## **CMS Draft Analysis Note**

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2012/10/16

Head Id: 152977

Archive Id: 152897:152979M Archive Date: 2012/10/15

Archive Tag: trunk

Background Combination Procedure for the searches for direct electroweak chargino and neutralino production with three or more leptons using 9.2 fb<sup>-1</sup> of  $\sqrt{s}$  = 8 TeV CMS data

SUSY Multilepton EWKino Working Group CERN

#### Abstract

We present the procedure used for the combination of background estimates in multilepton channels in the EWKino search. Because multiple groups are providing results in three and four lepton channels using a variety of different techniques, different methods are used to combine different types of backgrounds in each channel. The methods by which the background estimates are obtained are discussed elsewhere.

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PDFTitle: Background Combination Procedure for the searches for direct electroweak

chargino and neutralino production with three or more leptons using 9.2

fb''(-1'') of sqrt''(s'') = 8 TeV CMS data

PDFSubject: CMS

PDFKeywords: CMS, physics, supersymmetry

Please also verify that the abstract does not use any user defined symbols



## 1 Introduction

- <sup>2</sup> The search for direct electroweak production of SUSY particles in multilepton modes [1] com-
- bines search regions with two or more leptons with different requirements on  $E_{\rm T}^{\rm miss}$ ,  $M_{\rm T}$ , dijet
- 4 mass, lepton pair mass, and others to maximize sensitivity to different types of electroweak
- 5 production. In the three and four lepton channels, multiple groups have designed and vali-
- 6 dated a variety of methods to estimate and understand the various contributions to background
- 7 models. These methods are discussed elsewhere [2–5]. In this note, we restrict ourselves to
- 8 discussing the procedure for the combination of these estimates in a way that maximizes con-
- 9 sistency. We provide justification for assumptions and choices made to accomplish this goal.

#### 10 2 Rare Processes

- 11 The background estimates for rare processes are taken from the simulation samples described
- in Table 1 by all three groups. Therefore, the estimates from these sources are synchronized
- and the resulting contributions are used.

Table 1: Summary of the MC samples for rare Standard Model processes. All datasets are produced with the MADGRAPH generator.

DBS Name	σ (pb)
/DYJetsToLLM-50.TuneZ2Star.8TeV-madgraph-tarball/Summer12.DR53X-PU_S10.START53_V7A-v1	3532.8
/DYJetsToLLM-10To50filter.8TeV-madgraph/Summer12.DR53X-PU.S10.START53.V7A-v1	
/TTJets.MassiveBinDECAY.TuneZ2star.8TeV-madgraph-tauola/Summer12.DR53X-PUS10.START53_V7A-v1	225.2
/TTZJets.8TeV-madgraph.v2/Summer12.DR53X-PU.S10.START53.V7A-v1	0.208
/TTWJets.8TeV-madgraph/Summer12.DR53X-PU.S10.START53.V7A-v1	0.232
/TTGJets.8TeV-madgraph/Summer12DR53X-PUS10START53.V7A-v1	2.166
/ZZZNoGstarJets.8TeV-madgraph/Summer12.DR53X-PU.S10.START53.V7A-v1	0.01922
/WWWJets.8TeV-madgraph/Summer12.DR53X-PUS10.START53.V7A-v1	0.08217
/ZZJetsTo4LTuneZ2star.8TeV-madgraph-tauola/Summer12.DR53X-PUS10.START53_V7A-v1	0.1769
/ZZJetsTo2L2Q_TuneZ2star.8TeV-madgraph-tauola/Summer12_DR53X-PU_S10_START53.V7A-v1	2.4487
/ZZJetsTo2L2Nu.TuneZ2star.8TeV-madgraph-tauola/Summer12.DR53X-PU.S10.START53.V7A-v3	0.3648
/WZJetsTo3LNu.TuneZ28TeV-madgraph-tauola/Summer12DR53X-PUS10START53.V7A-v1	1.0575
/WZJetsTo2L2Q.TuneZ2star.8TeV-madgraph-tauola/Summer12.DR53X-PU.S10.START53.V7A-v1	2.206
/WWJetsTo2L2Nu.TuneZ2star.8TeV-madgraph-tauola/Summer12.DR53X-PU.S10.START53.V7A-v1	5.8123
/WJetsToLNu.TuneZ2Star.8TeV-madgraph-tarball/Summer12.DR53X-PU.S10.START53.V7A-v1	37509
/TTWWJets.8TeV-madgraph/Summer12.DR53X-PU.S10.START53.V7A-v1	0.002
/WWGJets.8TeV-madgraph/Summer12.DR53X-PU.S10.START53.V7A-v1	1.44
/WWZNoGstarJets.8TeV-madgraph/Summer12.DR53X-PU_S10.START53_V7A-v1	0.0633
/WZZNoGstarJets.8TeV-madgraph/Summer12.DR53X-PU_S10_START53.V7A-v1	0.01922

### 3 Internal Conversion and ZZ

- 15 The authors of [2] and [4] use separate data-driven methods to estimate the internal conver-
- sion background to light leptons and an official production of ZZJets to estimate the ZZ ir-
- 17 reducible background. The authors of [3] use a private simulation sample that combines the
- ZZ irreducible background with the internal conversion background. We take the data-driven
- estimate described in [4, 5] and a synchronized estimate to account for the ZZ irreducible back-
- 20 ground.

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#### 4 WZ

- 22 WZ production makes up the largest irreducible background for three lepton search regions.
- 23 The main source of the background estimation comes from the official WZ simulation sample.
- 24 However, two of the groups apply corrections to make sure that the MET distribution is mod-
- eled properly [2, 4]. The data-driven corrections are found to be small and the final result is
- 26 computed as a simple average of three predictions.

## <sub>27</sub> 5 Fake leptons

The three groups have a variety of ways to estimate the backgrounds from fake leptons (both light leptons and taus). Because one group [2] includes the contribution from tt in their data-driven estimate, we average the data-driven contributions summed with the tt simulation contribution used by the two groups in the three lepton channels. Because only two groups [3, 4] are providing the four lepton channels, we synchronize the tt contribution for the four lepton channels and average the data-driven backgrounds.

## <sub>34</sub> 6 Final Tables for Tri-lepton Analyses

- A graphical comparison of the background estimations to be averaged is shown in detail in appendix A. Here, the summarized results are presented in five tables:
  - Table 2 contains observed yields and background prediction for each search region in a tri-lepton channel with an opposite sign same flavor lepton pair present ( $3\ell$ ).
  - Table 3 for a channel without an opposite sign same flavor lepton pair (no OSSF).
  - Table 4 for a channel with a same sign di-lepton and a hadronically decaying tau  $(SS\tau)$ . This result is fully based on Ref. [2].
  - Table 5 for a channel with an opposite sign opposite flavor di-lepton and a hadronically decaying tau  $(e^{\pm}\mu^{\mp}\tau)$ . This result is provided only in Ref. [3].
  - Table 6 for a channel with an opposite sign same flavor di-lepton and a hadronically decaying tau (OSSF $\tau$ ). This result is based on Refs. [3] and [4].
- 46 The graphical representation of the combined results is shown in Figures 1, 2, 3, 4, 5.

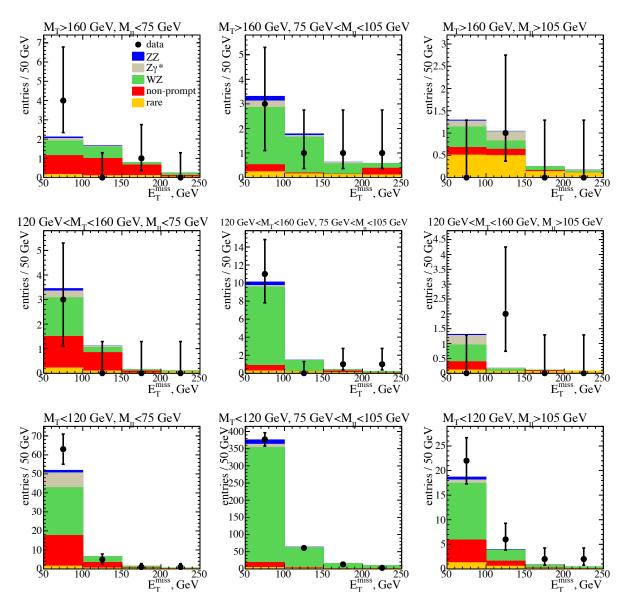


Figure 1: Observed yields and predicted backgrounds for a tri-lepton with an opposite sign same flavor lepton pair present in all defined search regions.

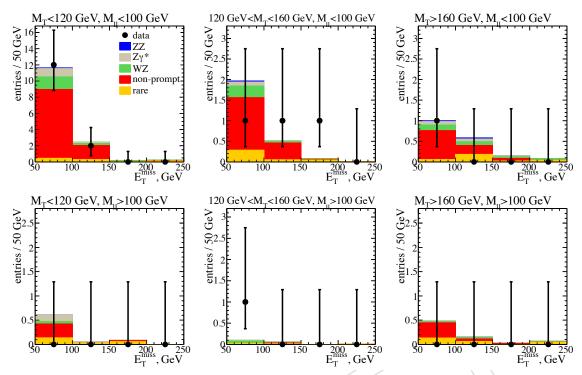


Figure 2: Observed yields and predicted backgrounds for a tri-lepton without an opposite sign same flavor lepton pair present in all defined search regions.

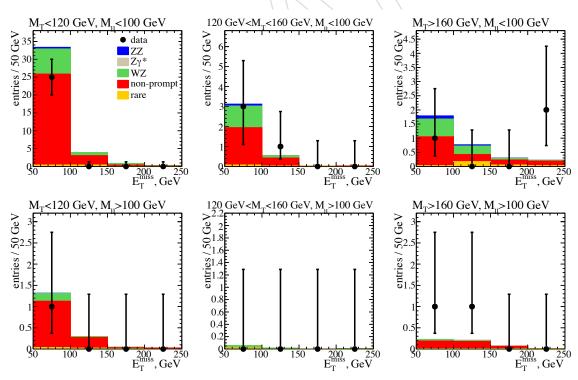


Figure 3: Observed yields and predicted backgrounds for a tri-lepton with a same sign dilepton and a hadronically decaying tau in all defined search regions.

