

Enriching continuous Lagrange finite element approximation spaces using neural networks

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Joint work with:

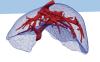
H. Barucq, F. Faucher, N. Victorion and V. Michel-Dansac.





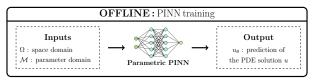
Scientific context

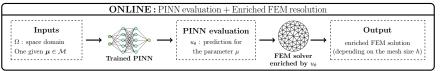
Context: Create real-time digital twins of an organ (e.g. liver).



Objective : Develop an hybrid finite element / neural network method.

accurate quick + parameterized





Problem considered

Stationary incompressible Navier-Stokes equations (with buoyancy and gravity):

We consider $\Omega = [-1,1]^2$ a squared domain and ${\bf e}_{\bf y} = (0,1)$.

Find the velocity ${m u}=(u_1,u_2)$, the pressure p and the temperature T such that

$$\begin{cases} (\textbf{\textit{u}}\cdot\nabla)\textbf{\textit{u}} + \nabla\rho - \mu\Delta\textbf{\textit{u}} - \textbf{\textit{g}}(\beta\textbf{\textit{T}}+1)\textbf{\textit{e}}_{\textbf{\textit{y}}} = 0 & \text{in }\Omega & \text{(momentum)} \\ \nabla\cdot\textbf{\textit{u}} = 0 & \text{in }\Omega & \text{(incompressibility)} \\ \textbf{\textit{u}}\cdot\nabla\textbf{\textit{T}} - k_{\textbf{\textit{f}}}\Delta\textbf{\textit{T}} = 0 & \text{in }\Omega & \text{(energy)} \\ + & \text{suitable BC} \end{cases}$$

with g=9.81 the gravity, $\beta=0.1$ the expansion coefficient, μ the viscosity and $k_{\rm f}$ the thermal conductivity. [Coulaud et al., 2024]

Problem considered

Objective: Simulation on a range of parameters $\mu = (\mu, k_f) \in \mathcal{M} = [0.01, 0.1]^2$.

Stationary incompressible Navier-Stokes equations (with buoyancy and gravity):

We consider $\mathbf{x} = (\mathbf{x}, \mathbf{y}) \in \Omega$ and $\mathbf{e}_{\mathbf{y}} = (0, 1)$. Find $\mathbf{U} = (\mathbf{u}, \mathbf{p}, \mathbf{T}) = (\mathbf{u}_1, \mathbf{u}_2, \mathbf{p}, \mathbf{T})$ such that

$$\begin{cases} \textit{R}_{\textit{mom}}(\textit{U}; \mathbf{x}, \boldsymbol{\mu}) = 0 & \text{in } \Omega & \text{(momentum)} \\ \textit{R}_{\textit{inc}}(\textit{U}; \mathbf{x}, \boldsymbol{\mu}) = 0 & \text{in } \Omega & \text{(incompressibility)} \\ \textit{R}_{\textit{ener}}(\textit{U}; \mathbf{x}, \boldsymbol{\mu}) = 0 & \text{in } \Omega & \text{(energy)} \\ + & \text{suitable BC} \end{cases}$$

with ${\it g}=9.81$ the gravity, ${\it \beta}=0.1$ the expansion coefficient, ${\it \mu}$ the viscosity and ${\it k_f}$ the thermal conductivity. [Coulaud et al., 2024]

Problem considered

Objective: Simulation on a range of parameters $\mu = (\mu, k_f) \in \mathcal{M} = [0.01, 0.1]^2$.

Stationary incompressible Navier-Stokes equations (with buoyancy and gravity):

We consider $\mathbf{x} = (x, y) \in \Omega$ and $\mathbf{e}_y = (0, 1)$.

