## Rfringe

November 5, 2003

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fillzm	Create a matrix of Zernike polynomial values

## Description

Creates a matrix of Zernike polynomials values from vectors  ${\tt rho}$  and  ${\tt theta}$ 

#### Usage

```
fillzm(rho, theta, phi = 0, zlist = zlist.qf)
```

## Arguments

rho	normalized radius, $0 \le rho \le 1$
theta	angular coordinate
phi	angular coordinate to rotate entire coordinate system
zlist	A list with named components n, m, t

fraunhofer

#### **Details**

The arguments rho and theta must be vectors of the same length.

The optional argument phi is an angle specified in degrees for rotation of the entire coordinate system. All angles are measured increasing counterclockwise.

#### Value

A matrix of size length(rho) rows by length(zlist\$n) columns.

#### Note

The main use for this function in Rfringe is creating a matrix of Zernikes to be used as the predictors in the least squares fit to fringe orders.

## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

## See Also

```
Zernike, makezlist, zlist.qf.
```

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 $Star\ test\ simulator$ 

## Description

A simple star test simulator, with optional MTF or wavefront plot.

#### Usage

```
fraunhofer(zcoef, zlist = zlist.qf, obstruct = 0, lambda = 1, defocus = 5,
    pupilsize = 255, npad=1024, gamma=2, psfmag = 2,
    displaymtf = TRUE, displaywf = FALSE, fileout = FALSE)
```

## Arguments

fileout

zcoef	Vector of Zernike coefficients
zlist	A list with named components ${\tt n}$ , ${\tt m}$ , ${\tt t}$ describing the contents of ${\tt zcoef}$
obstruct	central obstruction fraction
lambda	Wavelength. Defaults to 1 for a wavefront in waves.
defocus	Amount of defocus for side panes, in waves P-V.
pupilsize	Size of matrix representing pupil
npad	Size of the 0 padded matrix for fft calculations
gamma	gamma value for image display
psfmag	Magnification factor for in focus PSF display
displaymtf	Boolean: display MTF?
displaywf	Boolean: display Wavefront?

Boolean: Is the intended output a file?

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#### **Details**

Computes the intensity distribution from a monochromatic point source in the approximation of Fraunhofer diffraction theory. See *Suiter* (1994) for a full discussion of the application of the theory to practical star testing of small telescopes.

#### Note

The ratio of npad to pupilsize determines the size of the Airy disk in the "in focus" image. The default choices will produce an Airy disk with a diameter of approximately 10 pixels, and the amount of zero padding is sufficient to prevent significant aliasing for defocus values up to about 20 waves.

If memory limitations cause errors satisfactory images can still be obtained with, say, npad=512 and pupilsize=127.

## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

#### References

Born, M. and Wolf, E. 1999, *Principles of Optics, 7th Edition*, Cambridge University Press. Suiter, H. R., 1994, *Star Testing Astronomical Telescopes*, Willman-Bell, Inc.

#### See Also

```
Zernike, pupil.
```

## Examples

```
# a random, but probably almost diffraction limited, wavefront
fraunhofer(rnorm(length(zlist.qf$n), mean=0, sd=0.01), displaywf=TRUE)
```

gray256

8 bit Grayscale

### Description

A vector of gray scale levels

#### Usage

gray256

#### Value

```
Defined as gray256 <- grey(seq(0,1,length=256))
```

#### Author(s)

M.L. Peck (mpeck1@ix.netcom.com).

4 interferogram

interferogram	Interferogram	object
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#### Description

Creates an instance of an interferogram object.

#### Usage

interferogram(filename)

#### Arguments

filename The name of a graphics file in "Portable Anymap" format

#### **Details**

Provides an "object oriented" framework for analysis of interferograms. A call of the form .thisint <- interferogram("myinterferogram.ppm") creates an "instance" of an interferogram object. All functions for processing the interferometry data are returned with the function call.

The current version only supports graphics files in "Portable Anymap" format, and RGB files are converted to grayscale using the R channel only on the assumption that almost all interferograms will use a He-Ne source.

If interferogram() is called from the Rfringe GUI ImageMagick's (http://www.imagemagick.org) convert tool is used if necessary to convert from any graphics format supported by ImageMagick to pnm format. This means that convert must be present and on the search path if this feature is to be utilized. In Windows it also means that the path to the ImageMagick installation must be placed earlier in the path list than system files, since there is also a system convert command.

