

Conserved current and Noether theorem

Noether theorem: every continuous global symmetry of the action leads to a conserved current and thus a conserved charge, for solutions of the equations of motion.

Start from $S = \int d^4x \mathcal{L}[\varphi, \partial_\mu \varphi]$, $\mathcal{L} \equiv \mathcal{L}[\varphi(x), \partial_\mu \varphi(x)]$.

then $\delta S = \int d^4x \delta_0 \mathcal{L} = \int d^4x \left[\frac{\partial \mathcal{L}}{\partial \varphi} \delta_0 \varphi + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \delta_0 (\partial_\mu \varphi) \right] \dots \textcircled{1}$

here I write $\delta_0 \mathcal{L}$ instead of $\delta \mathcal{L}$ to emphasize the fact that S is not a function of x , and $\delta_0 \mathcal{L} = \mathcal{L}[\varphi'(x), \partial_\mu \varphi'(x)] - \mathcal{L}[\varphi(x), \partial_\mu \varphi(x)]$

just continue with $\textcircled{1}$, we get.

$\delta S = \int d^4x \left[\frac{\partial \mathcal{L}}{\partial \varphi} \delta_0 \varphi + \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \delta_0 \varphi \right) - \left(\partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \right) \delta_0 \varphi \right] \dots \textcircled{2}$

note that here the relation $\partial_\mu (\delta_0 \varphi) = \partial_\mu (\varphi'(x) - \varphi(x)) = \partial_\mu \varphi'(x) - \partial_\mu \varphi(x) = \delta_0 (\partial_\mu \varphi(x))$ is used.

the first and third term can be combined as $\left[\frac{\partial \mathcal{L}}{\partial \varphi} - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \right) \right] \delta_0 \varphi$.

Since we require that φ is solution of the equation of motion, this combination vanishes.

$\Rightarrow \delta S = \int d^4x \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \delta_0 \varphi \right) \dots \textcircled{3}$

"A global symmetry of the action" means that $\delta_0 \mathcal{L}$ is a total derivative (if $\delta_0 \mathcal{L}$ is not zero) that is,

$\delta_0 \mathcal{L} = \partial_\mu K^\mu$ $\dots \textcircled{4}$

\uparrow does not depend on spacetime \rightarrow total derivative.

Equating ③ and ④ gives $\delta W \int d^4x \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \frac{\delta \varphi}{\delta W} - K^\mu \right) = 0$.

Since δW is an arbitrary constant, then

$$\partial_\mu j^\mu = 0, \text{ where } j^\mu = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \frac{\delta \varphi}{\delta W} - K^\mu$$

Here, of course, by writing $\frac{\delta \varphi}{\delta W}$ we mean that the variation of φ is induced by δW .

If $\mathcal{L} = \mathcal{L}[\varphi_1, \varphi_2, \dots, \varphi_N, \partial_\mu \varphi_1, \partial_\mu \varphi_2, \dots, \partial_\mu \varphi_N]$,

then
$$j^\mu = \sum_{i=1}^N \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi_i)} \frac{\delta \varphi_i}{\delta W} - K^\mu$$

We can define $Q \equiv \int d^3\vec{x} j^0(t, \vec{x})$, then

$$\frac{dQ}{dt} = \int d^3\vec{x} \partial_0 j^0(t, \vec{x}) = \int d^3\vec{x} (\partial_\mu j^\mu - \vec{\nabla} \cdot \vec{j})$$

$$= - \int d^3\vec{x} \vec{\nabla} \cdot \vec{j}$$

\uparrow
 Gauss's theorem \rightarrow this is a surface integral.

If we assume that \vec{j} vanishes at the spatial boundary (actually, not necessarily be space infinity in practice, as long as the current \vec{j} does not leave our experimental apparatus) then

$\frac{dQ}{dt} = 0$, so we get a conserved charge Q (since it does not change with time).

