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Instituto de Física  
Universidade de São Paulo

# Magneto-hidrodinâmica

II Escola Jayme Tiomno de Física Teórica  
Prof. Rafael Rechiche de CAMPOS  
22 a 31 de julho de 2019  
Níckolas de Aguiar ALVES



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II Escola Jayme Tiomno de Física Teórica

**Professor:** Rafael Rechiche de CAMPOS, IAG-USP  
**Notas por:** Níckolas de Aguiar ALVES  
**Nível:** Graduação  
**Período:** 22 a 31 de julho de 2019



São Paulo  
22 a 31 de julho de 2019



II Escola Jayme Tiomno de Física Teórica  
Magne-to-hidrodinâmica - EJT2301  
Rafael Rechiche de Campos

Nicholas Alves  
2019



# Magneto-hidrodinâmica

## Referências

1. Fundamentals of Plasma Physics  
J. A. Bittencourt

2. Notas de Aula - Elizabete G. Dal Pino (Site)  
• Cálculo Numérico  
• MHD

Geodbloed  
Principles of MHD

Somov  
Plasma Astrophysics

## Introdução ao Espaço de Fase e a Equação de Boltzmann

Espaço de fase: espaço 6D  $\hat{x}, \hat{y}, \hat{z}, \hat{v}_x, \hat{v}_y, \hat{v}_z$

O elemento de volume no espaço de fase é

$$d^3r d^3v = d\vec{r} d\vec{v} = dx dy dz dv_x dv_y dv_z$$

Função de distribuição no EF  $\rightarrow$  Como se distribuem as partículas no EF

densidade das partículas

$$F_\alpha(\vec{r}, \vec{v}, t) = \frac{dN(\vec{r}, \vec{v}, t)}{d^3r d^3v}$$

número de partículas em um volume no espaço de fase

diferentes tipos de partícula

$$F_\alpha > 0$$

$$F_\alpha \rightarrow 0 \text{ para } \vec{v} \rightarrow +\infty$$

tratamento não-relativístico

do contrário,  $F_\alpha = 0$  para  $v > c$

Algumas grandezas macroscópicas que vem de  $F_\alpha$ :

$$n_\alpha(\vec{r}, t) = \int \frac{dN(\vec{r}, \vec{v}, t)}{d^3r} = \int \frac{F_\alpha(\vec{r}, \vec{v}, t) d^3r d^3v}{d^3r} = \int F_\alpha(\vec{r}, \vec{v}, t) d^3v$$

densidade numérica média

massa da partícula

$$m_\alpha n_\alpha(\vec{r}, t) = \rho_\alpha(\vec{r}, t)$$

densidade de massa

$$\begin{aligned} n_\alpha \cdot \vec{u}_\alpha(\vec{r}, t) &= \int \frac{\vec{v}_\alpha dN_\alpha}{d^3r} = \int F_\alpha(\vec{r}, \vec{v}, t) \vec{v}_\alpha d^3v \\ \vec{u}_\alpha(\vec{r}, t) &= \frac{\int \vec{v}_\alpha F_\alpha(\vec{r}, \vec{v}, t) d^3v}{\int F_\alpha(\vec{r}, \vec{v}, t) d^3v} \end{aligned}$$

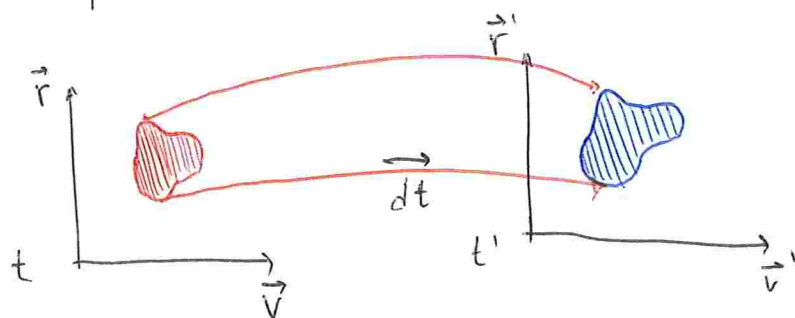
Equação de Boltzmann

descreve a evolução temporal de  $F_\alpha(\vec{r}, \vec{v}, t)$

3 variáveis nos interessam

$$\begin{cases} t' = t + dt \\ \vec{r}' = \vec{r} + \vec{v} dt \\ \vec{v}' = \vec{v} + \vec{a} dt \end{cases}$$

No espaço de fase, a evolução temporal pode ser vista na forma



O que se conserva na passagem de  $t$  para  $t'$ ?

$$dN(\vec{r}, \vec{v}, t) = dN'(\vec{r}', \vec{v}', t')$$

$$F'_\alpha = F_\alpha$$

ou seja

$$F_\alpha(\vec{r}, \vec{v}, t) d^3r d^3v = F'_\alpha(\vec{r}', \vec{v}', t') d^3r' d^3v'$$

Podemos dizer que

$$d^3r d^3v = \underbrace{|J|}_1 d^3r' d^3v'$$

A diferença das distribuições para diferentes tempos é então

$$[F_\alpha(\vec{r}, \vec{v}, t) - F_\alpha(\vec{r}', \vec{v}', t')] d^3r d^3v = 0$$

Abriremos  $F_\alpha(\vec{r}', \vec{v}', t')$  em Taylor series

$$\begin{aligned} F_\alpha(\vec{r}', \vec{v}', t') &= F_\alpha(\vec{r} + \vec{v} dt, \vec{v} + \vec{a} dt, t + dt) \\ &= F_\alpha(\vec{r}, \vec{v}, t) + \left[ \frac{\partial F_\alpha}{\partial t} + \frac{\partial F_\alpha}{\partial x} \frac{\partial x}{\partial t} + \frac{\partial F_\alpha}{\partial y} \frac{\partial y}{\partial t} + \frac{\partial F_\alpha}{\partial z} \frac{\partial z}{\partial t} \right. \\ &\quad \left. + \frac{\partial F_\alpha}{\partial v_x} \frac{\partial v_x}{\partial t} + \frac{\partial F_\alpha}{\partial v_y} \frac{\partial v_y}{\partial t} + \frac{\partial F_\alpha}{\partial v_z} \frac{\partial v_z}{\partial t} \right] dt \end{aligned}$$

