

Hawking Radiation as Seen by Observers

Bachelor thesis

Friedrich Hübner
Universität Bonn

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1 QFT in curved spacetime

- Scalar field
- Unruh detector
- Minkowski space

2 Static spacetime

- Properties of the Wightman function
- Observers
 - Static observer
 - General observer
 - Thermal

Massless scalar field in curved spacetime

- Spacetime metric: $g_{\mu\nu}$
- Lagrangian: $\mathcal{L} = -\frac{1}{2}\sqrt{|g|}g^{\mu\nu}\partial_\mu\phi\partial_\nu\phi$
- Klein-Gordon equation: $\nabla_\mu\nabla^\mu\phi = \frac{1}{\sqrt{|g|}}\partial_\mu\left(\sqrt{|g|}g^{\mu\nu}\partial_\nu\phi\right) = 0$
- Scalar product: $(\phi|\psi) := i\int_\Sigma dS^\mu\phi^*\nabla_\mu\psi - \psi\nabla_\mu\phi^*$
- Orthonormal basis: $(u_i|u_j) = \delta_{ij}$, $(u_i|u_j^*) = 0$, $(u_i^*|u_j^*) = -\delta_{ij}$
- Quantisation: $\phi(\mathbf{x}) = \sum_i u_i a_i + u_i^* a_i^\dagger$

State of the QFT

- Vacuum: $a_i |0\rangle = 0$
- Problem: $u_i \rightarrow v_j, a_i \rightarrow b_j: b_i |0\rangle \neq 0$
→ Need to guess state!
- Static spacetime: choose vacuum w.r.t. positive frequency modes:
 $u_i \sim e^{-i\omega t}$
- What does an observer see?

Greens functions

- Vacuum:

- Wightman function $D^+(\mathbf{x}, \mathbf{x}') := \langle 0 | \phi(\mathbf{x}) \phi(\mathbf{x}') | 0 \rangle$
- $iD(\mathbf{x}, \mathbf{x}') := [\phi(\mathbf{x}), \phi(\mathbf{x}')] = 2i \operatorname{Im} D^+(\mathbf{x}, \mathbf{x}')$
- $D^{(1)}(\mathbf{x}, \mathbf{x}') := \langle 0 | \{ \phi(\mathbf{x}), \phi(\mathbf{x}') \} | 0 \rangle = 2 \operatorname{Re} D^+(\mathbf{x}, \mathbf{x}')$

- Thermal:

- replace $\langle 0 | \dots | 0 \rangle$ by $\frac{1}{Z} \operatorname{Tr} e^{-\beta H} \dots$
- D is c-number: $D_\beta = D$
- $D_\beta^{(1)}(t, \vec{x}; t', \vec{x}') = \sum_n D^{(1)}(t - i\beta n, \vec{x}; t', \vec{x}')$

Unruh detector

- Detector model: $c \cdot m(\tau) \phi(\mathbf{x}(\tau))$, $c \ll 1$
- Transition amplitude: $Q_{|0,0\rangle \rightarrow |E,\psi\rangle}(\tau) \sim \int_{-\infty}^{\tau} e^{iE\tau'} \langle \psi | \phi(\mathbf{x}(\tau')) | 0 \rangle d\tau'$
- Transition rate: $\frac{dF_E}{d\tau} = 2\text{Re} \int_{-\infty}^0 d\tau' e^{-iE\tau'} D^+(\mathbf{x}(\tau + \tau'), \mathbf{x}(\tau))$
- For constant rate: $\frac{dF_E}{d\tau} = \int_{-\infty}^{\infty} d\tau e^{-iE\tau} D^+(\mathbf{x}(\tau), \mathbf{x}(0))$
- Interpretation: F_E is particle population for observer

Minkowski space

- Wightman function: $D^+(\mathbf{x}, \mathbf{x}') = -\frac{1}{4\pi^2} \frac{1}{(t-t'-i\varepsilon)^2 - |\vec{x}-\vec{x}'|^2}$
 - Static observer: $t(\tau) = \tau, \vec{x}(\tau) = \text{const}$
 - $D^+(\mathbf{x}(\tau), \mathbf{x}(0)) = -\frac{1}{4\pi^2} \frac{1}{(\tau-i\varepsilon)^2}$
 - Fourier transform: $\frac{dF_E}{d\tau} = 0$
- Inertial observer: vacuum contains no particles

image

Unruh effect

- Accelerating observer: $t(\tau) = 1/\alpha \sinh \alpha\tau$, $x(\tau) = 1/\alpha \cosh \alpha\tau$
 - $D^+(\mathbf{x}(\tau), \mathbf{x}(\tau')) = -\frac{\alpha^2}{16\pi^2} \frac{1}{\sinh^2 \frac{\alpha(\tau-\tau')}{2}}$
- Thermal state: $D_\beta^+(\mathbf{x}, \mathbf{x}') = -\frac{1}{4\beta^2} \frac{1}{\sinh^2 \left(\frac{\pi}{\beta} \sqrt{(t-t'-i\epsilon)^2 - |\vec{x}-\vec{x}'|^2} \right)}$
 - Static observer: $D_\beta^+(\mathbf{x}(\tau), \mathbf{x}(\tau')) = -\frac{1}{4\beta^2} \frac{1}{\sinh^2 \left(\frac{\pi}{\beta} (\tau-\tau') \right)}$
- Set $\beta = 2\pi/\alpha$
- Accelerating observer: vacuum is a thermal state