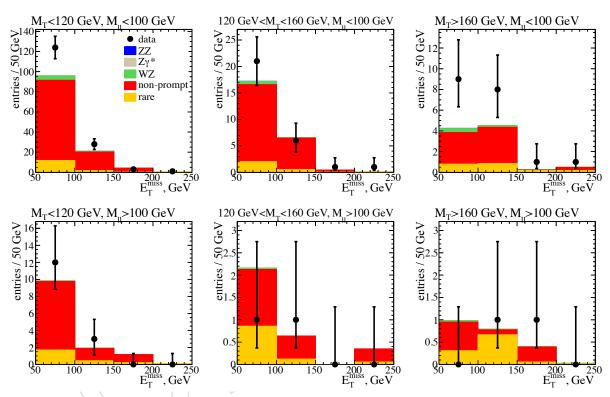


Figure 4: Observed yields and predicted backgrounds for a tri-lepton with an opposite sign opposite flavor di-lepton and a hadronically decaying tau in all defined search regions.

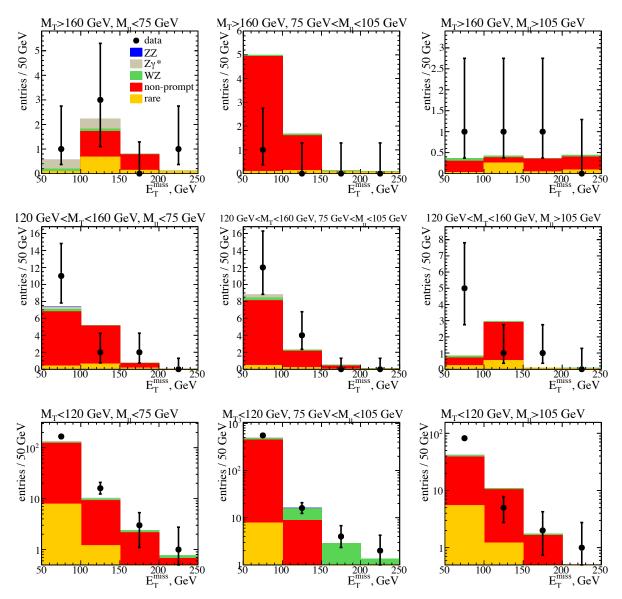


Figure 5: Observed yields and predicted backgrounds for a tri-lepton with an opposite sign same flavor di-lepton and a hadronically decaying tau in all defined search regions.

Table 2: The summary of the observed yields and predicted backgrounds for tri-lepton with opposite sign same flavor pair present.

$E_{\rm T}^{\rm miss}$ (GeV)	WZ	Non-Prompt	Rare SM	$Z\gamma^*$	ZZ	Total bkg	Observed
$M_{ m T} < 120~{ m Ge}$	${ m eV}$ , $M_{\ell\ell} < 75$ C	GeV					
50100	25±1.2	16±4.1	$1.4 \pm 0.84$	7.8±1	$1.3 \pm 0.15$	$52 \pm 4.5$	63
100150	$3.1 \pm 0.2$	$2.9 \pm 0.95$	$0.42{\pm}0.24$	$0.035 \pm 0.0044$	$0.1 {\pm} 0.014$	$6.5 \pm 1$	5
150200	$0.83 \pm 0.081$	$0.35 \pm 0.2$	$0.19\pm0.13$	$0.015\pm0.0019$	$0.018 \pm 0.0038$	$1.4 \pm 0.25$	1
200	$0.35 \pm 0.048$	$0.048 \pm 0.15$	$0.082 \pm 0.058$	$0.05 \pm 0.0063$	$0.0083 \pm 0.0024$	$0.54 \pm 0.17$	1
$120~{ m GeV} < \lambda$	$I_{\rm T} < 160~{ m GeV}$	$M_{\ell\ell} < 75 \mathrm{GeV}$					
50100	$1.6 \pm 0.12$	1.3±0.48	$0.22 \pm 0.14$	$0.27 \pm 0.035$	$0.1 {\pm} 0.014$	$3.4 \pm 0.51$	3
100150	$0.22 \pm 0.037$	$0.74 \pm 0.33$	$0.091 \pm 0.067$	$0.044{\pm}0.0056$	$0.015 \pm 0.0033$	$1.1 \pm 0.34$	0
150200	$0.06 \pm 0.019$	$0.08 \pm 0.16$	$0.0046 \pm 0.0052$	$0\pm0$	$0.00099 \pm 0.00075$	$0.15 \pm 0.16$	0
200	$0.041 \pm 0.016$	$0\pm0$	$0.044 \pm 0.044$	$0.025 \pm 0.0032$	$0.00078 \pm 0.00067$	$0.11 \pm 0.047$	0
$M_{\mathrm{T}} > 160~\mathrm{Ge}$	$eV$ , $M_{\ell\ell} < 75$ C	GeV					
50100	$0.76\pm0.077$	$0.98 \pm 0.42$	$0.16\pm0.19$	0.13±0.017	$0.081 \pm 0.011$	$2.1 \pm 0.46$	4
100150	$0.59 \pm 0.067$	$0.9 \pm 0.38$	$0.096 \pm 0.11$	$0.023\pm0.003$	$0.045 {\pm} 0.0071$	$1.7 \pm 0.4$	0
150200	$0.12 \pm 0.027$	$0.54 {\pm} 0.26$	$0.11 \pm 0.13$	0±0	$0.011 \pm 0.0028$	$0.78 \pm 0.3$	1
200	$0.13 \pm 0.029$	$0.073\pm0.19$	$0.033 \pm 0.038$	0±0	$0.0095 \pm 0.0025$	$0.25 \pm 0.2$	0
$M_{\mathrm{T}} < 120~\mathrm{Ge}$	eV, 75 GeV < 1	$M_{\ell\ell} < 105\mathrm{GeV}$					
50100	$335 \pm 14.6$	$14.2 \pm 3.38$	4.03±2.18	9.49±1.21	$13.1 \pm 1.49$	$376 \pm 15.3$	377
100150	$58 \pm 2.8$	$2.3 \pm 0.82$	1.8±1	$0.071\pm0.009$	$1.6 \pm 0.19$	$63 \pm 3.1$	61
150200	$14 \pm 0.76$	$0.38 \pm 0.17$	$0.63 \pm 0.35$	$0.0024 \pm 0.00031$	$0.37 \pm 0.044$	$16 \pm 0.86$	13
200	$8.7 \pm 0.49$	$0.027 \pm 0.084$	$0.55 {\pm} 0.3$	$0.031\pm0.0039$	$0.15 \pm 0.02$	$9.5 {\pm} 0.58$	3
$120~{ m GeV} < \lambda$	$I_{\rm T} < 160~{ m GeV}$	$75~{ m GeV} < M_{\ell\ell}$	< 105 GeV		1		
50100	$8.7 \pm 0.48$	$0.61 \pm 0.27$	$0.25 \pm 0.15$	$0.16\pm0.021$	$0.4 \pm 0.048$	$10 \pm 0.58$	11
100150	$1.2 \pm 0.11$	$0.02 {\pm} 0.14$	$0.17 \pm 0.11$	0±0	$0.032 \pm 0.0056$	$1.5 {\pm} 0.21$	0
150200	$0.16 \pm 0.031$	$0.17 {\pm} 0.41$	$0.071 \pm 0.054$	0±0	$0.0036 \pm 0.0015$	$0.4 {\pm} 0.42$	1
200	$0.13 \pm 0.028$	$0.029 \pm 0.092$	$0.014 \pm 0.014$	$0\pm0$	$0.0014 \pm 0.00089$	$0.17 \pm 0.097$	1

$E_{\rm T}^{\rm miss}$ (GeV)	WZ	Non-Prompt	Rare SM	$Z\gamma^*$	ZZ	Total bkg	Observed	
$M_{\mathrm{T}} > 160~\mathrm{Ge}$	$M_{\rm T} > 160~{ m GeV}, 75~{ m GeV} < M_{\ell\ell} < 105~{ m GeV}$							
50100	2.3±0.17	$0.28{\pm}0.4$	$0.24 \pm 0.27$	$0.27 \pm 0.034$	$0.19\pm0.023$	$3.3 \pm 0.51$	3	
100150	$1.5\pm0.12$	$0.033 \pm 0.1$	$0.15 \pm 0.17$	$0.074 \pm 0.0095$	$0.064 \pm 0.0094$	$1.8 \pm 0.23$	1	
150200	$0.43 \pm 0.054$	$0.015 \pm 0.048$	$0.12 \pm 0.14$	$0.051 \pm 0.0065$	$0.015 \pm 0.0034$	$0.63 \pm 0.16$	1	
200	$0.19 \pm 0.035$	$0.27 \pm 0.14$	$0.11 \pm 0.12$	0±0	$0.007 \pm 0.0021$	$0.58 \pm 0.19$	1	
$M_{\mathrm{T}} < 120~\mathrm{Ge}$	eV, $M_{\ell\ell} > 105$ Ge	V						
50100	12±0.62	$4.6{\pm}1.4$	$1.2 \pm 0.7$	$0.67 \pm 0.086$	$0.6 \pm 0.07$	19±1.7	22	
100150	$2.3 \pm 0.16$	$1\pm0.44$	$0.53 \pm 0.3$	$0.035 \pm 0.0045$	$0.091 \pm 0.012$	$4 \pm 0.56$	6	
150200	$0.6 \pm 0.067$	$0.099 \pm 0.23$	$0.078 \pm 0.047$	$0.05 \pm 0.0064$	$0.029 \pm 0.0052$	$0.86 {\pm} 0.25$	2	
200	$0.36 \pm 0.05$	$0.017 \pm 0.054$	$0.029 \pm 0.02$	) 0±0	$0.014 \pm 0.0031$	$0.42 {\pm} 0.076$	2	
120 GeV $< \Lambda$	$M_{\rm T} < 160$ GeV, $M_{\rm T}$	$I_{\ell\ell} > 105  \mathrm{GeV}$	$\neg \lceil //$					
50100	$0.57 \pm 0.065$	$0.28 \pm 0.16$	$0.11\pm0.071$	$0.32 \pm 0.041$	$0.039 \pm 0.0064$	$1.3\pm0.19$	0	
100150	$0.09 \pm 0.023$	$0.00033 \pm 0.00038$	$0.055 \pm 0.047$	$0.025 \pm 0.0032$	$0.0032 \pm 0.0014$	$0.17 \pm 0.053$	2	
150200	$0.027 \pm 0.013$	$0.029 \pm 0.092$	$0.054 \pm 0.048$	0±0	$0\pm0$	$0.11 \pm 0.1$	0	
200	$0.0018 \pm 0.0043$	$0\pm0$	$0.077 \pm 0.086$	0±0	$0\pm0$	$0.079 \pm 0.086$	0	
$M_{\mathrm{T}} > 160~\mathrm{Ge}$	eV, $M_{\ell\ell} > 105$ Ge	V						
50100	$0.46 {\pm} 0.058$	$0.17{\pm}0.54$	$0.5 \pm 0.66$	0.13±0.016	$0.025 \pm 0.0047$	$1.3 \pm 0.85$	0	
100150	$0.19 \pm 0.034$	$0.14 {\pm} 0.31$	$0.49 \pm 0.61$	$0.21 \pm 0.026$	$0.011 \pm 0.0028$	$1 \pm 0.69$	1	
150200	$0.083 \pm 0.022$	$0.022 {\pm} 0.068$	$0.14 {\pm} 0.17$	0±0	$0.0034 \pm 0.0014$	$0.25 {\pm} 0.18$	0	
200	$0.042 \pm 0.016$	0±0	$0.11 \pm 0.14$	$0.025 \pm 0.0032$	$0.0036\pm0.0015$	$0.18 \pm 0.14$	0	