$v_z$

$a_z$



$$F_\alpha(\vec{r}', \vec{v}', t') = F_\alpha(\vec{r}, \vec{v}, t) + \left[ \frac{\partial F_\alpha}{\partial t} + (\vec{v} \cdot \nabla) F_\alpha + (\vec{a} \cdot \nabla_{\vec{v}}) F_\alpha \right] dt$$

$$\hookrightarrow \nabla_{\vec{v}} = \hat{v}_x \frac{\partial}{\partial v_x} + \hat{v}_y \frac{\partial}{\partial v_y} + \hat{v}_z \frac{\partial}{\partial v_z}$$

Como  $F_\alpha(\vec{r}, \vec{v}, t) - F_\alpha(\vec{r}', \vec{v}', t') = 0$ , temos

$$\frac{\partial F_\alpha}{\partial t} + (\vec{v} \cdot \nabla) F_\alpha + (\vec{a} \cdot \nabla_{\vec{v}}) F_\alpha = 0$$

caso particular da Equação de Boltzmann

Equação de Boltzmann Não-Colisional

Outra forma de escrever a mesma equação é

$$\frac{D F_\alpha(\vec{r}, \vec{v}, t)}{Dt} = 0$$

$$\frac{D}{Dt} = \frac{\partial}{\partial t} + \vec{v} \cdot \nabla + \vec{a} \cdot \nabla_{\vec{v}}$$

A Equação de Boltzmann geral é

$$F_\alpha d^3r d^3v \neq F'_\alpha d^3r' d^3v'$$

$$\frac{\partial F_\alpha}{\partial t} + (\vec{v} \cdot \nabla) F_\alpha + (\vec{a} \cdot \nabla_{\vec{v}}) F_\alpha = \left[ \frac{\partial F_\alpha}{\partial t} \right]_{\text{col}}$$

Velocidades total, média e aleatória

$$\vec{v}_\alpha = \vec{u}_\alpha + \tilde{\vec{v}}_\alpha$$

Energia interna

$$E = \int \frac{\tilde{v}_\alpha^2}{2} f_\alpha(\vec{r}, \vec{v}, t) d^3v$$

densidade de energia por massa

Comentário

$$f_\alpha(\vec{r}, \vec{v}, t) d^3v = n(\vec{r}, t) e^{-\frac{m \tilde{v}^2}{2k_B T}} \left[ \frac{m}{2\pi k_B T} \right]^{3/2}$$

$$E = \int \frac{\tilde{v}_\alpha^2}{2} m f_\alpha(\vec{r}, \vec{v}, t) d^3v = \frac{3}{2} k_B T$$

Equações de Conservação e Momentos da Equação de Boltzmann

$$\int V_\alpha^k \left( \frac{\partial f_\alpha}{\partial t} + (\vec{v} \cdot \nabla) f_\alpha + (\vec{a} \cdot \nabla) f_\alpha \right) d^3v$$

O momento de ordem zero

$$\left[ \frac{\partial f_\alpha}{\partial t} \right]_{\text{col}} = 0$$

Termos

$$0 = \underbrace{\int \frac{\partial f}{\partial t} dv_i}_{\text{I}} + \underbrace{\int v_i \frac{\partial f}{\partial r_i} dv_i}_{\text{II}} + \underbrace{\int a_i \frac{\partial f}{\partial v_i} dv_i}_{\text{III}}$$

Parte 1:

$$\frac{\partial}{\partial t} \int f_\alpha dv_i = \frac{\partial n_\alpha(\vec{r}, t)}{\partial t}$$

Parte 2:

$$v_i \frac{\partial f}{\partial r_i} = \frac{\partial}{\partial r_i} (v_i f)$$

$$\frac{\partial}{\partial r_i} \int v_i f dv_i = \frac{\partial}{\partial r_i} (u_i n)$$

Parte 3:

$$\int a_i \frac{\partial f}{\partial v_i} dv_i = \int \frac{\partial (f a_i)}{\partial v_i} dv_i = [f a_i]_{-\infty}^{+\infty} = 0$$

Logo,

$$0 = \frac{\partial n_\alpha}{\partial t} + \frac{\partial}{\partial r_i} (n_\alpha u_i)$$

Em forma vetorial,

$$0 = \frac{\partial n}{\partial t} + \nabla \cdot (\vec{u} \cdot \vec{n})$$

termo  
fonte

ou  
servidouro

$$0 = \frac{\partial p}{\partial t} + \nabla \cdot (\vec{u} \cdot p)$$

O momento de ordem um

$$\int v_j \left[ \frac{\partial f_\alpha}{\partial t} + v_i \frac{\partial f_\alpha}{\partial r_i} + a_i \frac{\partial f_\alpha}{\partial v_i} \right] d^3 v = 0$$

$$\int v_j \frac{\partial f_\alpha}{\partial t} d^3 v + \int v_i v_j \frac{\partial f_\alpha}{\partial r_i} d^3 v + \int a_i v_j \frac{\partial f_\alpha}{\partial v_i} d^3 v = 0$$

(I)                      (II)                      (III)

Parte 1:

$$\int v_j \frac{\partial f_\alpha}{\partial t} d^3 v = \frac{\partial}{\partial t} \int v_j f_\alpha d^3 v = \frac{\partial}{\partial t} (u_j n)$$

