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## Project 2; Variational Monte Carlo studies of electronic systems

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Sean Bruce Sangolt Miller  
s.b.s.miller@fys.uio.no

Filip Henrik Larsen  
filiphenrikarsen@gmail.com

Roar Emaus  
roarem@fys.uio.no

Github repository:

<https://github.com/filiph1/FYS4411.git>

### Abstract

A variational Monte-Carlo (VMC) approach to the quantum dot with  $N = 2, 6, 12$  and 20 electrons was studied. The electrons were placed in a harmonic oscillator potential and assumed to interact with each other solely by the Coulomb interaction. This was an expansion of the previous project in which the ground state of a bosonic many-body system was studied. Specifically, the state wavefunction of a closed-shell model of the quantum dot was studied, where we approximated the system wavefunction by a Slater determinant (non-interacting) term and a Pade-Jastrow factor (correlation). We found such a simple model give ground state energies with, at most, 1.5 atomic units deviation from the results of [1]. Afterwards, we computed the so-called one-body densities from the radial distribution of particles. Finally, we studied the effects of efficiency and time-reduction obtained by parallelization and vectorization of the program.

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# 1 Introduction

Quantum dots are usually large atomic systems packed so closely together that the Pauli principle forces the electrons to be in different states. These systems can be made approximately 2-dimensional<sup>1</sup> by strings of quantum dots. Quantum dots are much larger than atoms, but possesses many of their characteristics, which means they are both easy and interesting to measure. Measurements provide means with which to test quantum theory and models.

With up to 20 electrons, it is reasonable to approximate the system Hamiltonian by placing all electrons in a harmonic oscillators and having coulomb interactions. If only closed-shell quantum dots are considered, a model that splits the wavefunction can be used, which greatly reduces the number of calculations. The purpose of this project is to see if this simple model can reproduce known results through variational Monte-Carlo (VMC) simulation by using a Slater determinant with a Pade-Jastrow factor.

First, a detailed calculation of the two-body quantum dot will be performed, followed by deduction of algorithms for a general many-body numerical approach to a system of  $N = \{2, 6, 12, 20\}$  electrons. Then, an explanation of the optimization of the variational parameters will be given, followed by a small explanation of the *blocking* method. Since the program script will be large, benchmarking is necessary, and therefore a few benchmarks will be mentioned. In order to reduce time consumption and increase efficiency, parallelization and vectorization will be used, respectively. Lastly, results will be presented and discussed in some detail.

## 2 Theory and Methods

### 2.1 Preliminary derivations

While performing VMC it is of course favourable to use analytical expressions, should they not demand an unacceptable increase in CPU time. We will therefore need to calculate the local energy  $E_L = \frac{1}{\Psi_T} H \Psi_T$  and the quantum force  $F = \frac{2}{\Psi_T} \nabla \Psi_T$ . The Hamiltonian  $H$  used will be:

$$H = H_0 + H_I = \sum_{i=1}^N \left( -\frac{1}{2} \nabla_i^2 + \frac{1}{2} \omega^2 r_i^2 \right) + \sum_{i < j} \frac{1}{r_{ij}} \quad (1)$$

I.e. a harmonic oscillator potential with Coulomb interactions. The Laplacian will be the most demanding quantity to calculate.

### 2.2 Singlet electron state

For an electron in a harmonic oscillator potential, the energy is given by  $\epsilon_n = \omega(n + 1)$ , where we have used natural units and  $n = n_x + n_y + \dots$ . For two non-interacting electrons, the energy is  $\epsilon_{n_1, n_2} = \omega(n_{x,1} + n_{y,1} + \dots + n_{x,2} + n_{y,2} + \dots + 2)$ . Obviously the energy is lowest for  $n_1 = n_2 = 0$ , giving  $\epsilon_{0,0} = 2\omega$ . Since  $n_1 = n_2 = 0$  means the two electron are in the same spatial wavefunction, they must have different spins. Since electrons are spin- $\frac{1}{2}$  particles, they combine to give total spin zero, i.e. they form the singlet state.

For the singlet electron state we will use the trial wavefunction:

$$\Psi_T(\mathbf{r}_1, \mathbf{r}_2) = C e^{-\frac{\alpha\omega}{2}(r_1^2 + r_2^2)} e^{\frac{\alpha\mathbf{r}_1 \cdot \mathbf{r}_2}{1 + \beta r_{12}}} \quad , \quad a = 1 \quad (2)$$

The Laplacian of which (for particle  $i$ ) is:

$$\nabla_i^2 \Psi_T = \nabla_i (\nabla_i \Psi_T) \quad (3)$$

We will use the following change of coordinates when it simplifies calculations:

$$\begin{aligned} \frac{\partial}{\partial r_{i,j}} &= \frac{\partial r_{12}}{\partial r_{i,j}} \frac{\partial}{\partial r_{12}} \\ \rightarrow \nabla_i &= \frac{(-1)^i}{r_{12}} (x_1 - x_2, y_1 - y_2) \frac{\partial}{\partial r_{12}} \\ &= \frac{(-1)^i}{r_{12}} \mathbf{r}_{12} \frac{\partial}{\partial r_{12}} \end{aligned} \quad (4)$$

where  $r_{i,j}$  is element  $j$  of  $\mathbf{r}_i$ . The gradient, which is needed for the quantum force as well, is then

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<sup>1</sup>I.e. all electron movement is confined to a plane. This does not mean the space is considered two-dimensional, so the 3-dimensional Coulomb potential will still be used.

$$\begin{aligned}\nabla_i \Psi_T &= -\alpha\omega \mathbf{r}_i \Psi_T + \frac{(-1)^i}{r_{12}} \mathbf{r}_{12} \left[ \frac{\partial}{\partial r_{12}} \left( \frac{ar_{12}}{1 + \beta r_{12}} \right) \right] \Psi_T \\ &= \left[ -\alpha\omega \mathbf{r}_i + \frac{(-1)^i}{r_{12}} \mathbf{r}_{12} \frac{a}{(1 + \beta r_{12})^2} \right] \Psi_T\end{aligned}\quad (5)$$

which means the Laplacian is

$$\begin{aligned}\nabla_i^2 \Psi_T &= [\nabla_i[\dots]] \Psi_T + [\dots] \nabla_i \Psi_T \\ &= [\nabla_i[\dots]] \Psi_T + [\dots]^2 \Psi_T\end{aligned}\quad (6)$$

where  $[\dots]$  is the last parenthesis in equation 5. The parenthesis in the first term above is

$$\begin{aligned}\nabla_i \left[ -\alpha\omega \mathbf{r}_i + \frac{(-1)^i}{r_{12}} \mathbf{r}_{12} \frac{a}{(1 + \beta r_{12})^2} \right] &= -2\alpha\omega + \frac{(-1)^i}{r_{12}} \left( \frac{(-1)^i 2ar_{12}}{(1 + \beta r_{12})^2} - \frac{(-1)^i 2a\beta r_{12}}{(1 + \beta r_{12})^3} - \frac{(-1)^i a}{r_{12}(1 + \beta r_{12})^2} \right) \\ &= -2\alpha\omega - \frac{a}{(1 + \beta r_{12})^2} \left( \frac{1}{r_{12}} - \frac{2}{r_{12}} + \frac{2\beta}{1 + \beta r_{12}} \right) \\ &= -2\alpha\omega + \frac{a}{r_{12}(1 + \beta r_{12})^2} - \frac{2a\beta}{(1 + \beta r_{12})^3}\end{aligned}\quad (7)$$

which gives

$$\nabla_i^2 \Psi_T = \left[ -2\alpha\omega + \frac{a}{r_{12}(1 + \beta r_{12})^2} - \frac{2a\beta}{(1 + \beta r_{12})^3} + \alpha^2 \omega^2 r_i^2 + \frac{a^2}{(1 + \beta r_{12})^4} - \frac{2\alpha\omega a(-1)^i}{r_{12}(1 + \beta r_{12})^2} \mathbf{r}_i \cdot \mathbf{r}_{12} \right] \Psi_T \quad (8)$$

We therefore have:

$$\sum_{i=1}^2 \frac{1}{\Psi_T} \nabla_i^2 \Psi_T = -4\alpha\omega + \frac{2a}{r_{12}(1 + \beta r_{12})^2} - \frac{4a\beta}{(1 + \beta r_{12})^3} + \alpha^2 \omega^2 (r_1^2 + r_2^2) + \frac{2a^2}{(1 + \beta r_{12})^4} - \frac{2\alpha\omega a}{(1 + \beta r_{12})^2} r_{12} \quad (9)$$

### 2.3 Closed, 2-dimensional shell states

If we again consider the non-interacting, harmonic oscillator confined electron system, we can increase the number of electrons beyond 2. If, in some wondrous universe, we had fermions with three different spins states (like a spin-1 boson), then for three electrons we could again set  $n_1 = n_2 = n_3 = 0$  and have an anti-symmetric spin state. However, in our world, we must go up in energy for more than two electrons.

The  $n = 0$  state was the ground state. The next state has degeneracy 2;  $(n_1, n_2) = (1, 0), (0, 1)$ . After that we have degeneracy 3;  $(n_1, n_2) = (2, 0), (1, 1), (0, 2)$ . Each of these states are doubly degenerate due to spin.

