Error modelling in difference imaging

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- 1 Variance of weighted sum (convolution)
- 2 Variance of photometry
- Covariance model
- Experiments with real data
- Conclusions

Weighted sum of random variables

The variance of a weighted sum is:

$$var(aX + bY) = a^{2}var(X) + b^{2}var(Y) + ab cov(X, Y),$$
 (1)

where X and Y are random variables. More generally:

$$var\left(\sum_{i} a_{i} X_{i}\right) = \sum_{i,j} a_{i} a_{j} cov(X_{i}, X_{j})$$
(2)

$$= \sum_{i} a_i^2 var(X) + \sum_{i \neq j} a_i a_j cov(X_i, X_j)$$
 (3)

Sum over 2D kernel

Now consider a multidimensional kernel applied over an intensity image I:

$$var\left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} I_{\vec{y}}\right) = \sum_{\vec{y}, \vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}} cov(I_{\vec{y}}, I_{\vec{z}})$$
(4)

$$= \sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^2 var(I_{\vec{y}}) + \sum_{\vec{y} \neq \vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}} cov(I_{\vec{y}}, I_{\vec{z}}), \qquad (5)$$

where now \vec{x} , \vec{y} and \vec{z} are positions in the image (e.g. pixels), $k_{\vec{y}}^{\vec{x}}$ is the kernel function centered in \vec{x} , evaluated at the position \vec{y} , and $I_{\vec{y}}$ is the image intensity at \vec{y} .

Approximations

Exact variance:

$$var\left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} I_{\vec{y}}\right) = \sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} var(I_{\vec{y}}) + \sum_{\vec{y} \neq \vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}} cov(I_{\vec{y}}, I_{\vec{z}})$$
(6)

No covariance approximation:

$$var\left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} I_{\vec{y}}\right) \approx \sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^2 var(I_{\vec{y}})$$
 (7)

Local variance approximation:

$$var\left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} I_{\vec{x}}\right) \approx \left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}}\right)^2 var(I_{\vec{x}}) \tag{8}$$

Convolved variance approximation:

$$var\left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} I_{\vec{y}}\right) \approx \left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}}\right) \sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} var(I_{\vec{y}}) \tag{9}$$

New approximation

$$var\left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} I_{\vec{y}}\right) = \left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}}\right)^{2} v\bar{a}r - \sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^{2} v\bar{a}r - \sum_{\vec{y} \neq \vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}} v\bar{a}r + \sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^{2} var(I_{\vec{y}}) + \sum_{\vec{y} \neq \vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}} cov(I_{\vec{y}}, I_{\vec{z}}),$$
(10)

Assuming

$$v\bar{a}r \equiv \frac{\sum_{\vec{y}\neq\vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}} cov(I_{\vec{y}}, I_{\vec{z}})}{\sum_{\vec{y}\neq\vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}}}$$
(11)

$$var\left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} I_{\vec{y}}\right) = \left(\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}}\right)^2 v\bar{a}r - \sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^2 v\bar{a}r + \sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^2 var(I_{\vec{y}}) \quad (12)$$

$$= \left(\sum_{\vec{v}} k_{\vec{y}}^{\vec{x}}\right)^2 v \bar{a} r + \sum_{\vec{v}} (k_{\vec{y}}^{\vec{x}})^2 \left[var(I_{\vec{y}}) - v \bar{a} r\right]$$
 (13)

Optimal photometry

Given that $Flux(\vec{x})k_{\vec{y}}^{\vec{x}} = I_{\vec{y}}$ for isolated stars, where now $k_{\vec{y}}^{\vec{x}}$ is the PSF centered in \vec{x} evaluated at position \vec{y} , we can estimate $Flux(\vec{x})$ given $I_{\vec{y}}$ and $k_{\vec{y}}^{\vec{x}}$:

$$Flux(\vec{x}) = \frac{\sum_{\vec{y}} (I_{\vec{y}}/k_{\vec{y}}^{\vec{x}})/var(I_{\vec{y}}/k_{\vec{y}}^{\vec{x}})}{\sum_{\vec{y}} 1/var(I_{\vec{y}}/k_{\vec{y}}^{\vec{x}})} = \frac{\sum_{\vec{y}} k_{\vec{y}}^{\vec{x}} I_{\vec{y}}/var(I_{\vec{y}})}{\sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^2/var(I_{\vec{y}})}$$
(14)

The we can write the variance of the flux as:

$$var[Flux(\vec{x})] = \frac{1}{\sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^2 / var(l_{\vec{y}})} \left\{ 1 + \frac{\sum_{\vec{y} \neq \vec{z}} \frac{k_{\vec{y}}^{\vec{x}}}{var(l_{\vec{y}})} \frac{k_{\vec{z}}^{\vec{x}}}{var(l_{\vec{z}})} cov(l_{\vec{y}}, l_{\vec{z}})}{\sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^2 / var(l_{\vec{y}})} \right\}$$
(15)