## Value

A list with the following components:

The environment of the interferogram() function call ev isInterferogram Identifies this object as an interferogram Data entry function - Basic image information image.info analysis.info Data entry function - Information required in wavefront analysis target.conic.info Data entry function - Target conic in single pass tests Image analysis function - Outline aperture edge circle.pars Image analysis function - Outline obstruction (perforation) obstruct.pars plot.fringes Plot function - Plots fringe trace autotrace Image analysis function - (Semi) automatic fringe tracing

clearpoints Image analysis function - Fringe editing
addpoints Image analysis function - Fringe editing

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clearfringe Image analysis function - Fringe editing retrace Image analysis function - Fringe editing Image analysis function - Fringe editing addsegment insertfringe Image analysis function - Fringe editing fitzernikes Image analysis function - Least squares fit to fringe centers Plot function - Synthetic interferogram plot.si plot.wf Plot function - Wavefront map plot.contour Plot function - Cross sections of wavefront along selected diameters Plot function - Interactive 3d plot of wavefront (requires package rgl) plot.wf3d plot.residuals Plot function - Some possibly useful diagnostic plots of residuals from least squares fit plot.startest Plot function - Star test simulation with optional MTF plot print.summary Output function - Prints basic summary statistics to the console window print.details Output function - Prints more details of estimated Zernike coefficients Output function - A printable detailed report, in pdf format (requires print.latex pdflatex)

#### Note

These functions will rarely be accessed directly at the command line if the GUI is used.

## Author(s)

M.L. Peck (mpeck1@ix.netcom.com). Thanks to Steven Koehler for valuable programming ideas, especially his approach to object instantiation in R.

## See Also

```
Rfringe, project, pixmap.
```

## Examples

```
## Don't run:
    .thisint <- interferogram("myinterferogram.ppm") # create an interferogram object
    .thisint$circle.pars() #interactively outline the aperture edge
    .thisint$obstruct.pars() #outline the obstruction
    .thisint$autotrace() #trace the fringes
    .thisint$fitzernikes() #do the least squares fit
    .thisint$plot.wf() #plot the wavefront
    .thisint$print.summary() #print some basic summary stats
## End Don't run</pre>
```

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listInterferograms

List Interferogram and Project Objects

## Description

Lists the objects identified as Interferogram or Project objects in the user's workspace.

## Usage

```
listInterferograms()
listProjects()
```

#### Value

Character vectors containing the names of objects identified as interferograms or interferogram projects.

#### Note

Interferogram objects are identified by having a named attribute isInterferogram. Similarly, project objects have a named attribute isIntProject.

## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

## See Also

```
interferogram, project.
```

 ${\tt makezlist}$ 

Make a Zernike polynomial list

## Description

Makes a complete list of Zernike polynomial indices (n, m, t) suitable for subsequent calls to fillzm, pupil, etc.

## Usage

```
makezlist(minorder = 2, maxorder = 12)
```

#### Arguments

minimum radial polynomial order (must be even)

maxorder maximum radial polynomial order (must be even)

project 7

#### Value

A list with the following components:

n Radial polynomial order

m Azimuthal order

t character for trig function: one of c("n", "c", "s")

#### Note

This is a popular one dimensional ordering of the indices (n,m) of Zernike polynomials.

The returned list is used in the functions fillzm, pupil, zmult, etc.

#### Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

project

Interfometry project

## Description

Creates an instance of an interferometry project.

## Usage

```
project(project.id, project.notes=NULL, project.tester=NULL, project.date=NULL)
```

## Arguments

```
project.id Character string - an identifier for the project

project.notes Optional character string - additional project notes

project.tester Optional character string - the tester (or any other useful information)

project.date Optional character string - date of the test (or other information)
```

#### **Details**

Provides an "object oriented" framework for the analysis of multiple interferograms as a group. A call of the form .thisproject <- project("myproject") creates an "instance" of a project object. All functions for processing the grouped data are returned with the function call.

## Value

A list with the following components:

```
ev The environment of the project() function call

isIntProject Identifies this object as an interferometry project

project.addto Utility function - copies data from an interferogram into project object
```

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project.removefrom Utility function - removes interferogram data from project object Plot function - Wavefront map plot.wf Plot function - Cross sections of wavefront along selected diameters plot.contour plot.startest Plot function - Star test simulation with optional MTF plot Plot function - Interactive 3d plot of wavefront (requires package rgl) plot.wf3d Plot function - Coplots of RMS, P-V, and Strehl from individual interferplot.spm ograms Output function - Prints basic summary statistics to the console window print.summary print.details Output function - Prints more details of estimated Zernike coefficients Output function - A printable detailed report, in pdf format (requires print.latex

#### Note

These functions will rarely be accessed directly at the command line if the GUI is used.