The reason in explaining this is because, assuming we have a so-called "closed shell" problem<sup>2</sup>, we can do some manipulations that greatly reduce the number of calculations necessary to perform VMC. Firstly, we need to rewrite the trial wavefunction.

As is already known, the true wavefunction is approximated by an analytical solution to some simpler problem, and a Jastrow factor. The analytical part can be written as a Slater determinant:

$$\Psi_D = \frac{1}{\sqrt{N}} \begin{vmatrix} \phi_1(\mathbf{r}_1) & \phi_2(\mathbf{r}_1) & \dots & \phi_{N-1}(\mathbf{r}_1) & \phi_N(\mathbf{r}_1) \\ \phi_1(\mathbf{r}_2) & \phi_2(\mathbf{r}_2) & \dots & \phi_{N-1}(\mathbf{r}_2) & \phi_N(\mathbf{r}_2) \\ \vdots & \vdots & \ddots & \vdots & \vdots \\ \phi_1(\mathbf{r}_N) & \phi_2(\mathbf{r}_N) & \dots & \phi_{N-1}(\mathbf{r}_N) & \phi_N(\mathbf{r}_N) \end{vmatrix} \quad (10)$$

and therefore our trial wavefunction will be:

$$\Psi_T = \Psi_D \Psi_C, \quad \Psi_C = \prod_{i < j}^N e^{f_{ij}} \quad , \quad f_{ij} \equiv \frac{a_{ij} r_{ij}}{1 + \beta r_{ij}} \quad (11)$$

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<sup>2</sup>For every new tier in energy, we fill it up with electrons. So all considered tiers, or "shells", are full (or "closed")

where  $\alpha, \beta$  are the variational parameters,  $a_{ij}$  is connected to particle spins, and  $N$  is the total number of particles. The single particle functions are solutions to the two dimensional, harmonic oscillator Schrödinger equation:

$$\phi_i(\mathbf{r}_i) = CH_{n_{x,i}}(\sqrt{\omega\alpha}x_i)H_{n_{y,i}}(\sqrt{\omega\alpha}y_i)e^{\frac{\omega\alpha}{2}r_i^2} \quad (12)$$

Obviously,  $\Psi_D$  is a time consuming object to calculate at every Metropolis step. We will therefore do some neat tricks that reduce the number of calculations.

The first is to rewrite  $\Psi_D$  by using that the Hamiltonian is spin-independent. If we now let  $\phi_n$  be the spatial-component of the wavefunction, then it causes no effect on the energy if we write:

$$\Psi_T = \Psi_{D+}\Psi_{D-}\Psi_C \quad (13)$$

where  $\Psi_{D+} \equiv |D_+|$  and  $\Psi_{D-} \equiv |D_-|$  are Slater determinants consisting only of spin-up and spin-down particles, respectively. This means  $|D_+|$  and  $|D_-|$  are  $\frac{N}{2}$ -dimensional, and we see this method only works for closed shell systems. We will let the first  $\frac{N}{2}$  particles be spin-up and the rest spin-down. That is,  $\mathbf{r}_1 - \mathbf{r}_{N/2}$  are the positions of the spin-up particles, while the rest are positions for spin-down particles. Obviously this is wrong, since we can't separate which particles are spin-up and which are spin-down in reality, due to the properties of identical particles. However, as stated, due to the spin-independent Hamiltonian, this causes no effect on the energy, which is all we're after.

## 2.4 The Metropolis ratio test

The basic principle behind the Metropolis algorithm is to make an assumption on the transition probability for a system to move from setting to another, as an exact, or even approximate, expression is lacking.

If the probability distribution for a state  $i$  is given by  $w_i$ , then from Markov chain theory the time derivative is:

$$\frac{\partial w_i(t)}{\partial t} = \sum_j W(j \rightarrow i)w_j(t) - W(i \rightarrow j)w_i(t) \quad (14)$$

where  $W_{i \rightarrow j}$  is the probability of moving from a state  $i$  to another state  $j$ , i.e the rate of change in  $w_i$  is given by the probability for a state  $j$  to go to  $i$  minus the probability of state  $i$  going to  $j$ , summed over all  $j$ . The most likely state will fulfil  $\frac{\partial w_i(t)}{\partial t} = 0$ , giving:

$$\begin{aligned} W(j \rightarrow i)w_j(t) &= W(i \rightarrow j)w_i(t) \\ \Rightarrow \frac{W(j \rightarrow i)}{W(i \rightarrow j)} &= \frac{w_i}{w_j} \end{aligned} \quad (15)$$

Since the transition probability  $W$  is unknown, we approximate it by guessing its form:

$$W(j \rightarrow i) = T(j \rightarrow i)A(j \rightarrow i) \quad (16)$$

where  $T$  is the transition moving probability, while  $A$  is the probability of accepting such a move. Furthermore, in brute force Metropolis, one guesses  $T_{i \rightarrow j} = T_{j \rightarrow i}$ . Therefore:

$$\frac{A_{j \rightarrow i}}{A_{i \rightarrow j}} = \frac{w_i}{w_j} \quad (17)$$

Since the probability densities are known, it is known whether or not this ratio is larger than one. If it's larger than one, then the acceptance probability from  $j$  to  $i$  is the biggest, i.e we are more likely to accept the move than not. Therefore, we simply say the move is accepted. However, if the probability ratio is smaller than one, then we are more likely to move from the new state to the one the system is currently in. To check whether or not we should accept the move, we may compare it to, say, a "coin toss"; if it's bigger, the move is accepted.

So, at each Metropolis step, we need the ratio of probabilities. We first define  $R \equiv \frac{\Psi_T^n}{\Psi_T^o}$ , where "n" means the new wavefunction and "o" means the old (or the current, but "c" could be confused with "correlation"). Written out, this is:

$$R = \frac{|D_+^n| |D_-^n| \Psi_C^n}{|D_+^o| |D_-^o| \Psi_C^o} \quad (18)$$

If we only move one position at a time, then only one row in either  $D_+$  or  $D_-$  will change. This means if we move a spin-up position, then  $|D_+^n| = |D_+^o|$ , so we need only consider one of the determinant fractions for each  $R$ .

Through some simple steps, one can show the determinant fraction ( $R_D$ ) reduces to:

$$R_D = \sum_{j=1}^{N/2} D_{ij}(\mathbf{r}^n) D_{ji}^{-1}(\mathbf{r}^o) \quad (19)$$

where  $D_{ij}^{-1}$  is element<sup>3</sup>  $ij$  of the inverse of  $D$ ,  $D$  is either the spin-up or spin-down determinant, and  $\mathbf{r}_i$  is the moved position. Since  $D_{ij}(\mathbf{r}) = \phi_j(\mathbf{r}_i)$ , the only difficulty remaining is to find the elements of the inverse matrix.

The elements of an inverse matrix are given by the Sherman-Morrison formula, which, when applied to the current case, gives:

$$D_{kj}^{-1}(\mathbf{r}^n) = \begin{cases} D_{kj}^{-1}(\mathbf{r}^o) - \frac{D_{ki}^{-1}(\mathbf{r}^o)}{R_D} \sum_{l=1}^{N/2} D_{il}(\mathbf{r}^n) D_{lj}^{-1}(\mathbf{r}^o) & \text{if } j \neq i \\ \frac{D_{ki}^{-1}(\mathbf{r}^o)}{R_D} & \text{if } j = i \end{cases} \quad (20)$$

The ratio for the correlation function has a rather nice expression:

$$\begin{aligned} \frac{\Psi_C^n}{\Psi_C^o} &= \prod_{i < j}^{N/2} e^{(f_{ij}^n - f_{ij}^o)} \\ &= \exp \left( \sum_{i < j}^{N/2} (f_{ij}^n - f_{ij}^o) \right) \\ &= \exp \left( \sum_{i=0}^{k-1} (f_{ik}^n - f_{ik}^o) + \sum_{j=k+1}^{N/2} (f_{kj}^n - f_{kj}^o) \right) \end{aligned} \quad (21)$$

where we used  $f_{ij}^n - f_{ij}^o = 0 \forall i, j \neq k$ , and  $k$  is the moved position. The first sum is then for  $j = k$  and the second sum is for  $i = k$ , with the restriction  $i < j$ .

The only step that remains is to square and multiply the two ratios<sup>4</sup>.

