



# X-RAY RESONANT RAMAN SCATTERING

SPECTRA SIMULATION FROM FIRST PRINCIPLES  
FOR COPPER BELOW IONIZATION THRESHOLD  
USING HIGH-PERFORMANCE COMPUTING

GONÇALO GARCÊS SOBREIRA RODRIGUES BAPTISTA

BSc in Physics Engineering

DOCTORATE IN PHYSICS ENGINEERING

NOVA University Lisbon  
September, 2023

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September, 2023

## **X-ray resonant Raman scattering**

### **Spectra simulation from first principles for Copper bellow ionization threshold using high-performance computing**

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# Acknowledgements

Acknowledgments are personal text and should be a free expression of the author.

However, without any intention of conditioning the form or content of this text, I would like to add that it usually starts with academic thanks (instructors, etc.); then institutional thanks (Research Center, Department, Faculty, University, FCT / MEC scholarships, etc.) and, finally, the personal ones (friends, family, etc.).

But I insist that there are no fixed rules for this text, and it must, above all, express what the author feels.

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”

*“Sometimes I’ll start a sentence, and I don’t even know where it’s going.  
I just hope I find it along the way.  
Like an improv conversation.  
An improversation.”*

— **Michael Scott**, *The Office*  
(Regional Manager of Dunder Mifflin Scranton)

---

”

*all work and no play makes Jack a dull boy*

[illegible]

— **Jack Torrance**, *The Shinning*  
(Caretaker of the Overlook)

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## Abstract

The work performed on this thesis comes as part of the effort to further understand the highly convoluted structure present on Copper's X-ray emission spectra, where, as with many other transition metals, a skewness can be observed on the  $K_{\alpha_{1,2}}$ ,  $K_{\beta}$  and L transition lines. These lines originate due to the radiative relaxation of the atom's electronic structure post-ionization of inner shell electrons. However, it is very likely that the observed skewness is due to copper's satellite states' transitions.

Throughout this thesis, a study will be performed for the satellite states formed by the excitation of the inner-shell electrons, where, as opposed to the ionization process, usually considered in X-ray calculations, a photoexcitation process occurs.

Multiple atomic structure calculations will be performed using the *ab initio* state of the art [Multiconfiguration Dirac-Fock General Matrix Elements \(MCDFGME\)](#) code for different excited states configurations.

The obtained results will then be used in the analysis of experimental data obtained from a High-Precision [Double Crystal Spectrometer \(DCS\)](#), using a synchrotron X-ray source.

Due to the complexity of the calculations, the process can become substantial in terms of computational power and time. Therefore, further similar and more complex studies will be performed by implementing and running a script in the *Oblivion* supercomputer located at the University of Évora.

**Keywords:** Atomic Excitation, X-ray lines, [MCDFGME](#), [DCS](#), High Performance Computing



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## Resumo

asdasdasdasd

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# Contents

List of Figures	xi
List of Tables	xii
Listings	xiii
Glossary	xiv
Acronyms	xv
<b>1 Theoretical Introduction</b>	<b>2</b>
1.1 Characteristic X-rays . . . . .	2
1.1.1 Ionization as a vacancy generator . . . . .	2
1.1.2 Transition notation . . . . .	4
1.1.3 Excitation as a vacancy generator . . . . .	4
1.2 Radiative transitions . . . . .	5
1.3 Solving the atomic many-body problem . . . . .	6
1.3.1 The non-relativistic Hamiltonian . . . . .	6
1.3.2 The Hartree-Fock Method . . . . .	7
1.3.3 The Dirac Equation . . . . .	8
1.4 Quantum Electrodynamics (QED) considerations . . . . .	11
1.4.1 The Multiconfiguration Dirac-Fock (MCDF) Method . . . . .	11
1.5 State of the Art . . . . .	12
1.5.1 Copper's characteristic X-rays . . . . .	12
1.5.2 MCDFGME capabilities . . . . .	13
<b>2 Atomic Structure Calculations</b>	<b>14</b>
2.1 The MCDFGME code's capabilities . . . . .	14
2.2 The atomic system at study . . . . .	14
2.2.1 The system at study . . . . .	15

2.3	Level Calculations . . . . .	15
2.4	Transition computations . . . . .	15
2.4.1	Diagram transitions . . . . .	15
2.4.2	Auger transitions . . . . .	15
2.4.3	Satellite transitions . . . . .	15
<b>3</b>	<b>A new High-Performance Computing code for the parallelization of atomic structure calculations</b>	<b>16</b>
3.1	A brief introduction on Message Passing Interface ( <i>MPI</i> ) . . . . .	16
3.2	Program advantages . . . . .	16
3.3	Program limitations . . . . .	16
<b>4</b>	<b>Fundamental atomic parameters calculation</b>	<b>17</b>
<b>5</b>	<b>Spectra simulation</b>	<b>18</b>
<b>6</b>	<b>Spectra analysis</b>	<b>19</b>
6.1	Photoexcitation cross-section estimation . . . . .	19
6.2	Photoionization cross-section computation . . . . .	19
<b>7</b>	<b>Comparison with experimental data</b>	<b>20</b>
<b>8</b>	<b>Next Steps</b>	<b>21</b>
<b>9</b>	<b>Final remarks and Conclusion</b>	<b>22</b>
9.1	Quantum states and properties . . . . .	22
9.2	State transitions . . . . .	22
9.2.1	Transition rates and widths . . . . .	22
9.2.2	Branching Ratios and Fluorescence Yields . . . . .	22
9.3	Spectra simulation . . . . .	22
<b>10</b>	<b>Spectra Analysis</b>	<b>23</b>
<b>11</b>	<b>Development of a High Performance Computing (HPC) script for the parallelization of Atomic Structure Calculations (ASC)</b>	<b>24</b>
11.1	blablabla . . . . .	24
11.2	Speedup . . . . .	24
	<b>Bibliography</b>	<b>25</b>
	<b>Appendices</b>	
<b>A</b>	<b>The Breit Hamiltonian Operators</b>	<b>29</b>
<b>B</b>	<b>Transition Diagram</b>	<b>31</b>

<b>C</b>	<b>QED considerations</b>	<b>32</b>
C.1	Self-Energy . . . . .	32
C.2	Vacuum Polarization . . . . .	32
 <b>Annexes</b>		
<b>I</b>	<b>Input File .f05 Example</b>	<b>33</b>

---

## List of Figures

1.1	Photoionization. . . . .	3
1.2	Principal atomic relaxation processes. . . . .	3
1.3	Resonant Photoexcitation . . . . .	5
1.4	HF method's block diagram. . . . .	8
1.5	Comparison of Hydrogen level diagrams. Adapted from [13] . . . . .	12
B.1	Transition notation scheme. Adapted from [2] . . . . .	31
C.1	QED Feynman Diagrams . . . . .	32

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## List of Tables

1.1 Siegbahn VS IUPAC notation. Adapted from [2]. . . . .	5
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# Listings