Table 3: The summary of the observed yields and predicted backgrounds for tri-lepton without opposite sign same flavor pair present.

E <sub>T</sub> <sup>miss</sup> (GeV)		Non-Prompt		$Z\gamma^*$	ZZ	Total bkg	Observed
$M_{\rm T} < 120~{ m Ge}$	${ m eV}$ , $M_{\ell\ell} < 100~{ m Ge}$	eV					
50100	$1.6 \pm 0.12$	8.5±2.3	$0.44 \pm 0.26$	1±0.13	$0.11 \pm 0.015$	12±2.4	12
100150	$0.27{\pm}0.042$	$1.7 \pm 0.6$	$0.29 \pm 0.18$	$0.18 \pm 0.023$	$0.01 {\pm} 0.0026$	$2.5 \pm 0.63$	2
150200	$0.079 \pm 0.022$	$1.7e-06\pm 2e-06$	$0.062 \pm 0.05$	$0.012 \pm 0.0019$	$0.0023 \pm 0.0012$	$0.16 \pm 0.055$	0
200	$0.031 \pm 0.013$	$0.053 \pm 0.17$	$0.13\pm0.14$	$0.016 \pm 0.002$	$0.00099 \pm 0.00075$	$0.23 \pm 0.22$	0
$120\mathrm{GeV} < \lambda$	$I_{\rm T} < 160 { m GeV}$ , N	$f_{\ell\ell} < 100  \mathrm{GeV}$					
50100	$0.28 \pm 0.041$	$1.3 \pm 0.51$	$0.28 \pm 0.22$	$0.098 \pm 0.013$	$0.029 \pm 0.0053$	$2 \pm 0.56$	1
100150	$0.041 \pm 0.015$	$0.41 \pm 0.63$	$0.051 \pm 0.036$	$0.019 \pm 0.0025$	$0.0042 \pm 0.0016$	$0.52 \pm 0.63$	1
150200	$0.015 \pm 0.0093$	$0.015 \pm 0.047$	$0.047 \pm 0.046$	$0\pm0$	$0\pm0$	$0.077 \pm 0.066$	1
200	$0.003 \pm 0.0039$	0±0	$0.0068\pm0.0062$	0±0	0±0	$0.0098 \pm 0.0073$	0
$M_{\mathrm{T}} > 160~\mathrm{Ge}$	${ m eV}$ , $M_{\ell\ell} < 100~{ m Ge}$	·V					
50100	$0.13 \pm 0.028$	$0.7 \pm 0.32$	$0.058\pm0.068$	$0.085 \pm 0.01$	$0.026 \pm 0.0048$	1±0.33	1
100150	$0.089 \pm 0.023$	$0.22 \pm 0.14$	$0.18\pm0.23$	$0.069 \pm 0.0085$	$0.028 {\pm} 0.0051$	$0.59 \pm 0.27$	0
150200	$0.038 \pm 0.015$	$0.063 \pm 0.17$	$0.022 \pm 0.026$	$0.026\pm0.0031$	$0.0037 \pm 0.0015$	$0.15 \pm 0.17$	0
200	$0.038 \pm 0.015$	0±0	$0.039\pm0.047$	0±0	$0.0043 \pm 0.0016$	$0.081 \pm 0.049$	0
$M_{\mathrm{T}} < 120~\mathrm{Ge}$	eV, $M_{\ell\ell} > 100$ Ge	·V	\ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \				
50100	$0.047 \pm 0.017$	$0.29 \pm 0.14$	$0.14\pm0.086$	$0.14 \pm 0.018$	$0.0021 \pm 0.0011$	$0.61 \pm 0.16$	0
100150	$0.0043 \pm 0.005$	$2.2e-06\pm7.5e-06$	$0.025 \pm 0.019$	$0.026\pm0.0033$	$0\pm0$	$0.055 {\pm} 0.02$	0
150200	$0\pm0$	$0.021 \pm 0.067$	$0.056\pm0.049$	$0.0083 \pm 0.0011$	$0.00011 \pm 0.00025$	$0.086 \pm 0.083$	0
200	0±0	0±0	$0.00021 \pm 0.00017$	\\0±0	0±0	$0.00021 \pm 0.00017$	0
$120  \mathrm{GeV} < \lambda$	$M_{\rm T} < 160 { m GeV}$ , $M_{\rm T}$	$ m M_{\ell\ell} > 100~GeV$					
50100	$0.059 \pm 0.018$	0±0	$0.023 \pm 0.019$	$0.022 \pm 0.0029$	$0.00088 \pm 0.00071$	$0.11 \pm 0.026$	1
100150	$0.0037 \pm 0.0083$	$0.021 {\pm} 0.067$	$0.014 \pm 0.014$	$0.025 \pm 0.0032$	0±0	$0.064 \pm 0.069$	0
150200	$0\pm0$	$0\pm0$	$0.00021 \pm 0.00019$	0±0	<u>0±0</u>	$0.00021 \pm 0.00019$	0
200	0±0	0±0	$0.004 \pm 0.0046$	0±0	0±0	$0.004 \pm 0.0046$	0
$M_{\rm T} > 160~{ m Ge}$	${ m eV}$ , $M_{\ell\ell} > 100~{ m Ge}$	eV					
50100	$0.025 \pm 0.012$	$0.31 \pm 0.75$	$0.13\pm0.16$	$0.016\pm0.0021$	$0.0012 \pm 0.00083$	$0.49 \pm 0.76$	0
100150	$0.03 \pm 0.013$	$0.053 \pm 0.12$	$0.06 \pm 0.076$	$0.012 \pm 0.0016$	$0.0026 \pm 0.0012$	$0.16 {\pm} 0.14$	0
150200	$0.0034 \pm 0.0067$	$0.021 \pm 0.067$	$0.00061 \pm 0.00073$	$0\pm0$	$0.0014 \pm 0.00089$	$0.027 \pm 0.067$	0
200	$0.017 \pm 0.0097$	$0\pm0$	$0.044 {\pm} 0.058$	$0.013 \pm 0.0016$	$0\pm0$	$0.073 \pm 0.059$	0

Table 4: The summary of the observed yields and predicted backgrounds for the channel with a same sign di-lepton and a hadronically decaying tau.