#### Author(s)

M.L. Peck (mpeck1@ix.netcom.com). Thanks to Steven Koehler for valuable programming ideas, especially his approach to object instantiation in R.

#### See Also

```
Rfringe, interferogram,
```

pdflatex)

## Examples

pupil

Create a circular pupil and fill it with a wavefront

#### Description

Creates a representation of a wavefront from a vector of Zernike polynomial values.

#### Usage

```
pupil(size = 255, obstruct = 0, zcoef = NULL, zlist = zlist.qf, phi = 0, piston = 0)
```

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#### Arguments

size	size of the returned matrix
obstruct	${\bf central\ obstruction}\ fraction$
zcoef	Vector of Zernike coefficients

zlist A list with named components n, m, t describing the contents of zcoef

phi angular coordinate to rotate entire coordinate system

piston Piston (constant) term to add to the wavefront

#### **Details**

The coordinate system is rotated *clockwise* by the angle **phi** specified in *degrees*. This is done to present consistent displays of rotated wavefronts in **Rfringe**.

#### Value

A size by size matrix of wavefront values. NA's are used to fill out the matrix outside the circular pupil and inside the obstruction.

#### Note

Most high level R graphics functions will handle NA's as intended.

This function can take a while if zcoef includes high order Zernikes. Decreasing the matrix size will help with speed, but may provide too low resolution for good graphical representations. No attempt is made to "anti-alias" the edges of the pupil.

## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

#### See Also

```
Zernike, makezlist, zlist.qf, fillzm, pupilrms, pupilpv.
```

## Examples

```
# A fairly typical use of this function:
wf <- pupil(zcoef=rnorm(length(zlist.qf$n)))
image(wf, col=topo.colors(256), asp=1)
contour(wf, add=TRUE)</pre>
```

Rfringe-internal  $Rfringe\ internal\ functions$ 

## Description

Internal functions called by the Rfringe GUI.

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#### Usage

```
aboutrfringe()
addtoproject()
analysisinfo()
autotrace()
autotrace.options()
batchadd()
changeint()
changeproject()
circlepars()
clearint()
clearproject()
closeRfringe()
closeRfringeandr()
editfringe()
fitzernikes()
helpbox(string)
helprfringe()
imageinfo()
maxorder()
newproject()
obstructpars()
openint()
opensavedint()
opensavedproject()
pdfreport()
pdfreport.project()
plotcontour()
plotcontour.project()
plotfringes()
plotresiduals()
plotspm.project()
plotstartest()
plotstartest.project()
plotwf()
plotwf.project()
printdetails()
printdetails.project()
printsummary()
printsummary.project()
printtoc.project()
prompts(string)
removefromproject()
saveint()
saveproject()
synthint()
targetconicinfo()
tclvar()
wf3d()
wf3d.project()
wf3d.vanilla()
wf3d.vanilla.project()
```

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#### **Details**

These are called by Rfringe() and are not to be called directly by the user. From the command line there are always functions stored within interferogram, project, or R to perform the actions carried out by these functions.

#### Author(s)

M.L. Peck (mpeck1@ix.netcom.com)

#### See Also

Rfringe, interferogram, project

Rfringe

Rfringe GUI

## Description

GUI front end for R fringe analysis.

#### Usage

Rfringe()

## **Arguments**

none

#### **Details**

A Tcl/Tk wrapper for the functions provided in interferogram and project.

#### Note

The full functionality of Rfringe will be described in a separate user guide. This note briefly describes how objects are created and manipulated within the GUI interface.

On startup Rfringe creates NULL valued objects named .thisint and .thisproject in the user's global environment. When the user subsequently creates a new interferogram object with the menu selection File|Load interferogram from image... make.names is run on the value entered in the image id field to create a new (syntactically correct) variable in the user's workspace. This variable is assigned the value of the call to interferogram() and .thisint is in turn evaluated to the newly created interferogram object. Subsequent commands operate directly on .thisint.

Similarly when a new project object is created with the menu selection File | Project | Create... a new variable is created from the Project ID entry, it is assigned the value of the call to project(), and .thisproject is evaluated to the newly created project object.

The objects created with make.names are deliberately *not* assigned unique names. This allows the user to replace an object simply by recycling her image or project id entries. The Rfringe GUI will warn if this is attempted.

