A practical walk through formal scattering theory

Connecting bound states, resonances, and scattering states in exotic nuclei and beyond

The Lippmann-Schwinger equation

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Scattering setup

- we consider two particles with masses m_1 and m_2 at positions ${\bf r}_1$ and ${\bf r}_2$
- we assume that the interaction does not depend on absolute particle positions
- then we can **neglect the overall center-of-mass motion** and work only with the relative coordinate ${\bf r}={\bf r}_1-{\bf r}_2$ and reduced mass $\mu=m_1m_2/(m_1+m_2)$
- for the two particles scattering off one another, we physically expect that the wavefunction describing their relative motion is given as a sum of an incoming plane wave and an outgoing spherical scattered wave:

$$\psi_{\mathbf{k}}^{(+)}(\mathbf{r}) = e^{i\mathbf{k}\cdot\mathbf{r}} + f_k(\theta) \frac{e^{ikr}}{r} \tag{1}$$

ullet all physics information is contained in the scattering amplitude $f_k(heta)$

Free Hamiltonian

- ullet the **free Hamiltonian** H_0 is an important reference operator
 - ullet differential operator $(\sim {
 m d}^2/{
 m d}r^2)$ in configuration space
 - ► diagonal multiplicative operator in momentum space
- it has a purely continuous spectrum of plane-wave states:

$$H_0 |\mathbf{k}
angle = rac{\mathbf{k}^2}{2\mu} |\mathbf{k}
angle \hspace{1cm} (2)$$

lacktriangle this is for a two-body system with relative momentum ${f k}$ and reduced mass μ

Note

- ullet the plane-wave states ${f k}$ are not elements of the physical Hilbert space
 - $\langle {f r}|{f k}
 angle={
 m e}^{{
 m i}{f k}\cdot{f r}}$ is not a normalizable wavefunction $(
 ot\in L^2(\mathbb{R}^3))$
- ullet strictly one should work with wave packets $|\phi
 angle\sim\int\mathrm{d}^3p\,g(\mathbf{p})|\mathbf{p}
 angle$
- nevertheless it is convenient and permissible to work with plane-wave states because every element of the Hilbert space can be expanded in them (Fourier transform!)

Interaction potential

- ullet we will consider generic potentials V written as operators
- ullet we assume these potentials to be time independent, but they may in principle depend on energy, V=V(E)
- taking matrix elements gives a concrete representation, for example:

$$V(\mathbf{r}',\mathbf{r}) = \langle \mathbf{r}' | V | \mathbf{r}
angle$$

• the most intuitive (and familiar) potentials are local and spherically symmetric in coordinate space: $V({\bf r}',{\bf r})=V(r)\delta^{(3)}({\bf r}-{\bf r}')$, with $r=|{\bf r}|$

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- the most intuitive (and familiar) potentials are local and spherically symmetric in coordinate space: $V({\bf r}',{\bf r})=V(r)\delta^{(3)}({\bf r}-{\bf r}')$, with $r=|{\bf r}|$
- in the following, we will work in **momentum space**
- then $V(\mathbf{p}',\mathbf{p}) = \langle \mathbf{p}'|V|\mathbf{p}\rangle$ is related to $V(\mathbf{r}',\mathbf{r})$ by Fourier transformation
- ullet we typically project onto fixed partial waves, denoted by l
- for local potentials, this leads to the following expression:

$$V_l(p',p) = 4\pi \int_0^\infty \mathrm{d}r \, r^2 j_l(p'r) V(r) j_l(pr) \,, \qquad \qquad (3)$$

where $j_l(z)$ it the *l*-th spherical Bessel function

Partial-wave projected potentials

- ullet numerically, it is better to use Riccati-Bessel functions $\hat{j}_l(z)=zj_l(z)$
- ullet this explicitly reflects the cancellation of the singularity at r=0

$$V_l(p',p) = rac{4\pi}{p'p} \int_0^\infty \mathrm{d}r\, \hat{j}_l(p'r) V(r) \hat{j}_l(pr)$$
 (4)

Python implementation

```
# lib/potential.py
from scipy.integrate import quad

class LocalPotential(Potential):
    # [...]

# Specific subclasses implement __call__(self, r)

def get(self, ell, p, q):
    return 4.0 * np.pi / (q * p) * quad( \
        lambda r: riccati_j(ell, p * r) * self(r) * riccati_j(ell, q * r), \
        0.0, np.inf \
        )[0]
```