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$E_{\rm T}^{\rm miss}$ (GeV)	WZ	Non-Prompt	Rare SM	$Z\gamma^*$	ZZ	Total bkg	Observed
$M_{\mathrm{T}} < 120~\mathrm{Ge}$	eV, $M_{\ell\ell} < 100$ Ge	V / \					
50100	7.1±3.5	25±4.6	$0.43 \pm 0.22$	0±0	$0.46 \pm 0.23$	$33 \pm 6.1$	25
100150	$0.87 \pm 0.43$	$2.6 \pm 0.68$	$0.41 {\pm} 0.36$	$0\pm0$	$0.04 {\pm} 0.02$	$3.9 \pm 1$	0
150200	$0.4{\pm}0.2$	$0.39 \pm 0.19$	$0.032 \pm 0.021$	$0\pm0$	$0.0063 \pm 0.0036$	$0.82 {\pm} 0.29$	0
200	$0.21 \pm 0.11$	$0.071 \pm 0.056$	$0.024 \pm 0.015$	0±0	$0.0033 \pm 0.002$	$0.31 {\pm} 0.14$	0
120 GeV $< \lambda$	$M_{\rm T} < 160$ GeV, $M_{\rm T}$	$I_{\ell\ell} < 100  \mathrm{GeV}$					
50100	$1.1 \pm 0.55$	$1.9 \pm 0.55$	$0.08 \pm 0.046$	0±0	$0.11 \pm 0.054$	$3.1 \pm 0.84$	3
100150	$0.12 \pm 0.062$	$0.39\pm0.19$	$0.027{\pm}0.018$	$0\pm0$	$0.0092 \pm 0.005$	$0.54 {\pm} 0.21$	1
150200	$0.02 \pm 0.014$	0±0	$0.0095 \pm 0.01$	$0\pm0$	$0.0015 \pm 0.0011$	$0.031 \pm 0.02$	0
200	$0.0054 \pm 0.0058$	$0.022 \pm 0.023$	$0.0035 \pm 0.0034$	$0\pm0$	$0.00048 \pm 0.00049$	$0.032 {\pm} 0.024$	0
$M_{\mathrm{T}} > 160~\mathrm{Ge}$	$eV$ , $M_{\ell\ell} < 100$ Ge	V //					
50100	$0.62 \pm 0.32$	1±0.37	$0.029 \pm 0.019$	0±0	$0.12 \pm 0.06$	$1.8 \pm 0.54$	1
100150	$0.3 \pm 0.15$	$0.25 {\pm} 0.14$	$0.17\pm0.11$	$0\pm0$	$0.05 {\pm} 0.025$	$0.77 \pm 0.31$	0
150200	$0.061 \pm 0.035$	$0.17 \pm 0.15$	$0.056 \pm 0.036$	$0\pm0$	$0.013 \pm 0.0071$	$0.3 \pm 0.17$	0
200	$0.029 \pm 0.019$	$0.16 {\pm} 0.12$	$0.028 \pm 0.019$	$0\pm0$	$0.0065 \pm 0.0036$	$0.22 {\pm} 0.12$	2
$M_{\rm T} < 120~{ m Ge}$	eV, $M_{\ell\ell} > 100$ Ge	V			\		
50100	$0.18\pm0.097$	$1.1 {\pm} 0.41$	$0.035\pm0.023$	0±0	$0.0072 \pm 0.0039$	$1.3 \pm 0.42$	1
100150	$0.025 \pm 0.017$	$0.25 {\pm} 0.11$	$0.02 \pm 0.014$	0±0	$0.0017 \pm 0.0012$	$0.29 \pm 0.11$	0
150200	$0.011 \pm 0.0091$	$0.022 \pm 0.023$	$0.0083 \pm 0.0065$	0±0	$0.0005 \pm 0.00051$	$0.042 {\pm} 0.027$	0
200	$0.0047 \pm 0.0053$	$0.022 \pm 0.023$	$0.00011 \pm 0.00011$	0±0	\\0±0	$0.027{\pm}0.024$	0
$120  \mathrm{GeV} < \Lambda$	$M_{\rm T} < 160$ GeV, $M_{\rm T}$	$I_{\ell\ell} > 100~{ m GeV}$					
50100	$0.035 \pm 0.022$	0±0	$0.02 \pm 0.015$	0±0	$0.0014 \pm 0.001$	$0.056\pm0.033$	0
100150	$0.013 \pm 0.01$	$0\pm0$	$0.00015 \pm 0.00013$	0±0	0±0	$0.013 \pm 0.01$	0
150200	0±0	$0\pm0$	0±0	0±0	$0\pm0$	$0\pm0$	0
200	$0.0065 \pm 0.0064$	0±0	0±0	$0\pm0$	0±0	$0.0065 \pm 0.0064$	0
$M_{\mathrm{T}} > 160~\mathrm{Ge}$	$eV$ , $M_{\ell\ell} > 100$ Ge	V					
50100	$0.034 \pm 0.021$	$0.18 \pm 0.12$	$0.009\pm0.0069$	0±0	$0.0021 \pm 0.0014$	$0.22 \pm 0.13$	1
100150	$0.025 \pm 0.017$	$0.17 \pm 0.15$	$0.0093 \pm 0.0069$	$0\pm0$	$0.0024 \pm 0.0015$	$0.21 {\pm} 0.15$	1
150200	0±0	$0.053 \pm 0.041$	$0.012 \pm 0.012$	$0\pm0$	$0.00048 \pm 0.00049$	$0.065 \pm 0.043$	0
200	0±0	$0\pm0$	$0.011 {\pm} 0.0085$	$0\pm0$	$0.00027 \pm 0.00035$	$0.012 \pm 0.0086$	0

Table 5: The summary of the observed yields and predicted backgrounds for the channel with a same sign opposite flavor di-lepton and a hadronically decaying tau.

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E <sub>T</sub> <sup>miss</sup> (GeV)	WZ	Non-Prompt	Rare SM	$Z\gamma^*$	ZZ	Total bkg	Observed		
$M_{\rm T} < 120~{\rm G}$	$eV$ , $M_{\ell\ell} < 100$ GeV		\						
50100	4.35±1.26	79.7±35.5	11.7±6.37	$0.373 \pm 0.321$	$0.0379\pm0.0432$	96.2±55	124		
100150	$0.66 \pm 0.2$	18±9	1.8±1	$0\pm0$	$0.14 {\pm} 0.12$	$21 \pm 13$	28		
150200	$0.15\pm0.052$	3.7±2	$0.37\pm0.26$	$0\pm0$	$0\pm0$	$4.2 \pm 3$	3		
200	$0.12 \pm 0.043$	0±0	$0.14\pm0.098$	$0\pm0$	$0\pm0$	$0.25 {\pm} 0.16$	1		
$120~{ m GeV} < M_{ m T} < 160~{ m GeV}, M_{\ell\ell} < 100~{ m GeV}$									
50100	$0.59 \pm 0.22$	15±8.1	2±1.3	$0.011 \pm 0.012$	$0.064 \pm 0.074$	17±13	21		
100150	$0.063 \pm 0.031$	$6\pm 3.7$	$0.51 {\pm} 0.33$	$0\pm0$	$0\pm0$	$6.5 {\pm} 5.4$	6		
150200	$0.031 \pm 0.019$	$0.38 \pm 0.44$	$0.035 \pm 0.029$	$0\pm0$	$0\pm0$	$0.45 {\pm} 0.63$	1		
200	$0.013 \pm 0.011$	$0\pm0$	$0.0072 \pm 0.0066$	$0\pm0$	$0\pm0$	$0.021 \pm 0.019$	1		
$M_{\rm T} > 160~{\rm G}$	eV, $M_{\ell\ell} < 100$ GeV	I /							
50100	$0.37 \pm 0.38$	3±3.6	$0.78\pm0.89$	0±0	$0.047 \pm 0.07$	$4.2 {\pm} 6.6$	9		
100150	$0.18 \pm 0.18$	$3.5 \pm 4.1$	$0.85 \pm 0.97$	0±0	$0\pm0$	$4.5 \pm 7.1$	8		
150200	$0.042 {\pm} 0.045$	$0.045 \pm 0.063$	$0.17 \pm 0.21$	$0\pm0$	$0\pm0$	$0.26 {\pm} 0.41$	1		
200	$0.026 \pm 0.03$	$0.29 \pm 0.44$	$0.17{\pm}0.21$	0±0	$0\pm0$	$0.49 {\pm} 0.86$	1		
$M_{\rm T} < 120~{\rm G}$	eV, $M_{\ell\ell} > 100$ GeV	I							
50100	$0.088 \pm 0.035$	8±4.7	1.7±0.96	0±0	0±0	9.8±6.9	12		
100150	$0.026 \pm 0.015$	$1.5 \pm 1.1$	$0.43 {\pm} 0.26$	\\ 0±0	$0\pm0$	$1.9 \pm 1.6$	3		
150200	1.7e-05±1.8e-05	$0.93 \pm 0.81$	$0.22 \pm 0.15$	\\0±0	0±0	$1.1 {\pm} 1.2$	0		
200	0±0	$0\pm0$	$0.095 \pm 0.073$	0±0	0±0	$0.095 {\pm} 0.1$	0		
120 GeV < 1	$M_{ m T} < 160$ GeV, $M_{\ell}$	$_{\ell\ell} > 100~{ m GeV}$							
50100	$0.035 \pm 0.019$	$1.3 \pm 1.1$	$0.85{\pm}0.61$	0±0	0±0	$2.2 {\pm} 1.9$	1		
100150	0±0	$0.52 \pm 0.57$	$0.12 \pm 0.091$	0±0	0±0	$0.64 {\pm} 0.83$	1		
150200	0±0	$0\pm0$	$6.6e-05\pm7.5e-05$	0±0	$0\pm0$	$6.6e-05\pm0.00011$	0		
200	0±0	$0.29 \pm 0.34$	$0.054 {\pm} 0.056$	0±0	$0\pm0$	$0.35 \pm 0.5$	0		
$M_{\rm T} > 160  {\rm G}$	$\overline{\mathrm{eV},M_{\ell\ell}>100\mathrm{GeV}}$	7							
50100	$0.033 \pm 0.036$	$0.64 \pm 0.85$	$0.31 \pm 0.36$	0±0	0±0	$0.98 \pm 1.6$	0		
100150	$0.013 \pm 0.015$	$0.11 {\pm} 0.17$	$0.66 {\pm} 0.87$	$0\pm0$	$0\pm0$	$0.79 \pm 1.4$	1		
150200	$0.008 \pm 0.011$	$0.33 \pm 0.5$	$0.055 {\pm} 0.072$	$0\pm0$	$0\pm0$	$0.4 {\pm} 0.77$	1		
200	$0.012 \pm 0.015$	$0\pm0$	$0.013 \pm 0.016$	$0\pm0$	$0\pm0$	$0.025 {\pm} 0.041$	0		

Table 6: The summary of the observed yields and predicted backgrounds for the channel with an OSSF di-lepton and a hadronically decaying tau.