2D Markov process correlation inversely proportional to distance

2D images are approximated by 2D Markov processes where the correlation is inversely proportional to the distance. For such a process the following relation holds:

$$cov(I_{\vec{x}},I_{\vec{y}}) = C \exp\left\{-\frac{||\vec{x}-\vec{y}||}{\lambda}\right\} = var_0 \exp\left\{-\frac{||\vec{x}-\vec{y}||}{\lambda}\right\}, \quad (16)$$

where \textit{var}_0 is the variance when $\vec{x} = \vec{y}$ and λ is some length scale, e.g. the FWHM. Then

$$v\bar{a}r \equiv \frac{\sum_{\vec{y}\neq\vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{z}} cov(I_{\vec{y}}, I_{\vec{z}})}{\sum_{\vec{y}\neq\vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{z}}} = \frac{\sum_{\vec{y}\neq\vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}} var_0 \exp\left\{-\frac{||\vec{x}-\vec{y}||}{\lambda}\right\}}{\sum_{\vec{y}\neq\vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{z}}}$$
(17)

Convolution with correlation inversely prop. to distance

Combining equations 11 and 16 we have:

$$v\bar{a}r \approx \frac{\sum_{\vec{y}\neq\vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}} var_0 \exp\left\{-\frac{||\vec{y}-\vec{z}||}{\lambda}\right\}}{\sum_{\vec{y}\neq\vec{z}} k_{\vec{y}}^{\vec{x}} k_{\vec{z}}^{\vec{x}}}$$
(18)

Photometry with correlation inversely prop. to distance

Combining equations 13 and 16 we have:

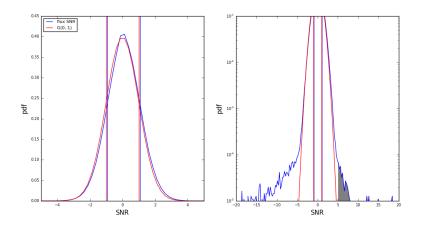
$$var[Flux(\vec{x})] \approx$$

$$\frac{1}{\sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^{2} / var(I_{\vec{y}})} \left\{ 1 + \frac{\sum_{\vec{y} \neq \vec{z}} \frac{k_{\vec{y}}^{\vec{x}}}{var(I_{\vec{y}})} \frac{k_{\vec{z}}^{\vec{x}}}{var(I_{\vec{z}})} var_{0} \exp\left\{-\frac{||\vec{y} - \vec{z}||}{\lambda}\right\}}{\sum_{\vec{y}} (k_{\vec{y}}^{\vec{x}})^{2} / var(I_{\vec{y}})} \right\}$$
(19)

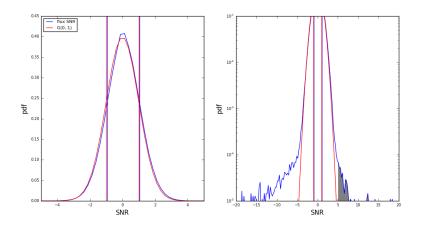
Test with real data

- Measure the distribution of $SNR(\vec{x}) \equiv \frac{Flux(\vec{x})}{\sqrt{var[Flux(\vec{x})]}}$, if errors are properly modeled we should see a standard normal distribution.
- Assuming correlation inversely proportional to distance, var_0 and λ in order to reproduce standard normal distribution ($var_0 \approx var_{\vec{x}}$ and $\lambda \approx FWHM$?).

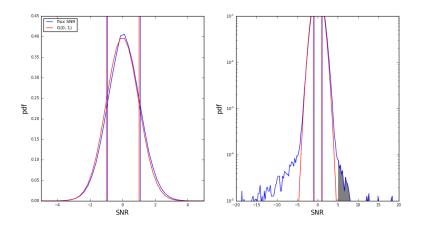
No covariance approximation, small kernel



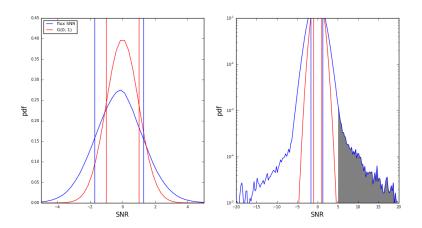
Local variance approximation, small kernel



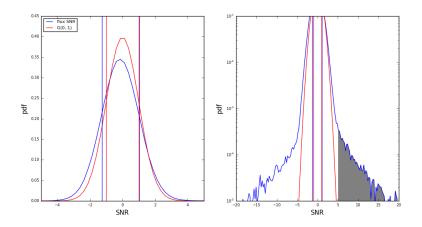
Convolved variance approximation, small kernel



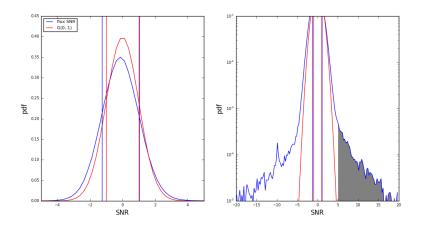
No covariance approximation, large kernel



Local variance approximation, large kernel



Convolved variance approximation, large kernel



Conclusions

- Exact expression for the variance of the convolution and photometry of an image difference were obtained, which depend on the full covariance matrix
- We expect non-diagonal terms of the covariance matrix to increase the variance of the convolved image and the photometry.
- The larger the convolution kernel, the larger the effect of the covariance matrix during convolution
- The larger the PSF and the smaller the variance in every pixel, the larger the effect of the covariance matrix during photometry
- An analytic approximation for the covariance matrix is proposed assuming that the correlation between pixels is inversely proportional to distance. The approximation has two free-parameters: a scaling factor and a length scale. We plan to test whether these parameters are related to the local variance and the FWHM of the PSF.