Plane-wave boundary condition

- ullet consider the stationary Schrödinger equation: $H|\psi
 angle=E|\psi
 angle$ with $H=H_0+V$
- ullet this equation alone does not specify a boundary condtion for solutions $\psi({f r})=\langle {f r}|\psi
 angle$
- ullet a scattering state should be such that for V o 0 , $|\psi
 angle o |{f k}
 angle$
- moreover, the state should be one that evolved from a free state in the infinite past

Interlude: time dependence

- ullet the full Schrödinger equation is $\mathrm{i}rac{\partial}{\partial t}|\psi(t)
 angle=H|\psi(t)
 angle=(H_0+V)|\psi(t)
 angle$
- Green's functions solve the time dependence:

$$G_0(t) = G_0^{(+)}(t) = \begin{cases} -\mathrm{i}\mathrm{e}^{-\mathrm{i}H_0t} & t > 0 \,, \\ 0 & t \le 0 \end{cases}$$
 (5)

- ▶ this is the free retarded Green's function
- ▶ advanced Green's functions vanish for $t \ge 0$
- full Green's functions are defined analogously with $H_0 \to H$ in Eq. (5)
- ullet sometimes these operators are denoted by $U_0(t)$ and U(t)
- ullet time dependence of waves is conventionally written $\sim {
 m e}^{-{
 m i}Et}$ $\hookrightarrow {
 m e}^{\pm {
 m i}kr}$ represent in- and outgoing spherical waves, respectively
- $G_0(z)$ is the Fourier tranform $G_0(E)$ of $G_0(t)$, analytically continued into the whole complex energy plane

Interlude: adiabatic switching

 define a free in state that was equal to the exact interacting state in the infinite past:

• this allows us to write down a formal solution using the full Green's function:

$$|\psi(t)
angle = |\psi_{
m in}(t)
angle + \lim_{arepsilon o 0^+} \int {
m d}t \, {
m e}^{-arepsilon t'} G^{(+)}(t-t') V |\psi_{
m in}(t)
angle \qquad (7)$$

- if the states monochromatic (fixed definite energy), the damping factor $e^{-\varepsilon t'}$ with $\varepsilon \to 0^+$ is required in Eq. (7)
- approximately this can be interpreted as multiplying the potential V with $e^{-\varepsilon t'}$, adibatically switching it off in the infinite past
- when wave packets are used to represent proper physical scattering states, such damping factors are not necessary; their presence otherwise permits the convenience of working with simple plane-wave states

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- moreover, the state should be one that evolved from a free state in the infinite past
- both conditions can be enforced with the ansatz

$$|\psi_{\mathbf{k}}^{(+)}\rangle = |\mathbf{k}\rangle + (E_{\mathbf{k}} - H_0 + \mathrm{i}\varepsilon)^{-1}V|\psi_{\mathbf{k}}^{(+)}\rangle$$
 (8)

with $E_{\mathbf{k}}=p^2/(2\mu)$

ullet this is the **Lippmann-Schwinger equation** for the scattering state $|\psi_{f k}^{(+)}
angle$

Notes

- ullet free Green's function $G_0(z)=(z-H_0)^{-1}$ appears in Eq. (8) with $z=E_{f k}+{
 m i}arepsilon$
- ullet arepsilon
 ightarrow 0 is implied in all equations, this implements the adiabatic switching
- linear equation becomes an integral equation when projected onto a concrete representation

The Lippmann-Schwinger equation

ullet consider further Eq. (8) and apply V from the left:

$$V|\psi_{\mathbf{k}}^{(+)}\rangle = V|\mathbf{k}\rangle + VG_0(E_{\mathbf{k}} + \mathrm{i}\varepsilon)V|\psi_{\mathbf{k}}^{(+)}\rangle$$
 (9)