## 2.5 Importance sampling

In importance sampling, each suggested move requires the calculation of the quantum force:

$$x_{new} = x_{old} + DF(x_{old})\Delta t + \xi\sqrt{\Delta t} \quad (22)$$

Solutions to the Fokker-Planck equation gives the transition probability, which must be multiplied with the probability density. The transition probability is therefore given by:

$$G(y, x, \Delta t) = \frac{1}{(4\pi D \Delta t)^{3N/2}} \exp \left\{ -\frac{(y - x - D \Delta t F(x))^2}{4D \Delta t} \right\} \quad (23)$$

where  $y$  is the new position and  $x$  the old, and the acceptance test becomes:

$$q(y, x) = \frac{G(x, y, \Delta t) |\Psi_T(y)|^2}{G(y, x, \Delta t) |\Psi_T(x)|^2} \quad (24)$$

The quantum force, given by  $F = 2 \frac{\nabla \Psi_T}{\Psi_T}$ , requires the gradient of  $\Psi_T$ . Obviously, the quantum force can be written<sup>5</sup>:

$$F = 2 \left( \frac{\nabla |D_+|}{|D_+|} + \frac{\nabla |D_-|}{|D_-|} + \frac{\nabla \Psi_C}{\Psi_C} \right) \quad (25)$$

Of course, when only position  $k$  is altered, only the  $k$ 'th gradient in  $\nabla$  changes (recall the definition  $\nabla \equiv (\nabla_1, \nabla_2, \dots, \nabla_N)$ , and needs to be re-evaluated. We know the determinant can be written:

<sup>3</sup>Where  $i$  is the row and  $j$  is the column.

<sup>4</sup>Or multiply and then square. Really, it's up to you.

<sup>5</sup>The superscript "o" has been dropped since there will be no "mix" of new and old coordinates for the rest of this subsection.

$$|D| = \sum_{j=1}^{N/2} D_{kj} C_{jk} \quad (26)$$

where  $C_{jk}$  are the cofactors of  $D$ , and is independent of the  $i$ 'th row in  $D$ , i.e. changing  $k$  does not change  $C_{jk}$ . It is therefore independent of the position change. Changing row  $i$  means all the other gradients in  $\nabla|D|$  are the same as before, and we only need to re-evaluate  $\nabla_k|D|$ . This means:

$$\begin{aligned} \frac{\nabla_k|D|}{|D|} &= \frac{\nabla_k \sum_{j=1}^{N/2} D_{kj} C_{jk}}{|D|} \\ &= \sum_{j=1}^{N/2} \frac{(\nabla_k D_{kj}) C_{jk}}{|D|} \\ &= \sum_{j=1}^{N/2} (\nabla_k D_{kj}) D_{jk}^{-1} \end{aligned} \quad (27)$$

which, from equation 20, means:

$$\frac{\nabla_k|D^n|}{|D^n|} = \frac{1}{R_D} \sum_{j=1}^{N/2} (\nabla_k D_{kj}^n) (D_{jk}^o)^{-1} \quad (28)$$

where the factor  $\nabla_k D_{kj} = \nabla_k \phi_j(\mathbf{r}_k)$  is:

$$\begin{aligned} \nabla_k D_{kj} &= A \left( H'_{n_{j,x}}(\sqrt{\omega\alpha}x_k) H_{n_{j,y}}(\sqrt{\omega\alpha}y_k) - \alpha\omega x_k H_{n_{j,x}}(\sqrt{\omega\alpha}x_k) H_{n_{j,y}}(\sqrt{\omega\alpha}y_k) \right. \\ &\quad \left. , H'_{n_{j,y}}(\sqrt{\omega\alpha}y_k) H_{n_{j,x}}(\sqrt{\omega\alpha}x_k) - \alpha\omega y_k H_{n_{j,x}}(\sqrt{\omega\alpha}x_k) H_{n_{j,y}}(\sqrt{\omega\alpha}y_k) \right) e^{-\frac{\alpha\omega}{2} r_k^2} \end{aligned} \quad (29)$$

The correlation function gradient can be expressed:

$$\begin{aligned} \frac{\nabla_k \Psi_C}{\Psi_C} &= \frac{1}{\Psi_C} \nabla_k e^{\sum_{i<j}^N f_{ij}} \\ &= \sum_{i=1}^{k-1} \nabla_k f_{ik} + \sum_{j=k+1}^N \nabla_k f_{kj} \end{aligned} \quad (30)$$

but since  $f_{ij}$  only depends on  $r_{ij}$ , it would preferable to express  $\nabla_k$  in terms of  $r_{ij}$ . In equation 4, we showed this change for a simpler system. Applied to this problem, we can derive:

$$\begin{aligned} \nabla_k &= \frac{1}{r_{ik}} \mathbf{r}_{ik} \frac{\partial}{\partial r_{ik}} \\ \text{or} \\ \nabla_k &= -\frac{1}{r_{jk}} \mathbf{r}_{jk} \frac{\partial}{\partial r_{jk}} \end{aligned} \quad (31)$$

which gives:

$$\frac{\nabla_k \Psi_C}{\Psi_C} = \sum_{i=1}^{k-1} \frac{\mathbf{r}_{ik}}{r_{ik}} \frac{\partial f_{ik}}{\partial r_{ik}} - \sum_{j=k+1}^N \frac{\mathbf{r}_{kj}}{r_{kj}} \frac{\partial f_{kj}}{\partial r_{kj}} \quad (32)$$

We now have all the necessary tools to perform importance sampling.

## 2.6 Local energy

Lastly, the local energy needs to be calculated. As usual, the Laplacian fraction  $\frac{\nabla^2 \Psi_T}{\Psi_T}$  is the most demanding object to calculate. the starting point is:

$$\frac{\nabla^2 \Psi_T}{\Psi_T} = \frac{\nabla^2 |D_+|}{|D_+|} + \frac{\nabla^2 |D_-|}{|D_-|} + \frac{\nabla^2 \Psi_C}{\Psi_C} + 2 \left( \frac{\nabla |D_+|}{|D_+|} + \frac{\nabla |D_-|}{|D_-|} \right) \cdot \frac{\nabla \Psi_C}{\Psi_C} \quad (33)$$

which looks easy enough. The last term contain vectors already known, while the first two are derived in the exact same manner as for the gradients, i.e.

$$\frac{\nabla_k^2 |D|}{|D|} = \sum_{j=1}^{N/2} (\nabla_k^2 D_{kj}) D_{jk}^{-1} \quad (34)$$

where (here we abbreviate  $H_{n_{j,z}}(\sqrt{\omega\alpha}z_k)$  by  $H_{n_j}(z)$  for compactness)

$$\begin{aligned} \nabla_k^2 D_{kj} = A & \left[ H_{n_j}''(x) H_{n_j}(y) + H_{n_j}''(y) H_{n_j}(x) \right. \\ & - 2\alpha\omega \left( x H_{n_j}'(x) H_{n_j}(y) + y H_{n_j}'(y) H_{n_j}(x) \right) \\ & \left. - \alpha\omega H_{n_j}(x) H_{n_j}(y) (d - \alpha\omega r_k^2) \right] e^{-\frac{\alpha\omega}{2} r_k^2} \end{aligned} \quad (35)$$

Unfortunately, the middle term is not so nice to find. The steps needed are many and tedious, but not difficult. They are therefore omitted and we show only the final result:

$$\frac{\nabla_k^2 \Psi_C}{\Psi_C} = \left( \frac{\nabla_k \Psi_C}{\Psi_C} \right)^2 + \sum_{i=1}^{k-1} \left[ \frac{d-1}{r_{ik}} \frac{\partial f_{ik}}{\partial r_{ik}} + \frac{\partial^2 f_{ik}}{\partial r_{ik}^2} \right] + \sum_{j=k+1}^N \left[ \frac{d-1}{r_{kj}} \frac{\partial f_{kj}}{\partial r_{kj}} + \frac{\partial^2 f_{kj}}{\partial r_{kj}^2} \right] \quad (36)$$

where  $d$  is the dimension we consider and:

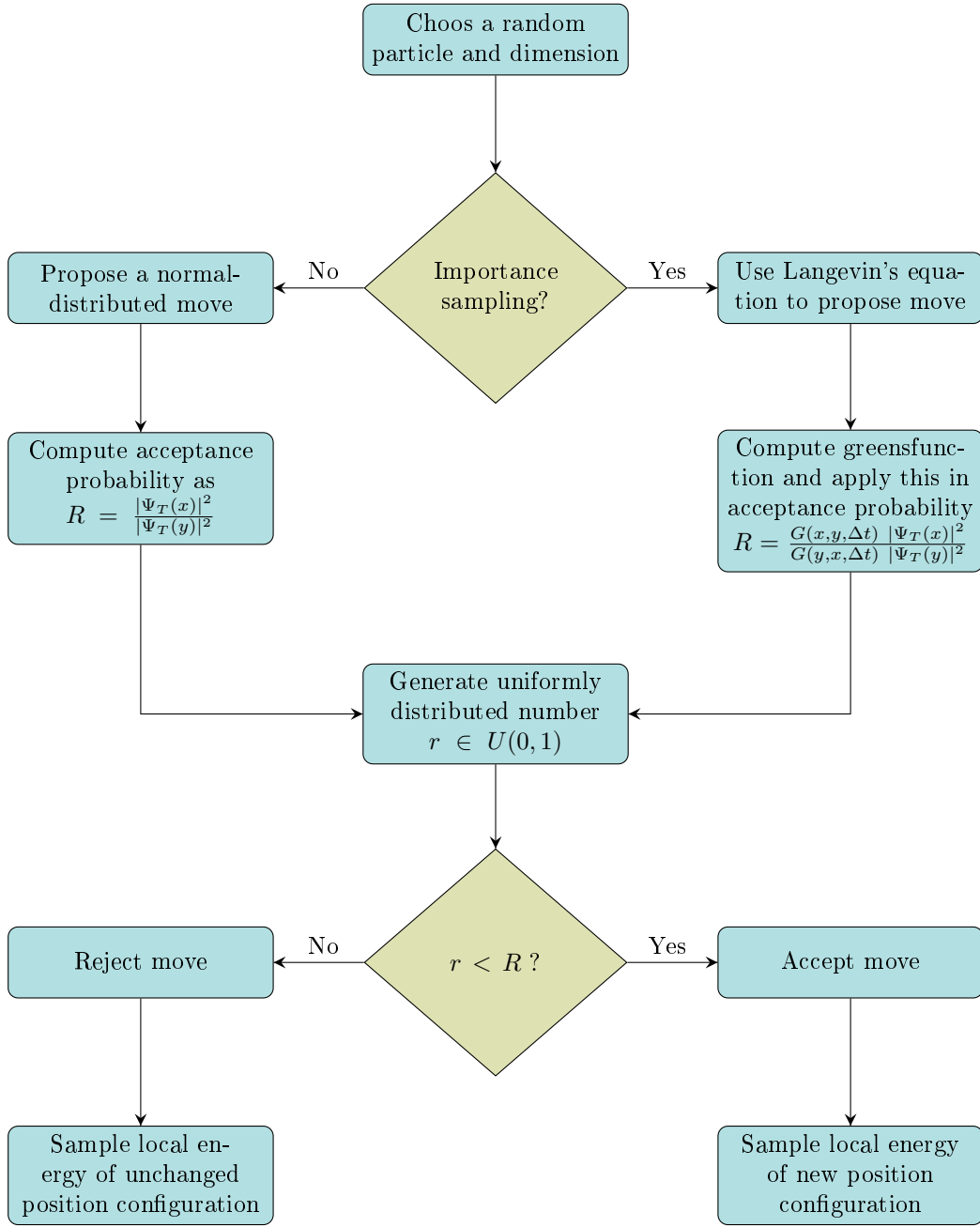
$$\frac{\partial f_{ij}}{\partial r_{ij}} = \frac{a_{ij}}{(1 + \beta r_{ij})^2} \quad (37)$$

$$\frac{\partial^2 f_{ij}}{\partial r_{ij}^2} = -\frac{2a_{ij}\beta}{(1 + \beta r_{ij})^3} \quad (38)$$

where  $a_{ij}$  equals 1 for anti-parallel spins and  $\frac{1}{3}$  for parallel.



## 2.7 Flowchart of metropolis algorithm



## 2.8 Optimizing parameters

Our trial wavefunctions, both for the two-body system and many-body system (equations 2 and 11), contain two variational parameters,  $\alpha$  and  $\beta$ . In the VMC approach, the minimal energy is sought. Therefore, the goal is to minimize  $\langle E_L \rangle$  with respect to these variational parameters. There are several ways to minimize a value with respect to some parameters, and here we will use the method of steepest descent (SD). The SD method, in algorithm form, is:

$$\vec{x}_{n+1} = \vec{x}_n - \gamma_n \nabla f \quad (39)$$

where  $\vec{x}$  is a vector containing the variables for which one wishes to find the minimum of  $f$ ,  $\gamma$  is a steplength and  $n$  is an index expressing the number of iterations. In application to the current problem,  $f = \langle E_L \rangle$ , and  $\vec{x} = (\alpha, \beta)$ . The above equation can thus be rewritten as

$$(\alpha, \beta)_{n+1} = (\alpha, \beta)_n - \gamma_n \left( \frac{\partial}{\partial \alpha}, \frac{\partial}{\partial \beta} \right) \langle E_L \rangle. \quad (40)$$

However, since  $\langle E_L \rangle$  is quite a time-consuming quantity we find numerically, and its derivatives ( $\bar{E}_\alpha \equiv \frac{d\langle E_L \rangle}{d\alpha}$  and  $\bar{E}_\beta \equiv \frac{d\langle E_L \rangle}{d\beta}$ ) even more so, an "analytical" expression is desirable. This can be found as follows:

$$\begin{aligned}\bar{E}_\alpha &= \frac{d}{d\alpha} \int dx P(x) E_L \\ &= \frac{d}{d\alpha} \int dx \frac{|\psi|^2}{\int dx' |\psi|^2} \frac{1}{\psi} H \psi \\ &= \frac{d}{d\alpha} \int dx \frac{\psi^* H \psi}{\int dx' |\psi|^2}\end{aligned}\tag{41}$$

Since the Hamiltonian is hermitian, one has  $\int dx \psi^* H \psi = \int dx H \psi^* \psi$ , giving:

$$\begin{aligned}\bar{E}_\alpha &= \frac{d}{d\alpha} \int dx \frac{H \psi^* \psi}{\int dx' |\psi|^2} \\ &= \left[ \int dx \frac{H \left( \psi^* \left( \frac{d\psi}{d\alpha} \right) + \left( \frac{d\psi^*}{d\alpha} \right) \psi \right)}{\int dx' |\psi|^2} \right] - \left[ \int dx \frac{H \psi^* \psi}{\left( \int dx' |\psi|^2 \right)^2} \int dx' \left( \psi^* \left( \frac{d\psi}{d\alpha} \right) + \left( \frac{d\psi^*}{d\alpha} \right) \psi \right) \right]\end{aligned}\tag{42}$$

Again one may use the hermiticity of the Hamiltonian to get  $\int dx H \psi^* \left( \frac{d\psi}{d\alpha} \right) = \int dx H \left( \frac{d\psi^*}{d\alpha} \right) \psi$ . So:

$$\begin{aligned}\bar{E}_\alpha &= 2 \left[ \int dx \frac{H \psi^* \frac{d\psi}{d\alpha}}{\int dx' |\psi|^2} \right] - 2 \left[ \int dx \frac{H \psi^* \psi}{\left( \int dx' |\psi|^2 \right)^2} \int dx' \psi^* \frac{d\psi}{d\alpha} \right] \\ &= 2 \left[ \int dx \frac{H \psi^* \frac{d\psi}{d\alpha}}{\int dx' |\psi|^2} - \int dx \frac{H \psi^* \psi}{\int dx' |\psi|^2} \int dx' \frac{1}{\int dx' |\psi|^2} \psi^* \frac{d\psi}{d\alpha} \right] \\ &= 2 \left[ \int dx \frac{\psi^* \left( \frac{E_L}{\psi} \frac{d\psi}{d\alpha} \right) \psi}{\int dx' |\psi|^2} - \int dx \frac{\psi^* E_L \psi}{\int dx' |\psi|^2} \int dx' \frac{\psi^* \left( \frac{1}{\psi} \frac{d\psi}{d\alpha} \right) \psi}{\int dx' |\psi|^2} \right] \\ \therefore \bar{E}_\alpha &= 2 \left( \left\langle \frac{\bar{\psi}_\alpha}{\psi} E_L \right\rangle - \left\langle \frac{\bar{\psi}_\alpha}{\psi} \right\rangle \langle E_L \rangle \right)\end{aligned}\tag{43}$$

where  $\bar{\psi}_\alpha \equiv \frac{d\psi}{d\alpha}$ . Obviously, the derivative with respect to  $\beta$  is the same. The result can also be written

$$\bar{E}_\alpha = 2 \left( \langle [\ln(\psi)]_\alpha E_L \rangle - \langle [\ln(\psi)]_\alpha \rangle \langle E_L \rangle \right)\tag{44}$$

due to the identity  $\frac{1}{x} \frac{d}{dx} f(x) = \frac{d}{dx} [\ln(f(x))]$ .

In order to find the optimal parameters, the ones that give minimal energy, we set a criteria that if the new parameters,  $\vec{x}_{n+1}$ , suggested by equation 40 has a shorter gradient than the current,  $\vec{x}_n$ , we accept this change, whereas if it doesn't we reset the parameters to the current and shorten the steplength by a factor 0.7. We are content when the length of the gradient is sufficiently short and store these values as the optimal parameters.

The optimal parameters produced are not trivial however. As illustrated in figure 1 the minimum value of the local energy depends much more strongly on the value of  $\alpha$  than the value of  $\beta$  in the two-electron system. Our approach will most likely produce a value of  $\alpha$  that is in agreement with the true minimum, but our value of  $\beta$  may differ somewhat. However, this difference does not seem to have much of a consequence for this system. For the case of the two-electron system the optimal parameters found were  $\alpha = 1.003$  and  $\beta = 0.3$ , which resulted in a local energy of  $\langle E_L \rangle = 3.003$ . This is sufficiently close to the real minimum of 3, and seems to be in agreement with the figure as well.

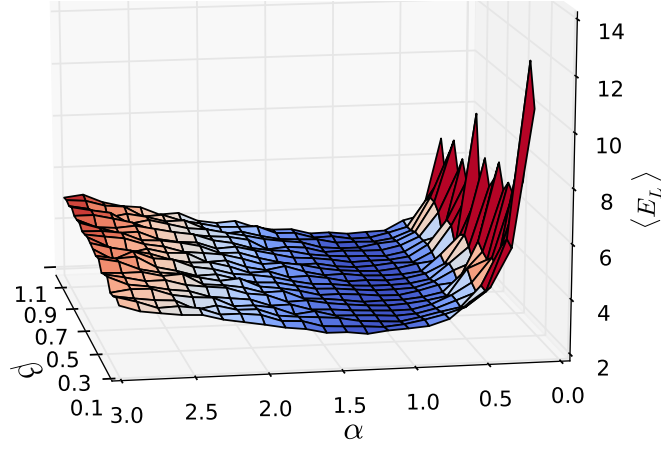


Figure 1: Local energy as a function of the parameters  $\alpha$  and  $\beta$  in the two-electron system. Red colored regions represents high expectation value of the local energy, while blue represents low values.