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## Glossary

<b>Feynman Diagram</b>	2D diagram depicting various physical interactions between elementary particles. Fermions are depicted as straight lines and bosons as wavy lines. (p. 32)
<b>four-vector</b>	Vector used in special relativity composed of 4 components, one scalar time-like, and three vectorial space-like. These vectors behave in special way, such as their norm being <a href="#">Lorentz invariant</a> . Can be written in covariant, $X_\mu$ , and contravariant form, $X^\mu$ , with the difference being the sign of the time-like components. Example of a contravariant four vector: $X^\mu = (X^0, X^1, X^2, X^3) = (X^0, \mathbf{X})$ (pp. <a href="#">xiv</a> , 9)
<b>Lorentz invariant</b>	A Lorentz invariant scalar, obtained, for example, from a <a href="#">Minkowski norm</a> , does not change when operated by a Lorentz Transformation. (pp. <a href="#">xiv</a> , 9)
<b>Minkowski norm</b>	Yields the Lorentz Invariant norm for a <a href="#">four-vector</a> : $p_\mu p^\mu$ . Equivalent to the dot product of a classical vector. (pp. <a href="#">xiv</a> , 9)
<b>self consistency method</b>	The electron wavefunctions go through a variational process in order to reach convergence for an energetic minimum. (p. 7)
<b>virtual photons</b>	While in reality, during a Coulomb interaction, 'real' particles are not exchanged, the electromagnetic field is still mediated by photons. This way virtual photons are tools used in order to better represent electromagnetic interactions. (p. 11)



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## Acronyms

<b><i>Grasp2k</i></b>	General-purpose Relativistic Atomic Structure Package 2k ( <i>p.</i> <a href="#">13</a> )
<b><i>MCDFGME</i></b>	Multiconfiguration Dirac-Fock General Matrix Elements ( <i>pp.</i> <a href="#">vi</a> , <a href="#">viii</a> , <a href="#">12–14</a> )
<b><i>MCDF</i></b>	Multiconfiguration Dirac-Fock ( <i>pp.</i> <a href="#">viii</a> , <a href="#">11</a> )
<b><i>MPI</i></b>	Message Passing Interface ( <i>pp.</i> <a href="#">ix</a> , <a href="#">16</a> )
<b>ASC</b>	Atomic Structure Calculations ( <i>pp.</i> <a href="#">ix</a> , <a href="#">24</a> )
<b>DCS</b>	Double Crystal Spectrometer ( <i>p.</i> <a href="#">vi</a> )
<b>EMF</b>	Electromagnetic Field ( <i>pp.</i> <a href="#">6</a> , <a href="#">11</a> )
<b>FAC</b>	Flexible Atomic Code ( <i>p.</i> <a href="#">13</a> )
<b>HF</b>	Hartree-Fock ( <i>p.</i> <a href="#">8</a> )
<b>HPC</b>	High Performance Computing ( <i>pp.</i> <a href="#">ix</a> , <a href="#">24</a> )
<b>QED</b>	Quantum Electrodynamics ( <i>pp.</i> <a href="#">viii</a> , <a href="#">2</a> , <a href="#">11</a> , <a href="#">14</a> )

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## Todo list

Figure: Maybe put a Thomas model plot? . . . . .	4
Explain antisymmetry and put it in the glossary . . . . .	7
Explain in the glossary . . . . .	7
Finish the explanation . . . . .	8
Try to explain why this might be a problem . . . . .	9
do the demonstration, dummy . . . . .	9
Falar aqui sobre as contribuições. Não tem que ser extensivo . . . . .	11
Ler bem o manual e explicar bem o método e todas as coisas diferentes que se podem fazer . . . . .	11
This really needs a re-do . . . . .	12
meter mais cenas . . . . .	14
Breit can be used as self consistent or perturbative . . . . .	15
Usar o Denitions aqui . . . . .	22
Compare with Amdahl's law or Gustafson's law. . . . .	24
Comparação com o do Jorge . . . . .	24

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# Theoretical Introduction

Throughout this thesis, different topics regarding Quantum Mechanics, Atomic Physics and spectroscopy will be approached, hence why a theoretical introduction is needed. In it, themes such as the characteristic X-rays of elements, the relativistic Dirac equation and QED corrections and methods for solving the many-body problem will be discussed.

## 1.1 Characteristic X-rays

When subjected to a high energy beam of particles (usually photons or electrons), an element may go through the process of the ionization of inner-core electrons, which is followed by the emission of radiation of its own, in the form of X-rays and/or Auger electrons. The radiation emitted throughout this process presents well-defined energy values and thus provides a way to ascertain if an element is present in a sample under study. Due to this reason, these characteristic emissions are of an extremely high importance in a wide range of scientific areas.

When an atomic system is in a bound state, the electrons orbiting the nucleus are occupying fixed quantum states, defined by their principal atomic number,  $n$ , angular momentum,  $l$ , and spin,  $s$ . Electrons are also fermions, and thus must respect Pauli's exclusion principle, each occupying a single state, only occupied by that single electron. These quantum numbers provide information about each electron's wavefunction, and the collection of all the electrons' wavefunctions can be used in order to describe the whole system. When operating the Hamiltonian on this collection of wavefunctions, the atomic system's state's energy is obtained. Besides all the occupied states, an infinite more number of possible eigenfunctions will compose that system's base.

### 1.1.1 Ionization as a vacancy generator

As previously mentioned, X-ray fluorescence spectroscopy has many uses and applications in a wide range of scientific areas. In this form of spectroscopy, the element at study, composed of a nucleus and  $N$  orbiting electrons, is bombarded with radiation leading to the ionization of inner-shell electrons while leaving a vacancy in their place (Figure 1.1).

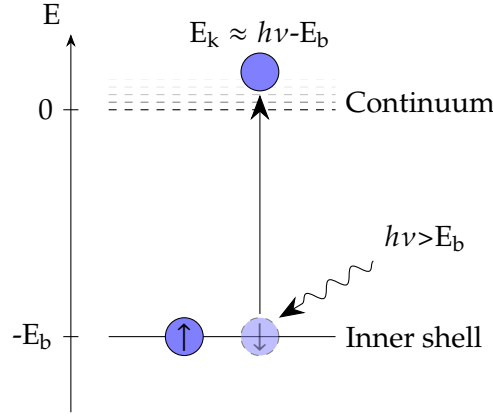


Figure 1.1: Photoionization.

The atomic structure, now composed of  $N - 1$  electrons, will be left in an energetically unstable state, due to there being other possible lower energy states. This will lead to various processes of atomic relaxation, where the system will rearrange itself in order to find a lower energy configuration.

The main processes for this rearrangement are two competing decay paths: radiative relaxation, and Auger electron emission.