E <sub>T</sub> miss (GeV)	WZ	Non-Prompt	Rare SM	$Z\gamma^*$	ZZ	Total bkg	Observed
	$V$ , $M_{\ell\ell} < 75 \mathrm{GeV}$	•		·			
50100	3.98±1.17	114±46	7.78±4.3	6.65±3.16	0.497±0.311	133±73.5	166
100150	$0.68\pm0.21$	$7.9 \pm 4.2$	$1.2 \pm 0.67$	$0.35\pm0.39$	$0\pm0$	$10 \pm 6.4$	16
150200	$0.2 \pm 0.068$	$2.1 {\pm} 1.4$	$0.064\pm0.049$	$0\pm0$	$0\pm0$	$2.3 \pm 2$	3
200	$0.088 \pm 0.034$	$0.64 {\pm} 0.51$	$0.028 \pm 0.024$	$0\pm0$	$0\pm0$	$0.75 \pm 0.74$	1
$120  \mathrm{GeV} < M$	$I_T < 160 \text{ GeV}, M_{\ell\ell}$	< 75 GeV					
50100	0.27±0.11	6.4±3.3	$0.38 \pm 0.25$	0.27±0.3	$0.068\pm0.079$	$7.4{\pm}5.4$	11
100150	$0.056\pm0.026$	$4.5{\pm}2.9$	$0.61 \pm 0.47$	$0\pm0$	$0\pm0$	$5.1 \pm 4.3$	2
150200	$0.039\pm0.022$	$0.56 \pm 0.51$	$0.13 \pm 0.1$	$0\pm0$	$0\pm0$	$0.72 \pm 0.76$	2
200	$0.0051\pm0.0054$	0±0	$0.019 \pm 0.021$	0±0	0±0	$0.024{\pm}0.031$	0
$M_{\rm T} > 160  {\rm Ge}$	$V$ , $M_{\ell\ell} < 75 \mathrm{GeV}$						
50100	0.1±0.11	0±0	$0.081 \pm 0.1$	$0.38 \pm 0.55$	0±0	$0.56 \pm 0.98$	1
100150	$0.097\pm0.1$	$1\pm1.1$	$0.67 \pm 0.81$	$0.41 {\pm} 0.6$	0±0	2.2±3.5	3
150200	$0.022 \pm 0.025$	$0.63 \pm 0.84$	$0.13\pm0.16$	$0\pm0$	0±0	$0.79 \pm 1.4$	0
200	$0.014\pm0.017$	0±0	$0.091\pm0.11$	0±0	0±0	0.1±0.18	1
	$V$ , 75 GeV $< M_{\ell\ell}$ $<$						
50100	29.2±8.46	438±130	$7.76 \pm 4.23$	8.17±4.08	$2.78\pm1.55$	486±198	554
100150	6.7±2	$7.9 \pm 3.5$	$0.85 {\pm} 0.48$	$0\pm0$	$0.58 \pm 0.36$	16±7	16
150200	$2.2\pm0.65$	$0.45 {\pm} 0.44$	$0.17\pm0.099$	0±0\	$0.016\pm0.018$	$2.9 \pm 1.2$	4
200	1.2±0.37	0±0	$0.092\pm0.054$	0±0	0±0	$1.3\pm0.54$	2
$120  \mathrm{GeV} < M$	$ m I_T < 160~GeV$ , 75 C						
50100	$0.33\pm0.13$	$7.6 \pm 3.4$	$0.45\pm0.29$	0.37±0.37	0±0	$8.8{\pm}5.6$	12
100150	$0.16\pm0.066$	$1.9 \pm 1.4$	$0.18\pm0.12$	0±0	\\0±0	$2.3 \pm 2$	4
150200	$0.085 \pm 0.038$	$0.34 \pm 0.4$	$0.04 \pm 0.026$	0±0	0±0	$0.47 \pm 0.58$	0
200	$0.03\pm0.017$	0±0	$0.044\pm0.036$	0±0	0±0	$0.074\pm0.061$	0
	$V$ , 75 GeV $< M_{\ell\ell}$ $<$	< 105 GeV			>		
50100	$0.05\pm0.053$	$4.8 \pm 4.6$	$0.087\pm0.1$	$0.00042 \pm 0.00062$	0±0	$5\pm7.6$	1
100150	$0.078\pm0.082$	$1.5\pm1.8$	$0.12\pm0.15$	0±0	$0\pm0$	$1.7 \pm 2.8$	0
150200	$0.039\pm0.043$	0±0	$0.078\pm0.094$	0±0	$0\pm0$	$0.12 \pm 0.18$	0
200	$0.029\pm0.032$	0±0	$0.059\pm0.068$	0±0	0±0	$0.087 \pm 0.13$	0
$M_{\rm T} < 120{ m Ge}$	$V$ , $M_{\ell\ell} > 105~{ m GeV}$						
50100	2.1±0.62	33±13	5.4±3	$1.1\pm0.71$	$0.23\pm0.17$	42±21	82
100150	$0.42\pm0.13$	9.1±5.3	$1.2 \pm 0.81$	$0\pm0$	$0.048 {\pm} 0.055$	$11\pm7.7$	5
150200	$0.1\pm0.038$	1.5±1.1	$0.11 \pm 0.082$	$0\pm0$	$0\pm0$	$1.7 {\pm} 1.6$	2
200	$0.06\pm0.025$	/ 0±0	$0.059 \pm 0.042$	0±0	0±0	$0.12\pm0.073$	1
	$M_{ m T} < 160~{ m GeV}$ , $M_{\ell\ell}$						
50100	0.12±0.051	$0.48 \pm 0.46$	$0.22 \pm 0.15$	0±0	0±0	$0.82{\pm}0.74$	5
100150	$0.047\pm0.023$	$2.4 \pm 1.6$	$0.54 \pm 0.35$	$0\pm0$	$0\pm0$	$2.9 \pm 2.4$	1
150200	$0.0075\pm0.0058$	$0\pm0$	$0.073 \pm 0.061$	$0\pm0$	$0\pm0$	$0.081 {\pm} 0.088$	1
200	$0.0099 \pm 0.01$	0±0	$0.052 \pm 0.06$	0±0	0±0	$0.062 \pm 0.087$	0
	$V$ , $M_{\ell\ell} > 105 \mathrm{GeV}$						
50100	$0.064\pm0.068$	$0.27{\pm}0.4$	$0.029\pm0.043$	0±0	0±0	$0.36 \pm 0.66$	1
100150	$0.038\pm0.041$	$0.12 \pm 0.17$	$0.25 \pm 0.3$	$0\pm0$	$0\pm0$	$0.41 {\pm} 0.65$	1
150200	$0.0089 \pm 0.011$	$0.3 {\pm} 0.46$	$0.042 \pm 0.059$	$0\pm0$	$0\pm0$	$0.35 \pm 0.7$	1
200	$0.038\pm0.042$	$0.31\pm0.47$	$0.087\pm0.11$	0±0	0±0	$0.43\pm0.79$	0

## **7 Final Tables for Four Lepton Analyses**

The summarized results are presented in three tables:

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- Table 7 contains observed yields and background prediction for each search region in the four lepton channels
- The graphical representation of the results is shown in Figures 6, 7, 8, 10, 9, 11, 12, and 13.

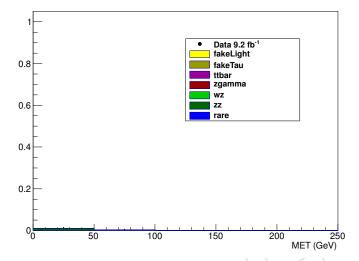


Figure 6: Observed yields and predicted backgrounds for four lepton events with no OSSF pairs and zero taus.

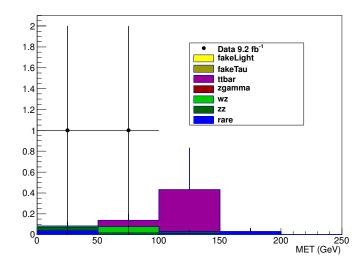


Figure 7: Observed yields and predicted backgrounds for four lepton events with no OSSF pairs and one tau.

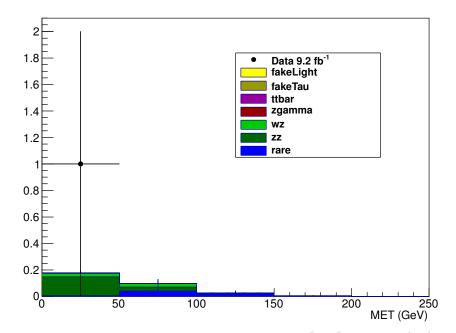


Figure 8: Observed yields and predicted backgrounds for four lepton events with one OSSF pair offZ and zero taus.

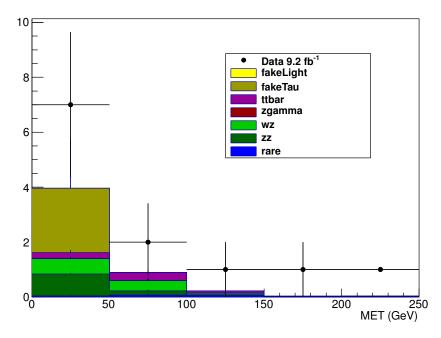


Figure 9: Observed yields and predicted backgrounds for four lepton events with one OSSF pair off Z and one tau.

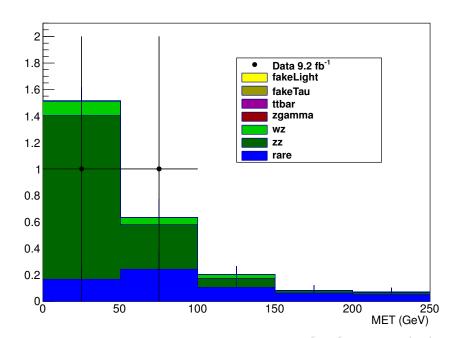


Figure 10: Observed yields and predicted backgrounds for four lepton events with one OSSF pair on Z and zero taus.

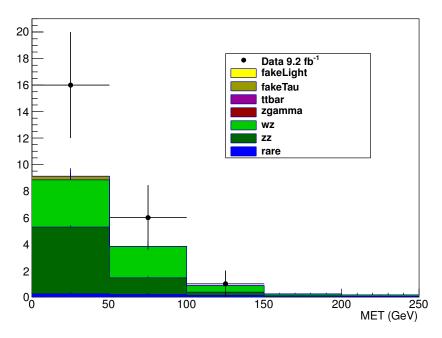


Figure 11: Observed yields and predicted backgrounds for four lepton events with one OSSF pair on Z and one tau.

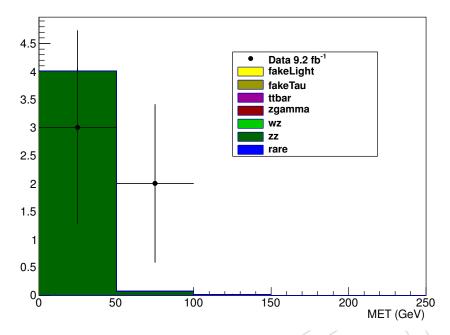


Figure 12: Observed yields and predicted backgrounds for four lepton events with two OSSF pairs offZ and zero taus.

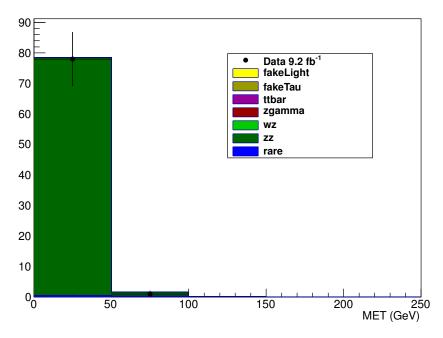


Figure 13: Observed yields and predicted backgrounds for four lepton events with two OSSF pairs on Z and zero taus.

