

## Active-Screen Gravity

### Conceptual Picture

We explore the possibility that gravity is not a fundamental constant interaction but an emergent response of microscopic degrees of freedom associated with a Planck-scale screen.

In this interpretation the Newton coupling becomes scale dependent  $G(\mu)$ . Matter fluctuations induce a renormalization flow described by

$$dG/d\ln\mu = a G^2$$

whose solution reads

$$G(\mu) = G_0 / (1 - a G_0 \ln(\mu/\mu_0)).$$

The coupling is therefore nearly constant in the infrared but grows toward ultraviolet scales.

### Vacuum Energy as Screen Tension

If the microscopic screen defines the gravitational response, its characteristic length obeys  $L^2 \propto 1/G$ .

The vacuum energy density becomes surface tension energy

$$\rho_\Lambda \sim 1/(L^2 G)$$

leading directly to  $\Lambda(\mu) \propto G(\mu)$ .

Hence the cosmological constant is not constant but follows the renormalization flow.

### Cosmic Expansion

Identifying  $\mu$  with the Hubble rate  $H$  gives  $\Lambda(H) = \Lambda_0 / (1 - a G_0 \ln(H/H_0))$ .

From energy conservation the equation of state becomes

$$w(z) \approx -1 + v \ln(1+z)$$

with  $v \approx a G_0 \approx 10^{-2}$ .

The universe therefore mimics  $\Lambda$ CDM at zeroth order while predicting a measurable drift.

### Inflation

At very early times  $H$  is large and the vacuum energy approaches a quasi constant value.

This generates a de-Sitter phase without introducing an inflaton field.

The slow-roll parameter  $\epsilon \approx aG$  naturally lies near  $10^{-3}$ – $10^{-2}$  producing  $N \approx 50$ – $70$  e-folds and spectral index  $n_s \approx 0.965$ .

### Black-Hole Horizons

In strong gravity the scale becomes local  $\mu(r) \approx (M/r^3)^{1/4}$ .

The running coupling modifies the metric near the horizon, preventing perfect absorption and introducing partial reflection.

The system behaves as a cavity between the photon sphere and the critical RG surface.

### Gravitational-Wave Echoes

The round-trip time yields  $\Delta t \approx 2r_s \ln(M/M_{\text{pl}})$ , about 0.1 s for stellar-mass black holes.

The reflection coefficient  $\sim 10^{-2}$  predicts echoes detectable by next-generation detectors.

### Unified Picture

A single renormalization flow connects inflation, late cosmic acceleration and horizon microphysics.

Each regime corresponds to a different scale of the same coupling rather than independent mechanisms.

## Appendix:

### Conceptual Picture

Gravity is treated as an emergent interaction with a scale-dependent Newton coupling  $G(\mu)$ .

Matter fluctuations induce RG flow  $dG/d\ln\mu = a G^2$  with solution  $G(\mu) = G_0/(1 - aG_0 \ln(\mu/\mu_0))$ .

### Cosmic Expansion

Identifying  $\mu \sim H$  yields  $\Lambda(H) \propto G(H)$  and  $w(z) \approx -1 + v \ln(1+z)$  with  $v \approx aG_0 \approx 10^{-2}$ .

### Inflation

At large  $H$  the vacuum energy becomes quasi-constant generating de Sitter expansion with  $n_s \approx 0.965$  and  $r \approx 0.01$ – $0.1$ .

### Black-Hole Echoes

Local scale  $\mu(r) \approx (M/r^3)^{1/4}$  modifies near-horizon absorption producing GW echoes with delay  $\sim 0.1$  s for stellar-mass BH.

## Appendix A — Detailed Derivations

### 1. RG Solution

Starting from beta function:

$$dG/d\ln\mu = -a G^2$$

Separate variables:

$$dG/G^2 = -a d\ln\mu$$

Integrate:

$$-1/G = -a \ln\mu + C$$

Using boundary  $G(\mu_0)=G_0$ :

$$C = -1/G_0 - a \ln\mu_0$$

Therefore:

$$G(\mu) = G_0 / (1 - a G_0 \ln(\mu/\mu_0))$$

### 2. $\Lambda(H)$ relation

Assume screen length  $L^2 \propto 1/G$ .

Vacuum energy density:

$$\rho_\Lambda \sim 1/(L^2 G) \rightarrow \Lambda \propto G.$$

### 3. Equation of state

$$w = -1 - (1/3) d \ln \rho_\Lambda / d \ln a$$

Since  $\rho_\Lambda \propto \Lambda(H)$ :

$$w(z) \approx -1 + v \ln(1+z), v = a G_0$$

### 4. Slow-roll parameters

$$\varepsilon = -d \ln H / dN \approx a G(H)$$

$$r = 16\varepsilon, n_s = 1 - 4\varepsilon$$

## 5. Echo delay

Cavity between photon sphere and RG surface:

$$\Delta t \approx 2r_s \ln(M/M_{\text{pl}})$$