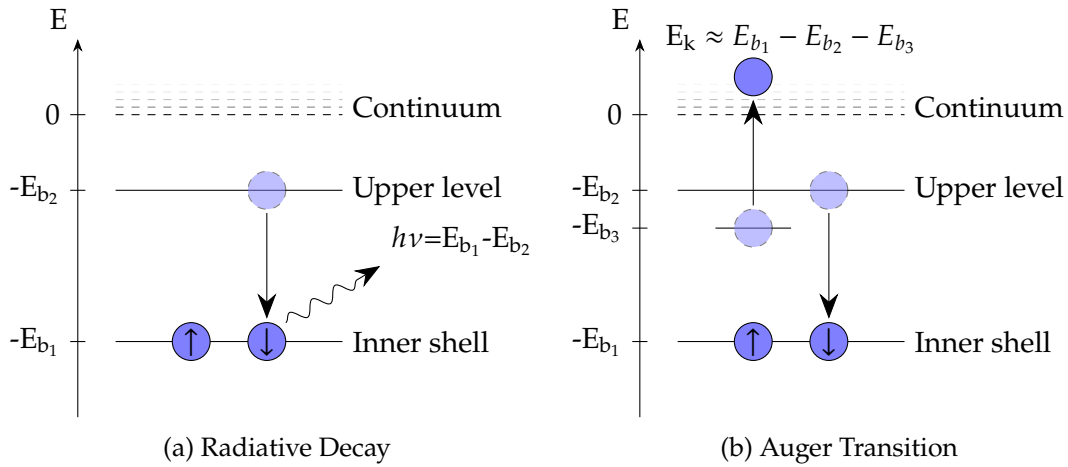


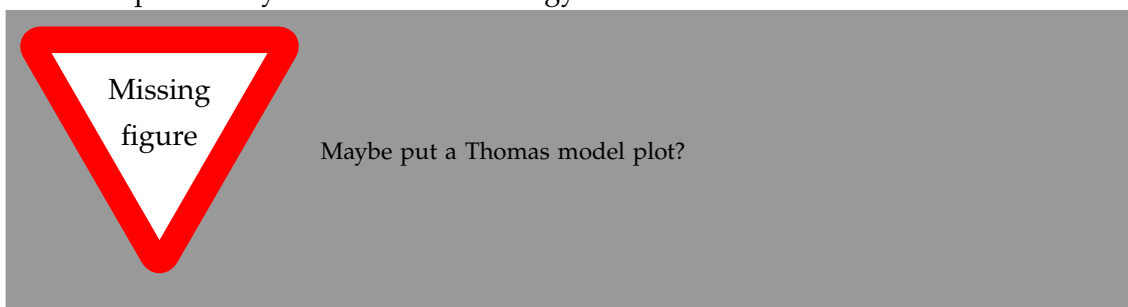
Figure 1.2: Principal atomic relaxation processes.

In the case of radiative relaxation, an upper shell electron will move and occupy the hole left in the inner shell. During this process, energy is released through the emission of photons with energy in the range of X-rays. The detection of these photons is what allows for an analysis and detection of the element at study.

On the other hand, an Auger emission occurs when in the process of an upper electron shifting to a shell hole, energy is released not in the form of photons, but by the atomic system going through the process of the ionization of a lesser bound electron.

In reality, when the initial vacancy is generated, two more processes can occur: shake-off and shake-up. Due to a sudden change in the potential felt by the remaining electrons,

whilst going through the ionization, one other electron may become unbound, leading to a second ionization, the shake-off process, or may be excited to an upper state, during shake-up. These processes' occurrence probabilities are related to the rate at which the electron leaves the system during the first ionization. For lower electron emission energies, where the ionization process is rather "slow", and there is not much energy for shake up/off to occur, the adiabatic regime reigns. In this regime, the shake probabilities are small, and increase with the beam's energy. For higher energies, the transition occurs in the sudden regime, where the electrons' exit can be considered instantaneous. Now, while the shake probabilities still increase with the energy of the beam, they quickly saturate for a maximum value of probability, which can be calculated by computing overlaps between the electrons' initial and final states. These processes, however, are out of scope for this thesis, but should the reader be interested, the model conceived by Thomas [1] predicts the shake probability as a function of energy.



### 1.1.2 Transition notation

The characteristic radiation measured from the radiative relaxation of a post-ionization unstable atomic system is one of the main ways of identifying an atomic element. This is due to the photons emitted possessing quantized values of energy, forming well-defined energy lines when observed in a spectrometer. In order to understand the change that occurred in the atomic system which lead to a specific emission, the spectral lines get labels based on a notation which usually takes into account the initial and final orbitals where the vacancy was present. A very illustrative diagram, exemplifying some transitions can be found in Appendix B.

Throughout this thesis, Siegbahn notation will be used, for the most part, but should the reader prefer IUPAC's, Table 1.1 has the conversion between notations.

### 1.1.3 Excitation as a vacancy generator

As previously mentioned, throughout this thesis, while the study is focused on the characteristic radiation emitted during an atomic relaxation process, the main vacancy generation method at study shall be the photoexcitation process (Figure 1.3), instead of ionization. The levels obtained after a photoexcitation of core shell electrons has occurred could be some of the many so-called satellite states, where the electronic configuration present during the relaxation process contains additional electrons or holes, or is simply not

Table 1.1: Siegbahn VS IUPAC notation. Adapted from [2].

Siegbahn	IUPAC	Siegbahn	IUPAC	Siegbahn	IUPAC
$K_{\alpha_1}$	$K - L_3$	$L_{\alpha_1}$	$L_3 - M_5$	$L_{\gamma_1}$	$L_2 - N_4$
$K_{\alpha_2}$	$K - L_2$	$L_{\alpha_2}$	$L_3 - M_4$	$L_{\gamma_2}$	$L_1 - N_1$
$K_{\beta_1}$	$K - M_3$	$L_{\beta_1}$	$L_2 - M_4$	$L_{\gamma_3}$	$L_1 - N_2$
$K_{\beta_2}^I$	$K - N_3$	$L_{\beta_2}$	$L_3 - N_5$	$L_{\gamma_4}$	$L_1 - O_3$
$K_{\beta_2}^{II}$	$K - N_2$	$L_{\beta_3}$	$L_1 - M_3$	$L_{\gamma_4}'$	$L_1 - O_2$
$K_{\beta_3}$	$K - M_2$	$L_{\beta_4}$	$L_1 - M_2$	$L_{\gamma_5}$	$L_2 - N_1$
$K_{\beta_4}^I$	$K - N_5$	$L_{\beta_5}$	$L_3 - O_{4,5}$	$L_{\gamma_5}$	$L_2 - O_4$
$K_{\beta_4}^{II}$	$K - N_4$	$L_{\beta_6}$	$L_3 - N_1$	$L_{\gamma_8}$	$L_2 - O_1$
$K_{\beta_4}^x$	$K - N_4$	$L_{\beta_7}$	$L_3 - O_1$	$L_{\gamma_8}'$	$L_2 - N_{5,6}$
$K_{\beta_5}^I$	$K - M_5$	$L_{\beta_8}$	$L_3 - N_{6,7}$	$L_{\eta}$	$L_2 - M_1$
$K_{\beta_4}^{II}$	$K - M_4$	$L_{\beta_9}$	$L_1 - M_5$	$L_l$	$L_3 - M_1$
		$L_{\beta_9}$	$L_1 - M_4$	$L_s$	$L_3 - M_3$
		$L_{\beta_9}$	$L_3 - N_4$	$L_t$	$L_3 - M_2$
		$L_{\beta_9}$	$L_2 - M_3$	$L_u$	$L_3 - N_{6,7}$
				$L_v$	$L_2 - N_{6,7}$

the standard configuration when talking about characteristic emissions. The characteristic radiation from transitions that originated from these states are one of the keys needed to fully comprehend and deconvolute an element's emission spectra.

