{\Delta r})^2} (c_0 + c_2 r^2) dr$$

$$= \frac{16}{\varepsilon_0} \frac{\rho_h}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 [\Omega^3 ln(1 - \frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}] \frac{p_0}{\Delta r^2} \{ -\frac{1}{2}\Delta r^2 e^{-(\frac{r}{\Delta r})^2} [c_2(\Delta r^2 + r^2) + c_0] \} |_0^{r_s}$$

$$= -\frac{8}{\varepsilon_0} \frac{\rho_h}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 [\Omega^3 ln(1 - \frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}] p_0 \{ e^{-(\frac{r}{\Delta r})^2} [c_2(\Delta r^2 + r^2) + c_0] \} |_0^{r_s}$$

$$= -\frac{8}{\varepsilon_0} \frac{\rho_h}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 [\Omega^3 ln(1 - \frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}] p_0 \{ e^{-(\frac{r}{\Delta r})^2} [c_2(\Delta r^2 + r^2) + c_0] \} |_0^{r_s}$$
(22)

$$\delta W_{k,s} = -\frac{8}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 \left[ \frac{\Omega}{1 - \Omega} + \Omega + \Omega^2 ln (1 - \frac{1}{\Omega}) \right] \int_0^{r_s} r p_h dr 
= -\frac{8}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 \left[ \frac{\Omega}{1 - \Omega} + \Omega + \Omega^2 ln (1 - \frac{1}{\Omega}) \right] p_0 \int_0^{r_s} r e^{-(\frac{r}{\Delta r})^2} dr 
= -\frac{8}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 \left[ \frac{\Omega}{1 - \Omega} + \Omega + \Omega^2 ln (1 - \frac{1}{\Omega}) \right] p_0 \left[ -\frac{1}{2} \Delta r^2 e^{-(\frac{r}{\Delta r})^2} \right] |_0^{r_s} 
= \frac{4}{n_0 T_0 C_p} (\varepsilon_0 \xi_s)^2 \left[ \frac{\Omega}{1 - \Omega} + \Omega + \Omega^2 lln (1 - \frac{1}{\Omega}) \right] p_0 \Delta r^2 \left[ e^{-(\frac{r}{\Delta r})^2} \right] |_0^{r_s} \tag{23}$$

So the normalized  $\delta W_k$  is :

 $\frac{4(\varepsilon_0\xi_0)^2p_0}{n_0T_0C_p}\left\{[\frac{\Omega}{1-\Omega}+\Omega+\Omega^2ln(1-\frac{1}{\Omega})]\Delta r^2[e^{-(\frac{r}{\Delta r})^2}]|_0^{r_s}-\frac{2\rho_h}{\varepsilon_0}[\Omega^3ln(1-\frac{1}{\Omega})+\Omega^2+\frac{1}{2}\Omega+\frac{1}{3}]\{e^{-(\frac{r}{\Delta r})^2}[c_2(\Delta r^2+r^2)+c_0]\}|_0^{r_s}\right\}$  Define  $W_a=[\frac{\Omega}{1-\Omega}+\Omega+\Omega^2ln(1-\frac{1}{\Omega})]\Delta r^2[e^{-(\frac{r}{\Delta r})^2}]|_0^{r_s}-\frac{2\rho_h}{\varepsilon_0}[\Omega^3ln(1-\frac{1}{\Omega})+\Omega^2+\frac{1}{2}\Omega+\frac{1}{3}]\{e^{-(\frac{r}{\Delta r})^2}[c_2(\Delta r^2+r^2)+c_0]\}|_0^{r_s}$ . Consider the  $\delta W_{k,0}=\pi^2a^2R_0n_0T_0$ , finally, in physical units:

$$\begin{split} \delta W_{k,0} \delta W_k &= \delta W_{k,0} \frac{4(\varepsilon_0 \xi_0)^2 p_0}{n_0 T_0 C_p} \qquad W_a \\ &= \pi^2 a^2 R_0 n_0 T_0 \frac{4(\varepsilon_0 \xi_0)^2 p_0}{n_0 T_0 C_p} \qquad W_a \\ &= \frac{4\pi^2 a^2 R_0 (\varepsilon_0 \xi_0)^2 p_0}{C_p} \qquad W_a \\ &= \frac{4\pi^2 a^2 R_0 (\varepsilon_0 \xi_0)^2 p_0}{C_p} \qquad \left\{ \left[ \frac{\Omega}{1-\Omega} + \Omega + \Omega^2 ln(1-\frac{1}{\Omega}) \right] \Delta r^2 [e^{-(\frac{r}{\Delta r})^2}]|_0^{r_s} \\ &- \frac{2\rho_h}{\varepsilon_0} [\Omega^3 ln(1-\frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}] \{e^{-(\frac{r}{\Delta r})^2} [c_2(\Delta r^2 + r^2) + c_0]\}|_0^{r_s} \right\} \end{split}$$

Then we can get:

$$i\frac{\Omega}{\Omega_A} = \frac{1}{4} \frac{1}{\varepsilon_s^2} \frac{\beta_0}{|\xi_0|^2} \delta W_k \tag{24}$$

$$i\frac{\Omega}{\Omega_{A}} = \frac{\varepsilon_{0}^{2}}{\varepsilon_{s}^{2}} \frac{p_{0}\beta_{0}}{n_{0}T_{0}C_{p}} \left\{ \left[ \frac{\Omega}{1-\Omega} + \Omega + \Omega^{2}ln(1-\frac{1}{\Omega}) \right] \Delta r^{2} \left[ e^{-(\frac{r}{\Delta r})^{2}} \right] \right|_{0}^{r_{s}} - \frac{2\rho_{h}}{\varepsilon_{0}} \left[ \Omega^{3}ln(1-\frac{1}{\Omega}) + \Omega^{2} + \frac{1}{2}\Omega + \frac{1}{3} \right] \left\{ e^{-(\frac{r}{\Delta r})^{2}} \left[ c_{2}(\Delta r^{2} + r^{2}) + c_{0} \right] \right\} \right|_{0}^{r_{s}} \right\}$$
(25)