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## Author(s)

M.L. Peck  $\langle mpeck1@ix.netcom.com \rangle$ . GUI based in part on the package Rcmdr by John Fox.

#### See Also

```
tcltk, interferogram, project
```

## Examples

```
# start the GUI
Rfringe()
```

rzernike

Radial Zernike Polynomial

## Description

Zernike's Radial Polynomials

## Usage

```
rzernike(rho, n, m)
odd(n,m)
```

## Arguments

rho normalized radius,  $0 \le rho \le 1$ 

n radial polynomial order

m azimuthal order

## Details

The arguments  ${\tt n}$  and  ${\tt m}$  must be relatively even.

#### Value

The value of the radial Zernike polynomial of order (n,m) at normalized radius rho. If rho is a vector, matrix, or higher order array the returned value is a vector or array of the same dimension.

The utility function odd(n,m) returns FALSE iff (n-m)%2 == 0.

### Note

This routine implements the recurrence relation given in Equation 12 of http://mathworld.wolfram.com/ZernikePolynomial.html.

In general you should call the higher level function Zernike instead of this.

## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

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#### References

```
http://mathworld.wolfram.com/ZernikePolynomial.html
```

#### See Also

Zernike.

Summarystats

 $Wave front\ summaries$ 

## Description

Estimate the RMS or P-V wavefront error over the pupil given in the argument.

## Usage

```
pupilrms(pupil)
pupilpv(pupil)
strehlratio(rms)
```

## Arguments

pupil is the matrix created by the call to pupil

rms The rms wavefront error

#### Value

An estimate of the RMS or P-V error of the wavefront, or Strehl ratio.

## Note

The function pupilrms simply returns the standard deviation of the defined values in pupil, which is a crude but usually good enough approximation to the properly defined integral over the aperture.

For a wavefront defined entirely in terms of a vector zcoef of Zernike coefficients rms <-sqrt(crossprod(zcoef)) is faster and more accurate.

pupilpv does the obvious. There is no analytical solution in general for P-V.

strehlratio computes Mahajan's approximation to the Strehl ratio.

## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

#### See Also

pupil.

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#### Examples

```
# A random vector of Zernike coefficients
zcoef <- rnorm(length(zlist.qf$n), mean=0, sd=0.01)
wf <- pupil(zcoef=zcoef)
image(wf, col=topo.colors(256), asp=1)
contour(wf, add=TRUE)
pupilrms(wf)
sqrt(crossprod(zcoef)) #should be the same to about 4 digits
pupilpv(wf)
strehlratio(sqrt(crossprod(zcoef))) #probably around 0.8</pre>
```

```
synth.interferogram Synthetic Interferogram
```

#### Description

Computes and displays a synthetic interferogram for a wavefront constructed from a vector of Zernike coefficients.

#### Usage

```
synth.interferogram(zcoef, zlist = zlist.qf, phi = 0,
    size = 255, obstruct = 0, iname = "")
```

### Arguments

zcoef	Vector of Zernike coefficients, with piston term as the first element
zlist	A list with named components n, m, t describing the contents of zcoef
phi	angular coordinate to rotate entire coordinate system
size	Size of matrix representing pupil
obstruct	central obstruction fraction
iname	short string for identification

## **Details**

It's important to note that zcoef is treated differently than in other functions that use the same variable name. The first element *must* be a piston (constant) term, which is stripped off and passed to pupil as the piston argument. The length of zcoef therefore should be one more than the length of zlist\$n.

#### Value

A size by size matrix of intensity levels in the simulated interferogram.

#### Note

The relationship between wavefront phase and intensity is iwf <- cos(2 \* pi \* wf + pi), which is the value returned by synth.interferogram. The plot routine in this function plots the image on a 256 level grayscale.

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## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

#### See Also

```
Zernike, pupil.
```

#### Examples

```
zcoef <- c(0, 3, 3, rnorm(length(zlist.qf$n)-2, mean=0, sd = .01))
temp <- synth.interferogram(zcoef, iname="Random wavefront")
# lets see what it looks like in a star test
zcoef <- zcoef[-1]
zcoef[1:2] <- 0
fraunhofer(zcoef, displaywf=TRUE)</pre>
```

wf.3dplot

Interactive 3D Wavefront plot

#### Description

Uses the rgl package for OpenGL graphics to produce an interactive 3D wavefront map.