## 2.9 Parallelization

Parallaizing codes can be very beneficial. It can utilize the entire workforce of a computer, or it can occupy a whole super cluster working towards the same goal. The parallalization of the kind needed for this project is fairly easy. There is no communication between processes during the run, and writing to files can be done separately. There is also vectorization of the code that can be seen as an internal parallazation, done by using the whole register, making use of ordering operations in a parallel order. The vectorization that is implemented in this project are only done by adding compiler flags. The change needed in the code to parallelize this problem can consist of only two MPI functions.

## 2.10 Blocking method

After having run a Monte Carlo simulation, the variation in estimated local energies are calculated as

$$\sigma = \sqrt{\frac{1}{n} (\langle E_L^2 \rangle - \langle E_L \rangle^2)}. \quad (45)$$

However, these values will be much too low. This is because we assume all data to be completely uncorrelated. Each energy is calculated by a small perturbation to the system setting<sup>6</sup>, which means each new setting is very dependant on the previous setting. After sufficiently many perturbations, though, the system at step  $i$  will be so different from that of step  $j$ , that  $\langle E_L \rangle_i$  is basically uncorrelated to  $\langle E_L \rangle_j$ , but not for all  $\langle E_L \rangle_k$  between  $i$  and  $j$ . Ideally, we would like to find a *correlation time*  $\tau$  such that  $i$  and  $j$  will be uncorrelated if a time greater than  $\tau$  has passed. If  $\Delta t$  is the time between two Metropolis steps, then we would like to find  $|i - j|$  in  $\tau = |i - j|\Delta t$ .

A method of dealing with this is the blocking technique. The set of  $\langle E_L \rangle$  measurements is grouped into blocks, each of which will give an average of average local energies. Afterwards, we can calculate the variance of these averages of averages. If then the standard deviation (equation above) is plotted as a function of the number of blocks<sup>7</sup>, we can find the lowest number of blocks where the curve approaches a plateau. Then we can calculate  $\tau$  and find the correlation time. The true standard deviation is then:

$$\sigma = \sqrt{\frac{1 + 2\tau/\Delta t}{n} (\langle E_L^2 \rangle - \langle E_L \rangle^2)} \quad (46)$$

Some examples of such plots are shown in figure 3. For the case of  $N = 2$  and  $\omega = 1$  shown in figure 3a we could choose  $\tau \approx 3000$ .

## 2.11 Benchmarking

While the flowchart seems nice and compact, there is quite a bit code that needs implementing and some benchmarks would be nice. The first is so check if the program reproduces the energy expected for non-interacting electrons in a harmonic oscillator potential. This means that if we remove the Coulomb potential

<sup>6</sup>In the program discussed here, only a single dimension of a single particle will be moved each step.

<sup>7</sup>Inversely proportional to the block size by block size =  $\frac{\text{number of samples}}{\text{number of blocks}}$

and set  $a_{ij} = 0$ , then the problem is simply several non-interacting harmonic oscillators and we are left with a simple harmonic oscillator with energies given as

$$E_{n_x, n_y} = \hbar\omega (n_x + n_y + 1). \quad (47)$$

In figure 2 these energies levels are shown. If sums of these values are reproduced, it means our Slater determinant expressions are correctly implemented.

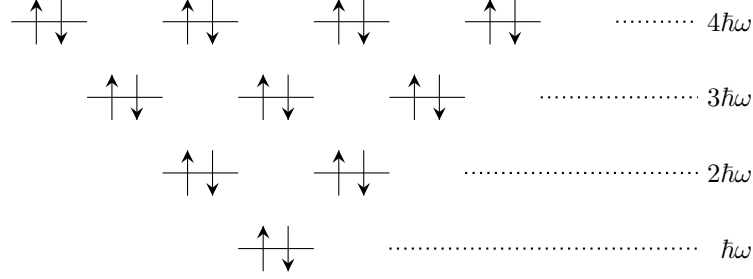


Figure 2: Illustrative figure of spin configurations in the closed shell system.

Another benchmark is to reproduce the results of our program for the two body quantum dot. This system was solved analytically, and thus would show that the many body program can reproduce analytical results for 2 particles. This means that the full problem (Coulomb interactions and Jastrow correlations) works, at least for the two body system.

### 3 Results

#### 3.1 Ground state energies, blocking, and benchmarks

Table 1: Expectation values of local energy, kinetic energy and potential energy for  $N = \{2, 6, 12, 20\}$  and oscillator frequencies  $\omega = \{1.0, 0.5, 0.1, 0.05, 0.01\}$ .  $\sigma$  is the standard deviation of  $\langle E_L \rangle$ , found by blocking.  $\alpha, \beta$  are the optimal parameters. Note that for the  $N = 2$  case, we used  $10^9$  Metropolis steps, while for  $N \geq 6$  we used  $10^7$  due to time shortage.

$N$	$\omega$	$\langle E_L \rangle$	$\langle E_K \rangle$	$\langle E_P \rangle$	$\sigma$	$\alpha$	$\beta$
2	1.00	3.0031	0.9017	2.1013	1.53360e-6	1.09975	0.29472
	0.50	1.6611	0.4559	1.2052	2.10014e-7	1.00015	0.29972
	0.10	0.4415	0.0964	0.3451	5.09917e-5	0.99534	0.16965
	0.05	0.2722	0.0480	0.2242	1.04670e-6	0.99591	0.22695
	0.01	0.0873	0.0107	0.0766	4.48831e-4	0.99777	0.19004
6	1.00	20.204	3.8059	16.398	1.43517e-3	1.00127	0.46939
	0.50	11.821	1.7797	10.042	1.60940e-3	0.97611	0.32474
	0.10	3.7466	0.2878	3.3077	2.22848e-3	0.99293	0.32392
	0.05	2.4792	0.2165	2.2627	3.35813e-3	0.96891	0.37989
	0.01	0.8972	0.0472	0.8500	1.84693e-3	1.00751	0.21825
12	1.00	65.776	8.7763	57.000	1.46457e-2	0.80173	0.80030
	0.50	39.223	4.0024	35.221	1.03759e-2	0.80003	0.50039
	0.10	14.683	1.1936	13.489	3.82364e-2	1.00071	0.69019
	0.05	7.6569	0.4908	7.1661	6.69810e-3	0.94721	0.15057
	0.01	3.2558	0.0937	3.1621	2.22688e-2	0.70992	0.34371
20	1.00	157.48	20.430	137.05	2.35715e-2	0.92930	0.80390
	0.50	93.981	7.9372	86.044	1.42877e-2	0.82140	0.47966
	0.10	30.157	1.1977	28.959	1.63797e-2	0.53094	0.39662
	0.05	18.667	0.5123	18.155	1.45351e-2	0.43081	0.31274
	0.01	6.2090	0.0821	6.1269	6.18577e-3	0.37267	0.12807

In order to give a correct value of the standard deviation, which also considers the covariance, we used the method of blocking described in section 2.10. A couple of the plots produced is shown in figure 3. These plots

show a trend in the behavior of the correlation length  $\tau$ . Comparing figures on the same row, we see that  $\tau$  increases as  $\omega$  decreases. Also, if we compare figures on the same column, we see that  $\tau$  increases as  $N$  increases.

The reason why  $\tau$  increases as  $N$  increases is intuitively simple to understand. If we have a system of many particles, more particles will be stationary at each time step. It will have a higher degree of correlation than that of a system of less particles. Thus, since we have to perform more moves in order to make a configuration that is basically uncorrelated with the initial, the correlation length is larger.

The other behaviour, that  $\tau$  increases as  $\omega$  decreases, is due to the fact that as  $\omega$  decreases, the particles are free to move farther away from the origin, as can be observed by comparing figure 4 to figure 5. Since the electrons will be more spread out in space the relative change in the configuration due to a move will be smaller. Therefore we must perform more moves to reach an approximately uncorrelated configuration.