- ullet define an operator T via $V|\psi_{f k}^{(+)}
 angle=T|{f k}
 angle$
 - ► for now, this merely assumes the existence of such an operator
 - ▶ it will be justified later
- using the definition of T, we can write:

$$T|\mathbf{k}\rangle = V|\mathbf{k}\rangle + VG_0(E_\mathbf{k} + i\varepsilon)T|\mathbf{k}\rangle$$
 (10)

ullet this represents the **Lippmann-Schrödinger equation** for the operator T

Notes

- since **k** is arbitrary in Eq. (10), we postulate at the operator level: $T = V + VG_0T$
- $T = T(E_{\mathbf{k}} + \mathrm{i}\varepsilon)$ carries an implicit energy dependence via G_0
- alternative form: $T = V + TG_0V$ (seen to be equivalent by iteration)

The T-matrix

- ullet in practice, we want an <code>explicit</code> representation of T
- ullet we apply $\langle {f p}|$ from the left and define $T(E_{f k}+{
 m i}arepsilon;{f p},{f k})=\langle {f p}|T(E_{f k}+{
 m i}arepsilon)|{f k}
 angle$
- this is called the **T-matrix**, and it satisfies

$$\langle \mathbf{p}|T|\mathbf{k}\rangle = \langle \mathbf{p}|V|\mathbf{k}\rangle + \langle \mathbf{p}|VG_0(E_\mathbf{k} + \mathrm{i}\varepsilon)T|\mathbf{k}\rangle \tag{11}$$

- ullet Eq. (11) involves the momentum-space potential $V(\mathbf{p},\mathbf{k})=\langle \mathbf{p}|V|\mathbf{k}
 angle$
- to solve Eq. (11), we need to fully write out the second term on the right
 - ▶ insert complete sets of momentum states
 - ullet note that Green's function is diagonal: $\langle {f q}|G_0(z)|{f q}'
 angle = G_0(z;{f q})\delta^{(3)}({f q}-{f q}')$

$$T(E_{\mathbf{k}} + i\varepsilon; \mathbf{p}, \mathbf{k}) = V(\mathbf{p}, \mathbf{k})$$

$$+ \int \frac{d^{3}q}{(2\pi)^{3}} V(\mathbf{p}, \mathbf{q}) G_{0}(E_{\mathbf{k}} + i\varepsilon; \mathbf{q}) T(E_{\mathbf{k}} + i\varepsilon; \mathbf{q}, \mathbf{k}) \quad (12)$$

• this integral equation can be solved numerically via discretization

Momentum discretization

Numerical quadrature

ullet a quadrature rule is a set of mesh points p_i together with associated weights w_i such that for a function f(p) it holds that

$$\int_a^b f(p) \, \mathrm{d}p \approx \sum_{i=1}^N w_i f(p_i) \tag{13}$$

- the interval boundaries may be among the p_i (closed rule) or not (open rule)
- increasing N improves the approximation in Eq. (13)

Mesh class

• a very common choice is **Gauss-Legendre quadrature**

```
from lib.mesh import *
mesh = GaulegMesh(16, 0.0, 1.0)
print(mesh.ps())
print(mesh.ws())
```

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Numerical integration

On-shell singularity

- consider now Eq. (10) written out and with explicit expression for G_0
- assume furthermore we have projected the equation on a single fixed partial wave:

$$T(E_k + i\varepsilon; p, k) = V(p, k) + \int_0^{\Lambda} \frac{dq \, q^2}{2\pi^2} \frac{V(p, q)T(E_k + i\varepsilon; q, k)}{k^2 + i\varepsilon - q^2}$$
(14)

- ullet for convenience, we have introduced a momentum cutoff Λ
- for $\varepsilon=0$, we would be integrating over the singularity at q=p, which can be isolated by writing

$$k^2 + i\varepsilon - q^2 = -(q - k - i\tilde{\varepsilon})(q + k), \qquad (15)$$

where $\lim_{\tilde{arepsilon} o 0}$ is equivalent to $\lim_{arepsilon o 0}$

this can now be dealt with using the principal value formula

$$\lim_{\varepsilon \to 0} \frac{1}{x \pm \mathrm{i}\varepsilon} = \mathrm{PV} \frac{1}{x} \mp \mathrm{i}\pi \delta(x) \tag{16}$$

• but how do we realize this with a numerical quadrature?