Find $\boldsymbol{u} = (\boldsymbol{u}, p, T) = (u_1, u_2, p, T)$ such that

$$\begin{cases} \textit{R}_{\textit{mom}}(\textit{U}; \textbf{\textit{x}}, \boldsymbol{\mu}) = 0 & \text{in } \Omega & \text{(momentum)} \\ \textit{R}_{\textit{inc}}(\textit{U}; \textbf{\textit{x}}, \boldsymbol{\mu}) = 0 & \text{in } \Omega & \text{(incompressibility)} \\ \textit{R}_{\textit{ener}}(\textit{U}; \textbf{\textit{x}}, \boldsymbol{\mu}) = 0 & \text{in } \Omega & \text{(energy)} \end{cases} \tag{\mathcal{P}}$$

with g=9.81 the gravity, $\beta=0.1$ the expansion coefficient, μ the viscosity and $k_{\!f}$ the thermal conductivity. [Coulaud et al., 2024]

Boundary Conditions:

- $\mathbf{u} = 0$ on $\partial \Omega$
- $\mathit{T}=1$ on the left wall ($\mathit{x}=-1$) and $\mathit{T}=-1$ on the right wall ($\mathit{x}=1$) $\frac{\partial \mathit{T}}{\partial \mathit{n}}=0$ on the top and bottom walls ($\mathit{y}=\pm 1$, denoted by $\Gamma_{\rm ad}$)

Physics-informed neural network (PINN)

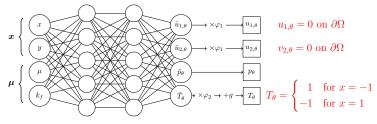
The PINN is parametrized by the μ parameter.

Neural Network considered

We consider a parametric NN with 4 inputs and 4 outputs, defined by

$$U_{\theta}(\mathbf{x}, \boldsymbol{\mu}) = (u_{1,\theta}, u_{2,\theta}, p_{\theta}, T_{\theta})(\mathbf{x}, \boldsymbol{\mu}).$$

The Dirichlet boundary conditions are imposed on the outputs of the MLP by a **post-processing** step. [Sukumar and Srivastava, 2022]



We consider two levelsets functions φ_1 and φ_2 , and the linear function ${\it g}$ defined by

$$\varphi_1(x,y) = (x-1)(x+1)(y-1)(y+1),$$

$$\varphi_2(x,y) = (x-1)(x+1) \quad \text{and} \quad g(x,y) = 1-(x+1).$$

PINN training

Approximate the solution of (\mathcal{P}) **by a PINN :** Find the optimal weights θ^{\star} , such that

$$\theta^* = \underset{\theta}{\operatorname{argmin}} \left(J_{\operatorname{inc}}(\theta) + J_{\operatorname{mom}}(\theta) + J_{\operatorname{ener}}(\theta) + J_{\operatorname{ad}}(\theta) \right), \tag{\mathcal{P}_{θ}}$$

where the different cost functions¹ are defined by

adiabatic condition

$$J_{ad}(heta) = \int_{\mathcal{M}} \int_{\Gamma_{\mathrm{ad}}} \left| rac{\partial au_{ heta}(\mathbf{x}, oldsymbol{\mu})}{\partial n}
ight|^2 d\mathbf{x} doldsymbol{\mu},$$

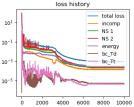
3 residual losses

$$\mathbf{J}_{\bullet}(\theta) = \int_{\mathcal{M}} \int_{\Omega} \left| \mathbf{R}_{\bullet}(\mathbf{U}_{\theta}(\mathbf{x}, \boldsymbol{\mu}); \mathbf{x}, \boldsymbol{\mu}) \right|^2 \mathrm{d}\mathbf{x} \mathrm{d}\boldsymbol{\mu},$$

with U_{θ} the parametric NN and • the PDE considered (i.e. *inc*, *mom* or *ener*).

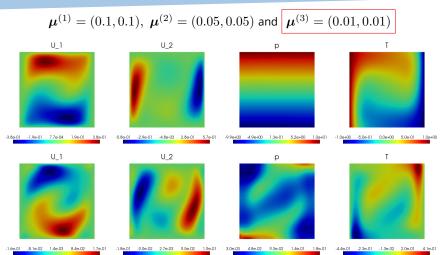
Network - MLP			
layers	40,60,60,60,40		
σ	sine		

Training (ADAM / LBFGs)				
lr	7e-3	$N_{ m col}$	40000	
$\overline{n_{epochs}}$	10000	$N_{ m bc}$	30000	



¹Discretized by a random process using Monte-Carlo method.