Parte 2:

$$\begin{aligned} \int v_i v_j \frac{\partial f_\alpha}{\partial r_i} d^3 v &= \frac{\partial}{\partial r_i} \int v_i v_j f d^3 v \\ &= \frac{\partial}{\partial r_i} \int (u_j + \tilde{v}_j)(u_i + \tilde{v}_i) f d^3 v \\ &= \frac{\partial}{\partial r_i} \left[ \int u_i u_j f d^3 v + \int (\tilde{v}_i u_j + u_j \tilde{v}_i) f d^3 v + \int \tilde{v}_i \tilde{v}_j f d^3 v \right] \end{aligned}$$

$\underbrace{\frac{1}{m} P_{ij}}_{\text{pressão}}$

Comentário

$P_{ij} = \int \tilde{v}_i \tilde{v}_j f d^3 v$  me  
se a pressão for isotrópica podemos dizer que  
os únicos termos não-nulos são  $i=j$

$$P = \sum_{ij} P_{ij}$$

$$P_{ii} = m \int \tilde{v}_i \tilde{v}_i f d^3 v = m \int \tilde{v}_i^2 f d^3 v = m \int \tilde{v}_3^2 f d^3 v$$

$$= m \frac{1}{3} \int \tilde{v}_i \tilde{v}_i f d^3 v = P_{ij} \delta_{ij} = P$$

$$P_{ii} = \int \frac{\tilde{v}_i^2}{3} m f d^3 v = \frac{2}{3} E m$$

A integral II se torna

$$\frac{\partial (P_{ij} \delta_{ij})}{\partial r_i} + \frac{\partial}{\partial r_i} \left[ \underbrace{\int u_i u_i f d^3 v}_{u_i u_i n} + \underbrace{\int (\tilde{v}_i u_i + u_i \tilde{v}_i) f d^3 v}_{2 u_i \int \tilde{v}_i f d^3 v = 0} \right]$$

$$\frac{\partial (P_{ij} \delta_{ij})}{\partial r_i} + \frac{\partial}{\partial r_i} (u_i u_i n)$$

Para a parte 3,

$$\int v_j a_i \frac{\partial f}{\partial v_i} d^3 v = a_i \int v_j \frac{\partial f}{\partial v_i} d^3 v = -a_i \delta_{ij} \cdot n$$

$$\frac{\partial (u v_i)}{\partial v_i} + \frac{\partial f}{\partial v_i} v_i + f \frac{\partial v_i}{\partial v_i}$$

Nova equação de conservação  $\rightarrow$  momento

$$\frac{\partial}{\partial t} (n u_i) + n \frac{\partial}{\partial r_i} (u_i n) = - \frac{\partial (P_{ij} \delta_{ij})}{\partial r_i} + n a_i \delta_{ij}$$

$$\frac{\partial}{\partial t} (\rho \vec{u}) + (\vec{u} \cdot \nabla) (\vec{u} \cdot \rho) = -\nabla P + \vec{a} \rho$$

Momento de Boltzmann de ordem dois

$$\int \frac{v_i^2}{2} \left[ \frac{\partial f}{\partial t} + v_i \frac{\partial f}{\partial r_i} + a_i \frac{\partial f}{\partial v_i} \right] d^3 v = 0$$

$$\underbrace{\frac{1}{2} \int v_i^2 \frac{\partial f}{\partial t} d^3 v}_{\text{I}} + \underbrace{\frac{1}{2} \int v_i^2 v_i \frac{\partial f}{\partial r_i} d^3 v}_{\text{II}} + \underbrace{\frac{1}{2} \int a_i v_i^2 \frac{\partial f}{\partial v_i} d^3 v}_{\text{III}} = 0$$

Para a parte I

$$\frac{1}{2} \int v_i^2 \frac{\partial f}{\partial t} d^3 v = \frac{1}{2} \frac{\partial}{\partial t} \int v_i^2 f d^3 v$$

$$\begin{aligned}
 \frac{1}{2} \int v_i^2 \frac{\partial f_\alpha}{\partial t} d^3 v &= \frac{1}{2} \frac{\partial}{\partial t} \int (u_i + \tilde{v}_i)(u_i + \tilde{v}_i) f_\alpha d^3 v \\
 &= \frac{1}{2} \frac{\partial}{\partial t} \int (u_i^2 + \overbrace{2 u_i \tilde{v}_i}^0 + \tilde{v}_i^2) f_\alpha d^3 v \\
 &= \frac{1}{2} \frac{\partial}{\partial t} [n u_i^2 + 2 E]
 \end{aligned}$$

Quanto a parte  $Z$ ,

$$\frac{1}{2} \int v_i^2 v_i \frac{\partial f_\alpha}{\partial r_i} d^3 v = \frac{1}{2} \frac{\partial}{\partial r_i} \int (\underbrace{u_i^2 v_i}_i + 2 \underbrace{u_i \tilde{v}_i v_i}_{ii} + \underbrace{\tilde{v}_i^2 v_i}_{iii}) f_\alpha d^3 v$$

Para a parte  $i$ ,

$$\frac{1}{2} \frac{\partial}{\partial r_i} \int u_i^2 v_i f d^3 v = \frac{1}{2} \frac{\partial}{\partial r_i} (u_i^2 u_i n)$$

Para a parte  $ii$ ,

$$\begin{aligned}
 \frac{2}{2} \frac{\partial}{\partial r_i} \int u_i \tilde{v}_i v_i f_\alpha d^3 v &= \frac{2}{2} \frac{\partial}{\partial r_i} \left( u_i \int (\tilde{v}_i u_i + \tilde{v}_i \tilde{v}_i) f d^3 v \right), \\
 &= \frac{2}{2} \frac{\partial}{\partial r_i} \left( u_i u_i \underbrace{\int \tilde{v}_i f d^3 v}_0 + u_i \underbrace{\int \tilde{v}_i \tilde{v}_i f d^3 v}_{\frac{1}{m} P_{ij}} \right), \\
 &= \frac{2}{2} \frac{\partial}{\partial r_i} \left( u_i P_{ij} \right).
 \end{aligned}$$

