

1. To observe the transit of a planet, it's inclination should be 90 degrees with respect to Earth. If we are not viewing it in such a way that the planet passes through our field of view of the star, we will not be able to view a transit
2. From Figure 1 we can see that the planet is transiting the star once every 5 days(approximately). So, this must be the period of the planet
3. We can estimate the radius by measuring the transit depth.

$$\Delta L/L = (R_{\text{planet}}/R_{\text{star}})^2$$

The relative flux when the planet passes in front of it is in between 0.9980 and 0.9975 so let's assume it to be 0.9977

$$\Delta L/L = (1 - 0.9975)/1 = 0.0025$$

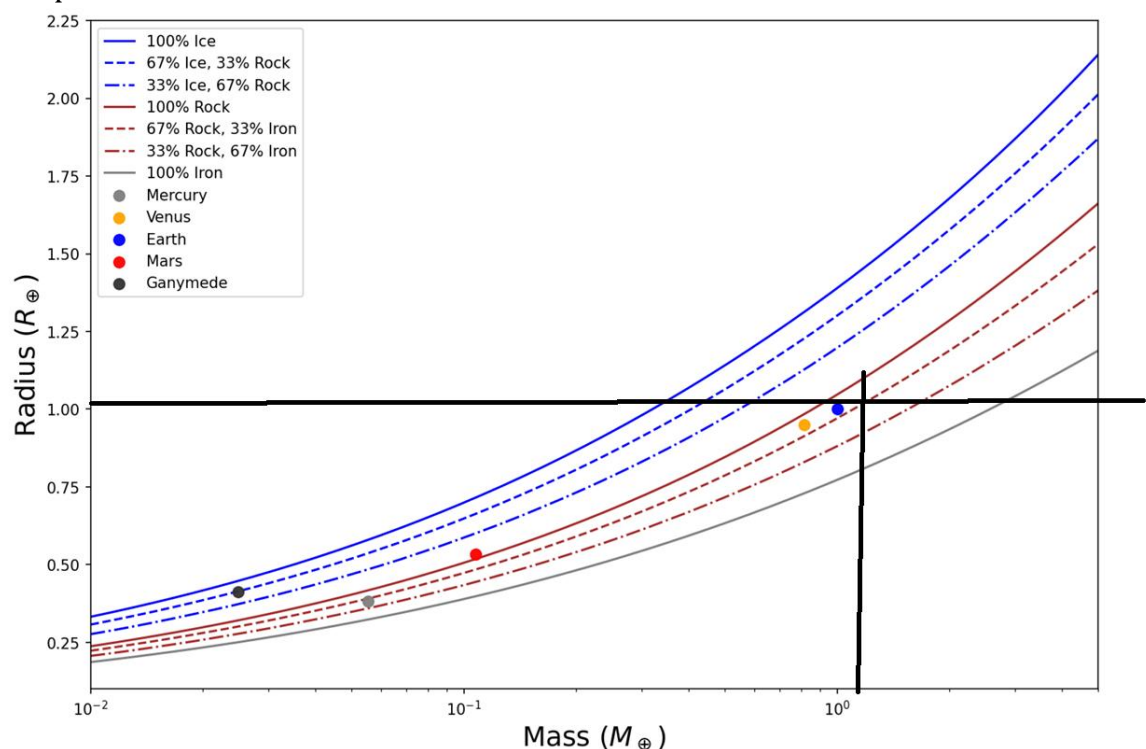
$$\text{Sqrt}(0.0025) * 0.2 * R_{\odot} = R_{\text{planet}}$$

So, the radius of the planet is  $0.01 R_{\odot}$

4. K is half of the difference between the crest and trough values of the radial velocity graph. In the case, it is  $(2 - (-2))/2 = 4/2 = 2 \text{ m/s}$
5.  $M_p = K / \sin(i) (2\pi G / P M_s^2)^{1/3}$   
 $M_p = 2 / \sin(90^\circ) ((2 * 3.14 * 6.67 * 10^{-11}) / (5 * 3600 * 24 * 2 * 1.98 * 10^{29} * 1.98 * 10^{30}))^{1/3}$   
 $M_p = 2 / \sin(90^\circ) ((23.04 * 10^{-11} / 3.38 * 10^{65}))^{1/3}$   
 $M_p \sim 1.17 * 10^{25} \text{ kg}$   
 $M_p \sim 1.97 M_e$
6. We have found out previously that  $R_p = 0.01 R_{\odot}$

Now,  $R_{\odot} = 109 R_e$

So  $R_p = 1.09 R_e$



The two lines intersect on the line depicting a composition of 67% rock and 33% iron. So, the composition of GJ 8999b must approximately be that