Numerical principal value integration

schematically we are dealing with an integral of the form

$$\lim_{\varepsilon \to 0} \int_0^{\Lambda} dq \, \frac{f(q)}{q - (k + i\varepsilon)} = PV \int_0^{\Lambda} dq \, \frac{f(q)}{q - k} + i\pi f(k) \tag{17}$$

where we have defined

$$f(q) = -\frac{q^2}{2\pi^2} \frac{V(p,q)T(E_k;q,k)}{q+k}$$
 (18)

ullet now, we add 0=f(k)-f(k) in Eq. $({f 17})$ and get:

$$PV \int_0^{\Lambda} dq \, \frac{f(q) - f(k)}{q - k} + PV \int_0^{\Lambda} dq \, \frac{f(k)}{q - k}$$
 (19)

- ▶ in the first term, the numerator cancels the singularity
- ► the integral in the second term can be solved analytically

Numerical principal value integration

overall, we have arrived at

$$PV \int_0^{\Lambda} dq \, \underbrace{\frac{f(q) - f(k)}{q - k}}_{\tilde{f}(q)} + PV \int_0^{\infty} dq \, \frac{f(k)}{q - k} + i\pi f(k)$$
(20)

- ullet to the first term, we can now apply a standard quadrature: $o \sum w_i ilde{f}\left(q_i
 ight)$
- ullet after this step, we can split up ilde f again because it's all finite sums!
- the second term is simply $f(k)\log\Bigl(rac{\Lambda-k}{k}\Bigr)$
- ullet overall, we arrive at the original integral (first part of ilde f), and a sum of three terms multiplying f(k):

$$R = \sum_{i} rac{w_i}{q_i - k} + \log\left(rac{\Lambda - k}{k}
ight) + \mathrm{i}\pi$$
 (21)

ullet this can be interpreted as adding an additional point $q_0=k$ with weight $w_0=R$ to the original quadrature mesh $\{q_i\}_{i=1}^N$

Partial-wave projection

ullet for V spherically symmetric, it makes sense to expand the T-matrix in partial waves:

$$T(E_{\mathbf{k}};\mathbf{p},\mathbf{k}) = \sum_{l=0}^{\infty} (2l+1)T_l(E_k;p,k)P_l(\cos heta) \hspace{1cm} (22)$$

- $\cos \theta$ here denotes the angle between **k** and **p**
- the l-th partial-wave projection is then given by

$$T_l(E_k; p, k) = rac{1}{2} \int \mathrm{d}\cos\theta \, P_l(\cos\theta) T(E_k; \mathbf{p}, \mathbf{k})$$
 (23)

this projection can be applied to the Lippmann-Schwinger equation as a whole:

$$egin{align} T_l(E_k+\mathrm{i}arepsilon;p,k) &= V_l(p,k) \ &+ \int rac{dq\,q^2}{2\pi^2} V_l(p,q) G_0(E_k+\mathrm{i}arepsilon;q) T_l(E_k+\mathrm{i}arepsilon;q,k) \end{array} \ \ (24)
onumber$$

ullet this involves the partial-wave projected potential $V_l(p,k)$

Scattering amplitude

ullet the **scattering amplitude** is now defined, for $|{f k}'|=|{f k}|=k$, as

$$f(\mathbf{k}', \mathbf{k}) = f_p(\theta) = -\frac{\mu}{2\pi} \langle \mathbf{k}' | V | \psi_{\mathbf{k}}^{(+)} \rangle = -\frac{\mu}{2\pi} \langle \mathbf{k}' | T | \mathbf{k} \rangle$$
 (25)

- physically, it describes the modulation of the outgoing scattered wave
- it has an expansion into partial waves analogous to the T-matrix
- the on-shell T-matrix is related to the partial-wave scattering amplitude, and also to the partial-wave S-matrix

$$f_l(k) = -\frac{\mu}{2\pi} T_l(E_k; k, k) = \frac{S_l(k) - 1}{2ik}$$
 (26)

Notes

- ullet conventions regarding the prefactors in Eq. (26) may differ in the literature
- in the three-dimensional vector representation, the on-shell point is defined as $T(E_{\bf k};{\bf k}',{\bf k})$ with $|{\bf k}'|=|{\bf k}|$

