

Fractal analysis of LIGO data a.k.a. How to characterize interferometric noise in low latency

Marco Cavaglià

MISSOURI S&T

INSTITUTE FOR MULTI-MESSENGER ASTROPHYSICS AND COSMOLOGY

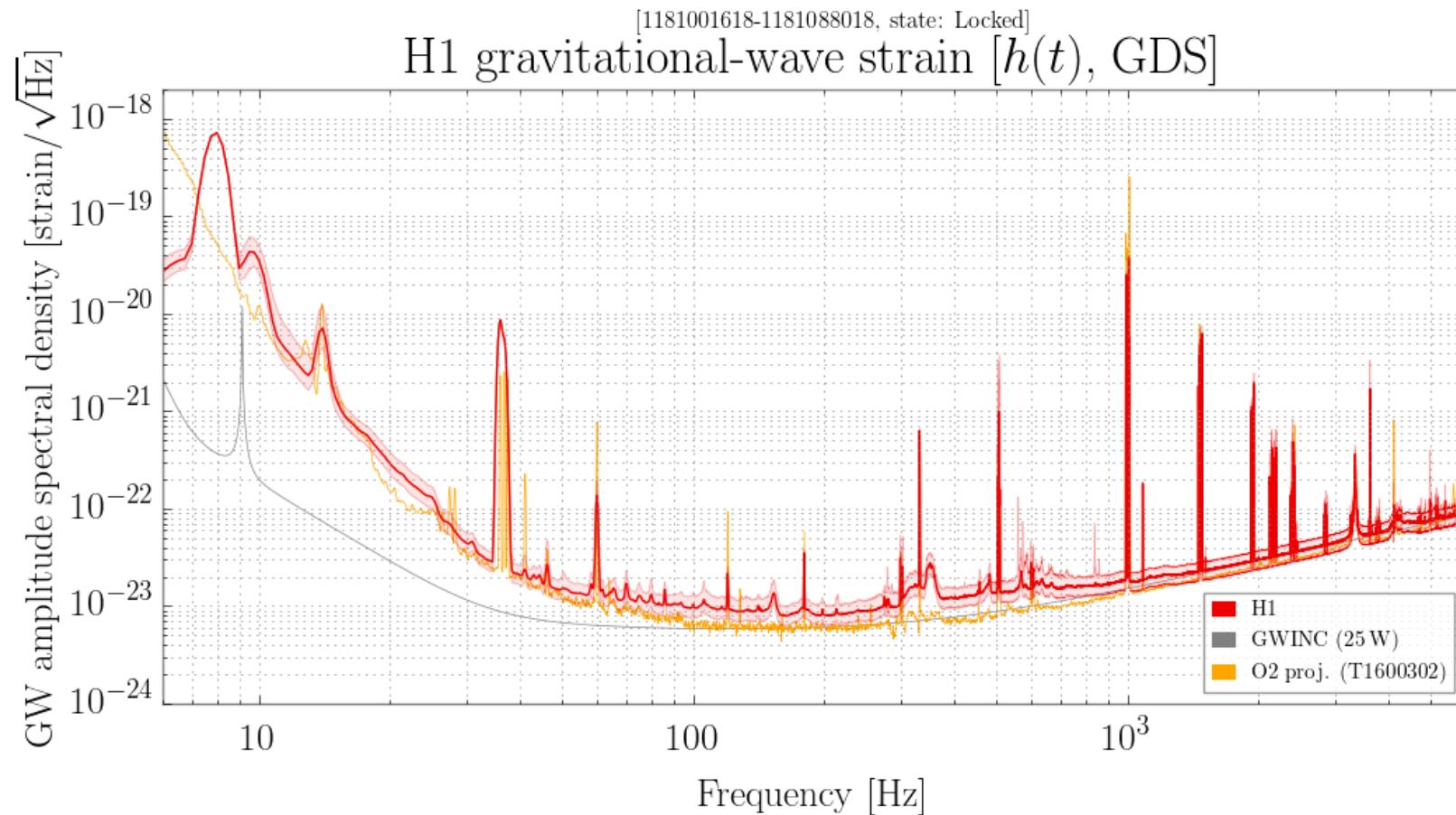
Today is Juneteenth!