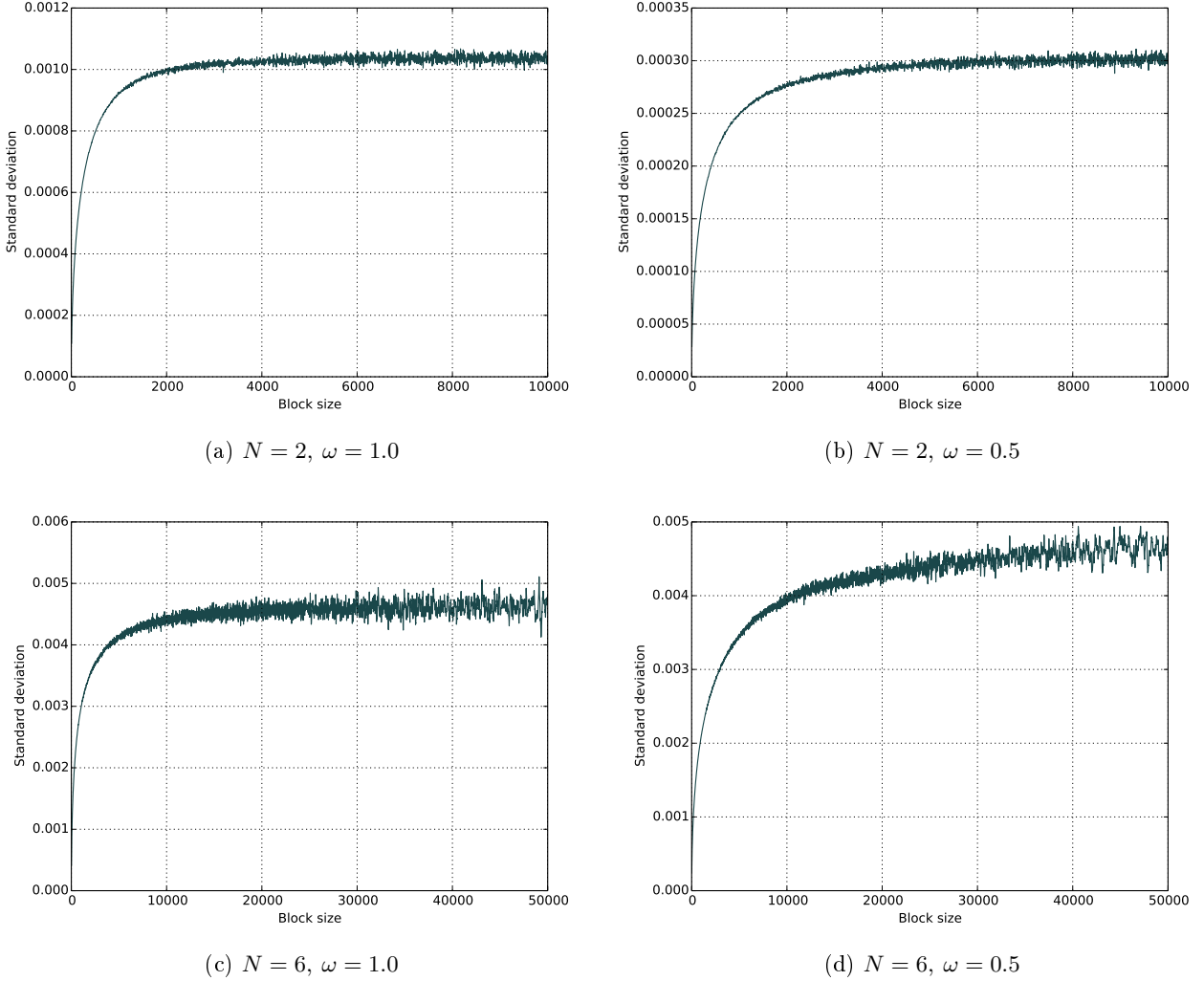


Figure 3: Plots produced by the blocking method for systems of  $N = \{2, 6\}$  and  $\omega = \{1, 0.5\}$ .

Table 2: Local energy computed for systems without the Jastrow factor and Coulomb potential. This resembles a pure harmonic oscillator and the results are dead on.

$N$	$E_L$	$\sigma$
2	2	0
6	10	0
12	28	0
20	60	0

This matches the sums of the energy levels shown in figure 2, and infer that our Slater determinant expres-

sions are implemented correctly.

In table 3, the energies for the two body system are presented, and were calculated using the analytical expressions from section 2.2. As is known, the energy for  $\omega = 1.0$  and  $N = 2$  is exactly 3, and the table confirms this to quite a high accuracy.

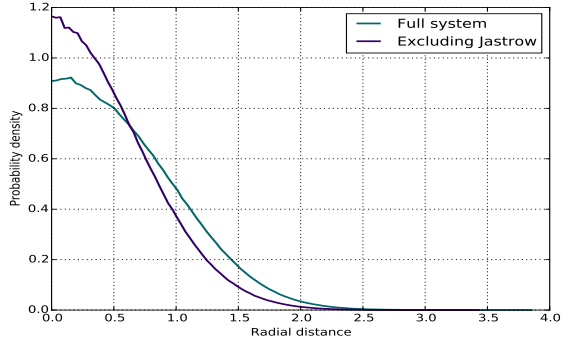
The average distance between the electrons for  $\omega = 1.0$  is found to be  $\langle r_{12} \rangle = 1.8$ .

Table 3: Expectation value of the kinetic energy and potential energy for several values of  $\omega$ .

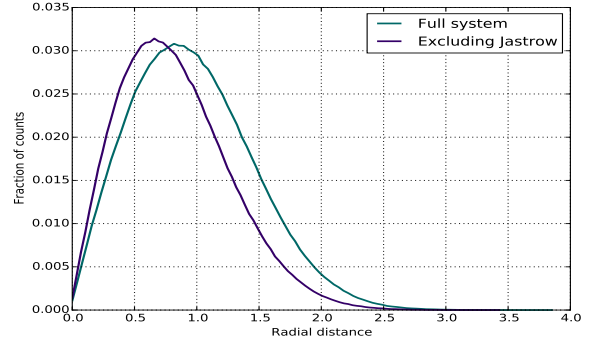
$\omega$	$\langle E_K \rangle$	$\langle E_P \rangle$	$\langle E_L \rangle$
0.01	0.0122	0.0969	0.0863
0.05	0.0511	0.2274	0.2769
0.1	0.0939	0.3544	0.4593
1.0	0.8983	2.1475	3.0002

Note, however, that each of the values listed in table 3 are from individual runs. The expectation values of the kinetic, potential and total (local) energy are thus not from the same simulation, and their value might not add up perfectly as  $E_L = E_K + E_P$ . The total energy is only shown to give a sense of its magnitude.

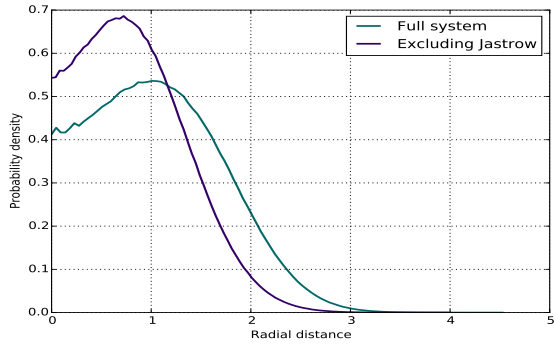
### 3.2 One-body densities and radial distributions