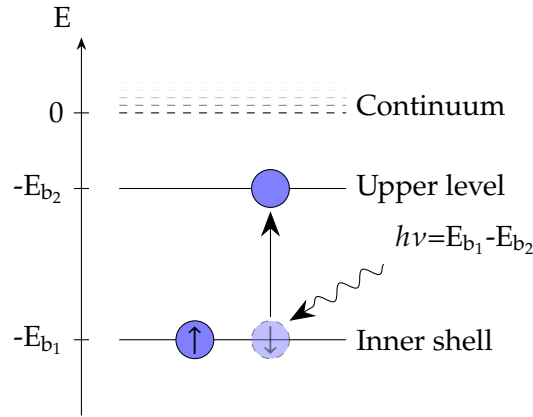


Figure 1.3: Resonant Photoexcitation

## 1.2 Radiative transitions

In terms of quantum mechanics, the properties of a system's change of state can be derived using perturbation theory, where a change in energy can be treated as perturbation, and takes into account the conservation laws physical properties, such as the angular momentum of the system as basis for the selection rules.

The transition's radiation type can be of two main flavors, Electric,  $E_k$ , or Magnetic,  $M_k$ ,

with  $k$  representing its multiplicity. These labels indicate which of the [Electromagnetic Field \(EMF\)](#) component had the strongest influence on the transition's occurrence. For both these types of transitions, the system's total angular momentum is allowed a change of  $\Delta J = 0, \pm k$ , and so does its projection,  $M_J$ . However, different transitions lead to different changes of parity in the system and to different selection rules.

It should also be of note that, usually, for the same multiplicity, an  $E$  transition is more intense than an  $M$ , and that it is also possible for rare 2-photon transitions to take place, where a combination of different transition types can occur.

The intensity of a transition is proportional to the squared norm of the perturbation's matrix element involving the initial and final state ( $\Gamma_{if} \propto |\langle \psi_i | H' | \psi_f \rangle|^2$ ). It should also be noted that the electron's initial state population will serve as a scaling factor for the transitions' rate. For example, while the transition rate for a  $2p_{1/2} \rightarrow 1s$  and  $2p_{3/2} \rightarrow 1s$  should be about the same, the orbital  $2p_{3/2}$  has double the population  $2p_{1/2}$  has, hence why the  $K_{\alpha_2}$  line has about half the  $K_{\alpha_1}$  line's intensity.

One should also mention that monochromatic transitions do not exist. While the transition might have a well-defined energy, calculated by the difference in energy between the initial and final levels, due to Heisenberg's uncertainty principle,  $\Delta E \Delta t > \frac{\hbar}{2}$ , there will exist a natural energy broadening, the transition's natural width. The shapes representing these transitions are given by a Lorentzian distribution.

## 1.3 Solving the atomic many-body problem

When studying a system composed of multiple charged bodies, one must consider all the existing interactions. Whilst there are known analytical solutions for 2-bodies Hydrogenoid systems, with the presence of more non spatially-bound particles, the Coulomb interaction pairs lead to the impossibility of finding a set of analytical wavefunctions which are a part of the Hamiltonian's eigenset. Consecutively, the need for a numerical method that is able to compute solutions for these complex systems arose.

### 1.3.1 The non-relativistic Hamiltonian

The first approach used in order to solve the many-bodies problem used a non-relativistic consideration. This way, the Hamiltonian consists on the sum of the system's non-relativistic momentum-related energies and the energy of Coulomb interactions between all the bodies in the system, while considering the nucleus as fixed in space, due to it being thousands of times more massive than the orbiting electrons.

Essentially, and in atomic units:

$$\underbrace{\sum_i^N \left( \overbrace{\frac{1}{2} \nabla_i^2}^{E_1} - \overbrace{\frac{Z}{r_i}}^{E_2} \right)}_{\text{Individual Hamiltonian}} + \underbrace{\sum_{i<j}^j \frac{1}{r_{ij}}}_{\text{Pair repulsion}} \quad E_3$$

$$E_1 \rightarrow \text{Momentum} \quad E_2 \rightarrow e^- \text{ nuc. Coulomb attraction} \quad E_3 \rightarrow e^- e^- \text{ Coulomb repulsion} \quad (1.1)$$

### 1.3.2 The Hartree-Fock Method

This numerical method is one of the staple and most enduring procedures for solving the problem associated with a many-body system. As a side note, an in-depth explanation on all the intricacies of this method can be found in many of the literature, but as for the writing of this thesis, the works [3–6] were the ones consulted.

Hartree developed an iterative method, further enhanced by Fock and Slater, based on the field's [self consistency method](#). In this approach, when studying a multi-electronic system, such as an atom, each electron's wavefunction is composed as a product of a spacial part,  $\psi$ , and one related to the electron's spin,  $\chi$ , as to be able to account for Pauli's exclusion principle and, if so desired, relativistic effects.

$$u = \psi \chi \quad (1.2)$$

The wavefunction capable of describing the whole system,  $\Psi$ , should be somewhat of a product of all the wavefunctions describing each individual electron. However, one must not forget the need for this wavefunction to respect the antisymmetry principle, due to the electron's fermionic nature. In order to achieve this,  $\Psi$  is to be composed of a Slater determinant:

$$\Psi = \frac{1}{\sqrt{N!}} \begin{vmatrix} u_1(x_1) & u_2(x_1) & \cdots & u_N(x_1) \\ u_1(x_2) & u_2(x_2) & \cdots & u_N(x_2) \\ \vdots & \vdots & \ddots & \vdots \\ u_1(x_N) & u_2(x_N) & \cdots & u_N(x_N) \end{vmatrix} \quad (1.3)$$

Explain anti-symmetry and put it in the glossary

Explain in the glossary

It is of high importance that the set of basis wavefunctions respects orthonormality. These are to be initialized as trial wavefunctions for the numerical method.

The main goal for this algorithm is to, as per the [self consistency method](#), follow the variational principle with the goal of minimize a functional, as a way of reaching an energetic minimum. This optimal, yet unknown energy,  $E_0$  (calculated by operating the Hamiltonian on the system's optimal wavefunctions), will attempt to be reached by the variation of the trial wavefunctions that provide a non-minimized solution. Whilst the optimal wavefunctions will never be reached, the purpose of this method is to reach a



solution that yields an energetic value as close as possible to  $E_0$ , while always yielding  $\langle \Psi | H | \Psi \rangle \geq E_0$ , due to the numerical nature of the method.

The computational method consists on starting with the previously mentioned trial wavefunctions and employing them in the Hartree-Fock's potential calculation through the [Hartree-Fock \(HF\)](#) equations.

In an over-simplistic scheme, the self-consistent Hartree-Fock computational method can be represented by the block diagram in Figure 1.4.