JUNETEENTH

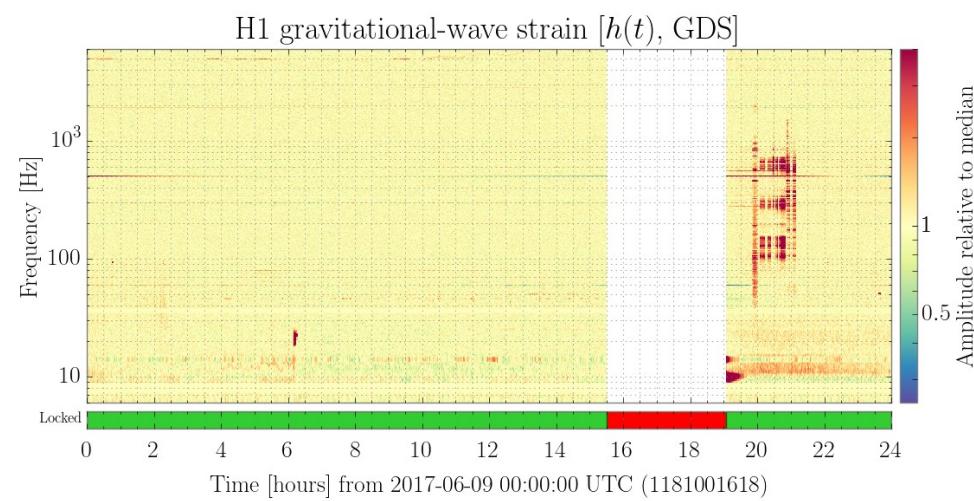
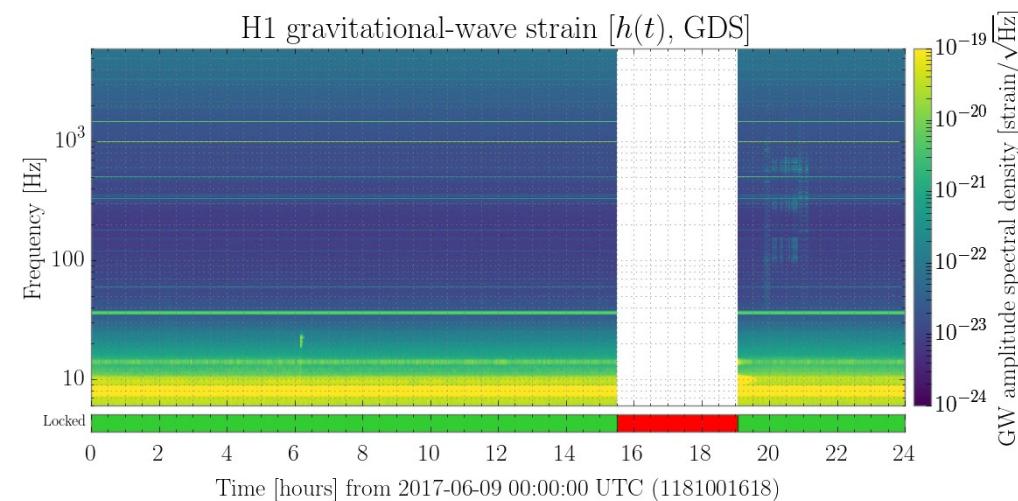
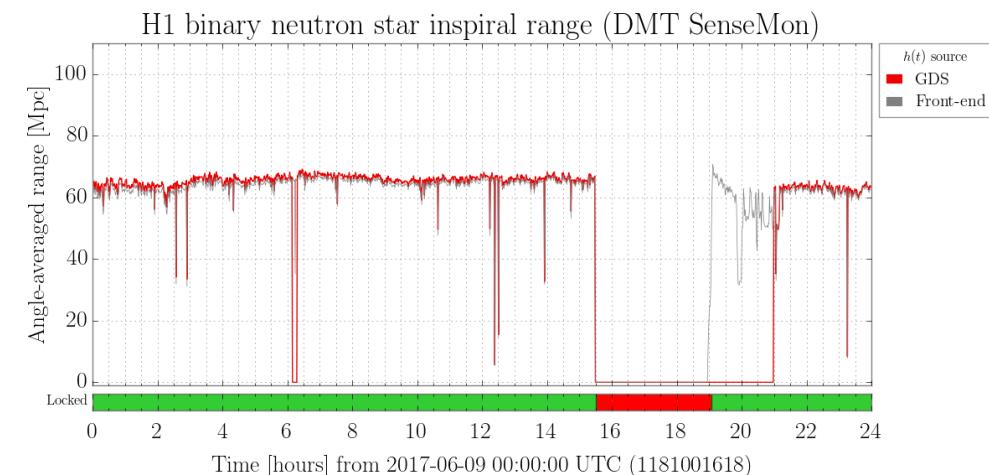
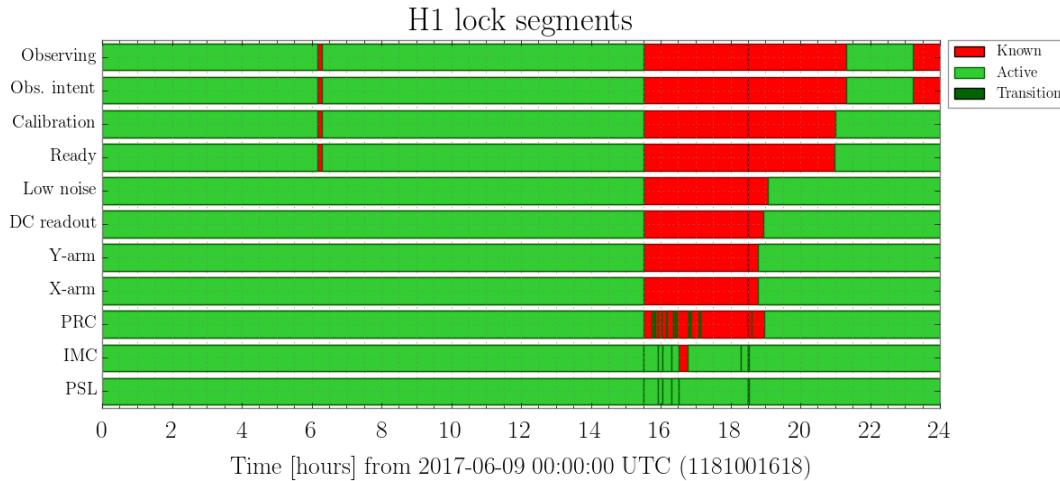


Juneteenth is the oldest nationally celebrated commemoration of the ending of slavery in the United States. Dating back to 1865, it was on June 19th that the Union soldiers, led by Major General Gordon Granger, landed at Galveston, Texas with news that the war had ended and that the enslaved were now free. It is now a federal holiday in the U.S.

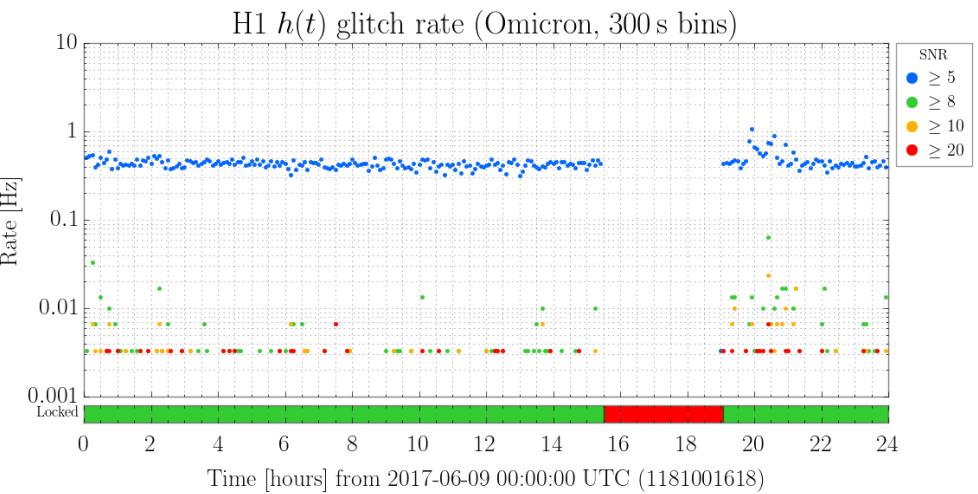
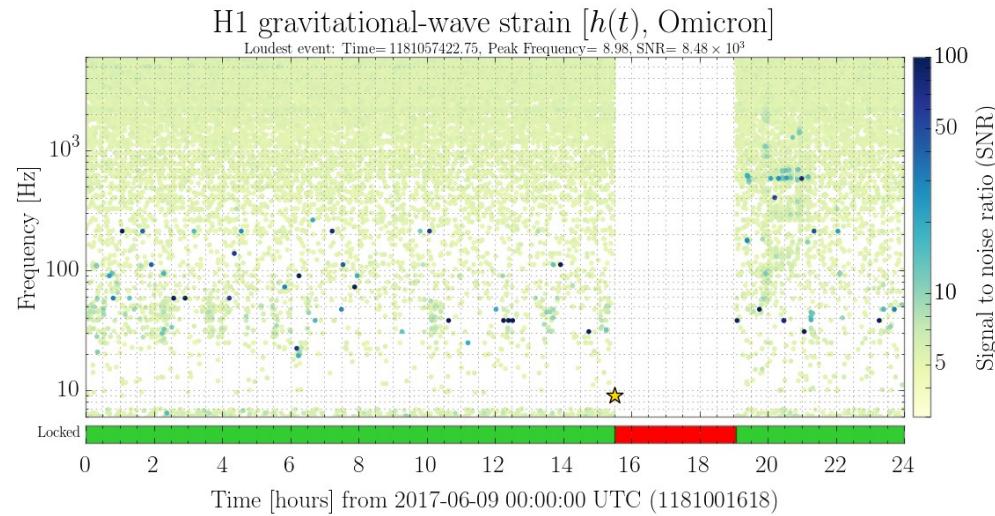
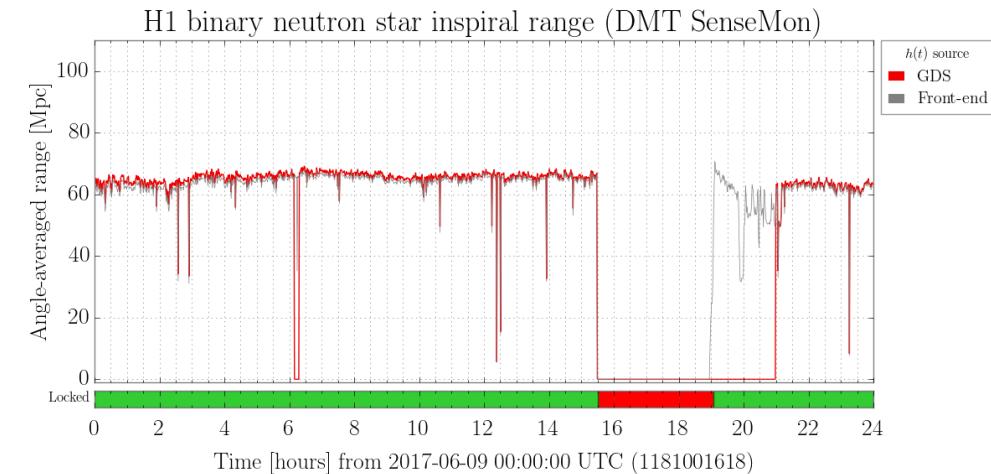
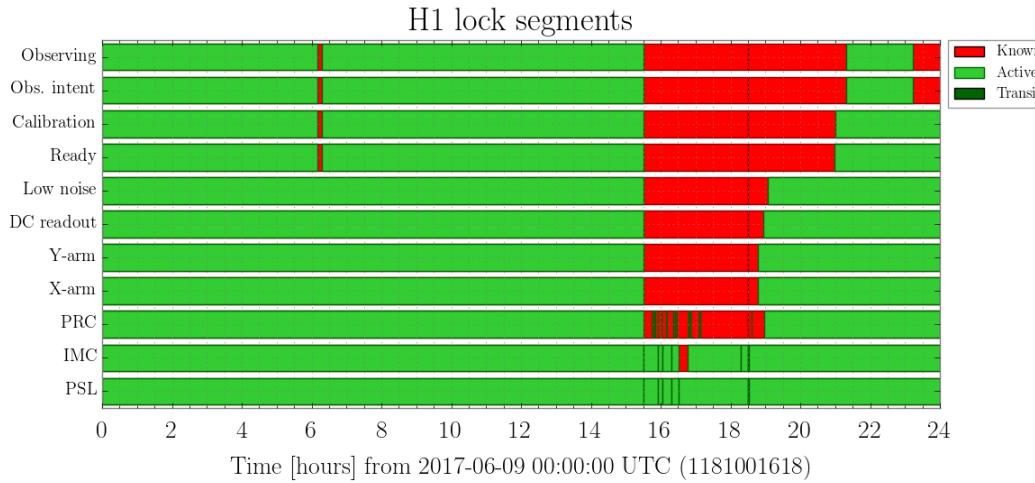
A typical day in LIGO