Para a parte  $iii$ ,

$$\begin{aligned}
 \frac{1}{2} \frac{\partial}{\partial r_i} \int \tilde{v}_i^2 v_i f_\alpha d^3 v &= \frac{1}{2} \frac{\partial}{\partial r_i} \left( u_i \int \tilde{v}_i^2 f_\alpha d^3 v + \overbrace{\int \tilde{v}_i^2 \tilde{v}_i f_\alpha d^3 v}^0 \right) \\
 &= \frac{1}{2} \frac{\partial}{\partial r_i} (\tilde{v}_i^2 u_i n)
 \end{aligned}$$

também podemos escrever como  $E$

Finalmente, quanto à parte 3

$$\frac{1}{2} \int a_i v_i^2 \frac{\partial \rho}{\partial v_i} d^3v$$

$$v_i^2 = v_i v_i \quad v_i^2 = v_i v_i = v_i^2$$

$$\frac{\partial v_i^2}{\partial v_i} = 2 v_i$$

$$\frac{\partial \rho}{\partial v_i} v_i^2 = \frac{\partial (\rho v_i^2)}{\partial v_i} - \rho \frac{\partial v_i^2}{\partial v_i}$$

$$\frac{1}{2} \int a_i \left( \underbrace{\frac{\partial (\rho v_i^2)}{\partial v_i}}_0 - 2 \rho v_i \right) d^3v = - \frac{1}{2} a_i u_i n$$

Assim, a equação de conservação de energia fica

$$\frac{\partial}{\partial t} \left( \frac{u^2 n}{2} + E \right) + \frac{\partial}{\partial r_i} \left( n u_i \left( \frac{u^2}{2} + \frac{\tilde{v}_i^2}{2} \right) \right) = a_i n u_i - \frac{\partial}{\partial r_i} \left( u_i \frac{P_{ij}}{m} \right)$$

$$\frac{\partial}{\partial t} \left( \rho \left( \frac{u^2}{2} + E \right) \right) + \frac{\partial}{\partial r_i} \left( \rho u_i \left( \frac{u^2}{2} + E \right) \right) = a_i \rho u_i - \frac{\partial}{\partial r_i} \left( u_i P_{ij} \right)$$

$$\rightarrow E \equiv \rho \left( \frac{u^2}{2} + E \right)$$

$$\frac{\partial E}{\partial t} + \nabla \cdot (\vec{u} E) = \vec{a} \cdot \vec{u} \rho - \nabla \cdot (\vec{u} P)$$

Comentário: Referenciais de Euler e Lagrange

derivada material

$$\frac{\partial}{\partial t} + \vec{u} \cdot \nabla = \frac{D}{Dt} = \frac{d}{dt}$$

visão Euleriana

visão Lagrangeana

referencial do laboratório

referencial do fluido



Exemplo: eq. de conservação de massa

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\vec{u} \rho) = \frac{\partial \rho}{\partial t} + \rho \nabla \cdot \vec{u} + \cancel{\rho (\vec{u} \cdot \nabla) \rho}$$

$$\rightarrow (\vec{u} \cdot \nabla) \rho = \vec{u} \cdot \nabla \rho$$

$$\frac{\partial \rho}{\partial t} + (\vec{u} \cdot \nabla) \rho = - \rho \nabla \cdot \vec{u}$$

$$\frac{d\rho}{dt} = - \rho \nabla \cdot \vec{u} \rightarrow \text{forma Lagrangeana}$$

Comentário: forças no plasma  $\rightarrow$  ênfase em plasmas astrofísicos

$$\vec{F} = \begin{cases} -\nabla \Phi \cdot \rho & \rightarrow \text{gravidade} \\ \frac{\vec{J} \times \vec{B}}{4\pi e} & \rightarrow \text{Lorentz} \\ + \rho \nu \left( \nabla^2 \vec{u} + \frac{1}{3} \nabla (\nabla \cdot \vec{u}) \right) & \rightarrow \text{viscosidade} \end{cases}$$

am geral, o próprio plasma compensa efeitos elétricos

$\hookrightarrow$  vamos ignorar

A equação para o momento se torna então

$$\frac{\partial}{\partial t} (\vec{u} \rho) + (\vec{u} \cdot \nabla) (\vec{u} \rho) = - \nabla P - \rho \nabla \Phi + \frac{\vec{J} \times \vec{B}}{4\pi e}$$

Já a equação para a energia

$$\frac{\partial \mathcal{E}}{\partial t} + \nabla \cdot (\vec{u} (\mathcal{E} + P)) = \vec{F} \cdot \vec{u}$$

$\Downarrow$  considerando  $\vec{E}$  e  $\vec{B}$

densidade de energia:

$$\frac{E^2}{8\pi}, \frac{B^2}{8\pi}$$

fluxo de energia:

$$\frac{c}{4\pi} (\vec{B} \times \vec{E})$$

$$\frac{\partial \mathcal{E}}{\partial t} \left( \mathcal{E} + \frac{E^2}{8\pi} + \frac{B^2}{8\pi} \right) + \nabla \cdot \left( \vec{u} [\mathcal{E} + P] + \frac{c}{4\pi} (\vec{B} \times \vec{E}) \right) = \vec{F} \cdot \vec{u}$$

$$\vec{F} = e \left( \vec{E} + \frac{\vec{u} \times \vec{B}}{c} \right) \cdot \frac{1}{V} \rightarrow \text{volume}$$

$$\frac{e}{V} \cdot \vec{u} \cdot \left( \vec{E} + \frac{\vec{u} \times \vec{B}}{c} \right) = \frac{e (\vec{u} \cdot \vec{E})}{V}$$

