

# Gravity @ ICTP

## Wald

- recap:

- view stationary BH as body in thermal equilibrium

• 0<sup>th</sup> law -  $\kappa$  locally constant

• 1<sup>st</sup> law -  $\delta M = \frac{1}{8\pi} \kappa \delta A + \Omega_H \delta J + \varphi_H \delta Q$

• 2<sup>nd</sup> law -  $\delta A \geq 0$

## BH and Black Brane stability

- by black brane we mean a  $(D+p)$ -dim spacetime metric of the form  $d\tilde{s}_{D+p}^2 = ds_D^2 + \sum_{i=1}^p dz_i^2$  where  $ds_D^2$  is  $D$ -dim BH metric

- we define the "canonical energy"  $\mathcal{E}$  of a perturbation  $\gamma_{ab}$  and show its positivity is necessary for stability of BH

- we view entropy as  $S = k \ln V$  where  $V$  is a volume of phase space on a shell of constant energy and some other parameters  $\{X_i\}$

- naturally, starting from a small volume, we evolve to a distinguishable volume which extremises  $S$

- we could evolve out of it, after a long time, but not generically

- here we consider a "semi ergodic" system, where orbits fill the shell s.t. the time spent in a volume is prop. to the volume

- now  $\delta E = T \delta S + \sum \gamma_i \delta x_i$ , where this should be viewed as defining  $\delta S$  w.r.t deformations of the shell

- thermodynamic stability, i.e. entropy being maximum at 2<sup>nd</sup> order,

demands  $\delta^2 E - T \delta^2 S - \sum \gamma_i \delta^2 x_i > 0$

where  $(E, x_i)$  are fixed at 1<sup>st</sup> order

for only extended, homogeneous systems

- alternatively, if  $H_S = \begin{pmatrix} \partial_B^2 S & \partial_{x_i} \partial_B S \\ \partial_B \partial_{x_i} S & \partial_{x_i} \partial_{x_j} S \end{pmatrix}$

has a positive eigenvalue, the system is unstable, which

corresponds to a negative heat capacity

→ this is true for a homogeneous, only extended system, but not necessarily for a finite one.

- BH's and branes have

$$\begin{array}{ccc} E & & M \\ S & \longleftrightarrow & A \\ x_i & & J_i, Q_i \end{array}$$

so for  $\delta E = \delta A = 0$  we want

$$\delta^2 M - \frac{K}{8\pi} \delta^2 A - \sum \Omega_i \delta^2 J_i > 0$$

and since BH's are only extended,  $H_A$  needs to lack a pos eigenv.

→ Schwarzschild has  $A = 16\pi^2 M^2 \Rightarrow \frac{\partial^2 A}{\partial M^2} > 0$   
but it's finite

- but this implies black strings built from  
Schw. aren't stable (Schw.  $\times \mathbb{R}$ )