A typical day in LIGO

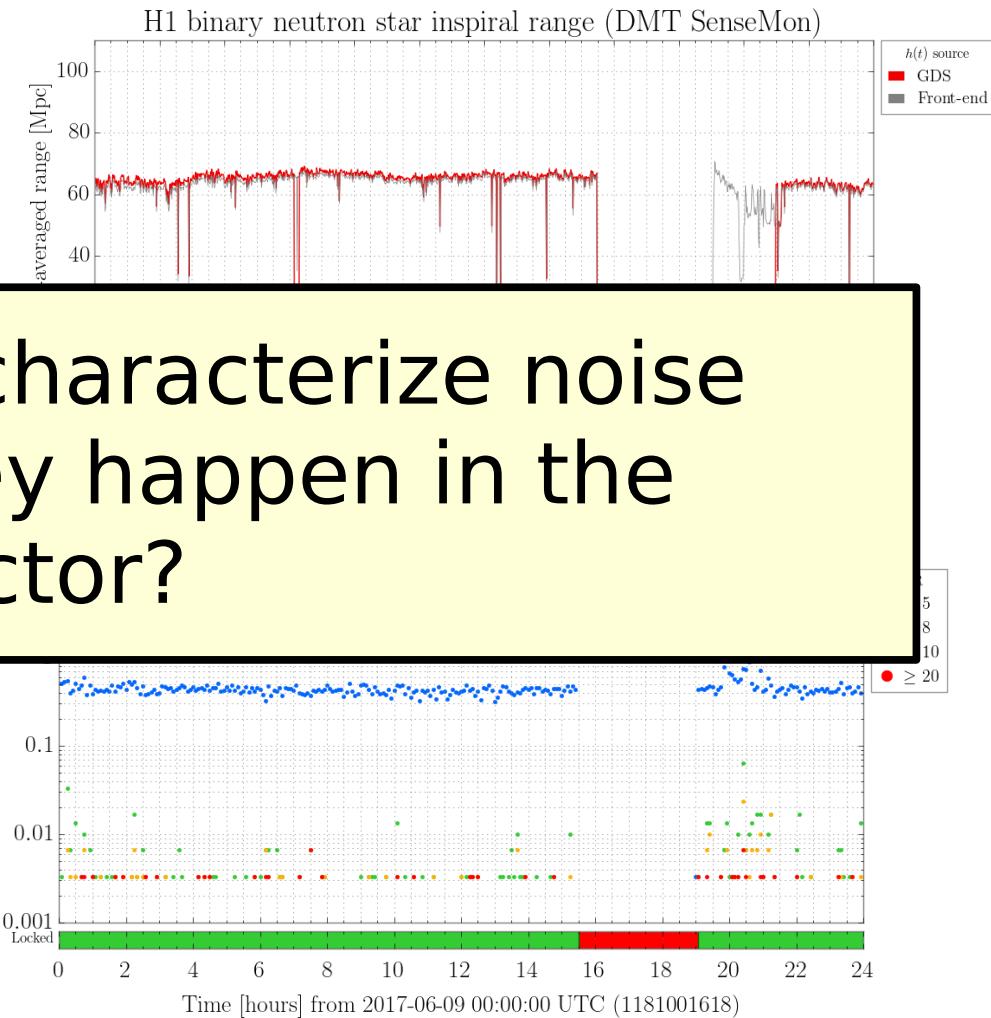
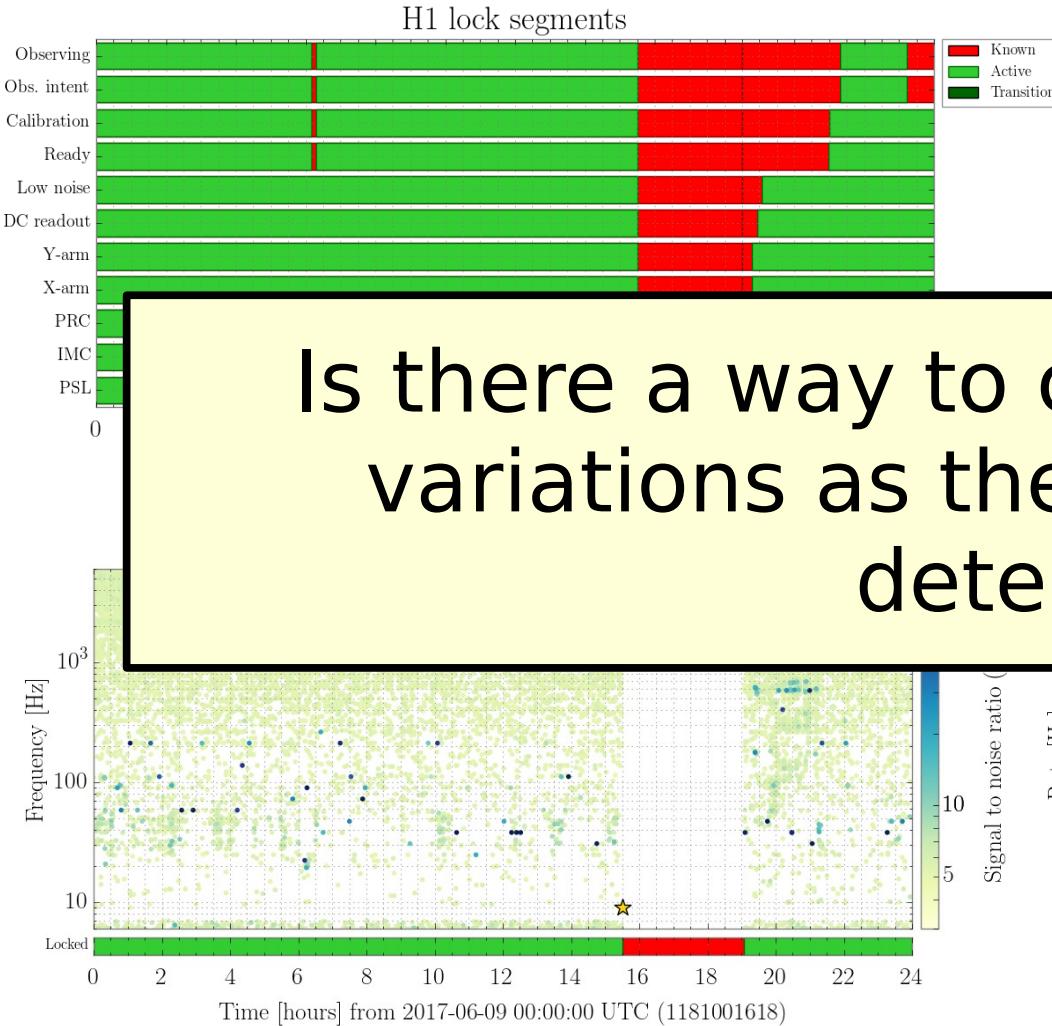