Vamos analisar as forças sobre os elétrons no sistema

$$n_e m_e \frac{d\vec{u}_e}{dt} = -\nabla p_e + n_e m_e \vec{g} - n_e e (\vec{E} + \frac{\vec{u} \times \vec{B}}{c}) + \vec{P}_{ei}$$

↳ termo associado a pressão colisional entre elétrons e íons

$$\frac{d\vec{u}_e}{dt} \approx \vec{0} \quad \text{e} \quad n_e m_e \vec{g} \approx \vec{0}$$

$$\vec{0} = -\nabla p_e - n_e e (\vec{E} + \frac{\vec{u} \times \vec{B}}{c}) + \vec{P}_{ei}$$

$$\vec{E} = \frac{-\nabla p_e}{n_e e} - \frac{\vec{u} \times \vec{B}}{c} + \frac{\vec{P}_{ei}}{n_e e}$$

Abrindo  $\vec{P}_{ei}$ :

$$\vec{P}_{ei} = n_e n_i \sigma_{ei} u_{th} (\vec{u}_i - \vec{u}_e) m_e$$

Lembrando que

$$\vec{J} = \underbrace{\sum n_i e \vec{u}_i}_{q_i} - \underbrace{n_e e \vec{u}_e}_{q_e}$$

Assumimos que há neutralidade de carga

$$\sum n_i e = n_e e$$

$$\therefore \vec{J} = \sum n_i e (\vec{u}_i - \vec{u}_e)$$

$$\vec{u}_i - \vec{u}_e = \frac{\vec{J}}{\sum n_i e}$$

$$\vec{P}_{ei} = \frac{n_e n_i \sigma_{ei} u_{th} m_e \vec{J}}{\sum n_i e}$$

$$\frac{\vec{P}_{ei}}{n_e e} = \underbrace{\frac{\sigma_{ei} u_{th} m_e}{\sum e^2}}_{\eta} \cdot \vec{J}$$



$$\vec{E} = \eta \vec{J} \quad \rightarrow \quad 1^{\text{a}} \text{ Lei de Ohm}$$

Logo,

$$\vec{E} = -\frac{\nabla p_e}{n_e e} + \eta \vec{J} - \frac{\vec{u} \times \vec{B}}{c}.$$

A equação de energia se torna

$$\frac{\partial}{\partial t} \left( \mathcal{E} + \frac{\vec{E}^2}{8\pi} + \frac{\vec{B}^2}{8\pi} \right) + \nabla \cdot \left( \vec{u}(\mathcal{E} + p) + \frac{c}{4\pi} (\vec{E} \times \vec{B}) \right) = \vec{J} \cdot \left( \frac{\nabla p_e}{n_e e} + \eta \vec{J} - \frac{\vec{u} \times \vec{B}}{c} \right)$$

A equação de indução do campo magnético

$$-\frac{\partial \vec{B}}{\partial t} = \nabla \times \vec{E} = \nabla \times \left[ -\frac{\nabla p_e}{n_e e} + \eta \vec{J} - \frac{\vec{u} \times \vec{B}}{c} \right]$$

$$-\frac{\partial \vec{B}}{\partial t} = \nabla \times \left[ -\frac{\vec{u} \times \vec{B}}{c} + \eta \vec{J} \right]$$

$$\frac{\partial \vec{B}}{\partial t} = \nabla \times \left[ \frac{\vec{u} \times \vec{B}}{c} - \eta \vec{J} \right]$$

Temos então as quatro equações de MHD

$$\bullet \quad \frac{\partial p}{\partial t} + \nabla \cdot (p \vec{u}) = 0$$

$$\bullet \quad \frac{\partial}{\partial t} (p \vec{u}) + (\vec{u} \cdot \nabla) (p \vec{u}) = -\nabla p + \nabla \Phi - \frac{\vec{J} \times \vec{B}}{4\pi}$$

$$\bullet \quad \frac{\partial}{\partial t} \left( \mathcal{E} + \frac{\vec{E}^2}{8\pi} + \frac{\vec{B}^2}{8\pi} \right) + \nabla \cdot \left[ \vec{u}(\mathcal{E} + p) + \frac{c}{4\pi} (\vec{B} \times \vec{E}) \right] = \vec{J} \cdot \left( \frac{\nabla p_e}{n_e e} + \eta \vec{J} - \frac{\vec{u} \times \vec{B}}{c} \right)$$

$$\bullet \quad \frac{\partial \vec{B}}{\partial t} = \nabla \times \left( \frac{\vec{u} \times \vec{B}}{c} \right) - c \nabla \times (\nabla \times (\vec{B} \eta))$$

Na equação de energia, podemos fazer

$$\begin{aligned}\vec{u}_e &= \vec{u}_i - \vec{u}_i + \vec{u}_e \\ &= \vec{u}_i - \frac{\vec{J}}{n_e e}\end{aligned}$$

$$\vec{J} \cdot (\vec{u}_e \times \vec{B}) = \vec{J} \cdot \left( \vec{u}_i \times \vec{B} - \frac{\vec{J} \times \vec{B}}{n_e e} \right)$$

Para reescrever  $\vec{J} \times \vec{B}$ , podemos usar a conservação do momento

$$\rho \frac{d\vec{u}_i}{dt} = -\nabla p_i - \rho \nabla \Phi + \frac{1}{c} \vec{J} \times \vec{B}$$

$$\vec{J} \times \vec{B} = c \left[ \rho \frac{d\vec{u}_i}{dt} + \nabla p_i + \rho \nabla \Phi \right]$$