Example (a) Internal transformation of a complex scalar field.

$$\mathcal{L} = \partial_\mu \phi^* \partial^\mu \phi - m^2 \phi^* \phi - \lambda (\phi^* \phi)^2$$

$$\phi(x) \rightarrow \phi'(x) = e^{-i\delta\alpha} \phi(x) \approx \phi(x) - i\delta\alpha \phi(x)$$

$$\phi^*(x) \rightarrow \phi'^*(x) = e^{i\delta\alpha} \phi^*(x) \approx \phi^*(x) + i\delta\alpha \phi^*(x)$$

where $\delta\alpha$ is an infinitesimal real parameter, independent of x

$$\begin{aligned} \text{Since } \delta_0 \mathcal{L} &= [\partial_\mu \phi'^*(x) \partial^\mu \phi'(x) - m^2 \phi'^*(x) \phi'(x) - \lambda (\phi'^*(x) \phi'(x))^2] \\ &\quad - [\partial_\mu \phi^*(x) \partial^\mu \phi(x) - m^2 \phi^*(x) \phi(x) - \lambda (\phi^*(x) \phi(x))^2] \\ &= [e^{i\delta\alpha} \partial_\mu \phi^*(x) e^{-i\delta\alpha} \partial^\mu \phi(x) - m^2 e^{i\delta\alpha} \phi^*(x) e^{-i\delta\alpha} \phi(x) \\ &\quad - \lambda (e^{i\delta\alpha} \phi^*(x) e^{-i\delta\alpha} \phi(x))^2] \\ &\quad - [\partial_\mu \phi^*(x) \partial^\mu \phi(x) - m^2 \phi^*(x) \phi(x) - \lambda (\phi^*(x) \phi(x))^2] \\ &= [\partial_\mu \phi^*(x) \partial^\mu \phi(x) - m^2 \phi^*(x) \phi(x) - \lambda (\phi^*(x) \phi(x))^2] \\ &\quad - [\partial_\mu \phi^*(x) \partial^\mu \phi(x) - m^2 \phi^*(x) \phi(x) - \lambda (\phi^*(x) \phi(x))^2] \\ &= 0 \end{aligned}$$

$$\Rightarrow K^\mu = 0$$

$$\text{Since } \delta_0 \phi = \phi'(x) - \phi(x) = -i\delta\alpha \phi(x)$$

$$\delta_0 \phi^* = \phi'^*(x) - \phi^*(x) = +i\delta\alpha \phi^*(x)$$

$$\begin{aligned} \text{Then let } \delta\omega &= \delta\alpha \Rightarrow \frac{\delta_0 \phi}{\delta\omega} = -i\phi(x) \text{ and } \frac{\delta_0 \phi^*}{\delta\omega} = i\phi^*(x) \\ \Rightarrow j^\mu &= \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \frac{\delta_0 \phi}{\delta\omega} + \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi^*)} \frac{\delta_0 \phi^*}{\delta\omega} \\ &= \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} (-i\phi) + \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi^*)} (i\phi^*) \\ &= (\partial^\mu \phi^*) (-i\phi) + (\partial^\mu \phi) (i\phi^*) = i[(\partial^\mu \phi) \phi^* - (\partial^\mu \phi^*) \phi] \end{aligned}$$

$$\Rightarrow Q = \int d^3\vec{x} j^0 = i \int d^3\vec{x} (\dot{\phi} \phi^* - \dot{\phi}^* \phi)$$

Example (b). Translation of a generic field.

Consider infinitesimal translation $x^\mu \rightarrow x'^\mu = x^\mu - \delta a^\mu$, so that $\delta x^\mu = x'^\mu - x^\mu = -\delta a^\mu$. note that δa^μ is independent of x .

For a scalar field, we have shown that

$$\delta_0 \phi(x) = (\delta a^\mu) \partial_\mu \phi(x)$$

In fact, for a generic field $\psi(x)$, we have
 \rightarrow scalar, vector, spinor etc.

$$\psi'(x') = \psi'(x - a) = \psi(x)$$

This relation expresses the fact that the field has the same value at the same spacetime point, which is (or, each of the field components, for labeled as x in one coordinate and x' in the other coordinate, and the difference of the two coordinates is merely a translation.

$$\Rightarrow \delta_0 \psi(x) = \psi'(x) - \psi(x) = \psi(x+a) - \psi(x) = (\delta a^\mu) \partial_\mu \psi(x)$$

Furthermore, the Lagrangian density itself is a scalar field, that is, $\mathcal{L}'[\psi'(x'), \partial'_\mu \psi'(x')] = \mathcal{L}[\psi(x), \partial_\mu \psi(x)]$, where x' and x are related by a translation, and therefore.