A typical day in LIGO



A typical day in LIGO

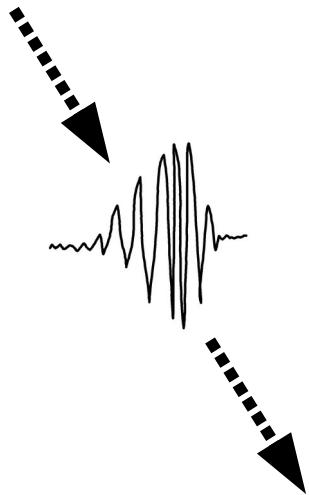
Is there a way to characterize noise variations as they happen in the detector?



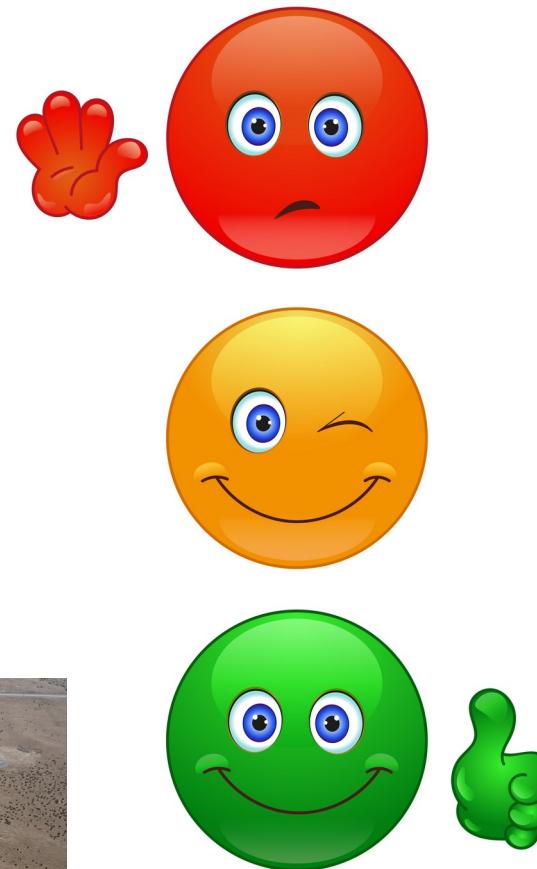
Noise monitor concept



Credit:SSU - Aurore Simonnet



Credit: LIGO Laboratory



Bad data

Please double
check...

Data OK!

Low-latency data quality assessment

Typically a lot of works goes in Data Quality assessment

| Superevent Information | |
|------------------------|---|
| Superevent ID | S200311bg <i>EM advocate says event is okay.</i> |
| Category | Production Added by: Geoffrey Mo Added: March 11, 2020, 12:15 p.m. |
| Labels | DQOK EM_READY ADVOK EM_Selected EMBRIGHT_READY PASTRO_READY SKYMAP_READY GCN_PRELIM_SENT PE_READY |
| t _{start} | 1267963150.37 |
| t ₀ | 1267963151.40 |
| t _{end} | 1267963152.44 |
| Submitted ▾ | 2020-03-11 11:59:09 UTC |
| Links | Data |

Rapid Response Team

Analyst Comments Data Quality Sky Localization External Coincidence EM Followup

Analyst Comments

Log Comment

RRT minutes as PDF, updated to include chat text. ([S200311bg RRT Minutes.pdf](#))

— Submitted by Jenne Driggers on March 11, 2020 17:04:24 UTC

Log Comment

RRT minutes: <https://docs.google.com/document/d/1TFNX3n7ZfaXmZzTuundu5rYGhmw5pFRhDtM6jcaPuYo/edit>
([S200311bg_RRT.mp3](#))

— Submitted by Geoffrey Mo on March 11, 2020 12:34:14 UTC

Log Comment

Online DQ report - H1 and L1 both look clean around the time of the event. There is some slight noise at 1.2s (below 100Hz) after the event at H1 in case anyone uses data after the event time.

— Submitted by Laura Nuttall on March 11, 2020 12:16:24 UTC

Rapid Response Team

| | Analyst Comments | Data Quality | Sky Localization | External Coincidence | EM Followup |
|---------------------------------------|---|--------------|------------------|----------------------|-------------|
| Analysis | | | | | |
| Localization | | | | | |
| Redundancy | | | | | |
| Localization | | | | | |
| Redundancy | | | | | |
| Localization | | | | | |
| Redundancy | | | | | |
| Ongoing | | | | | |
| atmospheric | | | | | |
| – | | | | | |
| Metrics for the Detchar RRT During O3 | | | | | |
| 2.1 | Did the candidate occur at a suspicious GPS time? | | | | |
| 2.1.1 | Existing items included within this item | | | | |
| 2.1.2 | Technical solutions | | | | |
| 2.2 | Is there a high probability that a glitch was present near the candidate based on statistical inference of auxiliary information? | | | | |
| 2.2.1 | Existing items included within this item | | | | |
| 2.2.2 | Technical solutions | | | | |
| 2.3 | Are known sources of noise with auxiliary witnesses active near the candidate? | | | | |
| 2.3.1 | Existing items | | | | |
| 2.3.2 | Technical solutions | | | | |
| 2.4 | Are known sources of noise without auxiliary witnesses active near the candidate? | | | | |
| 2.4.1 | Existing items | | | | |
| 2.4.2 | Technical solutions | | | | |
| 2.5 | Are environmental monitors active near the candidate? | | | | |
| 2.5.1 | Existing items | | | | |
| 2.5.2 | Technical solutions | | | | |
| 2.6 | Was the detector in a nominal state? | | | | |
| 2.6.1 | Existing items | | | | |
| 2.6.2 | Technical solutions | | | | |