The image part of  $\Omega^3 \log(1-\frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}$  is  $\left[(\Omega_r^3 - 3\Omega_r\Omega_i^2)\pi + (3\Omega_i\Omega_r^2 - \Omega_i^3)\log\left|1 - \frac{1}{\Omega}\right| + 2\Omega_r\Omega_i + \frac{1}{2}\Omega_i\right]$ , and the image part of  $-\frac{\Omega}{\Omega-1} + \Omega + \Omega^2 log(1-\frac{1}{\Omega})$  is  $\frac{\Omega_i}{(\Omega_r-1)^2 + \Omega_i^2} + \Omega_i + (\Omega_r^2 - \Omega_i^2)\pi + 2\Omega_i\Omega_r log\left|1 - \frac{1}{\Omega}\right|$ . With  $\Omega_i = 0$ , it becomes  $\pi\Omega_r^3$  and  $\pi\Omega_r^2$ . The image part of the dispersion relation is:

$$\frac{\Omega_r}{\Omega_A} = \frac{\varepsilon_0^2}{\varepsilon_s^2} \beta_0 \Omega_r^2 \pi \{ \Delta r^2 [e^{-(\frac{r}{\Delta r})^2}] |_0^{r_s} - \frac{2\rho_h}{\varepsilon_0} \Omega_r \{ e^{-(\frac{r}{\Delta r})^2} [c_2(\Delta r^2 + r^2) + c_0] \} |_0^{r_s} \}$$
(26)

The critical  $\beta_0^{crit}$  given by

$$\beta_0^{crit} = \frac{\varepsilon_s^2}{\pi \Omega_A \Omega_r \varepsilon_0^2} \frac{1}{\Delta r^2 [e^{-(\frac{r}{\Delta r})^2}]_0^{r_s} - \frac{2\rho_h}{\varepsilon_0} \Omega_r \{e^{-(\frac{r}{\Delta r})^2} [c_2(\Delta r^2 + r^2) + c_0]\}_0^{r_s}}$$
(27)

The real part of  $\Omega^3 \log(1-\frac{1}{\Omega}) + \Omega^2 + \frac{1}{2}\Omega + \frac{1}{3}$  is the dispersion relation is  $:(\Omega_r^3 - 3\Omega_r\Omega_i^2)ln(\frac{1}{\Omega_r} - 1) - \pi(3\Omega_r^2\Omega_i - \Omega_i^3) + \Omega_r^2 - \Omega_i^2 + \frac{1}{2}\Omega_r + \frac{1}{3}$ , and the real part of  $-\frac{\Omega}{\Omega-1} + \Omega + \Omega^2 log(1-\frac{1}{\Omega})$  is  $\frac{-\Omega_r^2 + \Omega_r - \Omega_i^2}{(\Omega_r - 1)^2 + \Omega_i^2} + \Omega_r + (\Omega_r^2 - \Omega_i^2)ln(\frac{1}{\Omega_r} - 1) - 2\pi\Omega_r\Omega_i$ . With  $\Omega_i = 0$ , The read part of dispersion relation is:

$$0 = \left[\frac{-\Omega_r^2 + \Omega_r}{(\Omega_r - 1)^2} + \Omega_r + \Omega_r^2 ln(\frac{1}{\Omega_r} - 1)\right] \Delta r^2 \left[e^{-(\frac{r}{\Delta r})^2}\right] \Big|_0^{r_s} - \frac{2\rho_h}{\varepsilon_0} \left[\Omega_r^3 ln(\frac{1}{\Omega_r} - 1) + \Omega_r^2 + \frac{1}{2}\Omega_r + \frac{1}{3}\right] \left\{e^{-(\frac{r}{\Delta r})^2} \left[c_2(\Delta r^2 + r^2) + c_0\right]\right\} \Big|_0^{r_s}$$
(28)

Using  $q = 0.8 + 1.385r^2$ ,  $\Delta r = 0.2$ , and  $\beta = 0.01$ ,  $p_0 = 2.8083 \times 10^3 pascal$ , and  $\xi_0/a = 0.01$  compare the analytical results to dwk + + results (shown in Fig. 1). For  $imag(\Omega) = 0.005$ , the real frequency can be found by the code:  $\Omega_r = 0.8472, (0.29137\omega_A, 160.965kHz), \beta_{0,crit} = 0.0402$  (without  $\omega$  term,  $\Omega_r = 0.8901, \beta_{0,crit} = 0.0333$ ). And by Eq. 27,  $\beta_{0,crit} = 0.04130$  (without  $\omega$  term,  $\beta_{0,crit} = 0.03442$ ).

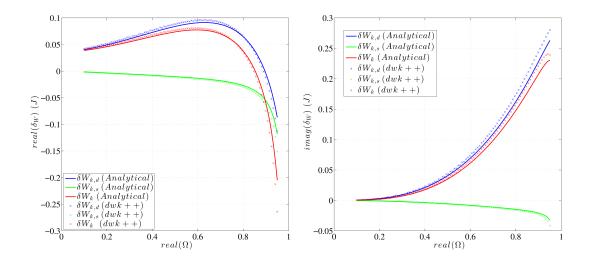


Figure 1: Comparison between dwk + + results and analysical's.

# 5.2 Appendix

For more details about formula derivation, please read  $\lim_{k \to \infty} \frac{1}{k} \int_{\mathbb{R}^n} \frac{dx}{dx} dx dx$ .