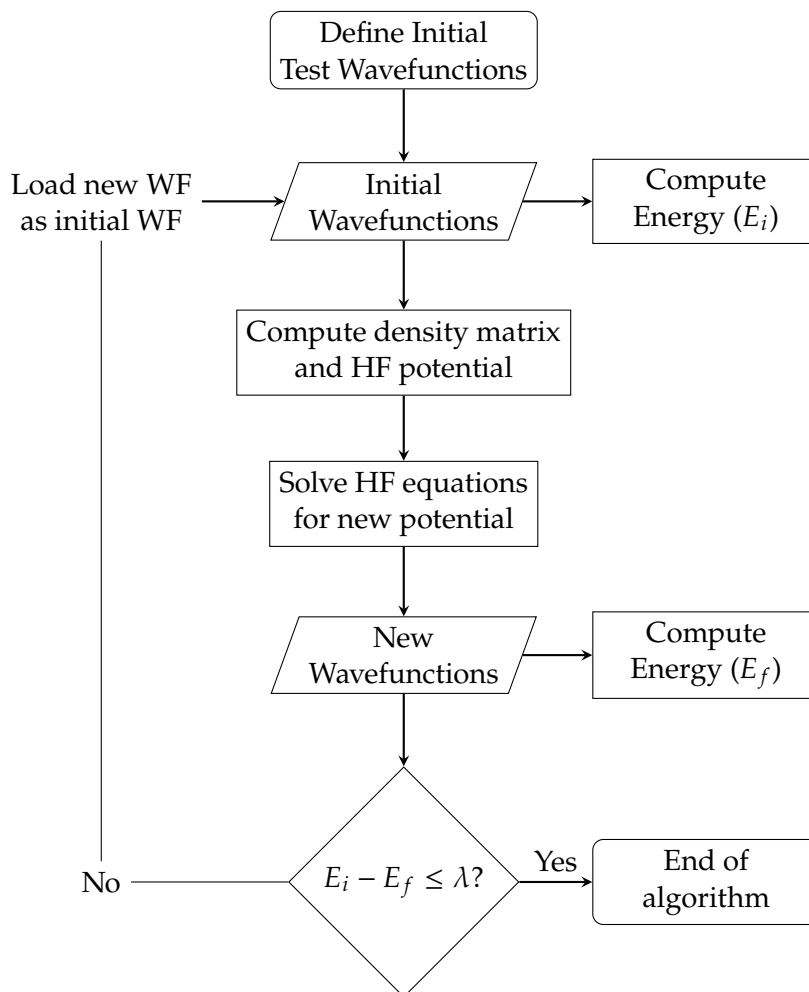


Figure 1.4: HF method's block diagram.

### 1.3.3 The Dirac Equation

While Schrödinger's equation may be one of the most significant and impactful equations in Modern Physics, it is also not free of its limitations. The fact that it does not account for the existence of the electron's spin and the lack of consideration of relativistic effects are some of the most impactful setbacks. Consulting the works in [7–10], this work will follow an exploration of different approaches that followed in order to solve the aforementioned problem.

Many scientists, such as Klein, Gordon and later Fock, had already conceived a relativistic correction to Schrödinger's equation, where the free-particle energy makes use of the relativistic momentum-energy relation, as displayed in equation (1.4).

$$E = \sqrt{c^2 \mathbf{p}^2 + m^2 c^4} \quad (1.4)$$

This notable definition can be derived from the **Lorentz invariant** scalar produced by **Minkowski norm** of the momentum **four-vector** (1.5).

$$p^\mu p_\mu = m^2 c^2 \Leftrightarrow \frac{E^2}{c^2} - \mathbf{p}^2 = m^2 c^2 \Leftrightarrow \frac{E^2}{c^2} = \mathbf{p}^2 + m^2 c^2 \quad (1.5)$$

Now, inputting this new energy operator into Schrödinger's equation, yields the Klein-Gordon equation (1.6), allowing for Schrödinger's equation to now be Lorentz-invariant (proof can be found in Strickland's book [11]).

$$-\hbar^2 \frac{\partial^2}{\partial t^2} \psi = (-c^2 \hbar^2 \nabla^2 + m^2 c^4) \psi \quad (1.6)$$

This new approach was, however, still limited, as it only described spin 0 particles (e.g., some mesons), and made use of a second order derivative in the time-like component.

Came 1928 and a new equation was developed by Paul Dirac [12], one taking now into account not the classical 3 dimensional space components, but the relativistic four components.

Dirac started by rewriting the energy-momentum relation, ending up with an equivalent equation (1.7), employing  $4 \times 4$  matrices, due to the 4 relativistic dimensions at play, and incorporating spins into the equation by making use of the Pauli matrices (1.9).

$$E = c \boldsymbol{\alpha} \cdot \mathbf{p} + \beta m c^2, \quad \boldsymbol{\alpha} = (\alpha_1, \alpha_2, \alpha_3) \quad (1.7)$$

$$\alpha_i = \begin{pmatrix} 0 & \sigma_i \\ \sigma_i & 0 \end{pmatrix} \quad \beta = \begin{pmatrix} I_2 & 0 \\ 0 & -I_2 \end{pmatrix} \quad I_2 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \quad (1.8)$$

$$\sigma_1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad \sigma_2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \quad \sigma_3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \quad (1.9)$$

In order to fully comprehend this shift of notation, one should equate the square of the two equations, (1.4) and (1.7), and confirm its validity.

$$c^2 \mathbf{p}^2 + m^2 c^4 = c^2 \boldsymbol{\alpha}^2 \mathbf{p}^2 + 2 m c^3 \boldsymbol{\alpha} \cdot \mathbf{p} + \beta^2 m^2 c^4 \quad (1.10)$$

In order for this equation to make sense, the following conditions must be true (which in fact, they are):

Try to explain why this might be a problem

do the demonstration, dummy

$$\begin{cases} c^2 p^2 = c^2 \alpha^2 p^2 & \Leftrightarrow \alpha^2 = 1 \\ 0 = 2mc^3 p \alpha \beta & \Leftrightarrow \alpha \beta = 0 \\ m^2 c^4 = \beta^2 m^2 c^4 & \Leftrightarrow \beta^2 = 1 \end{cases} \quad (1.11)$$

Taking the previous considerations into account, one can now construct Dirac's free-particle equation (1.12):

$$i\hbar \frac{\partial}{\partial t} \psi = (c \alpha \cdot p + \beta mc^2) \psi = \begin{pmatrix} mc^2 I_2 & -i\hbar c \sigma \cdot \nabla \\ -i\hbar c \sigma \cdot \nabla & -mc^2 I_2 \end{pmatrix} \cdot \begin{pmatrix} \psi_1 \\ \vdots \\ \psi_4 \end{pmatrix} \quad (1.12)$$

This equation, however, as mentioned above, can only describe a single particle present in a field-free region. In order to account for the existence of a field, such as the electromagnetic field, derived from the four-potential  $A^\mu$ , composed by the electric scalar potential field,  $A^0 = \phi$ , and the vector potential,  $(A^1, A^2, A^3) = \mathbf{A}$ , the following change on the momentum four-vector must be made:

$$p^\mu \rightarrow p^\mu - eA^\mu, \quad A^\mu = (\phi, \mathbf{A}) \quad (1.13)$$

The Hamiltonian can now be rewritten as to account for the presence of the electromagnetic field (1.14). This way it is possible to include, for example, the electron-nucleus Coulomb attraction.

$$H_D = -e\phi + \beta mc^2 + \alpha(cp + eA) \quad (1.14)$$

For a central fixed potential, as is the one generated by the nuclear charge, the 3 space-like components from the four-potential are null, and the time-like component,  $\phi = \frac{Ze}{r}$ . The Hamiltonian gains now a more recognizable form:

$$H_D = -\frac{e^2 Z}{r} + \beta mc^2 + \alpha \cdot pc \quad (1.15)$$

A very interesting fact about Dirac's equation is that it yields, in fact, two sets of solutions: the large component (positive energy values), for particles, and the small component (negative energy values), for antiparticles.