Detector state

Analyst Comments Data Quality Sky Localization External Coincidence EM Followup

Metrics for the Detchar RRT During O3

| | | |
|-----------------|--|-----|
| Analysis | 2.1 Did the candidate occur at a suspicious GPS time? 2.1.1 Existing items included within this item 2.1.2 Technical solutions | |
| Localization | 2.2 Is there a high probability that a glitch was present near the candidate based on statistical inference of auxiliary information? 2.2.1 Existing items included within this item 2.2.2 Technical solutions | |
| Reference | 2.3 Are known sources of noise with auxiliary witnesses active near the candidate? 2.3.1 Existing items 2.3.2 Technical solutions | |
| Observation | 2.4 Are known sources of noise without auxiliary witnesses active near the candidate? 2.4.1 Existing items 2.4.2 Technical solutions | |
| Environmental | 2.5 Are environmental monitors active near the candidate? 2.5.1 Existing items 2.5.2 Technical solutions | +z) |
| Detector Status | 2.6 Was the detector in a nominal state? 2.6.1 Existing items 2.6.2 Technical solutions | |

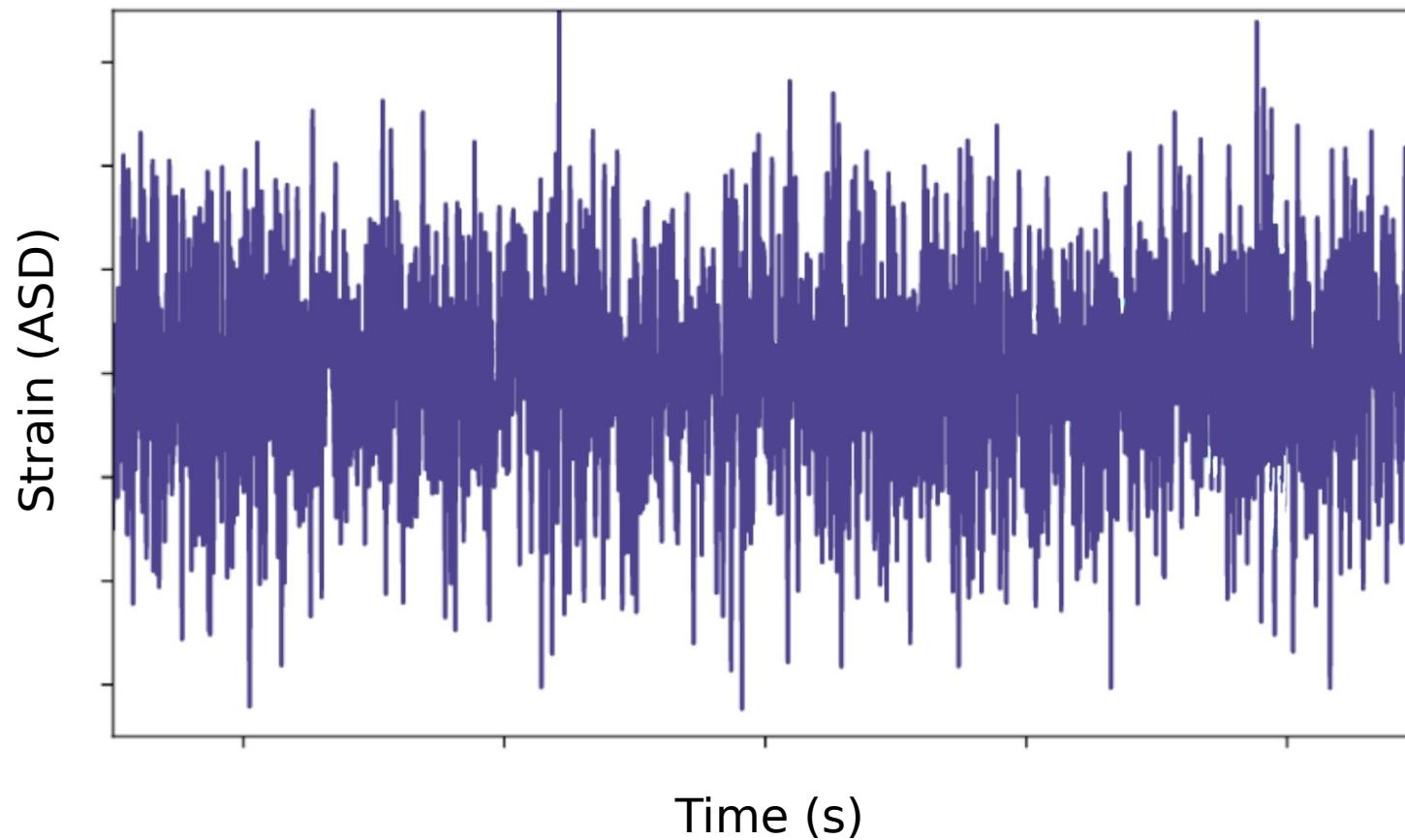
Detector state

Analyst Comments Data Quality Sky Localization External Coincidence EM Followup

Metrics for the Detchar RRT During O3

| | | |
|---------|---|---|
| Analyst | 2.1 Did the candidate occur at a suspicious time? | • Identify nearest hardware injection |
| Log | 2.1.1 Existing items included within the analysis | |
| R | 2.1.2 Technical solutions | |
| Log | 2.2 Is there a high probability that a given candidate based on statistical inference is a true signal? | • BruCO https://git.ligo.org/gabriele-vajente/bruco |
| R | 2.2.1 Existing items included within the analysis | |
| (S) | 2.2.2 Technical solutions | |
| Log | 2.3 Are known sources of noise within the detector near the candidate? | • Stochmon https://dcc.ligo.org/LIGO-T1400205/public |
| R | 2.3.1 Existing items | |
| (S) | 2.3.2 Technical solutions | |
| Log | 2.4 Are known sources of noise without near the candidate? | • Stationarity |
| R | 2.4.1 Existing items | |
| (S) | 2.4.2 Technical solutions | |
| Log | 2.5 Are environmental monitors active near the candidate? | • Check calibration kappas |
| O | 2.5.1 Existing items | |
| at | 2.5.2 Technical solutions | |
| Log | 2.6 Was the detector in a nominal state? | • Calibration state vector check |
| O | 2.6.1 Existing items | |
| at | 2.6.2 Technical solutions | |

Building a nominal state metric from strain data



Noise and fractals



Benoit Mandelbrot

“In 1961 IBM was involved in transmitting computer data over phone lines, but a kind of white noise kept disturbing the flow of information—breaking the signal—and IBM looked to Mandelbrot to provide a new perspective on the problem.”