Scattering phase shift

ullet unitarity of the partial-wave S-matrix, $|S_l(k)|=1$, implies that

$$S_l(k) = e^{2i\delta_l(k)} \tag{27}$$

- this defines the **scattering phase shift** $\delta_l(k)$
- the factor of two in the exponent is a (convenient) convention at this point
- from this form we obtain further useful ways to express the scattering amplitude:

$$f_l(k) = rac{\mathrm{e}^{2\mathrm{i}\delta_l(k)} - 1}{2\mathrm{i}k} = rac{\mathrm{e}^{\mathrm{i}\delta_l(k)}\sin\delta_l(k)}{k} = rac{1}{k\cot\delta_l(k) - \mathrm{i}k}$$
 (28)

 the final form in Eq. (28) is particularly useful to calculate the phase shift from the T-matrix:

$$\delta_l(k) = \operatorname{arccot}\left(-\frac{2\pi}{\mu k}T(E_k; k, k)^{-1} + \mathrm{i}\right)$$
 (29)

• the ik in the Eq. (28) is directly related to the i π in the principal-value formula

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Solution of Lippmann-Schwinger equation

Effective range expansion

• the scattering amplitude can be expanded in p, and it can be shown that for the l-th partial wave, $f_l(k)=\mathcal{O}(k^{2l})$, so:

$$f_l(p) = \frac{1}{k \cot \delta_l(k) - \mathrm{i}k} = \sum_{n=2l}^{\infty} c_{l,n} k^n$$
 (30)

- this implies a hierarchy of partial waves
 - ► larger *l* become subsequently importent as *p* increases
 - only S-waves (l=0) contribute for k o 0
- odd powers in Eq. (30) actually only arise from the unitarity cut, -ik
- ullet all nontrivial physics information is contained in $k\cot\delta_l(k)$, and this quantity is analytic in $k^2\sim E$
- conventionally one defines the **effective range expansion** as

$$k^{2l+1}\cot\delta_l(k) = -\frac{1}{a_l} + \frac{r_l}{2}k^2 + \mathcal{O}(k^4)$$
 (31)

ullet from this one can infer that $c_{l,0}=a_l$

Scattering wavefunctions

- from the half off-shell T-matrix, we can also obtain scattering wavefunctions in momentum space
- recall the initial form of the Lippmann-Schwinger equation, Eq. (8):

$$|\psi_{\mathbf{k}}^{(+)}
angle = |\mathbf{k}
angle + (E_{\mathbf{k}} - H_0 + \mathrm{i}arepsilon)^{-1}V|\psi_{\mathbf{k}}^{(+)}
angle$$

ullet with $V|\psi_{f k}^{(+)}
angle=T|{f k}
angle$, we obtain directly:

$$\langle \mathbf{q} | \psi_{\mathbf{k}}^{(+)} \rangle = \langle \mathbf{q} | \mathbf{k} \rangle + \langle \mathbf{q} | (E_{\mathbf{k}} - H_0 + i\varepsilon)^{-1} T | \mathbf{k} \rangle$$
 (32)

- from the first term $\langle {\bf q}|{\bf k}\rangle=(2\pi)^3\delta^{(3)}({\bf q}-{\bf k})$ it is clear that this is a distribution, not an ordinary function
- the second term, with $\varepsilon \to 0$ implied, contains a smooth part as well as a pole contribution (from the on-shell point):

$$\langle \mathbf{q} | (E_{\mathbf{k}} - H_0 + i\varepsilon)^{-1} T | \mathbf{k} \rangle = \frac{2\mu T(E_{\mathbf{k}}; \mathbf{q}, \mathbf{k})}{\mathbf{k}^2 - \mathbf{q}^2 + i\varepsilon}$$
(33)

Scattering wavefunctions

altogether we have

$$\psi_{\mathbf{k}}^{(+)}(\mathbf{q}) = \langle \mathbf{q} | \psi_{\mathbf{k}}^{(+)} \rangle = (2\pi)^3 \delta^{(3)}(\mathbf{q} - \mathbf{k}) + \frac{2\mu T(E_{\mathbf{k}}; \mathbf{q}, \mathbf{k})}{\mathbf{k}^2 - \mathbf{q}^2 + i\varepsilon}$$
(34)