### 1.3.3.1 The Dirac-Breit Equation

Once again, when considering a system composed of many bodies, one must consider all the present interactions, namely, the electron-electron repulsion in an atom. Breit, in 1929, had created a relativistic approach to treat the electron-electron interactions, consisting on a set of equations building upon the classical non-relativistic Hamiltonian from equation (1.1), which can be consulted in Appendix A. Breit's equations are able to

account for angular momenta couplings and estimate level energy splittings, the change of a particle's apparent mass as a function of velocity, and even include the interaction of an applied external magnetic field [10].

It is quite obvious Breit's equations introduce a great complexity in the search of the new Hamiltonian's eigenfunctions. Nonetheless, when trying to include an approximation of Breit's considerations into Dirac's equation, one must add the following operator to the one present in equation (1.15):

$$H_B = \sum_{i>j} \frac{e^2}{r_{ij}} - e^2 \left( \frac{\alpha_i \alpha_j}{r_{ij}} + \frac{(\alpha_i \nabla_i)(\alpha_j \nabla_j) r_{ij}}{2} \right) \quad (1.16)$$

This set of terms will account for the fact that Coulomb interactions, mediated by the electric field, and therefore, **virtual photons**, cannot occur at instantaneous velocities, but at the speed of light.

## 1.4 QED considerations

Lorem ipsum dolor sit amet, consectetur adipiscing elit. Etiam ut justo justo. Cras pulvinar massa sollicitudin ligula faucibus, in vulputate magna viverra. Nulla risus ante, maximus ut nunc sit amet, condimentum convallis mi. Phasellus et mi aliquet, ornare sapien egestas, vehicula orci. Donec ac massa tempus, iaculis tortor vitae, tempus tortor. Donec nec justo eros. Sed mauris purus, facilisis eu tempus quis, mattis et erat. Sed et congue metus, in venenatis nibh. Nullam ipsum ex, scelerisque non nunc sit amet, ultrices commodo felis. Proin sodales turpis nulla, quis tincidunt leo auctor id. Ut ac nulla quis felis molestie ullamcorper. Nulla tristique dui velit, sit amet rutrum libero dignissim id.

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Should the atomic system be subjected to an external **EMF**, the level structure would change due to Zeeman's and Stark's effects, allowing for the probing of the atom's hyperfine structure.

### 1.4.1 The **MCDF** Method

As previously mentioned in section 1.3.2, there is a need for a numerical method in order to compute and find the eigenfunctions for a many-body Hamiltonian. While the Hartree-Fock method was able to reasonably solve the non-relativistic problem, now,

Falar aqui sobre as contribuições. Não tem que ser extensivo

Ler bem o manual e explicar bem o método e todas as coisas diferentes que se podem fazer

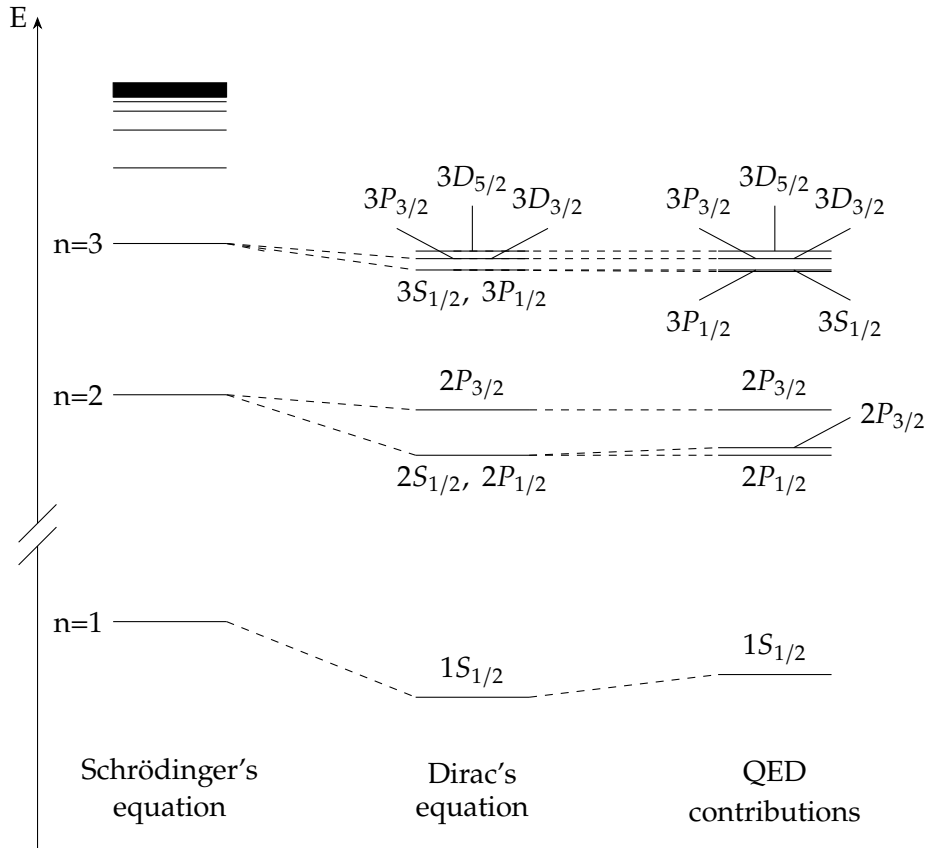


Figure 1.5: Comparison of Hydrogen level diagrams. Adapted from [13]

while considering the Dirac-Breit Hamiltonian from equations (1.15) and (1.16), there is a need for a new method.

Hence, the state of the art *MCDFGME* arises. This self-consistent iterative method, based on the same method present in section 1.3.2, is able to solve and find eigenfunctions for a multielectronic system, now taking into account the Dirac-Breit Hamiltonian. Moreover, it is also capable of incorporating electron correlation and many QED effects not yet considered in the relativistic equation, such as the Lamb-shift, vacuum polarization, and the electron's self energy. A brief description of these contributions can be found in appendix C

## 1.5 State of the Art

### 1.5.1 Copper's characteristic X-rays

Copper is a dominant element in today's society. While most of its uses are day-to-day related, it also has a high prevalence in many physical areas, namely, copper's  $K_\alpha$  transitions [14]. While these radiative transitions have been measured countless times, with very well recorded energy values [15–19]. It is also common knowledge Copper's  $K_\alpha$

This really  
needs a re-do

lines do not have a symmetric distribution, since both  $K_{\alpha_1}$  and  $K_{\alpha_2}$  line display a negative (left-tailed) skewness. Due to this fact, most of the fitting models used in order to analyze both Copper's  $K_{\alpha_1}$  and  $K_{\alpha_2}$  transitions involve the usage of a Lorentzian doublet, for each line [20, 21]. While many associate these asymmetries due to satellite states formed by shake processes [14, 22], with some theoretical studies having been performed [23], some authors note it could be due to X-ray resonant Raman scattering [22]. This effect occurs when a sample is exposed to energies under to near the ionization threshold, when a bound electron is excited to an upper state [24]. While there are some studies exploring this topic, most are focused on the cascade of low energy transitions that follow the post-scattering excitation [25]. Nonetheless, a previous experimental study has been able to show that for Copper exposed to synchrotron radiation tuned to energies near K-shell's ionization energy, some  $K_{\alpha_1}$  transitions demonstrated to be narrower than expected [26].