$$\delta_0 \mathcal{L}[\psi(x), \partial_\mu \psi(x)] = \delta a^\mu \partial_\mu \mathcal{L}[\psi(x), \partial_\mu \psi(x)] \equiv \delta a^\mu \partial_\mu \mathcal{L}$$

note that the dependence of \mathcal{L} on x is achieved through the dependence of ψ on x .

$$\text{Let } \delta w = \delta a^\nu$$

$$\Rightarrow K^\mu_\nu = \frac{\partial \mathcal{L}}{\partial a^\nu} \dot{a}^\mu = \dot{a}^\mu_\nu \mathcal{L}$$

since each $\dot{a}^0, \dot{a}^1, \dot{a}^2$ and \dot{a}^3 are independent then $\frac{\partial \mathcal{L}}{\partial a^\nu}$ can be understood as $\frac{\partial \mathcal{L}}{\partial a^\nu} = \dot{a}^\mu_\nu$

$$\frac{\partial_0 \mathcal{L}}{\partial \omega} = \frac{\partial \mathcal{L}}{\partial a^\nu} \partial_\nu \varphi(x) = \dot{a}^\mu_\nu \partial_\nu \varphi(x) = \partial_\nu \varphi(x)$$

$$\Rightarrow j^\mu_\nu = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \partial_\nu \varphi(x) - \delta^\mu_\nu \mathcal{L}$$

Here we notice that the additional index ν for the Noether current is introduced by $\partial \omega$.

For $\mathcal{L} = \mathcal{L}[\varphi_1, \varphi_2, \dots, \varphi_N, \partial_\mu \varphi_1, \partial_\mu \varphi_2, \dots, \partial_\mu \varphi_N]$, we get

$$j^\mu_\nu = \sum_{i=1}^N \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi_i)} \partial_\nu \varphi_i - \delta^\mu_\nu \mathcal{L}$$

$$\Rightarrow Q_\nu = \int d^3\vec{x} j^0_\nu = \int d^3\vec{x} \left(\sum_{i=1}^N \frac{\partial \mathcal{L}}{\partial (\partial_0 \varphi_i)} \partial_\nu \varphi_i - \delta^0_\nu \mathcal{L} \right)$$

Recall that $\pi_i = \frac{\partial \mathcal{L}}{\partial (\partial_0 \varphi_i)}$

$$\text{then } Q_\nu = \int d^3\vec{x} \left[\sum_{i=1}^N \pi_i \partial_\nu \varphi_i - \delta^0_\nu \mathcal{L} \right]$$

$$\text{where } Q_0 = \int d^3\vec{x} \left[\sum_{i=1}^N \pi_i \partial_0 \varphi_i - \mathcal{L} \right]$$

$$\stackrel{?}{=} \int d^3\vec{x} \mathcal{H} = H$$

Recall that $\mathcal{H}[\pi_1, \pi_2, \dots, \pi_N, \varphi_1, \varphi_2, \dots, \varphi_N] = \sum_{i=1}^N \pi_i \partial_0 \varphi_i - \mathcal{L}$

$$Q_i = \int d^3\vec{x} \sum_{j=1}^N \pi_j \partial_i \varphi_j$$

Since Q_0 is the Hamiltonian H , and (Q^0, Q^1, Q^2, Q^3) form a four-vector, it is natural to call $Q_i \equiv P_i$, and j^μ_ν is actually just the energy-momentum tensor, usually denoted as T^μ_ν .

$\frac{dQ_\mu}{dt} = 0$ means we get energy and momentum conservation from the translation symmetry of the action.

$Q_0 \qquad Q_i$

In the case of a free real scalar field,

$$\mathcal{L} = \frac{1}{2} \partial_\nu \phi \partial^\nu \phi - \frac{1}{2} m^2 \phi^2$$

the energy-momentum tensor T^μ_ν is

$$\begin{aligned} T^\mu_\nu &= \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \partial_\nu \phi - \delta^\mu_\nu \mathcal{L} \\ &= \partial^\mu \phi \partial_\nu \phi - \delta^\mu_\nu \left(\frac{1}{2} \partial_\alpha \phi \partial^\alpha \phi - \frac{1}{2} m^2 \phi^2 \right) \end{aligned}$$

$$\Rightarrow T^{\mu\nu} = g^{\nu\beta} T^\mu_\beta = \partial^\mu \phi \partial^\nu \phi - g^{\mu\nu} \left(\frac{1}{2} \partial_\alpha \phi \partial^\alpha \phi - \frac{1}{2} m^2 \phi^2 \right)$$