Juntando à equação de energia,

$$\frac{\rho}{2} \frac{da^2}{dt} + \vec{u} \cdot \nabla p - \rho \nabla \phi - \frac{\vec{J} \cdot \nabla p_e}{n_e e} + \eta J^2$$

$$\rho \frac{dp}{dt} = \frac{\partial(\rho p)}{\partial t} + \nabla \cdot (\rho \vec{u} p)$$

$$\frac{\partial}{\partial t} \left( \rho \left( \frac{u^2}{2} + \phi + \frac{p}{\rho} \right) \right) + \nabla \cdot \left( \rho \vec{u} \left( \frac{u^2}{2} + \phi + \frac{p}{\rho} \right) \right) - \rho \frac{dV}{dt} - \frac{\vec{J} \cdot \nabla p_e}{n_e e} + \eta J^2$$

$$V = \frac{1}{\rho}$$

$$P_i \approx P$$

~~Princípio da~~  
Primeira Lei  
da Termodinâmica

$$dU = dQ - PdV$$

$$\cancel{dQ} = \cancel{dU} + PdV$$

$$dQ = Tds$$

~~Princípio da~~  
Segunda Lei da Termodinâmica

Tomando a variação temporal

$$\frac{dU}{dt} = T \frac{dS}{dt} - P \frac{dV}{dt}$$

Definição de entalpia

$$H = U + P.$$

Aplicando a primeira lei da termodinâmica na relação para a energia

$$\frac{\partial}{\partial t} \left( \rho \left( \frac{1}{2} u^2 + \phi + H \right) \right) + \nabla \cdot \left( \rho \vec{u} \left( \frac{1}{2} u^2 + \phi + H \right) \right) - \rho T \frac{dS}{dt} - \frac{\vec{J} \cdot \nabla P_e}{n_e e} + \eta J^2.$$

A equação de energia total fica

$$\frac{\partial}{\partial t} \left( \frac{E^2}{8\pi} + \frac{B^2}{8\pi} + \rho \left( \frac{1}{2} u^2 + \phi + H \right) \right) + \nabla \cdot \left( \frac{c}{4\pi} (\vec{E} \times \vec{B}) + \rho \vec{u} \left( \frac{1}{2} u^2 + \phi + H \right) \right) = \rho T \frac{dS}{dt} - \eta J^2 + \frac{\vec{J} \cdot \nabla P_e}{n_e e}$$

Tomando que

$$U = \frac{3}{2} \frac{k_B T}{m} \quad e \quad H = \frac{3k_B T}{2m} + \frac{k_B T}{m}$$

$$\frac{\partial}{\partial t} \left( \frac{E^2}{8\pi} + \frac{B^2}{8\pi} + \rho \left( \frac{1}{2} u^2 + \phi + \frac{3k_B T}{2m} \right) \right) + \nabla \cdot \left( \frac{B^2 \vec{u}_e}{4\pi} + \frac{\vec{u}_e \cdot \vec{B}}{4\pi} \vec{B} + \frac{c n}{4\pi} \vec{J} \times \vec{B} \right) + \rho \vec{u} \left( \frac{1}{2} u^2 + \phi + \frac{5k_B T}{2m} \right) - \frac{\vec{J} \cdot \nabla P_e}{n_e e} - \rho T \frac{dS}{dt} + \eta J^2 = 0$$

Para chegarmos em  $\textcircled{I}$

$$\frac{c}{4\pi} \vec{E} \times \vec{B} = \frac{1}{4\pi} \left( -\vec{a}_e \times \vec{B} + \eta \vec{J} \times \vec{B} \right) \times \vec{B}$$

A equação de aquecimento  $\rightarrow$  ~~libera~~ energia irradiada

$$\rho \frac{dT}{ds} = \eta J^2 + K \nabla^2 T - \rho L(\rho, T)$$

$$s = \frac{1}{m} k_B \log \left( \frac{T^{\frac{1}{r-1}}}{\rho} \right)$$

A equação do calor assume a forma

$$\rho T \frac{d}{dt} \left( \frac{k_B}{m} \log \left( \frac{T^{\frac{1}{r-1}}}{\rho} \right) \right) = K \nabla^2 T + \eta J^2 - \rho L(\rho, T)$$

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \vec{u} \cdot \nabla$$

Temos assim

$$0 = \frac{k_B}{m} \left( \frac{1}{r-1} \rho \frac{\partial T}{\partial t} + \frac{1}{r-1} \rho (\vec{u} \cdot \nabla) T - T \frac{\partial \rho}{\partial t} - T (\vec{u} \cdot \nabla) \rho \right) - K \nabla^2 T - \eta J^2 - \rho L(\rho, T)$$

Perturbando as equações  $\rightarrow \vec{B}, T, \vec{u} \propto \rho$

As perturbações são da ~~forma~~ forma

$$\rho = \rho_0 + \rho_1 \rightarrow \rho_1 = \rho_1(t, x, y, z)$$

constante

$$\vec{B} = \vec{B}_0 + \vec{B}_1$$

$$T = T_0 + T_1$$

$$\rho = \rho_0 + \rho_1$$

$$\vec{u} = \vec{u}_0 + \vec{u}_1$$

Conservação de massa

$$\frac{\partial}{\partial t} (\rho_0 + \rho_1) + \nabla \cdot ((\rho_0 + \rho_1) \vec{u}_1) = 0$$

$\rho_0$  é constante  
 $\rightarrow$  ao ser derivado se anula

Equação de Indução

$$\frac{\partial}{\partial t} (\vec{B}_0 + \vec{B}_1) = \nabla \times (\vec{u} \times \vec{B}) - c \nabla \times (\nabla \times \vec{B} \eta)$$

$$\frac{\partial \vec{B}_1}{\partial t} - (\vec{B}_0 \cdot \nabla) \vec{u}_1 + \vec{B}_0 \cdot \nabla \vec{u}_1 - c \eta \nabla^2 \vec{B}_1 = 0$$