<https://www.ibm.com/ibm/history/ibm100/us/en/icons/fractal/>

Fractal analysis

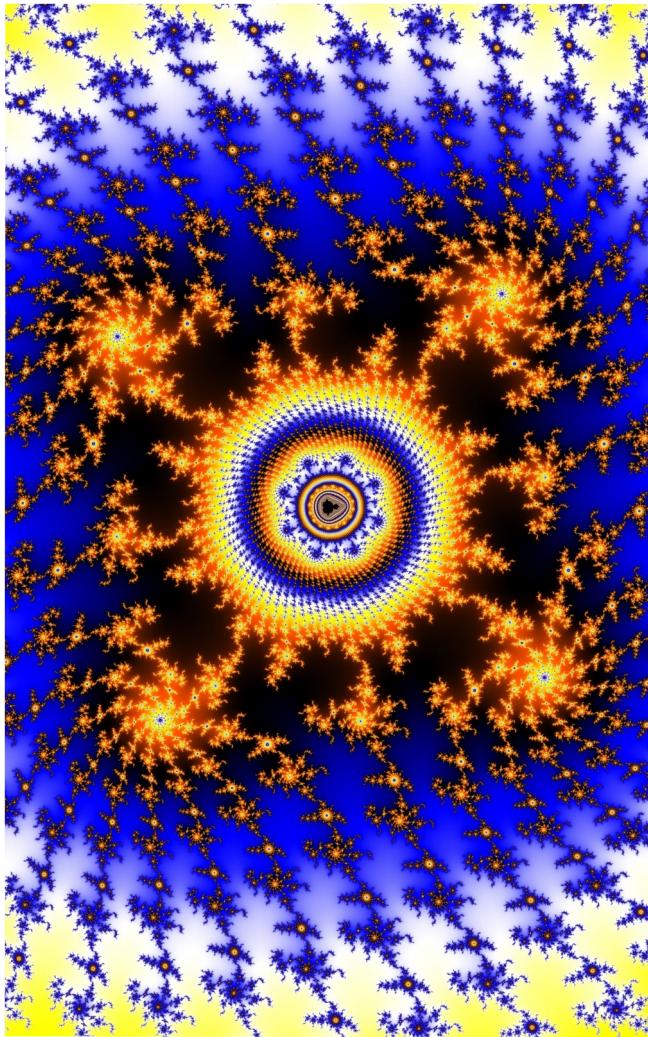


Figure credit: <https://wallup.net/>

- Fractal analysis can be used to characterize the degree of complexity of a set.
- The concept has been applied to different physical phenomena met in various fields from materials science to chemistry, biology, etc.

Topological dimension of a set

A topological space is an ordered pair (X, τ) , where X is a set and τ is a collection of subsets of X , satisfying the following axioms:

- The empty set and X itself belong to τ .
- Any arbitrary (finite or infinite) union of members of τ belongs to τ .
- The intersection of any finite number of members of τ belongs to τ .

The elements of τ are called open sets and the collection τ is called a topology on X .

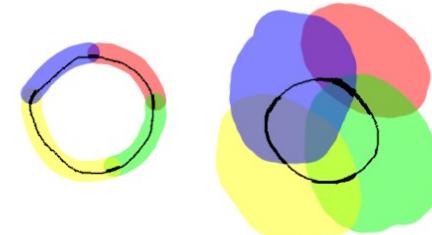
A topological space can be covered by open sets. The topological (Lebesgue covering) dimension is the smallest number n such that for every cover, there is a refinement in which every point in X lies in the intersection of no more than $n + 1$ covering sets. The topological dimension is an integer number that does not change as the space is continuously deformed under an homeomorphism.

M. Yamaguti et al. "Mathematics of Fractal" Mathematical monographs, vol.167 AMS 1997s

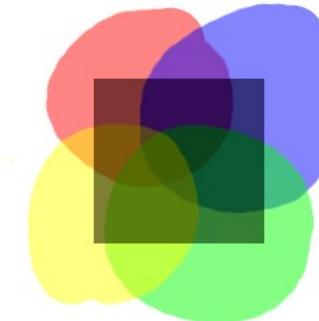
Topological dimension of a set

Examples:

A circle requires a set of two (or more) open arcs to be covered. The topological dimension of the circle is one.



A disk requires at least three open sets to be covered. The topological dimension of the disk is two.



In general an n -dimensional Euclidean space has topological dimension n .



Fractal dimension

A fractal is a subset of an Euclidean space with a fractal dimension that strictly exceeds its topological dimension.

The fractal dimension D is defined with the number of covering areas necessary to cover the fractal structure F .

Box-counting

$$D = \frac{\log [N(\epsilon)]}{\log [1/\epsilon]}$$

ϵ = radius of the covering balls

Hausdorff dimension

$$D = \sup \{ d \in \mathbb{R}^+ : m_d(F) = \infty \}$$

m_d = s -dimensional Hausdorff-outer measure

D is a measure of roughness.

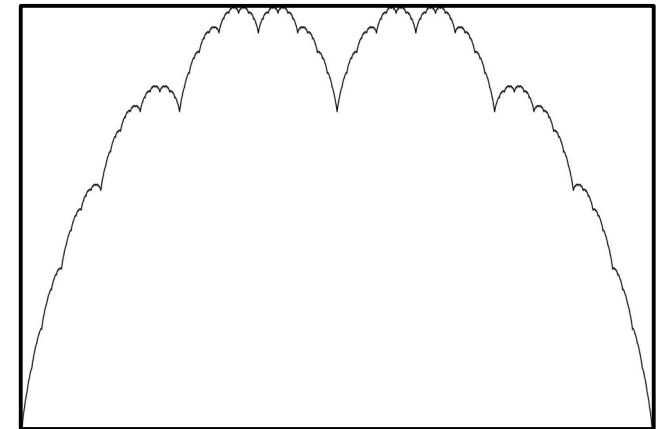
For an n -dimensional Euclidean space, the Hausdorff dimension coincides with the topological dimension: Single point = 0, line segment = 1, square = 2, etc.

Examples

Takagi-Landsberg curve

$$T(x) = \sum_{n=0}^{\infty} w^n s(2^n x) \quad s(x) = \min_{n \in \mathbb{Z}} |x - n|$$