Apparently, $T^{\mu\nu} = T^{\nu\mu}$

However, it's not always automatically that $T^{\mu\nu} = T^{\nu\mu}$, which is a property we need in the case for example in General Relativity ($R^{\mu\nu} - \frac{1}{2} g^{\mu\nu} R = 8\pi G T^{\mu\nu}$). Nevertheless, there is typically a way to massage the energy-momentum tensor of any theory into a symmetric form by adding an extra term

$$T^{\mu\nu} \longrightarrow T^{\mu\nu}_{\text{new}} = T^{\mu\nu} + \partial_\rho N^{\rho\mu\nu},$$

where $N^{\rho\mu\nu}$ is some function of the fields that is anti-symmetric in the first two indices so $N^{\rho\mu\nu} = -N^{\mu\rho\nu}$, and therefore

$$\partial_\mu \partial_\rho N^{\rho\mu\nu} = -\partial_\mu \partial_\rho N^{\mu\rho\nu} = -\partial_\rho \partial_\mu N^{\mu\rho\nu} = -\partial_\rho \partial_\rho N^{\rho\mu\nu} = 0$$

$$\Rightarrow \partial_\mu T^{\mu\nu}_{\text{new}} = \partial_\mu T^{\mu\nu} + \partial_\mu \partial_\rho N^{\rho\mu\nu} = 0 + 0 = 0.$$

So, $T^{\mu\nu}_{\text{new}}$ is also conserved, i.e., $\frac{d}{dt} \int d^3\vec{x} T^{\mu\nu}_{\text{new}} = 0$

Example (c). Lorentz transformation of a free real scalar field.

As shown before, $\chi'^{\mu} = \Lambda^{\mu}_{\nu} \chi^{\nu} = (\delta^{\mu}_{\nu} + \epsilon^{\mu}_{\nu}) \chi^{\nu}$

$$\Rightarrow \delta \chi^{\mu} = \epsilon^{\mu}_{\nu} \chi^{\nu} = \epsilon^{\mu\nu} \chi_{\nu}$$

$$\phi'(x') = \phi(x) \Rightarrow \delta_0 \phi(x) = \phi'(x) - \phi(x) = -\epsilon^{\mu}_{\nu} \chi^{\nu} \partial_{\mu} \phi(x) \\ = -\epsilon^{\mu\nu} \chi_{\nu} \partial_{\mu} \phi(x)$$

Since the Lagrangian density itself is a scalar field, then

$$\delta_0 \mathcal{L}(x) = -\epsilon^{\mu\nu} \chi_{\nu} \partial_{\mu} \mathcal{L}(x)$$

where the dependence of \mathcal{L} on x is achieved through the dependence of ϕ on x . In fact, the relation $\delta_0 \mathcal{L}(x) = -\epsilon^{\mu\nu} \chi_{\nu} \partial_{\mu} \mathcal{L}(x)$ is valid for any Lagrangian as a function of generic fields (scalar, vector, spinor etc.) for Lorentz transformation, because Lagrangian ^{density} as a whole behaves as a scalar field, which is just a single number for a spacetime point.

Since $\partial_{\mu} \chi_{\nu} = g_{\mu\nu}$, we have

$$\delta_0 \mathcal{L}(x) = -\left[\epsilon^{\mu\nu} \partial_{\mu} (\chi_{\nu} \mathcal{L}(x)) - (\epsilon^{\mu\nu} \partial_{\mu} \chi_{\nu}) \mathcal{L}(x) \right] \\ = -\epsilon^{\mu\nu} \partial_{\mu} (\chi_{\nu} \mathcal{L}(x)) + \underbrace{\epsilon^{\mu\nu} g_{\mu\nu} \mathcal{L}(x)}_0 \\ = -\epsilon^{\mu\nu} \partial_{\mu} (\chi_{\nu} \mathcal{L}(x))$$