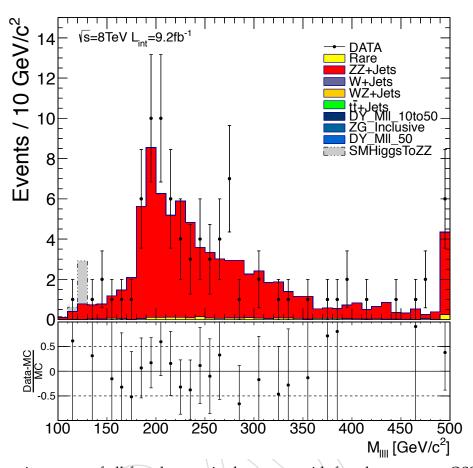


Figure 14: Invariant mass of all four leptons in the events with four leptons, two OSSF pairs, and zero leptons. For comparison, we show the Higgs to ZZ to 4L expected signal.

## 8 Statistical Procedure

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To build a statistical model for this analysis we consider following sources of systematics shared (100% correlated between channels):

- uncertainties in luminosity, jet energy scale, trigger efficiency
- uncertainty of backgrounds: WZ, ZZ,  $Z\gamma$ ,  $t\bar{t}$ , rare processes
- uncertainty of electron reconstruction, identification, selection efficiencies
- uncertainty of muon reconstruction, identification, selection efficiencies
- uncertainty of tau reconstruction, identification, selection efficiencies

Every channel also has two nuisances uncorrelated with other channels. These uncertainties account for statistical fluctuations affecting background estimations, and signal efficiency calculation. As the total number of nuisances is proportional to number of channels used, and therefore necessary computing resources rise exponentially with number of combined channels, we have to keep number of channels actually used for the analysis limited. For the given point in the model parameter space we select a predefined number (currently 10) of the most sensitive channels. The sensitivity of the channel is defined as an expected limit on the total model cross section obtained by using this channel only. This analysis is a typical multi-channel counting experiment. Higgs group combination tool, LandS [6], is technically used to obtain

Table 7: Four lepton backgrounds and observed yields

E <sub>T</sub> <sup>miss</sup> (GeV)	WZ	tŧ	Non-prompt	$Z\gamma^*$	ZZ	Rare SM	Total Bkg	Observed
4L MET dist o	offZ ossf0 tau0							
50–100	$0.00 \pm 0.00$	0						
100-150	$0.00 \pm 0.00$	0						
150-200	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.00 \pm 0.00$	0
200	$0.00\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.00 \pm 0.00$	0
4L MET dist o	offZ ossf0 tau1							
50-100	$0.06 \pm 0.02$	$0.06 \pm 0.06$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.02 \pm 0.01$	$0.00 \pm 0.00$	$0.14 \pm 0.06$	1
100-150	$0.00\pm0.00$	$0.40\pm0.40$	$0.00 \pm 0.00$	$0.00\pm0.00$	$0.02\pm0.01$	$0.02\pm0.01$	$0.43 \pm 0.40$	0
150-200	$0.00 \pm 0.00$	$0.00\pm0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00\pm0.00$	$0.03 \pm 0.03$	$0.03 \pm 0.03$	0
200	$0.00 \pm 0.00$	0						
4L MET dist o	offZ ossf1 tau0							
50-100	$0.03 \pm 0.01$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.03 \pm 0.00$	$0.04 \pm 0.03$	$0.10 \pm 0.03$	0
100-150	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00\pm0.00$	$0.01 \pm 0.00$	$0.02 \pm 0.02$	$0.03 \pm 0.02$	0
150-200	$0.00 \pm 0.00$	0						
200	$0.00 \pm 0.00$	0						
4L MET dist o	offZ ossf1 tau1					$\wedge$		
50-100	$0.37 \pm 0.06$	$0.30 \pm 0.04$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.16 \pm 0.01$	$0.07 \pm 0.04$	$0.90 \pm 0.09$	2
100-150	$0.06 \pm 0.01$	$0.08 \pm 0.02$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.03 \pm 0.00$	$0.05 \pm 0.04$	$0.23 \pm 0.05$	1
150-200	$0.00 \pm 0.00$	$0.01\pm0.01$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.01 \pm 0.00$	$0.00\pm0.00$	$0.03 \pm 0.01$	1
200	$0.01 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.01 \pm 0.00$	$0.01 \pm 0.01$	$0.03 \pm 0.01$	1
4L MET dist o	offZ ossf2 tau0							
50-100	$0.01 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.06 \pm 0.01$	$0.00 \pm 0.00$	$0.07 \pm 0.01$	2
100-150	$0.00 \pm 0.00$	$0.00\pm0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.01 \pm 0.00$	$0.01 \pm 0.01$	$0.02 \pm 0.01$	0
150-200	$0.00 \pm 0.00$	$0.00\pm0.00$	$0.00 \pm 0.00$	0				
200	$0.00 \pm 0.00$	0						
4L MET dist o	onZ ossf1 tau0							
50-100	$0.05 \pm 0.02$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.34 \pm 0.01$	$0.24 \pm 0.14$	$0.63 \pm 0.14$	1
100-150	$0.03 \pm 0.01$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.07 \pm 0.01$	$0.11 \pm 0.06$	$0.20 \pm 0.06$	0
150–200	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.02 \pm 0.00$	$0.06 \pm 0.04$	$0.08 \pm 0.04$	0
200	$0.01 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.01 \pm 0.00$	$0.05 \pm 0.03$	$0.07 \pm 0.03$	0
4L MET dist o	onZ ossf1 tau1							
50-100	$2.35 \pm 0.34$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$1.27 \pm 0.03$	$0.19 \pm 0.11$	$3.81 \pm 0.36$	6
100-150	$0.48 \pm 0.07$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00\pm0.00$	$0.24 \pm 0.01$	$0.13 \pm 0.08$	$0.85 \pm 0.11$	1
150-200	$0.12 \pm 0.03$	$0.00 \pm 0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.06\pm0.01$	$0.04\pm0.03$	$0.22 \pm 0.04$	0
200	$0.08 \pm 0.02$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.03 \pm 0.00$	$0.04\pm0.03$	$0.15 \pm 0.03$	0
4L MET dist c	onZ ossf2 tau0							
50-100	$0.03 \pm 0.01$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$1.30 \pm 0.03$	$0.32 \pm 0.18$	$1.65 \pm 0.18$	1
100-150	$0.01\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.08\pm0.01$	$0.09\pm0.05$	$0.17 \pm 0.05$	0
150-200	$0.00\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.00\pm0.00$	$0.03\pm0.00$	$0.06\pm0.04$	$0.09 \pm 0.04$	0
200	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.00 \pm 0.00$	$0.02 \pm 0.00$	$0.06 \pm 0.04$	$0.08 \pm 0.04$	0

limits. We calculate "LHC style" CLs limitwhich effectively means using frequentist CLs with one-sided profiled likelihood test statistics.

## 8.1 Limits on SMS from the search with three leptons and from same-sign dilepton searches

Figure 15 displays, for  $x_{\tilde{\ell}}=0.5$ , the results of the three-lepton search using  $E_{\rm T}^{\rm miss}$ ,  $M_{\rm T}$  and  $M_{\ell\ell}$ . The figure depicts 95% CL upper limit on the cross section times branching fraction in the  $m_{\tilde{\chi}_1^0}$  versus  $m_{\tilde{\chi}_2^0} \, (=m_{\tilde{\chi}_1^\pm})$  plane in the flavor-democratic scenario. The contour bounds the excluded region in the plane assuming the NLO cross section calculation and a 50% branching fraction, as appropriate for the visible decay products in this scenario. Figure 16 shows the results for two values  $x_{\tilde{\ell}}$  (0.05 and 0.95) obtained after combination of three-lepton and the SS di-lepton searches [7].

Figure 17 presents the corresponding limits for the  $\tau$ -enriched scenario and Figure 18 for the  $\tau$ -dominated scenario. As the SS di-lepton search does not have sensitivity for  $x_{\tilde{\ell}}=0.50$ , there is no limit curve for this search for Figures 15, 17b and 18b. In the other limit curves in both Figures 16, 17 and 18, the increase in the combined mass limit from incorporation of the SS di-lepton search ranges up to approximately 20 GeV.

# 8.2 Limits on SMS with on-shell W and Z from WZ + $E_{\rm T}^{\rm miss}$ and three-lepton analyses

For limits on the SMS with on-shell W and Z bosons, we show the results of the WZ +  $E_T^{miss}$  analysis [8], the three-lepton analysis as well as the combined result.

In the combination, the common signal-related systematic uncertainties for luminosity, jet energy scale, lepton identification, trigger efficiency, and misidentification of light-flavor jets as b jets are considered to be 100% correlated. For backgrounds, the only common systematic uncertainty is that for the WZ simulation, which is treated as 100% correlated. No events in the data pass both signal selections. For the backgrounds, the overlap in the control sample is less than 1%. Thus the two selections are treated as independent.

Figure 19 displays the observed limits for the two individual analyses and the combination.

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This analysis note presents a summary for searches for supersymmetric charginos and neutralinos in three and four leptons final states which were developed and presented in several analysis notes. The searches featured here explore the final states with exactly three or exactly four leptons using transverse mass, missing transverse energy and lepton-pair invariant mass. No excesses above the standard model expectations are observed. The results are used to exclude a range of chargino and neutralino masses for the models considered in the dedicated analyses where these particles have large branching fractions to leptons and vector bosons.