PINN solution



TODO : renommer figure u_{θ} ... (solutions et erreurs) + ajouter erreurs L2

Finite element method (FEM)

The μ parameter is fixed in the FE resolution.

Discrete weak formulation

We consider a mixed finite element space $M_h = [V_h^0]^2 imes Q_h imes W_h$ and

with
$$W = \{ w \in H^1(\Omega), w|_{x=-1} = 1, w|_{x=1} = -1 \}.$$

Discrete weak formulation

We consider a mixed finite element space $M_h = [V_h^0]^2 imes Q_h imes W_h$ and

with $W = \{ w \in H^1(\Omega), w|_{x=-1} = 1, w|_{x=1} = -1 \}.$

$$\begin{aligned} \text{Weak problem : Find } U_h &= (\textbf{\textit{u}}_h, p_h, T_h) \in \textit{M}_h \text{ s.t., } \forall (\textbf{\textit{v}}_h, q_h, w_h) \in \textit{M}_h^0, \\ &\int_{\Omega} (\textbf{\textit{u}}_h \cdot \nabla) \textbf{\textit{u}}_h \cdot \textbf{\textit{v}}_h \, d\textbf{\textit{x}} + \mu \int_{\Omega} \nabla \textbf{\textit{u}}_h : \nabla \textbf{\textit{v}}_h \, d\textbf{\textit{x}} \\ &- \int_{\Omega} p_h \, \nabla \cdot \textbf{\textit{v}}_h \, d\textbf{\textit{x}} - g \int_{\Omega} (1 + \beta T_h) \textbf{\textit{e}}_y \cdot \textbf{\textit{v}}_h \, d\textbf{\textit{x}} = 0, \quad \text{(momentum)} \\ &\int_{\Omega} q_h \, \nabla \cdot \textbf{\textit{u}}_h \, d\textbf{\textit{x}} + 10^{-4} \int_{\Omega} q_h \, p_h \, d\textbf{\textit{x}} = 0, \quad \text{(incompressibility + pressure penalization)} \\ &\int_{\Omega} (\textbf{\textit{u}}_h \cdot \nabla T_h) \, w_h \, d\textbf{\textit{x}} + \int_{\Omega} k_f \nabla T_h \cdot \nabla w_h \, d\textbf{\textit{x}} = 0, \quad \text{(energy)} \end{aligned}$$

where $\mathbf{M}_h^0 = [\mathbf{V}_h^0]^2 \times \mathbf{Q}_h \times \mathbf{W}_h^0$ with $\mathbf{W}_h^0 \subset \{\mathbf{w} \in \mathbf{H}^1[\Omega], \ \mathbf{w}|_{\mathbf{x}=\pm 1} = 0\}.$

Newton method

We consider the following three parameters:

$$\boldsymbol{\mu}^{(1)} = (0.1, 0.1), \ \boldsymbol{\mu}^{(2)} = (0.05, 0.05) \text{ and } \boldsymbol{\mu}^{(3)} = (0.01, 0.01).$$

Denoting N_h the dimension of M_h , we want to solve the non linear system:

$$F(\vec{U}_k) = 0$$

with $F: \mathbb{R}^{N_h} \to \mathbb{R}^{N_h}$ a non linear operator and $\vec{U}_{\nu} \in \mathbb{R}^{N_h}$ the unknown vector associated to the *k*-th parameter $\mu^{(k)}$ (k=1,2,3). Appendix 1

Algorithm 1: Newton algorithm [Aghili et al., 2025]

Initialization step: set
$$\vec{U}_k^{(0)} = \vec{U}_{k,0}$$
;

for
$$n \geq 0$$
 do

Solve the linear system
$$F(\vec{U}_k^{(n)}) + F'(\vec{U}_k^{(n)}) \delta_k^{(n+1)} = 0$$
 for $\delta_k^{(n+1)}$; Update $\vec{U}_k^{(n+1)} = \vec{U}_k^{(n)} + \delta_k^{(n+1)}$;

end

Newton method

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$$\boldsymbol{\mu}^{(1)} = (0.1, 0.1), \ \boldsymbol{\mu}^{(2)} = (0.05, 0.05) \text{ and } \boldsymbol{\mu}^{(3)} = (0.01, 0.01).$$