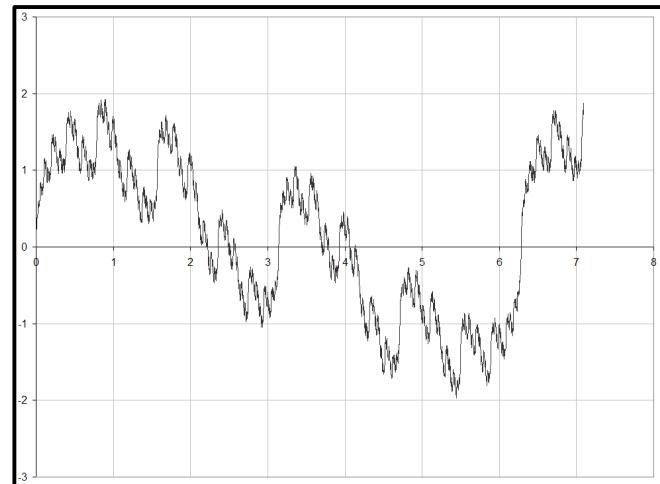
$$D = 2 + \log_2(w)$$



Weierstrass function

$$f(x) = \sum_{k=1}^{\infty} a^{-k} \sin(b^k x), \quad 1 < a < 2, \quad b > 1$$

$$D = 2 - \log_b(a)$$



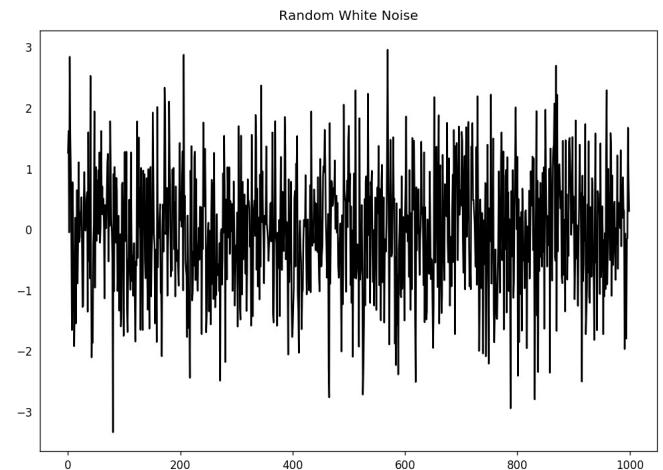
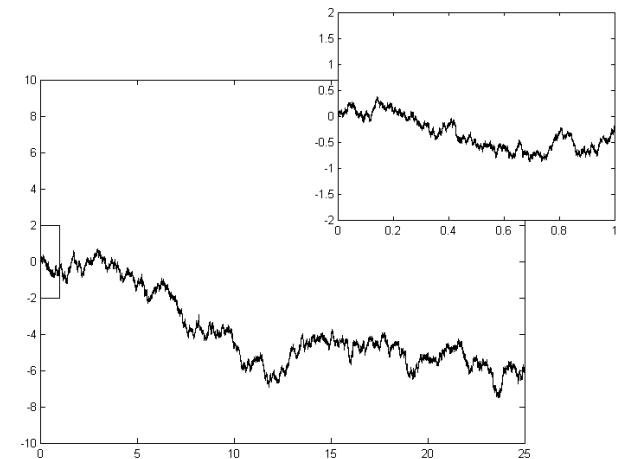
Examples

Brownian motion (Wiener process)

$$D = 3/2$$

White noise

$$D = 2$$



Calculation of fractal dimension

ANAM method

$$M_\tau^\alpha(f, x) = \left[\frac{1}{\tau^2} \int_{t_1=0}^{\tau} \int_{t_2=0}^{\tau} |f(x + t_1) - f(x - t_2)|^\alpha dt_1 dt_2 \right]^{1/\alpha}$$
$$K_\tau^\alpha(f, a, b) = \frac{1}{b-a} \int_{x=a}^{x=b} \left[\frac{1}{\tau^2} \int_{t_1=0}^{\tau} \int_{t_2=0}^{\tau} |f(x + t_1) - f(x - t_2)|^\alpha dt_1 dt_2 \right]^{1/\alpha} dx$$

$$D = \lim_{\tau \rightarrow 0} \left(2 - \frac{\log K_\tau^\alpha(f, a, b)}{\log \tau} \right)$$

Calculation of fractal dimension

Variation (VAR) method

$$f : [a, b] \rightarrow \mathbb{R},$$

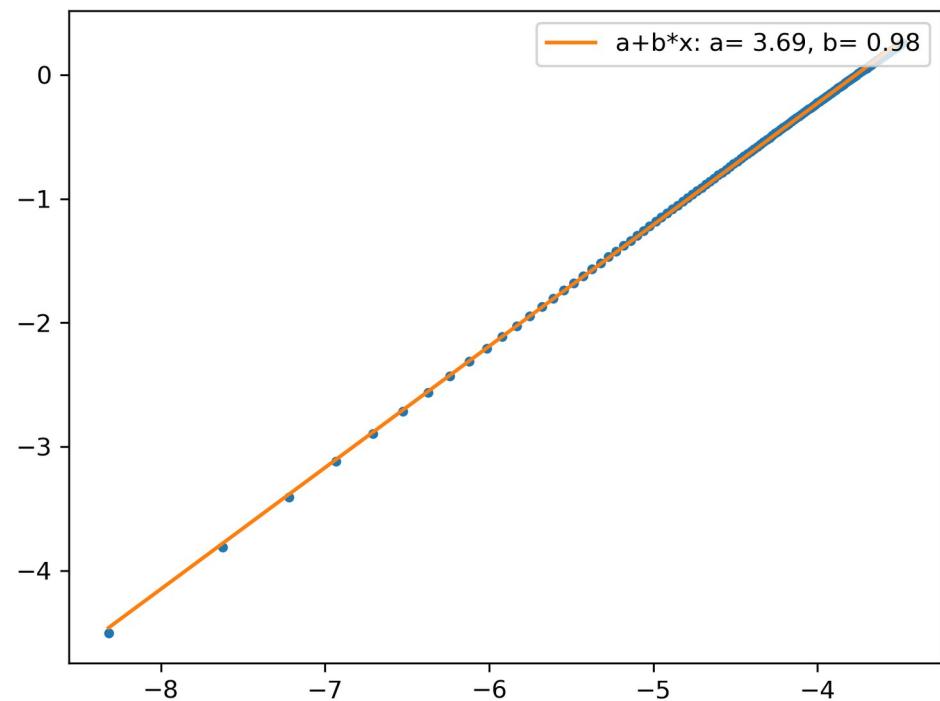
$$\text{OSC}_\tau(f, x) = \left| \max_{|x-t|<\tau} (f(t)) - \min_{|x-t|<\tau} (f(t)) \right|$$

$$\text{VAR}_\tau(f, a, b) = \frac{1}{b-a} \int_a^b \text{OSC}_\tau(f, x) \, dx$$

$$D = \lim_{\tau \rightarrow 0} \left(2 - \frac{\log \text{VAR}_\tau(f, a, b)}{\log \tau} \right)$$

Algorithm

- Discretize either the $\text{VAR}_\tau(f, a, b)$ or the $K_\tau^\alpha(f, a, b)$ estimator
- The fractal dimension of the time series is the slope of $2 - \log(\text{estimator}) / \log(\tau)$
- Theoretically this is a straight line. In practice, do a fit of the $\log(\text{estimator})$ as function of $\log(\tau)$.