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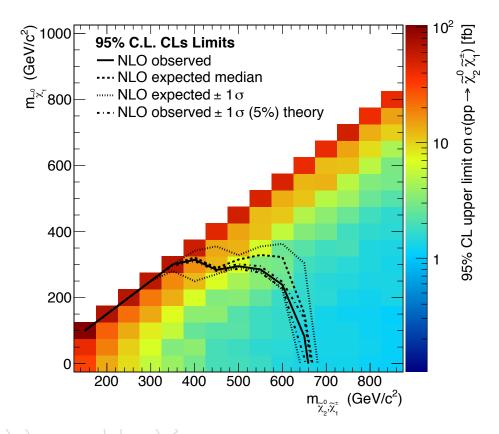


Figure 15: The shading in the  $m_{\widetilde{\chi}_1^0}$  versus  $m_{\widetilde{\chi}_2^0}$  (=  $m_{\widetilde{\chi}_1^\pm}$ ) plane indicates the 95% CL upper limit on the chargino-neutralino production NLO cross section times branching fraction in the flavor-democratic scenario, for the three-lepton search. The contours bound the mass regions excluded at 95% CL for a branching fraction of 50%, as appropriate for the visible decay products in this scenario. The contours based on the observations are shown; in addition, the expected bound is shown.

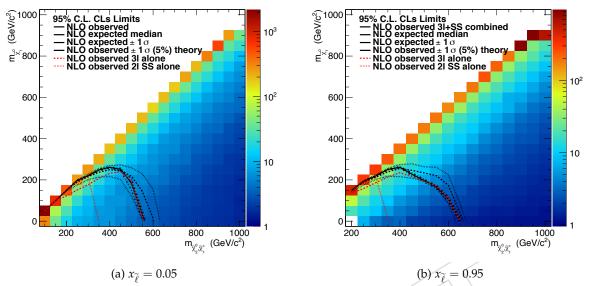


Figure 16: The shading in the  $m_{\tilde{\chi}_1^0}$  versus  $m_{\tilde{\chi}_2^0}$  (=  $m_{\tilde{\chi}_1^\pm}$ ) plane indicates the 95% CL upper limit on the chargino-neutralino production NLO cross section times branching fraction in the flavor-democratic scenario, for the combined analysis of the three-lepton search and the same-sign dilepton search. The contours bound the mass regions excluded at 95% CL for a branching fraction of 50%, as appropriate for the visible decay products in this scenario. The contours based on the observations are shown for the combination; in addition, the expected combined bound is shown. Red contours demonstrate separate mass exclusions for the three-lepton search and the same-sign di-lepton search alone.

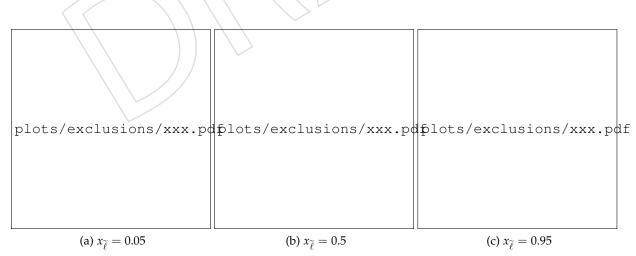


Figure 17: The exclusion contours for  $\tau$ -enriched scenario corresponding to results in Figure 15: (a) and (c) combination of 3-lepton searches with SS d-lepton analysis, (b) 3 lepton searches.

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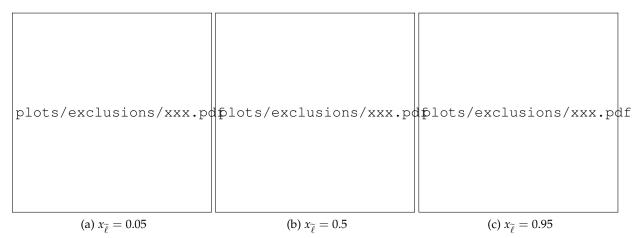


Figure 18: The exclusion contours for  $\tau$ -dominated scenario corresponding to results in Figure 15: (a) and (c) combination of 3-lepton searches with SS d-lepton analysis, (b) 3 lepton searches.

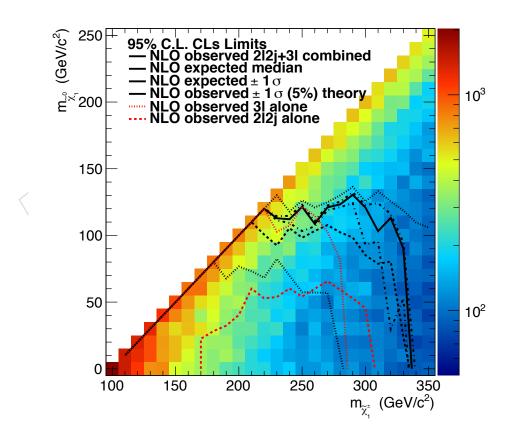


Figure 19: Interpretation of the WZ +  $E_{\rm T}^{\rm miss}$  and three-lepton results in the context of the WZ SMS. The WZ+  $E_{\rm T}^{\rm miss}$  observed, three-lepton observed, combined observed, and combined expected contours are indicated.

## A Graphical Comparison of the Background Estimations

The efforts described in works [2–4] are summarised for a review in the plots presented in this section. If not stated explicitly, the black color in the plots corresponds to a Ref. [2], the red color to a Ref. [3], and the blue color to a Ref. [4]. While for the tri-lepton processes without a tau three results are compared, for channels with a tau present contribution were made by two out of three works which are presented in the plots.

## References

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- [1] CMS Collaboration, "Search for direct EWK production of SUSY particles in multilepton modes with 8TeV data",.
- [2] M. Chen et al., "Search for Direct Electroweak Production of Charginos and Neutralinos with the Tri-lepton Plus Missing Energy Final State",.
- 115 [3] W. de Boer et al., "Search for SUSY using signatures with at least three leptons and missing transverse energy and/or high jet activity at  $\sqrt{s}$ =8 TeV",.
- III [4] S. Arora et al., "A Search for Direct Chargino Neutralino production with three or more leptons using 9.5 fb<sup>-1</sup> of  $\sqrt{s}$ =8 TeV CMS data",.
- [5] S. Arora et al., "Background and Efficiency Determination Methods for Multilepton
   Analyses",.
- [6] M. Chen., "LandS: A statistical tool for calculating limits and significance".

  https://mschen.web.cern.ch/mschen/LandS/index.html, April, 2011.
- 123 [7] M. Dünser, S. Folgueras, F. Moortgat et al., "Search for Direct Production of Charginos and Neutralinos in the Same-Sign Di-Lepton Channel Using 8TeV pp Collisions",.
- 125 [8] D. Barge, C. Campagnari, D. Kovalskyi et al., "Search for beyond-the-standard model physics in events with a Z boson, jets, and missing transverse energy",.

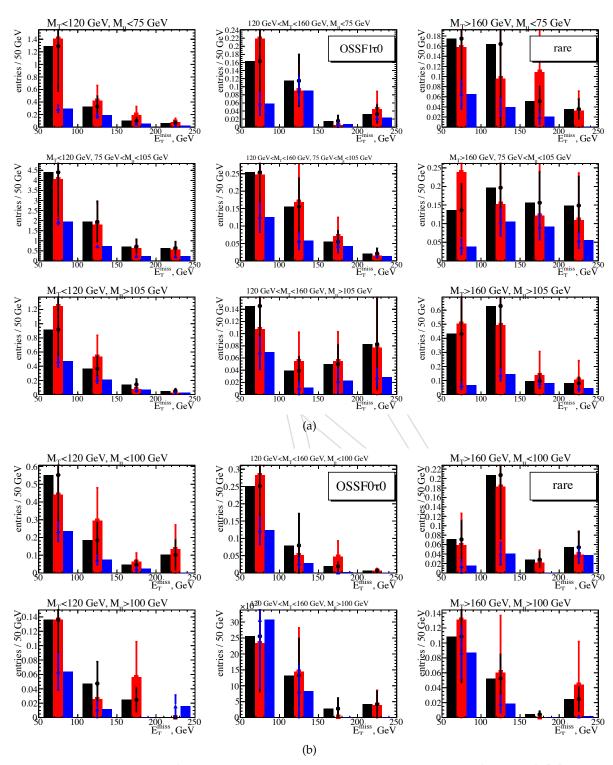


Figure 20: Comparison of the Rare SM contribution. Blue histograms from Ref. [4] do not fully take into account all the processes therefore their contribution is excluded from the final combination.

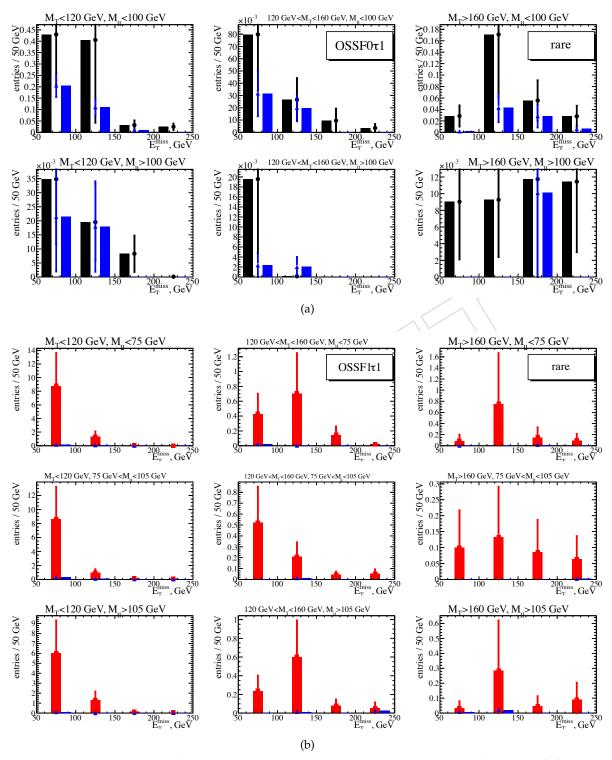


Figure 21: Comparison of the Rare SM contribution. Blue histograms from Ref. [4] do not fully take into account all the processes therefore their contribution is excluded from the final combination.