Static spacetimes

- Metric: $ds^2 = -\beta(\vec{x}) dt^2 + g_{ij}(\vec{x}) dx^i dx^j$
- Positive frequency solutions: $u_i(t, \vec{x}) = \frac{1}{\sqrt{2\omega_i}} e^{-i\omega_i t} A_i(\vec{x})$
- State of QFT – vacuum: $a_i |0\rangle = 0$
- Normal coordinates ($a = \beta(0)$):

$$D^+(\mathbf{x}, 0) = -\frac{1}{4\pi^2} \frac{1}{a(t - i\epsilon)^2 - |\vec{x}|^2} + \mathcal{O}(x^2)$$

Pole at $\mathbf{x} = \mathbf{x}'$

- Consider trajectory $\mathbf{x}(\tau)$, $\mathbf{x}(0) = 0$
- D^+ has second order pole at $\tau = 0$: $\frac{1}{a(t(\tau) - i\varepsilon)^2 - |\vec{x}(\tau)|^2}$
- ε shift pole to upper half:
 - $a(t(\tau_\varepsilon) - i\varepsilon)^2 - |\vec{x}(\tau_\varepsilon)|^2 = 0$
 - $\delta\tau = \left. \frac{d\tau_\varepsilon}{d\varepsilon} \right|_{\varepsilon=0} = iat(0) \pm \sqrt{-a^2\dot{t}(0)^2 + 1}$
 $\rightarrow \delta\tau$ has positive imaginary part
- Only poles in the lower half contribute
 \rightarrow drop pole at $\tau = 0$

Other singularities of D^+

- $D^+(\mathbf{x}, \mathbf{x}') = \langle 0 | \phi(\mathbf{x}) \phi(\mathbf{x}') | 0 \rangle$ satisfies $\nabla_\mu \nabla^\mu D^+(\mathbf{x}, \mathbf{x}') = 0$
- Define $A = 1/D^+ \Rightarrow A \nabla_\mu \nabla^\mu A = 2 \nabla_\mu A \nabla^\mu A$
- $D^+ = \infty \Rightarrow A = 0 \Rightarrow \nabla_\mu A \nabla^\mu A = 0$
 $\rightarrow D^+ = \infty$ is a lightlike hypersurface
- D^+ is singular on the lightcone
- Assume no more singularities

image

- \rightarrow no singularities on timelike trajectories
- $D = [\phi(\mathbf{x}), \phi(\mathbf{x}')] is only non zero on lightcone$
 $\rightarrow D^+$ is real for all trajectories

Static observers

Lemma:

box

In a static spacetime a static observer does not observe any particles

- $t(\tau) = \frac{\tau}{\sqrt{a}}$
- $u_i = \frac{1}{\sqrt{2\omega_i}} e^{-i\omega_i t} A_i(\vec{x})$
- $D^+(\mathbf{x}(\tau), \mathbf{x}(0)) = \sum_i \frac{1}{2\omega_i} e^{-i\omega_i \tau / \sqrt{a}} A_i(\vec{x}_0) A_i^*(\vec{x}_0)$

$$\begin{aligned} \frac{dF_E}{d\tau} &= \sum_i \frac{1}{2\omega_i} A_i(\vec{x}_0) A_i^*(\vec{x}_0) \left(\int_{-\infty}^{\infty} d\tau e^{-iE\tau} e^{-i\omega\tau/\sqrt{a}} \right) \\ &= \sum_i \frac{1}{2\omega_i} A_i(\vec{x}_0) A_i^*(\vec{x}_0) \delta(E + \omega/\sqrt{a}) = 0 \end{aligned}$$

General observers

- Transition rate not constant:

$$\frac{dF_E}{d\tau} = 2\text{Re} \int_{-\infty}^0 d\tau' e^{-iE\tau'} D^+(\mathbf{x}(\tau + \tau'), \mathbf{x}(\tau))$$

→ can not use residuum theorem

- How to handle pole at $\tau' = 0$?
- Expansion: $D^+(\mathbf{x}(\tau'), 0) = -\frac{1}{4\pi^2\tau'^2} + W(\tau')$
- $\frac{1}{\tau'^2}$ term does not contribute
- $W(\tau')$ is non singular

Thermal case

- Thermal Wightman function:

$$\begin{aligned}
 D_{\beta}^{+}(t(\tau'), \vec{x}(\tau'); 0) &= \sum_{n=-\infty}^{\infty} D^{+}(t(\tau') - i\beta n, \vec{x}(\tau'); 0) \\
 &= \sum_{n=-\infty}^{\infty} -\frac{1}{4\pi^2(\tau' - i\beta\sqrt{a}n)^2} + \sum_{n=-\infty}^{\infty} W(\tau'(t - i\beta n)) \\
 &= -\frac{1}{4\beta^2 a} \frac{1}{\sinh^2\left(\frac{\pi}{\beta\sqrt{a}}\tau'\right)} + W_{\beta}(\tau')
 \end{aligned}$$

- Static observers: $T = \frac{T_0}{\sqrt{a}}$ (Tolman effect)
- General: Corrections from $W_{\beta}(\tau')$
 - shift temperature?
 - non thermal?

Quellen

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