A equação de momento

$$\frac{\partial}{\partial t} [(\rho_0 + \rho_1) \cdot \vec{u}_1] + (\vec{u}_1 \cdot \nabla)(\rho_0 + \rho_1) \vec{u}_1 = \frac{k_B}{m} \nabla [(\rho_0 + \rho_1)(T_0 + T_1)]$$

$$+ \rho \nabla \phi + \frac{\vec{B}_1 \times (\vec{B}_0 + \vec{B}_1)}{4\pi}$$

$$0 = \frac{\partial u}{\partial t} + \frac{k_B \nabla T}{m} + \frac{k_B T_0 \nabla \rho_1}{\rho_0} + \frac{1}{4\pi \rho_0} \vec{B}_0 \times (\vec{B}_0 \times \vec{B}_1) + \nabla \phi$$

Abrindo  $L(\rho, T)$  em série

$$\rho L(\rho, T) = (\rho_0 + \rho_1) \left[ L(\rho_0, T_0) + \frac{\partial L}{\partial T} \Big|_{\rho_0} (T - T_0) + \frac{\partial L}{\partial \rho} \Big|_{T_0} (\rho - \rho_0) \right]$$

$$\rho L(\rho, T) = (\rho_0 + \rho_1) [L_T(T_1) + L_\rho(\rho_1)]$$

$$\rho L(\rho, T) = \rho_0 [L_T T_1 + L_\rho \rho_1]$$

Assim a equação de energia perturbada é:

$$0 = \frac{k_B}{m} \left( \frac{\rho_0}{\gamma - 1} \frac{\partial T_1}{\partial t} - T_0 \frac{\partial \rho_1}{\partial t} \right) + \rho_0 (T_1 L_T + \rho_1 L_\rho) - K \nabla^2 T_1$$

A perturbação de forma geral

$$f_1 = f_1^* \cdot e^{i(\omega t + \vec{h} \cdot \vec{r})}$$

ERRATA

$$f_1 = f_1^* e^{i\omega t + i\vec{h} \cdot \vec{r}}$$

$$\frac{\partial f_1}{\partial t} = i\omega f_1$$

etc

Teoremas

$$\frac{\partial f_1}{\partial t} = i\omega f_1$$

$$\nabla f_1 = i\vec{h} f_1$$

$$\nabla \cdot \vec{f}_1 = i\vec{h} \cdot \vec{f}_1$$

$$\nabla \times \vec{f}_1 = i\vec{h} \times \vec{f}_1$$

$$\nabla^2 f_1 = -h^2 f_1$$

Aplicando a definição de  $f_1$  nas equações teoremas

• Conservação de massa

$$\omega \frac{f_1}{\rho_0} + i\vec{h} \cdot \vec{u} = 0$$

• Equação de indução

$$\omega \vec{B}_1 - i(\vec{h} \cdot \vec{B}_0) \vec{u} + i\vec{B}_0(\vec{h} \cdot \vec{u}) + c\mu h^2 \vec{B}_1 = 0$$

• Equação de calor

$$\frac{k_B}{m} \left( \frac{\rho_0}{\gamma-1} i\omega T_1 - i\omega T_0 \rho_1 \right) + \rho_0 \left( \rho_1 L_p + T_1 L_T \right) + K T_1 h^2 = 0$$

$$\left[ \frac{k_B i\omega \rho_0}{m(\gamma-1)} + L_T \rho_0 + K h^2 \right] T_1 - \frac{T_0 \rho_1 k_B i\omega}{m} + \rho_0 \rho_1 L_p = 0$$

$$\frac{T_1}{T_0} = \frac{\rho_1}{\rho_0} \left[ \frac{\frac{k_B i\omega}{m} - \frac{\rho_0 L_p}{T_0}}{\frac{k_B i\omega}{m(\gamma-1)} + L_T \rho_0 + \frac{K h^2}{\rho_0}} \right]$$



Lembrando  $\bar{P} = \rho \frac{k_B T}{m}$

$$\frac{d\bar{P}}{dt} = \frac{k_B}{m} \left[ T \frac{d\rho}{dt} + \rho \frac{dT}{dt} \right]$$

$$= \frac{k_B}{m} \left[ T_0 \rho_1 i\omega + \rho_0 T_1 i\omega \right]$$

$$i\omega \bar{P}_1 = \frac{i\omega k_B}{m} [T_0 \rho_1 + \rho_0 T_1]$$

$$\bar{P}_1 = \frac{k_B}{m} \left[ \rho_1 + \rho_0 \frac{T_1}{T_0} \right] T_0$$

$$= \frac{\rho_0 k_B T_0}{m} \left[ \frac{\rho_1}{\rho_0} + \frac{T_1}{T_0} \right]$$

$$\frac{\bar{P}_1}{\bar{P}_0} = \frac{\rho_1}{\rho_0} + \frac{T_1}{T_0} \quad \alpha(\omega, h)$$

$$\frac{\bar{P}_1}{\bar{P}_0} = \frac{\gamma \rho_1}{\rho_0} \left[ \frac{\frac{k_B \omega}{m(\gamma-1)} + \frac{1}{\gamma} (L_T + L_P \frac{\rho_0}{T_0}) + \frac{K h^2}{\gamma \rho_0}}{\frac{k_B \omega}{m} \frac{1}{\gamma-1} + L_T + \frac{K h^2}{\rho_0}} \right]$$

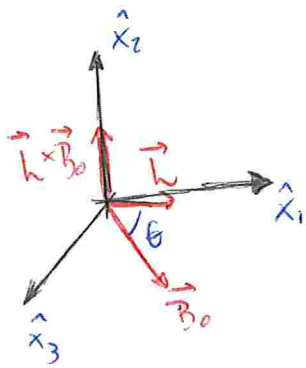
• Equações de momento

$$\omega \vec{u} + i \gamma \frac{\bar{P}_0}{\rho_0} \alpha \frac{\vec{h}}{h} \frac{\rho_1}{\rho_0} + \frac{i}{4\pi \rho_0} \vec{B}_0 \times (\vec{h} \times \vec{B}_1) = 0$$