Let $\delta\omega = \epsilon^{\rho\sigma}$, then

$$K^{\mu}_{\rho\sigma} = \frac{-\epsilon^{\mu\nu} \chi_{\nu} \mathcal{L}}{\epsilon^{\rho\sigma}} = -(\delta^{\mu}_{\rho} \delta^{\nu}_{\sigma} - \delta^{\mu}_{\sigma} \delta^{\nu}_{\rho}) \chi_{\nu} \mathcal{L} \\ \uparrow \text{again, } \frac{\epsilon^{\mu\nu}}{\epsilon^{\rho\sigma}} \text{ is understood as } \frac{\partial \epsilon^{\mu\nu}}{\partial \epsilon^{\rho\sigma}} \\ = -(\delta^{\mu}_{\rho} \chi_{\sigma} - \delta^{\mu}_{\sigma} \chi_{\rho}) \mathcal{L}$$

$$\frac{\delta_0 \phi}{\delta\omega} = -(\delta^{\mu}_{\rho} \delta^{\nu}_{\sigma} - \delta^{\mu}_{\sigma} \delta^{\nu}_{\rho}) \chi_{\nu} \partial_{\mu} \phi = -(\delta^{\mu}_{\rho} \chi_{\sigma} - \delta^{\mu}_{\sigma} \chi_{\rho}) \partial_{\mu} \phi \\ = \chi_{\rho} \partial_{\sigma} \phi - \chi_{\sigma} \partial_{\rho} \phi$$

$$\Rightarrow j^{\mu}_{\rho\sigma} = (\partial^{\mu} \phi)(\chi_{\rho} \partial_{\sigma} \phi - \chi_{\sigma} \partial_{\rho} \phi) + (\delta^{\mu}_{\rho} \chi_{\sigma} - \delta^{\mu}_{\sigma} \chi_{\rho}) \left(\frac{1}{2} \partial_{\lambda} \phi \partial^{\lambda} \phi - \frac{1}{2} m^2 \phi^2 \right)$$

$$= \chi_e T^\mu_\sigma - \chi_\sigma T^\mu_e = -j^\mu_{\sigma e}$$

$$\Rightarrow Q_{e\sigma} = \int d^3\vec{x} j^\sigma_{e\sigma} = \int d^3\vec{x} (\chi_e T^\sigma_\sigma - \chi_\sigma T^\sigma_e) = -Q_{\sigma e}$$

$$\Rightarrow Q^{ij} = \int d^3\vec{x} (\chi^i T^{oj} - \chi^j T^{oi}) = -Q^{ji}$$

$$\text{Since } \int d^3\vec{x} T^{oi} = p^i,$$

then $\frac{dQ^{ij}}{dt} = 0$ means orbital angular momentum conservation.

$$\text{Also, } Q^{oi} = \int d^3\vec{x} (\chi^o T^{oi} - \chi^i T^{oo}) = \int d^3\vec{x} T^{oi} - \int d^3\vec{x} \chi^i T^{oo} \\ = p^i - \int d^3\vec{x} \chi^i T^{oo} = -Q^{io}$$

$$\Rightarrow 0 = \frac{dQ^{oi}}{dt} = p^i + \underbrace{\frac{dp^i}{dt}}_0 - \frac{d}{dt} \int d^3\vec{x} \chi^i T^{oo}$$

$$\Rightarrow \text{constant} = p^i = \frac{d}{dt} \int d^3\vec{x} \chi^i T^{oo}$$

Since $\int d^3\vec{x} T^{oo} = H$, then the above expression means that the center of energy of the field travels with a constant velocity, i.e., we can define the center of energy as

$$\chi^i_{cm} = \frac{\int d^3\vec{x} \chi^i T^{oo}}{\int d^3\vec{x} T^{oo}} = \frac{\int d^3\vec{x} \chi^i T^{oo}}{H}$$

$$\Rightarrow \frac{d\chi^i_{cm}}{dt} = v^i_{cm} = \frac{\frac{d}{dt} \int d^3\vec{x} \chi^i T^{oo}}{H} = \frac{p^i}{H} = \text{constant}$$

This is kind of like a field theoretic version of Newton's first law.

In sum, for translation, we get for generic fields,

$$T^\mu_\nu = j^\mu_\nu = \sum_{i=1}^N \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_i)} \partial_\nu \phi_i - \delta^\mu_\nu \mathcal{L}$$

for Lorentz transformation, we get for generic fields,

$$j^\mu_{\sigma\sigma} = \sum_{i=1}^N \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_i)} \frac{\partial \phi_i}{\partial x^\sigma} + (\delta^\mu_\sigma \chi_\sigma - \delta^\mu_\sigma \chi_\sigma) \mathcal{L}, \text{ where } \delta x^\mu = \epsilon^{\mu\nu} \chi_\nu.$$