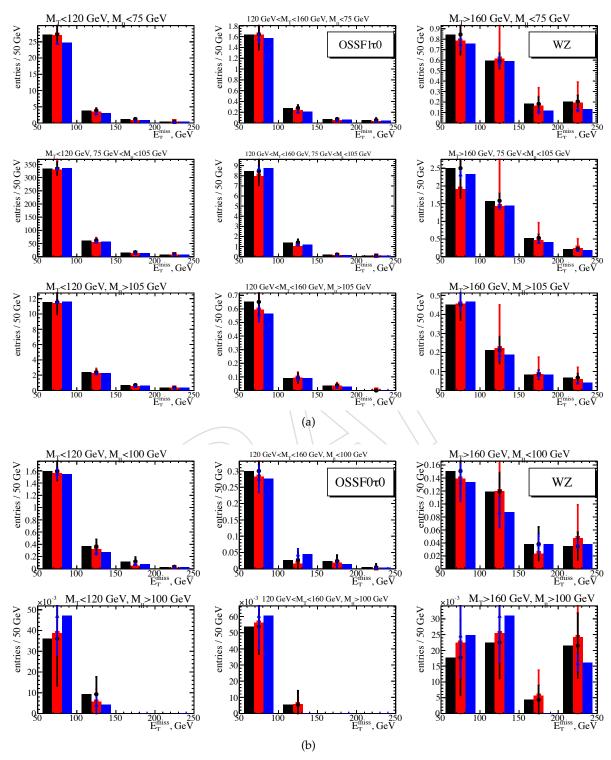


Figure 22: Comparison of the WZ backgrounds.

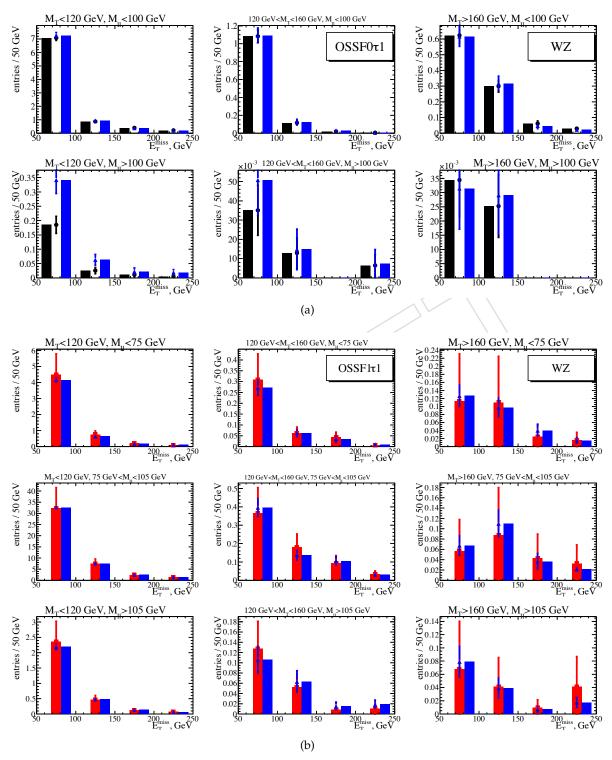
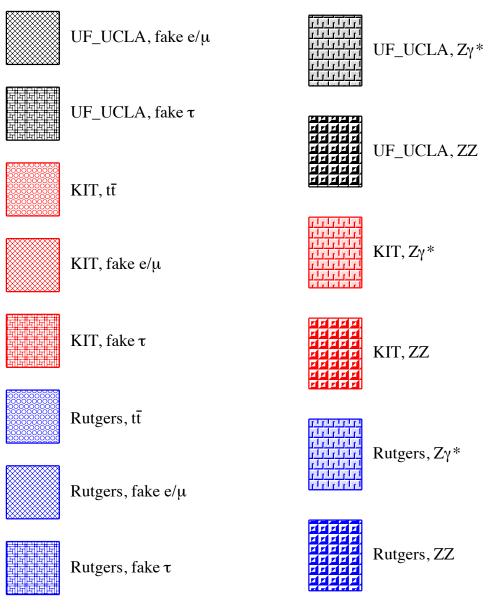


Figure 23: Comparison of the WZ backgrounds.



(a) Non-prompt leptons legend.

(b) ZZ and  $Z\gamma^*$  legend.

Figure 24: Legends for the following plots.

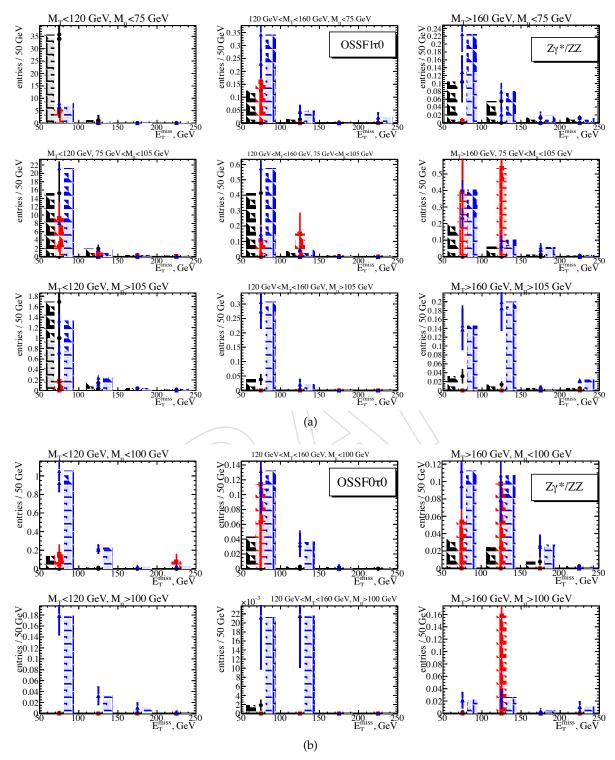


Figure 25: Comparison of the ZZ and  $Z\gamma^*$  contribution.

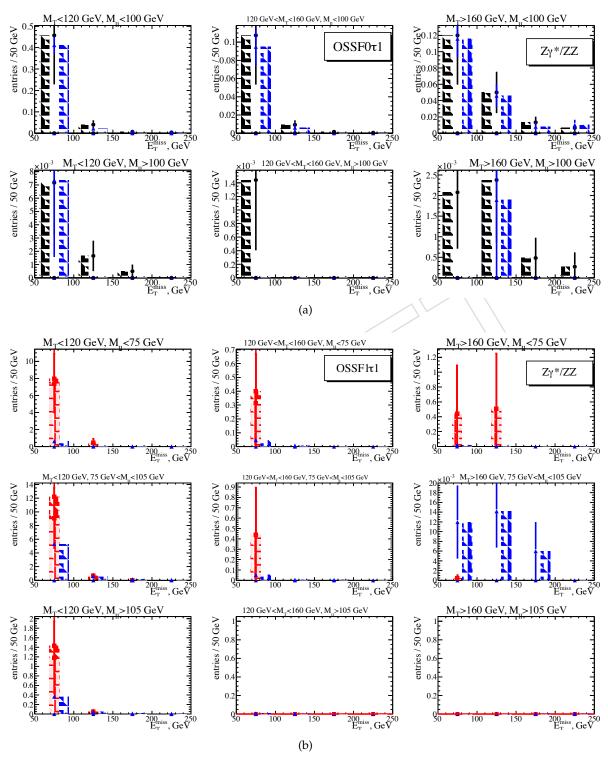


Figure 26: Comparison of the ZZ and  $Z\gamma^*$  contribution.

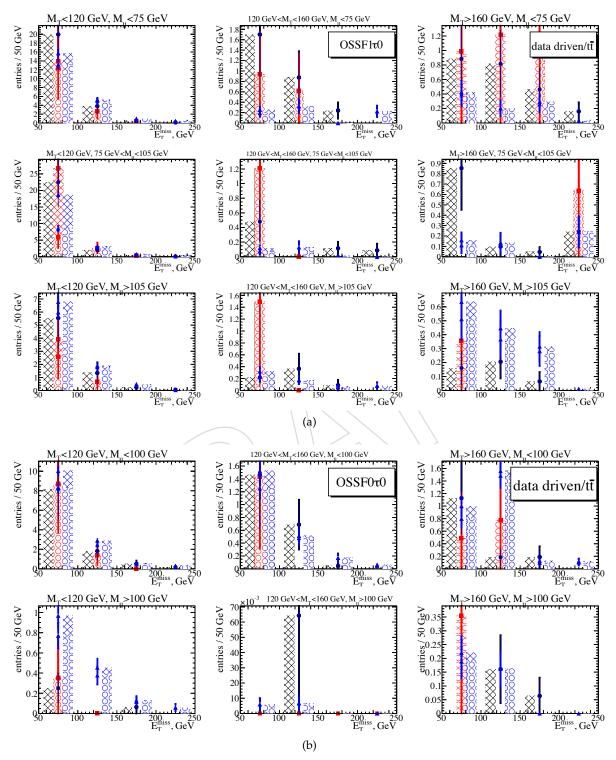


Figure 27: Comparison of the non-prompt leptons contribution.

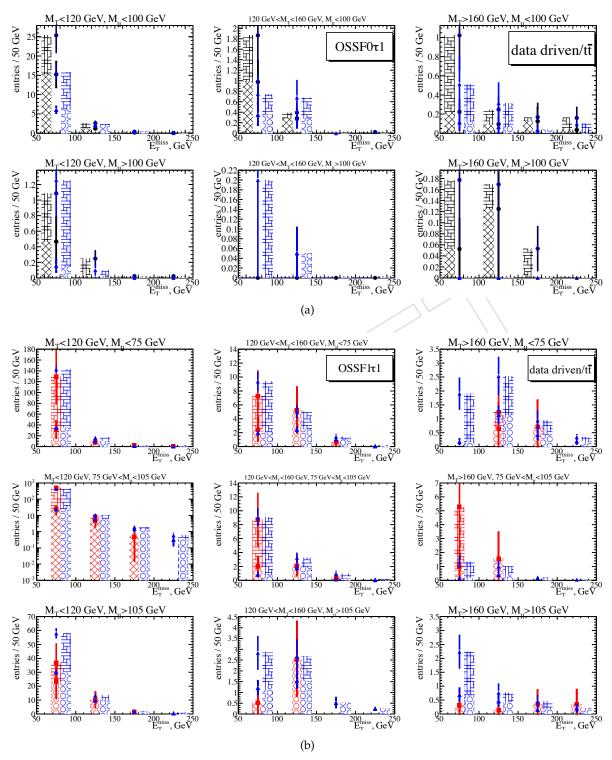


Figure 28: Comparison of the non-prompt leptons contribution.