### 1.5.2 MCDFGME capabilities

It has been noted multiple times that *MCDFGME* code excels in atomic structure calculations of super-heavy elements and highly-charged ions, where relativistic and QED effects are in prevalence [27–29]. However, it has also been proven to be an excellent tool for the calculation of less ionized and lighter atomic systems [30].

In addition, the *MCDFGME* code is able to calculate radiative and auger transition rates for the calculated configurations, which can be used in the simulation of theoretical spectra, due to being able to compute the transition's intensity and natural width. Since it is able to perform calculations, even for exotic atoms, it can be used to further understand many QED phenomena, further exploring the limits of our theory, and it's comparison to experimental data [31].

It should also be of note that there are many other code alternatives. While *MCDFGME*, which is a close-source project, provides a very high precision in the performed calculations at a high computational cost, *Flexible Atomic Code (FAC)*, is an open source code which requires less computational time for the calculations, however, it lacks *MCDFGME*'s precision, since it only is able to consider all the spin-orb couplings, but does not mix the possible configurations originating them. It can, however, calculate other collisional processes, such as electronic impact excitation cross-sections [32]. *General-purpose Relativistic Atomic Structure Package 2k (Grasp2k)* [33], and *AUTOSTRUCTURE* [34] are other codes with some of the same capabilities.

# Atomic Structure Calculations

In this chapter, the procedure that follows a standard atomic structure calculation will be discussed and explained in detail. Topics ranging from the usage of the *MCDFGME* code to compute quantities such as energy levels, orbital wavefunctions and transition rates, to the manner in which these parameters can be used in order to simulate a theoretical spectrum will be explored. All the information present in this chapter was obtained after a thorough study of *MCDFGME*'s manual [35].

## 2.1 The *MCDFGME* code's capabilities

As previously stated, *MCDFGME* is a program that allows for not only solving the many-body problem for an atomic system while making the proper QED energy corrections, but also for the computation of a great deal of atomic parameters. Consulting [36], one can see that these include, but are not limited to:

- Energy level calculations.
- Multipole radiative transition probabilities.
- Auger transition probabilities.
- Photoionization cross-sections.
- Electronic impact excitation cross-sections.
- Orbital wavefunction overlaps between same and different atomic systems.

## 2.2 The atomic system at study

Before proceeding with the explanation behind every step of the calculations performed, a previous discussion on what the objective for the following calculations are should be had.

### 2.2.1 The system at study

The main purpose of this thesis is that of simulating a theoretical spectrum for Copper's emission lines when subjected to a near ionization threshold x-ray source.

At this energy range, two main processes will be responsible for an electron moving out a core-shell: Resonant photoexcitation, and ionization.

While the simulation of the theoretical spectra for ionized Copper would be quite straightforward (due to the low shake probabilities at near ionization threshold energies, transitions for satellite states were not considered), the more extensive calculation is that of the resonant photoexcitation.

In order to fulfill the aim of simulating the theoretical spectra for this last phenomenon, multiple atomic structure calculations were computed for the first excited state configurations for Copper.

## 2.3 Level Calculations

The first step needed to be performed in order to simulate theoretical spectra is the calculation of the level structure for all given configurations.

Breit can be used as self consistent or perturbative

## 2.4 Transition computations

### 2.4.1 Diagram transitions

### 2.4.2 Auger transitions

### 2.4.3 Satellite transitions

#### 2.4.3.1 Rate Matrices as a calculation quality comparison tool



# A new High-Performance Computing code for the parallelization of atomic structure calculations

## 3.1 A brief introduction on *MPI*

## 3.2 Program advantages

## 3.3 Program limitations

Blocking communication

CPU overheat

No buffering for now

# Fundamental atomic parameters calculation

asdas

## Spectra simulation

asdas

## **Spectra analysis**

### **6.1 Photoexcitation cross-section estimation**

### **6.2 Photoionization cross-section computation**

## Comparison with experimental data

## Next Steps

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## Final remarks and Conclusion

Usar o Deni-  
tions aqui

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### 9.1 Quantum states and properties

### 9.2 State transitions

#### 9.2.1 Transition rates and widths

#### 9.2.2 Branching Ratios and Fluorescence Yields

### 9.3 Spectra simulation

## Spectra Analysis



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## Development of a HPC script for the parallelization of ASC

### 11.1 blablabla

### 11.2 Speedup

Compare with  
Amdahl's law  
or Gustafson's  
law.

Comparação  
com o do Jorge

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## The Breit Hamiltonian Operators

Note: This operators are valid for the electrons in an atom.

The free particle energy:

$$H_0 = \sum_i^N \frac{p_i^2}{2m_e} \quad (\text{A.1})$$

The electron-nucleus Coulomb attraction:

$$H_1 = \sum_i^N -\frac{e^2 Z}{r_i} \quad (\text{A.2})$$

The electron-electron Coulomb repulsion:

$$H_2 = \sum_{i < j} \frac{e^2}{r_{ij}} \quad (\text{A.3})$$

Incorporates the relativistic apparent mass - velocity dependance:

$$H_3 = -\frac{1}{8m_e^3 c^2} \sum_i^N p_i^4 \quad (\text{A.4})$$

Electric field retardation and magnetic dipole interaction:

$$H_4 = -\frac{e^2}{2m_e^2 c^2} \sum_{i < j} \left[ \frac{\mathbf{p}_i \cdot \mathbf{p}_j}{r_{ij}} + \frac{(\mathbf{r}_{ij} \cdot \mathbf{p}_{ij})(\mathbf{r}_{ij} \cdot \mathbf{p}_j)}{r_{ij}^3} \right] \quad (\text{A.5})$$

Darwin's term, accounts for the electron's quantum fluctuation motion:

$$H_5 = \frac{\pi e \hbar^2}{2m_e^2 c^2} \sum_{i < j} \frac{Z}{2} [\delta(\mathbf{r}_i) + \delta(\mathbf{r}_j)] + \delta(\mathbf{r}_{ij}) \quad (\text{A.6})$$

And the last two operators, for the consideration of spin orbit interactions:

$$H_6 = \frac{e^2 \hbar^2}{2m_e^2 c^2} \sum_{i < j} \left( Z \frac{\mathbf{r}_i \times \mathbf{p}_i}{r_i^3} - \frac{\mathbf{r}_{ij} \times \mathbf{p}_i}{r_{ij}^3} + 2 \frac{\mathbf{r}_{ij} \times \mathbf{p}_j}{r_{ij}^3} \right) \mathbf{s}_i + \left( Z \frac{\mathbf{r}_j \times \mathbf{p}_j}{r_j^3} - \frac{\mathbf{r}_{ji} \times \mathbf{p}_j}{r_{ij}^3} + 2 \frac{\mathbf{r}_{ji} \times \mathbf{p}_i}{r_{ij}^3} \right) \mathbf{s}_j \quad (\text{A.7})$$

$$H_7 = \frac{e^2 \hbar^2}{m_e^2 c^2} \sum_{i < j} \left( -\frac{8\pi}{3} \mathbf{s}_i \cdot \mathbf{s}_j \delta(\mathbf{r}_{ij}) + \frac{1}{r_{ij}^3} \left[ \mathbf{s}_i \cdot \mathbf{s}_j - \frac{3(\mathbf{s}_i \cdot \mathbf{r}_{ij})(\mathbf{s}_j \cdot \mathbf{r}_{ij})}{r_{ij}^2} \right] \right) \quad (\text{A.8})$$