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Algorithm 1: Newton algorithm [Aghili et al., 2025]

Initialization step: set
$$\vec{U}_k^{(0)} = \vec{U}_{k,0}$$
; for $n > 0$ do

Solve the linear system
$$F(\vec{U}_k^{(n)}) + F'(\vec{U}_k^{(n)}) \delta_k^{(n+1)} = 0$$
 for $\delta_k^{(n+1)}$; Update $\vec{U}_k^{(n+1)} = \vec{U}_k^{(n)} + \delta_k^{(n+1)}$;

end

How to initialize the Newton solver?

- · Natural initialization:
- DeepPhysics initialization :
- · Incremental initialization.

 Natural initialization: Using constant or linear function. Considering a fixed parameter with $k \in \{1, 2, 3\}$, we can use the following initialization:

$$ec{U}_{k,0} = \left(\mathbf{0}_{N_u}, \mathbf{0}_{N_u}, \mathbf{0}_{N_p}, ec{T}_0
ight)$$

where for $i = 1, \ldots, N_T$,

$$(\vec{\tau}_0)_i = g(\mathbf{x}^{(i)}) = 1 - (\mathbf{x}^{(i)} + 1)$$

with $\mathbf{x}^{(i)} = (\mathbf{x}^{(i)}, \mathbf{y}^{(i)})$ the *i*-th dofs coordinates of W_h .

- · DeepPhysics initialization:
- Incremental initialization.

- **Natural initialization :** Using constant or linear function.
- DeepPhysics initialization: Using PINN prediction [Odot et al., 2021]. Considering a fixed parameter with $k \in \{1, 2, 3\}$, we can use the following initialization for $i = 1, \ldots, N_h$,

$$\left(\vec{U}_{k,0}\right)_i = U_{\theta}(\mathbf{x}^{(i)}, \boldsymbol{\mu}^{(k)})$$

with $\mathbf{x}^{(i)} = (\mathbf{x}^{(i)}, \mathbf{y}^{(i)})$ the *i*-th dofs coordinates of M_h and U_{θ} the PINN.

Incremental initialization.

- **Natural initialization :** Using constant or linear function.
- **DeepPhysics initialization:** Using PINN prediction [Odot et al., 2021].
- **Incremental initialization.** Using a coarse FE solution of a simpler parameter.
 - We consider a fixed parameter with $k \in \{2, 3\}$.
 - We consider a coarse grid (16×16 grid) and compute the FE solution of (\mathcal{P}_h) for the parameter $\mu^{(k-1)}$.
 - We interpolate the coarse solution to the current mesh.
 - We use it as an initialization for the Newton method, i.e.

$$\vec{U}_{k,0} = (\vec{u}_{k-1}, \vec{v}_{k-1}, \vec{p}_{k-1}, \vec{T}_{k-1})$$

where \vec{u}_{k-1} , \vec{v}_{k-1} , \vec{p}_{k-1} and \vec{T}_{k-1} are the FE solutions for the parameter $\mu^{(k-1)}$.

Enriched space using PINN

Considering the PINN prior $U_{\theta} = (\mathbf{u}_{\theta}, p_{\theta}, T_{\theta})$, we define the mixed finite element space additively enriched by the PINN as follows:

$$M_h^+ = \{ U_h^+ = U_\theta + C_h^+, \quad C_h^+ \in M_h^0 \}$$

with
$$M_h^0 = [V_h^0]^2 \times Q_h \times W_h^0$$
, $U_h^+ = (\boldsymbol{u}_h^+, \rho_h^+, T_h^+) \in M_h^+$ and $C_h^+ = (\boldsymbol{C}_{h,\boldsymbol{u}}^+, C_{h,p}^+, C_{h,T}^+)$.