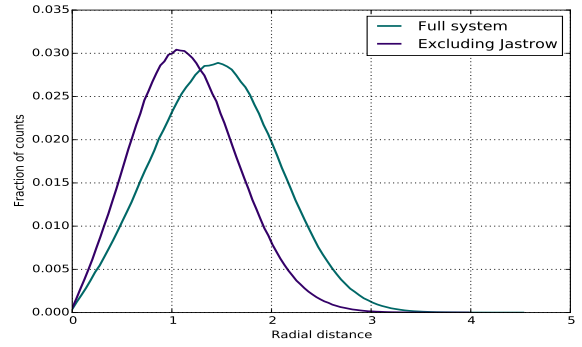
(a)  $N = 2, \omega = 1.0$



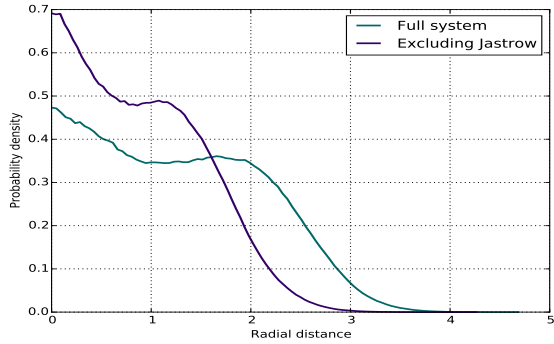
(b)  $N = 2, \omega = 1.0$



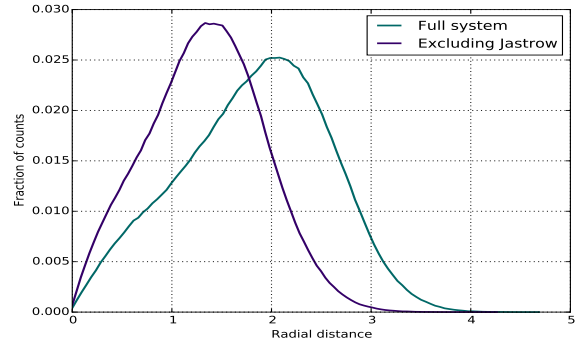
(c)  $N = 6, \omega = 1.0$



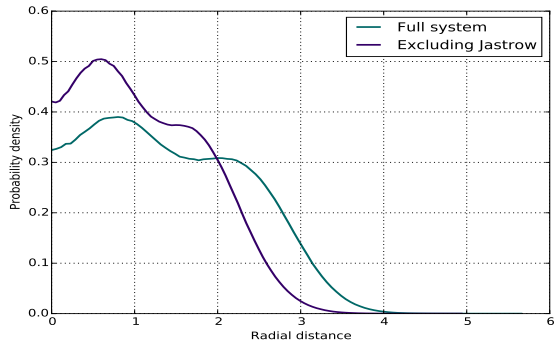
(d)  $N = 6, \omega = 1.0$



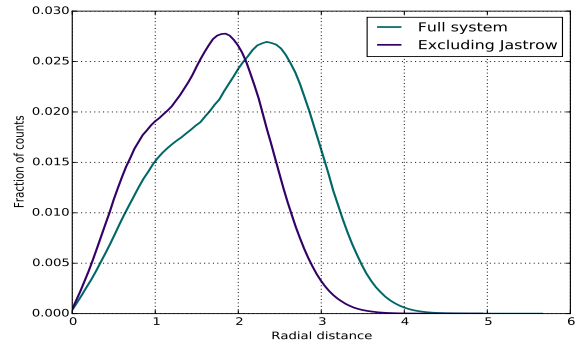
(e)  $N = 12, \omega = 1.0$



(f)  $N = 12, \omega = 1.0$

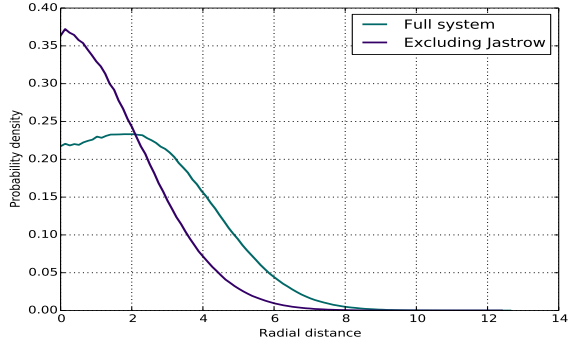


(g)  $N = 20, \omega = 1.0$

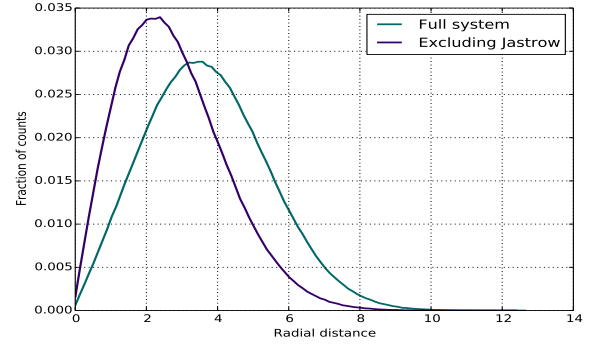


(h)  $N = 20, \omega = 1.0$

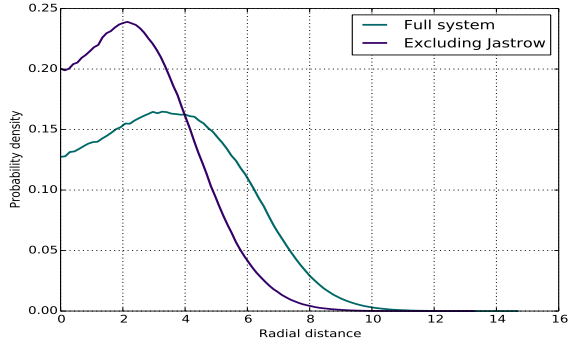
Figure 4: One-body densities (left) and radial distributions (right) for  $N = \{2, 6, 12, 20\}$  and  $\omega = 1.0$ .



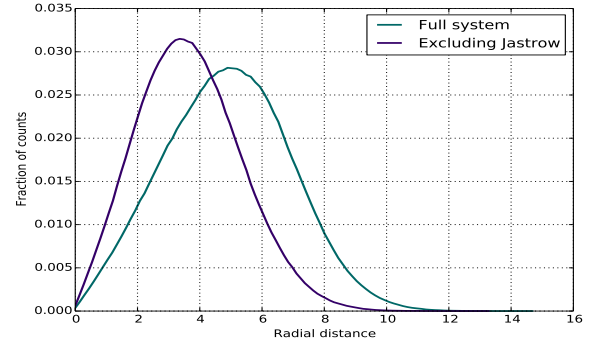
(a)  $N = 2$ ,  $\omega = 0.1$



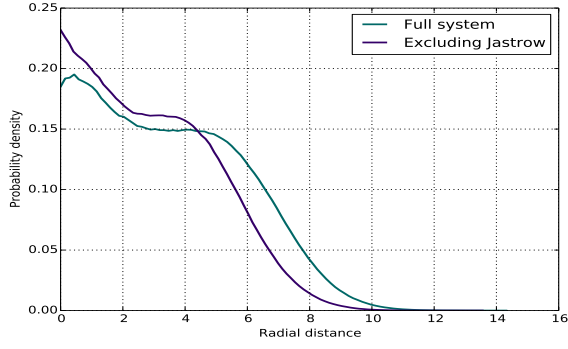
(b)  $N = 2$ ,  $\omega = 0.1$



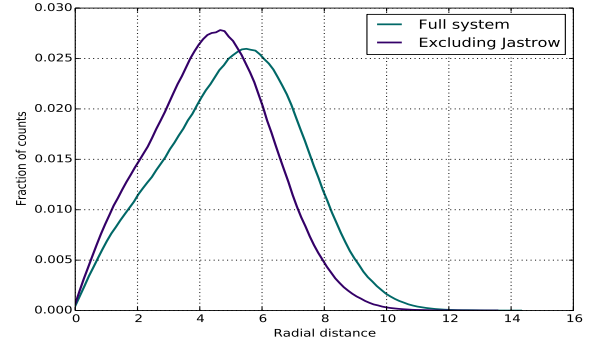
(c)  $N = 6$ ,  $\omega = 0.1$



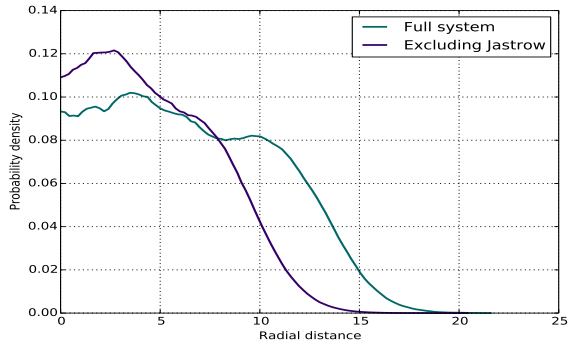
(d)  $N = 6$ ,  $\omega = 0.1$



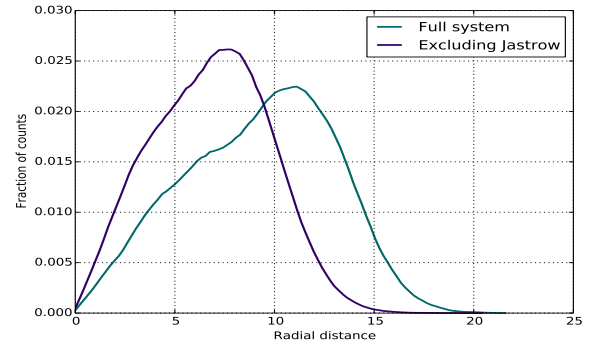
(e)  $N = 12$ ,  $\omega = 0.1$



(f)  $N = 12$ ,  $\omega = 0.1$



(g)  $N = 20$ ,  $\omega = 0.1$



(h)  $N = 20$ ,  $\omega = 0.1$

Figure 5: One-body densities (left) and radial distributions (right) for  $N = \{2, 6, 12, 20\}$  and  $\omega = 0.1$ .



### 3.3 Parallelization and vectorization

$p$	$N/p$	Unvectorized		Vectorized O2		Vectorized O3	
		$T$	$T_s/T_p$	$T$	$T_{O0}/T_{O2}$	$T$	$T_{O0}/T_{O3}$
1	8e6	757.0298	1	220.0905	3.4396	216.6741	3.4939
2	4e6	374.8545	2.0195	110.5357	3.3913	109.6561	3.4185
4	2e6	187.8710	4.0295	56.0112	3.3542	55.4134	3.3904
8	1e6	95.0406	7.9653	28.6783	3.3140	28.2144	3.3685
16	5e5	48.7444	15.5306	14.6023	3.1154	15.6096	3.1327

Table 4: Benchmark of program with timing starting before the MPI initialization and ending after the MPI finalization. For one process it is without MPI. Each process runs  $N/p$  number of steps where  $N = 8e1$  and  $p$  is the number of processes. All compiled on smaug and run on smaug-c with different vectorization flags in the mpicxx compiler.