## Transition Diagram

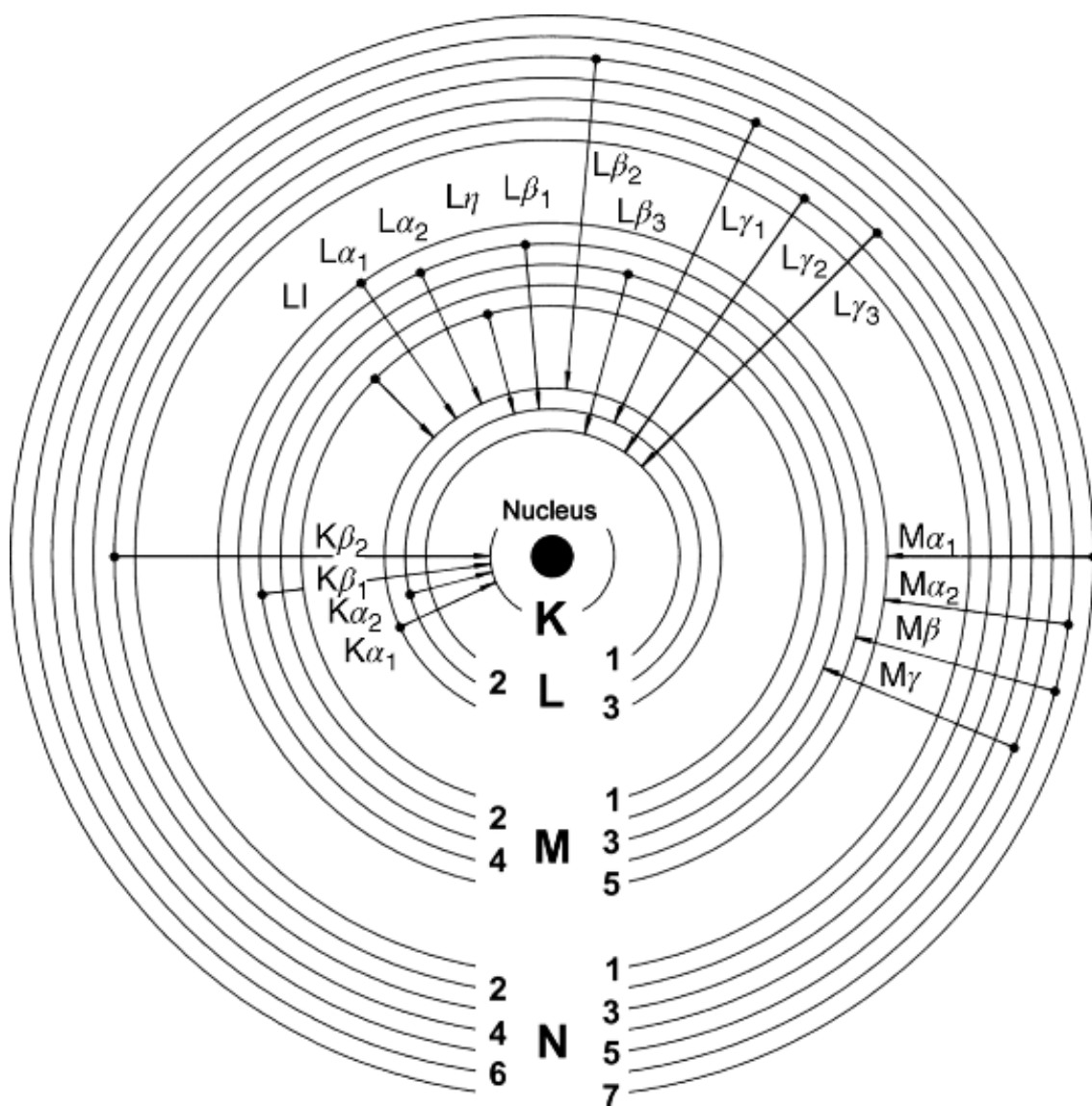


Figure B.1: Transition notation scheme. Adapted from [2]



## QED considerations

It should be apparent by now that studying Atomic systems call not only for relativistic effects and corrections, but also for QED ones.

One of the most famous cases where QED came to light was the discovery of the Lamb Shift [37], when it was discovered Hydrogen's  $2s_{1/2}$  and  $2p_{1/2}$  levels were in fact, not degenerate (did not have the same energy), contrary to what was expected from solving Dirac's equation. This difference in energy came to be known as the Lamb Shift, only explained by QED effects.

### C.1 Self-Energy

The self-energy represents a particle's emission and reabsorption of virtual photon, present in the particle's own generated field. This interaction has the most impact in the Lamb Shift effect and when performing energy corrections. One of its [Feynman Diagrams](#) can be seen in Figure C.1a.

### C.2 Vacuum Polarization

As previously stated, electromagnetic fields, such as the Coulomb field generated by the nucleus, are mediated by virtual photons. These photons can lead to the creation of electron-positron pairs which create screening effects. Pair annihilation will follow, leading to the production of another virtual photon (Figure C.1b).



Figure C.1: QED Feynman Diagrams

## Input File .f05 Example

```

1  program_year=2019 program_version=1
2  * 1 Z=29 (1s)1 (2s)2 (2p)6 (3s)2 (3p)6 (3d)10 (4s)2 2J=1 neig=[2S1/2]#1
   LS1
3      scfmdf max :
4      mod_lightspeed=n
5      nz=29
6      mdf opt_ener=todo modfilename_ener=n modfilename_wf=n do_scf=y
7      Breit=full mag_scf=y ret_scf=y
8      qedstpg_n4=n
9      vacpol_scf=y
10     energy
11     # use_mcdfener=y
12     # opt_relax=y
13     ret_Lorentz=y
14     opt_qedel=y :
15     # ":" above is for ilams taken to be the default value
16     mod_mesh=n
17     # hx=0.012 r(1)=0.001 amesh=0.01 :
18     exotic=n
19     use_nms=y
20     mod_nuc=n
21     project=n
22     # cgt_order_vint=y order=6
23     cgt_order_vint=n
24     def_config=given
25     nbel=29 jjt=1 :
26     c 1 (1s)1 (2s)2 (2p)6 (3s)2 (3p)6 (3d)10 (4s)2 :
27     end
28     # initial state parameters
29     neigv=1 icmul=0 iprfgr=0
30     norbsc=00 ndep=0 nlec=0 nec=1 :
31     nstep=0
32     lregul=n modtest=n
33     modsolv_orb=n
34     mod_odlm=n

```

## ANNEX I. INPUT FILE .F05 EXAMPLE

---

```
35  # data for uwfrdf
36      :
37  **
```