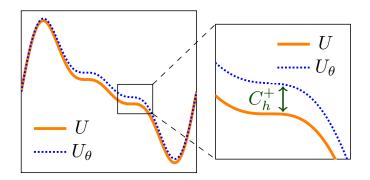
We can then define the three finite element subspaces of M_h^+ as follows:

$$\mathbf{V}_{h}^{+} = \left\{ \mathbf{u}_{h}^{+} = \mathbf{u}_{\theta} + \mathbf{C}_{h,\mathbf{u}}^{+}, \ \mathbf{C}_{h,\mathbf{u}}^{+} \in [V_{h}^{0}]^{2} \right\},
Q_{h}^{+} = \left\{ p_{h}^{+} = p_{\theta} + C_{h,p}^{+}, \ C_{h,p}^{+} \in Q_{h} \right\},
W_{h}^{+} = \left\{ T_{h}^{+} = T_{\theta} + C_{h,T}^{+}, \ C_{h,T}^{+} \in W_{h}^{0} \right\},$$

where $C_{h,u}^+$, $C_{h,n}^+$ and $C_{h,\tau}^+$ becomes the unknowns of the problem.

à ajouter : dans quoi vit U_{θ} ?

Schématisation en 1D de la correction... à voir comment présenter.



Weak formulation - Additive approach

Weak problem : Find
$$C_h^+ = (\boldsymbol{C}_{h,\boldsymbol{u}}^+, C_{h,p}^+, C_{h,T}^+) \in M_h^0$$
 s.t., $\forall (\boldsymbol{v}_h, q_h, w_h) \in M_h^0$,

$$\begin{split} \int_{\Omega} \left[(\boldsymbol{u}_{\theta} \cdot \nabla) \boldsymbol{u}_{\theta} + (\boldsymbol{u}_{\theta} \cdot \nabla) \boldsymbol{c}_{h,\boldsymbol{u}}^{+} + (\boldsymbol{c}_{h,\boldsymbol{u}}^{+} \cdot \nabla) \boldsymbol{u}_{\theta} + (\boldsymbol{c}_{h,\boldsymbol{u}}^{+} \cdot \nabla) \boldsymbol{c}_{h,\boldsymbol{u}}^{+} \right] \cdot \boldsymbol{v}_{h} \, d\boldsymbol{x} \\ &- \mu \left(\int_{\Omega} \Delta \boldsymbol{u}_{\theta} \boldsymbol{v}_{h} \, d\boldsymbol{x} - \int_{\Omega} \nabla \boldsymbol{c}_{h,\boldsymbol{u}}^{+} : \nabla \boldsymbol{v}_{h} \, d\boldsymbol{x} \right) + \left(\int_{\Omega} \nabla p_{\theta} \cdot \boldsymbol{v}_{h} \, d\boldsymbol{x} - \int_{\Omega} C_{h,\rho}^{+} \nabla \cdot \boldsymbol{v}_{h} \, d\boldsymbol{x} \right) \\ &- g \int_{\Omega} (1 + \beta (T_{\theta} + C_{h,T}^{+})) \boldsymbol{e}_{\boldsymbol{y}} \cdot \boldsymbol{v}_{h} \, d\boldsymbol{x} = 0, \text{ (momentum)} \\ \int_{\Omega} q_{h} \left[\nabla \cdot \boldsymbol{u}_{\theta} + \nabla \cdot \boldsymbol{c}_{h,\boldsymbol{u}}^{+} \right] d\boldsymbol{x} + 10^{-4} \int_{\Omega} q_{h} \left(p_{\theta} + C_{h,\rho}^{+} \right) d\boldsymbol{x} = 0, \text{ (incompressibility + penal)} \\ \int_{\Omega} \left[\boldsymbol{u}_{\theta} \cdot \nabla T_{\theta} + \boldsymbol{u}_{\theta} \cdot \nabla C_{h,T}^{+} + \boldsymbol{c}_{h,\boldsymbol{u}}^{+} \cdot \nabla T_{\theta} + \boldsymbol{c}_{h,\boldsymbol{u}}^{+} \cdot \nabla C_{h,T}^{+} \right] \boldsymbol{w}_{h} \, d\boldsymbol{x} \\ - k_{f} \left(\int_{\Omega} \Delta T_{\theta} \boldsymbol{w}_{h} \, d\boldsymbol{x} - \int_{\Omega} \nabla C_{h,T}^{+} \cdot \nabla \boldsymbol{w}_{h} \, d\boldsymbol{x} + \int_{\Gamma_{ad}} \frac{\partial C_{h,T}^{+}}{\partial n} \, \boldsymbol{w}_{h} \, d\boldsymbol{s} \right) = 0, \text{ (energy)} \end{split}$$

with $U_{ heta} = (oldsymbol{u}_{ heta}, p_{ heta}, T_{ heta})$ the PINN prior.