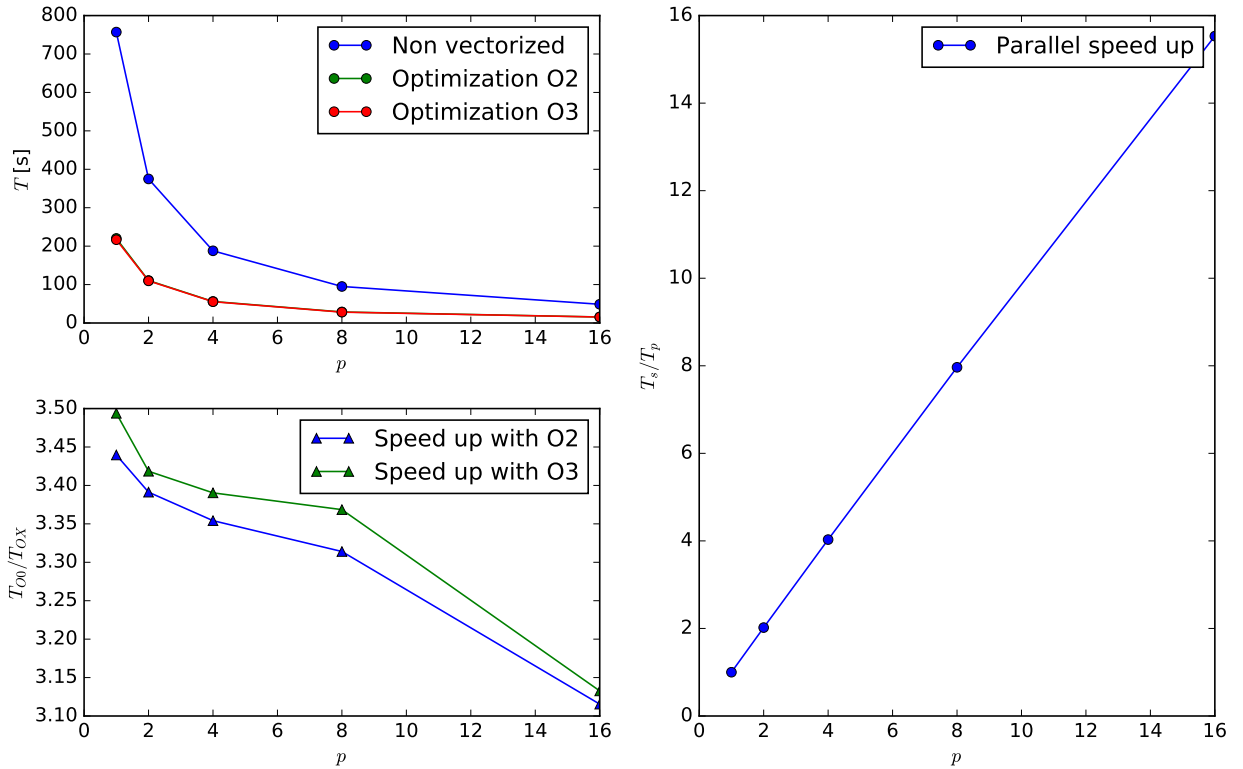


Figure 6: The upper left graph is world time  $T$  with different amount of processors  $p$ , with and without vectorization flags when compiling, running a system with a total of  $8e6$  metropolis cycles. The graph to the lower left shows the effectiveness of using the mpicxx compiler vectorization flags O2 and O3 compared to the default flag O0. And to the right is the serial time  $T_s$  compared to parallel time  $T_p$  for different amount of processors.

## 4 Comments

### 4.1 Energies

The two-body problem was already known since it was solved analytically above, and running the two-body Monte-Carlo code gave a result 3.0002 a.u. for  $\hbar\omega = 1$  a.u. (table 3). In table 1 the results from the many-body code are presented. As we see, the 2-body energy is reproduced down to the third decimal, with an error of  $\sim 1.53e-6$ . The result isn't as good as for the two-body specific code, but acceptable for a numerical estimation of this quality.

In regard to the quantum virial theorem,

$$2\langle T \rangle = \sum_n \left\langle X_n \frac{dV}{dX_n} \right\rangle, \quad (48)$$

which, in the case of the pure harmonic oscillator potential, reduces to  $2\langle T \rangle = \langle V_{\text{tot}} \rangle \equiv \sum_n \langle V_n \rangle$ , where  $n$  indexes particle positions. We can see that it fits, very roughly<sup>8</sup>, for 2 particles and  $\omega > 0.1$ . When reducing the "strength" of the harmonic oscillator, i.e. reducing  $\omega$ , the Coulomb interaction effects will start to dominate the potential energy. Approximately, we see  $3\langle T \rangle \sim \langle V_{\text{tot}} \rangle$  for  $\omega = 0.1$ ,  $4\langle T \rangle \sim \langle V_{\text{tot}} \rangle$  for  $\omega = 0.05$ , and  $7\langle T \rangle \sim \langle V_{\text{tot}} \rangle$  for  $\omega = 0.01$ .

The energies in table 1 match quite well with [1], page 114. There is a small difference, usually in the second or third decimal, but it may simply be due to lack of statistics.

## 4.2 Distributions and densities

Figures 4 and 5 present the radial distributions and one-body densities for  $\omega = 1.0$  and  $\omega = 0.1$ , respectively. Except for the  $N = 12$  case, the full system and the system without Jastrow factor will move closer together when increasing  $\omega$ . This seems natural since the graphs should overlap perfectly in the pure harmonic oscillator potential, and increasing the frequency means dwarfing Coulomb effects. Therefore, the behaviour of the  $N = 12$  system seems strange. There is no apparent reason as to why this case is special when looking at the trial wavefunction, nor that the program should not work properly for 12 particles.

The one-body densities, which are tied to the total system wavefunctions, show that the spatial wavefunction is circularly symmetric<sup>9</sup>. Since we chose our total spin-function to be anti-symmetric, a symmetric spatial wavefunction is precisely what we would expect (fermions require anti-symmetric wavefunctions). This means it is more likely to find a particle in the centre of the system than further out, despite what the radial distributions suggest at a first glance.

It is also interesting to note that for each new full shell, there is another extrema in the wavefunction. In other words, we can see the particle "shells".

Finally, the Jastrow corrections seem to "push" the densities away from the centre when compared to the system without the Jastrow factor. The Coulomb interactions are repulsive between particles of same electric charge. Therefore, it seems only natural that when introducing this modification to the Slater determinant, it should spread the particles further out, which is indeed the case here. As we see, the total particle density will decrease.

## 4.3 Parallelization and vectorization

In column number 4 of table 4 the ratio shows a linear dependence between number of processes and achieved speed up, which can be seen more clearly in the graph below. Column 6 and 9 shows the speed up due to vectorization flags. There is not a large difference in speed up so the justification to use O3 would be for heavier runs than what were tested here. And since the O3 flag appends more compile options it could be more invasive and create errors in the compiled code. Hence we are better off using O2 in the small scale calculations.

From the graph in figure 6 to the right there is a clear linear dependence on the number of processes. Due to the simple nature of this kind of parallelizations there is very little time consumption regarding overhead and therefore linearity is to be expected. Also there is a danger of taking too many processes when not using enough metropolis cycles considering the number of cycles are divided to each process and therefore can reach too few cycles to have a good result. The difference in speed up due to vectorization O2 and O3 is about 0.05 independent of number of processes, but they both go down as the number increases. The exact nature of this is not known, but a possible explanation is that the amount of time spent going through loops on each process effects vectorization speed up.

<sup>8</sup>Which can mean the harmonic oscillator potential greatly dominates the total potential energy. The deviation assumably originates from Coulomb interactions.

<sup>9</sup>The plots presented here alone do not suggest spherical symmetry, but with the knowledge that the plots have no angular dependence we can say so (a function  $f(r, \phi) = f(r)$  is circularly symmetric).

## 5 Conclusions

We set out to expand upon the VMC code for bosonic systems written in the first project to one for fermionic systems. This mainly meant using an anti-parallel wavefunction so as to respect the Pauli exclusion principle, and was done by using a Slater determinant. The Slater determinant would greatly increase the number of necessary calculations if we didn't split the trial wavefunction, which can be done in closed shell systems. This was done by placing the first half of the particles in an up-spin state and the other half in a down-spin state. This configuration made it possible to only calculate one of the spin matrices per cycle. Furthermore, by the use of some neat relations, only certain parts of the matrices, and their inverse matrices, had to be updated with a new position.

The ground-state energies were found for  $N = 2, 6, 12$ , and 20 electrons, and fit very well with results already derived in [1]. Plotting the one-body densities showed that the shells appear in the system wavefunction, and how the Jastrow factor affected the particle positions. It confirmed expectations since the Jastrow factor was intended to compensate our ignorance of the system, which it did for repulsive interactions.

We saw that the virial theorem still holds true, but has a very different nature for many particles in a two-body Coloumb interaction. Whether it confirms expectations was a bit hard to say, due to the complex form of the full potential.

Adding parallalization increased the speed up of the program linearly as a function of number of processes. And vectorizing it through the mpicxx compiler increased the efficiency by between 3.10 and 3.50 with a decreasing factor  $\propto 1/p$ . In conclusion, running a VMC code without vectorization nor parallelization, means a completely impractical world time. Furthermore, we saw that the difference between  $O2$  and  $O3$  vectorization is negligible.

## References

- [1] J. Høgberget, "*Quantum Monte-Carlo Studies of Generalized Many-Body Systems*", 2013, UiO
- [2] M. Hjorth-Jensen, "*Computational Physics, Lecture Notes Fall 2015*", UiO