$c_s^2 \rightarrow$  velocidade do som

Aplicando as eqs. de massa e indução perturbadas, temos

$$0 = \vec{u} \omega + \frac{\gamma \alpha c_s^2}{\omega} \vec{h} \cdot (\vec{h} - \vec{u}) - \frac{\vec{B}_0 \times [(\vec{h} \times \vec{u})(\vec{h} \cdot \vec{u}) - (\vec{h} - \vec{u})(\vec{h} \times \vec{B}_0)]}{4\pi \rho_0 (\omega + c_s^2 h^2)}$$



Tomando que

$$\vec{h} \cdot \vec{u} = h u_1$$

$$\vec{h} \times \vec{u} = u_2 \hat{x}_3 - u_3 \hat{x}_2$$

$$\vec{h} \cdot \vec{B}_0 = h B_0 \cos \theta$$

$$\vec{h} \times \vec{B}_0 = h B_0 \sin \theta \hat{x}_2$$

$$u_A = \frac{B_0}{\sqrt{4\pi\rho}}$$

velocidade  
Alfvén

$$\textcircled{A} \quad \omega u_1 + \frac{\gamma c_s^2 h}{\omega} u_1 + \frac{h^2 u_A^2 \sin \theta}{\omega + c\eta h^2} [\sin \theta u_1 - \cos \theta u_3] = 0$$

$$\textcircled{B} \quad \omega u_2 + \frac{h^2 u_A^2}{\omega + c\eta h^2} \cos^2 \theta u_2 = 0$$

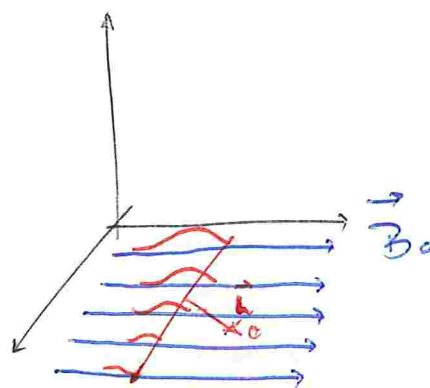
$$\textcircled{C} \quad \omega u_3 + \frac{h^2 u_A^2}{\omega + c\eta h^2} \cos \theta [\cos \theta u_3 - \sin \theta u_1] = 0$$

Eq. (B)

Primeiramente  $\eta = 0$

$$\omega^2 + h^2 u_A^2 \cos^2 \theta = 0$$

$$\omega = \pm i h u_A \cos \theta$$



Onda Alfvén

$$u_d = \frac{i\omega}{h} = u_A \cos \theta$$

Força Restauradora

$$\frac{(\vec{B} \cdot \nabla) \vec{B}}{4\pi}$$

Observamos

$$\vec{J} \times \vec{B} = \frac{(\nabla \times \vec{B})}{4\pi} \times \vec{B} = \frac{(\vec{B} \cdot \nabla) \vec{B}}{4\pi} + \nabla \left( \frac{B^2}{8\pi} \right)$$



Eq. (A) e (C)

$$\begin{cases} (\omega^2 + \gamma h^2 c_s^2 + h^2 u_A^2 \sin^2 \theta) u_1 - (h^2 u_A^2 \sin \theta \cos \theta) u_3 = 0 \\ (\omega^2 + h^2 u_A^2 \cos^2 \theta) u_3 - (h^2 u_A^2 \sin \theta \cos \theta) u_1 = 0 \end{cases}$$

Para que o determinante das equações seja 0, é preciso que

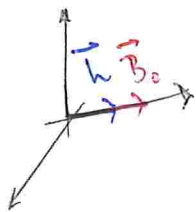
$$\textcircled{D} \quad \omega^4 + h^2 \omega^2 \underbrace{(u_A^2 + \gamma c_s^2)}_{u_{ms}^2 \rightarrow \text{velocidade magneto-sônica}} + \gamma h^4 u_A^2 c_s^2 \cos^2 \theta = 0$$

Como  $u_\phi = \frac{i\omega}{h}$ , obtidas ao dividir por  $h^4$

$$u_\phi^2 - (u_A^2 + \gamma c_s^2) u_\phi^2 + \gamma c_s^2 u_A^2 \cos^2 \theta = 0$$

Caso em que  $\theta = 0$

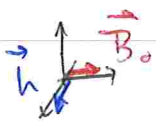
$$0 = u_\phi^2 - (u_A^2 + \gamma c_s^2) u_\phi^2 + \gamma c_s^2 u_A^2$$



$$0 = (u_\phi^2 - u_A^2)(u_\phi^2 - \gamma c_s^2)$$

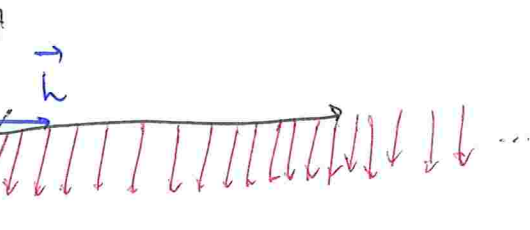
$$\begin{cases} u_\phi = \pm u_A \rightarrow \text{onda Alfvén} \\ u_\phi = \pm \sqrt{\gamma} c_s \rightarrow \text{onda acústica} \end{cases}$$

Caso em que  $\theta = \frac{\pi}{2}$



$$u_\phi (u_\phi - (u_A^2 + \gamma c_s^2)) = 0$$

$$\begin{cases} u_\phi = 0 \\ u_\phi = \pm \sqrt{u_A^2 + \gamma c_s^2} \end{cases}$$



ondas magnetoacústicas