#### Usage

```
wf.3dplot(wf, zoom.wf = 1)
```

## Arguments

wf A wavefront matrix as returned by pupil zoom.wf Zoom factor to stretch wavefront heights

## Details

The rgl package is available at http://wsopuppenkiste.wiso.uni-goettingen.de/~dadler/rgl/. The README file in the source distribution contains installation instructions.

## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

#### References

```
http://wsopuppenkiste.wiso.uni-goettingen.de/~dadler/rgl/
```

## See Also

pupil

16 Zernike

wf.persp	3D Wavefront plot
----------	-------------------

## Description

An alternative 3D Wavefront plot using the R base package plotting function persp.

## Usage

```
wf.persp(wf, zoom.wf = 1, theta=0, phi=30, ...)
```

## Arguments

wf A wavefront matrix as returned by pupil
zoom.wf Zoom factor to stretch wavefront heights
theta Value of theta to pass to persp

phi Value of theta to pass to persp

Value of phi to pass to persp

... Additional parameters for call to persp

## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

## See Also

```
pupil, persp, wf.3dplot.
```

Zernike

Zernike Polynomials

## Description

Routines for creating and manipulating Zernike polynomials.

#### Usage

```
Zernike(rho, theta, n, m, t)
```

#### Arguments

rho normalized radius,  $0 \le rho \le 1$ 

theta angular coordinaten radial polynomial order

m azimuthal order

t character for trig function: one of c("n", "c", "s")

## **Details**

The arguments  ${\tt n}$  and  ${\tt m}$  must be relatively even.

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#### Value

The value of the Zernike polynomial of order (n, m) at polar coordinates (rho, theta). The arguments rho and theta may be vectors, matrices, or higher order arrays, in which case the returned value is a vector or array of the same dimension.

#### Note

This function returns Zernikes scaled such that they form an orthonormal basis set for the space of functions defined on the unit circle. Note that this is not the most commonly used definition (as given e.g. in *Born and Wolf*). The definition I use is often associated with *Noll* (1976).

The otherwise unused function **zmult** can be used to convert between normalized and conventionally defined vectors of Zernike coefficients.

#### Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

#### References

Born, M. and Wolf, E. 1999, *Principles of Optics, 7th Edition*, Cambridge University Press, chapter 9 and appendix VII.

Noll, R.J. 1976, **Zernike polynomials and atmospheric turbulence**, *J. Opt. Soc. Am.*, Vol. 66, No. 3, p. 207.

```
http://wyant.opt-sci.arizona.edu/zernikes/zernikes.htm
http://mathworld.wolfram.com/ZernikePolynomial.html
```

## See Also

```
rzernike, makezlist, zlist.qf, zmult, fillzm, pupil, pupilrms, pupilpv, strehlratio.
```

#### Examples

```
Zernike(1, 0, 4, 0, "n") # == sqrt(5)

# A slightly more complex example

rho <- seq(0, 1, length = 101)
theta <- rep(0, 101)

plot(rho, Zernike(rho, theta, 6, 0, "n"), type="1",
   ylim=c(-3.5,3.5), main="Some 6th order Zernike Polynomials")
lines(rho, Zernike(rho, theta, 5, 1, "c"), lty=2)
lines(rho, Zernike(rho, theta, 4, 2, "c"), lty=3)
lines(rho, Zernike(rho, theta, 3, 3, "c"), lty=4)</pre>
```

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zlist.qf

Fringe set of Zernike Polynomials

## Description

List of the 'Fringe' set of Zernike polynomials

## Usage

```
zlist.qf
```

## **Details**

The "Fringe" aka "QuickFringe" set of Zernike polynomials.

## Value

A list with the following components:

n Radial polynomial order

m Azimuthal order

t character for trig function: one of c("n", "c", "s")

#### Note

This is the default list of Zernike polynomial indices used by all functions that work with lists of Zernikes. It includes all Zernikes from 2nd through 10th order, plus the 12th order spherical term.

## Author(s)

M.L. Peck (mpeck1@ix.netcom.com)

zmult

Zernike coefficient multipliers

## Description

Conversion factors between normalized and conventional Zernike polynomials.

#### Usage

```
zmult(zlist = zlist.qf)
```

## Arguments

zlist

A list with named components n, m, t

## Details

The list indicates the Zernike polynomial orders to use, in the form returned by makezlist and zlist.qf.

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## Value

A vector the same length as the components of zlist.

## Note

This function is not actually used in the current version of Rfringe, but is included for user convenience.

## Author(s)

```
M.L. Peck (mpeck1@ix.netcom.com)
```

## See Also

```
Zernike, makezlist, zlist.qf.
```

## Examples

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