Parler des BC modifiés pour l'approche add?

Newton method - Additive approach

We want to solve the non linear system:

$$F_{\theta}(\vec{c}) = 0$$

with $F_{\theta}: \mathbb{R}^{N_h} \to \mathbb{R}^{N_h}$ the non linear operator associated to the weak problem (\mathcal{P}_h^+) and $\vec{C} \in \mathbb{R}^{N_h}$ the correction vector (unknown).

Algorithm 2: Newton algorithm [Aghili et al., 2025]

Initialization step: set $\vec{c}^{(0)} = 0$;

for $n \ge 0$ do

Solve the linear system $F_{\theta}(\vec{C}^{(n)}) + F'_{\theta}(\vec{C}^{(n)})\delta^{(n+1)} = 0$ for $\delta^{(n+1)}$;

Update $\vec{c}^{(n+1)} = \vec{c}^{(n)} + \delta^{(n+1)}$;

end

Advantage compared to DeepPhysics¹: Appendix 2

 u_{θ} is not required to live in the same discrete space as C_h^+ .

¹The additive approach is exactly the same as DeepPhysics if we take U_{θ} in the same space as C_h^+ .

Numerical results

TODO

Conclusion

TODO

Parler du papier en linéaire et dire que dans ce cadre on a des résultats théoriques de convergence.

References

- J. Aghili, E. Franck, R. Hild, V. Michel-Dansac, and V. Vigon. Accelerating the convergence of newton's method for nonlinear elliptic pdes using fourier neural operators. 2025.
- Guillaume Coulaud, Maxime Le, and Régis Duvigneau. Investigations on Physics-Informed Neural Networks for Aerodynamics, 2024.
- A. Odot, R. Haferssas, and S. Cotin. Deepphysics: a physics aware deep learning framework for real-time simulation, 2021.
- N. Sukumar and A. Srivastava. Exact imposition of boundary conditions with distance functions in physics-informed deep neural networks. 2022.

Appendix 1 : Finite element method (FEM)

A1 – Construction of the unknown vector

Considering $(\phi_i)_{i=1}^{N_u}$, $(\psi_j)_{j=1}^{N_p}$ and $(\eta_k)_{k=1}^{N_\tau}$ the basis functions of the finite element spaces V_h^0 , Q_h and W_h respectively, we can write the discrete solutions as:

$$\boldsymbol{u}_h(\boldsymbol{x}) = \sum_{i=1}^{N_u} \begin{pmatrix} u_i \\ v_i \end{pmatrix} \phi_i(\boldsymbol{x}), \quad \rho_h(\boldsymbol{x}) = \sum_{j=1}^{N_p} \rho_j \psi_j(\boldsymbol{x}) \quad \text{and} \quad T_h(\boldsymbol{x}) = \sum_{k=1}^{N_T} T_k \eta_k(\boldsymbol{x}),$$

with the unknown vectors for velocity, pressure and temperature defined by

$$\vec{u} = (u_i)_{i=1}^{N_u} \in \mathbb{R}^{N_u}, \quad \vec{v} = (v_i)_{i=1}^{N_u} \in \mathbb{R}^{N_u},$$

$$\vec{p} = (p_j)_{i=1}^{N_p} \in \mathbb{R}^{N_p} \text{ and } \vec{T} = (T_k)_{k=1}^{N_T} \in \mathbb{R}^{N_T}.$$

Considering $N_h = 2N_u + N_p + N_T$, we can define the global vector of unknowns as:

$$\vec{U} = (\vec{u}, \vec{v}, \vec{p}, \vec{T}) \in \mathbb{R}^{N_h}$$
.

and $F: \mathbb{R}^{N_h} \to \mathbb{R}^{N_h}$ the nonlinear operator associated to the weak formulation (\mathcal{P}_h).

Appendix 2 : DeepPhysics / Additive approach

A2 - ??