• Fourier-transform yields wavefunctions in configuration space:

$$\psi_{\mathbf{k}}^{(+)}(\mathbf{r}) = e^{i\mathbf{k}\cdot\mathbf{r}} + 2\mu \int \frac{d^3q}{(2\pi)^3} e^{i\mathbf{q}\cdot\mathbf{r}} \frac{T(E_{\mathbf{k}}; \mathbf{q}, \mathbf{k})}{\mathbf{k}^2 - \mathbf{q}^2 + i\varepsilon}$$
(35)

Notes

- this $\psi^{(+)}_{\mathbf{p}}(\mathbf{r})$ is an ordinary function, but it is still not normalizable and therefore not an element of the Hilbert space $L^2(\mathbb{R}^3)$
- ullet the integral in Eq. (35) goes across the pole at ${f q}={f k}$, so it is again defined as a principal value
- numerically this is handled by exactly the same modified quadrature rule as before

Scattering eigenstates

• the Schrödinger equation asserts that

$$(H_0 + V)|\psi_{\mathbf{k}}^{(+)}\rangle = E_p|\psi_{\mathbf{k}}^{(+)}\rangle \tag{36}$$

- in configuration space, where $\psi_{\mathbf{k}}^{(+)}(\mathbf{r})$ is an ordinary function and H is a differential operator, we can test this by looking at the ratio $(H\psi_{\mathbf{k}}^{(+)})(\mathbf{r})/\psi_{\mathbf{k}}^{(+)}(\mathbf{r})$
- in momentum space, the analog of this is not well defined because we are dealing with distributions:

$$\psi_{\mathbf{k}}^{(+)}(\mathbf{q}) = (2\pi)^3 \delta^{(3)}(\mathbf{q} - \mathbf{k}) + \cdots$$
(37)

- ullet nor can we take matrix elements of $|\psi_{f k}^{(+)}
 angle$
- the resolution is that scattering states are really **generalized eigenstates** of the Hamiltonian, in the sense that for any test function $\phi(\mathbf{q}) = \langle \phi | \mathbf{q} \rangle$ it holds that

$$\langle \phi | (H_0 + V) | \psi_{\mathbf{k}}^{(+)} \rangle = E_k \langle \phi | \psi_{\mathbf{k}}^{(+)} \rangle \tag{38}$$

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Eigenstate verification

Bound states

- we will formally derive later that bound states correspond to poles of the S-matrix at negative energies
- ullet since, schematically, $S={f 1}+T$, these poles really need to be in the T-matrix
- moreover, it can be shown that these poles are necessarily simple poles
- ullet therefore, assuming the existence of a bound state at energy E=-B, we can parametrize the T-matrix in the vicinity of the pole as follows:

$$T(E;p,p') \sim rac{Z(p,p')}{E+B} \; ext{ for } \; E
ightarrow -B$$

- ullet finally, it can be shown that the residue Z(p,p') factorizes: $Z(p,p')=B^*(p)B(p')$
- we call B(p) the vertex fuction associated with the bound state, and we will later derive its relation to the momentum-space wavefunction of the bound state
- in abstract operator notation, we write Eq. (39) as

$$T(E) \sim rac{|B
angle\langle B|}{E+B} \; ext{ for } \; E
ightarrow -B$$

Bound-state equation

• inserting the factorized pole form into the Lippmann-Schwinger equation gives:

$$\frac{B^*(p)B(p')}{E+B} = V(p,p') + \int \frac{dq \, q^2}{2\pi^2} V(p,q) G_0(E;q) \frac{B^*(q)B(p')}{E+B} \tag{41}$$

- ullet we have used here the partial-wave projected form and dropped the subscript l
- ullet next, we multiply through with (E+B) and take the limit E
 ightarrow -B
 - ullet the potential does not have any poles, so $\lim_{E o -B}V(p,p')=0$
 - ullet the free Green's function is also regular for E < 0

$$B^*(p)B(p') = \int \frac{\mathrm{d}q\,q^2}{2\pi^2} V(p,q)G_0(-B;q)B^*(q)B(p')$$
 (42)

• finally, we can drop the common factor B(p') on both sides and use that both V and G_0 are real to obtain the **Schrödinger equation for the vertex function**:

$$B(p) = \int \frac{\mathrm{d}q \, q^2}{2\pi^2} V(p, q) G_0(-B; q) B(q) \tag{43}$$