Simple tests

Data length: 1 s, sampling rate: 4096 Hz. Decimate factor: 1/128 (VAR/ANAM).

| | Theory | VAR | ANAM |
|-------------------------------------|--------|---------------|---------------|
| Simple analytic function $f(x)$ | 1.000 | 1.020 (+2%) | 1.002 (+0.2%) |
| Takagi-Landberg curve ($w=0.7$) | 1.485 | 1.448 (-2.5%) | 1.503 (+1.2%) |
| Takagi-Landberg curve ($w=0.8$) | 1.678 | 1.748 (+4.2%) | 1.668 (-0.6%) |
| Weierstrass curve ($a=0.7, b=9$) | 1.838 | 1.747 (-5.0%) | 1.804 (-1.9%) |
| Weierstrass curve ($a=0.5, b=12$) | 1.721 | 1.638 (-4.8%) | 1.662 (-3.4%) |

VAR

Faster

Less accurate

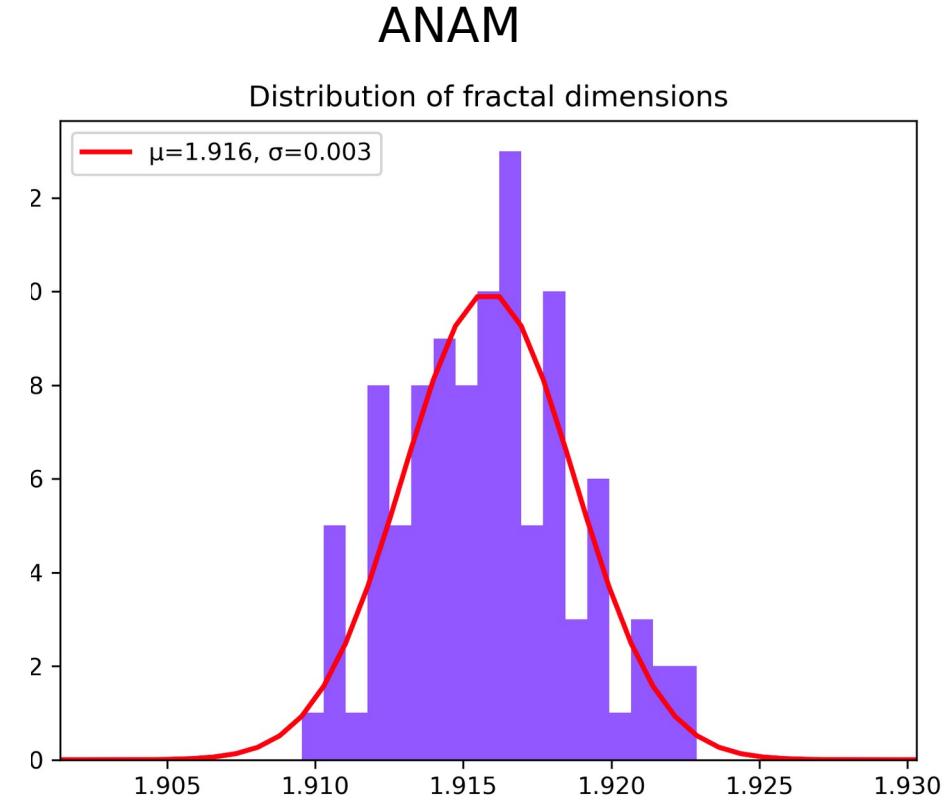
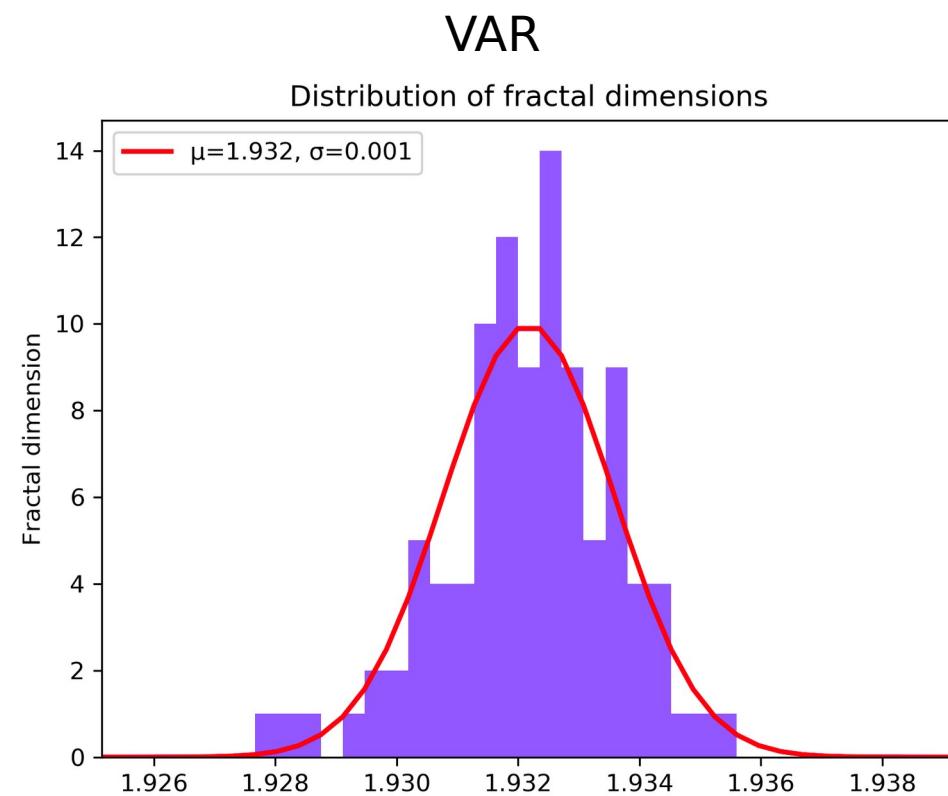
ANAM

Slower ($\times 100$ VAR)

More accurate

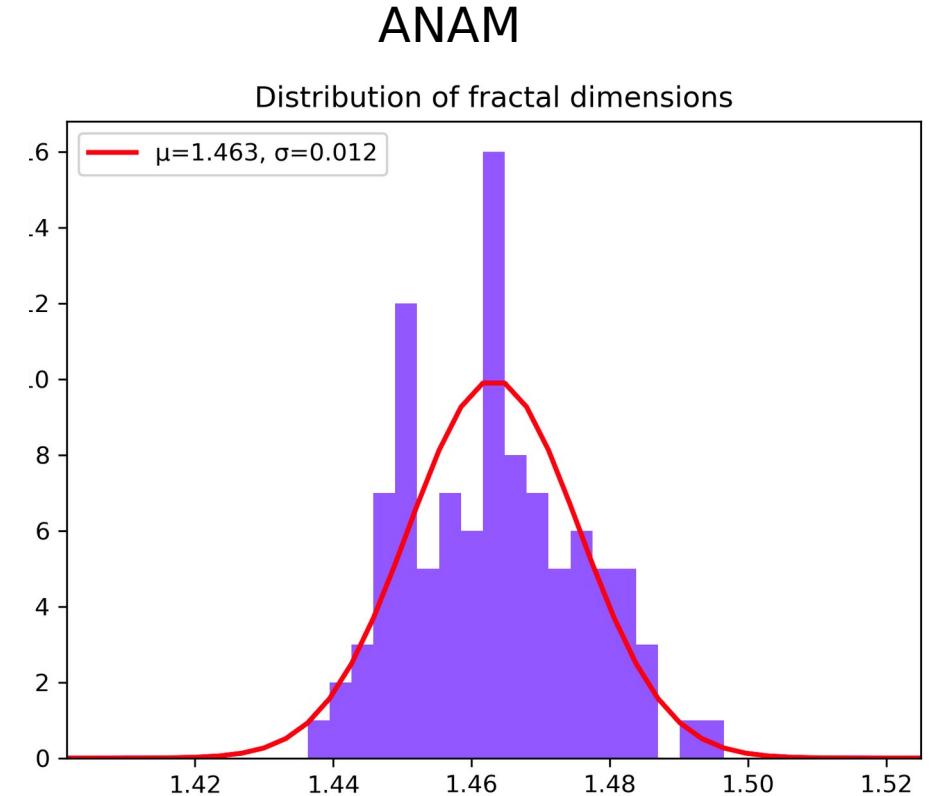
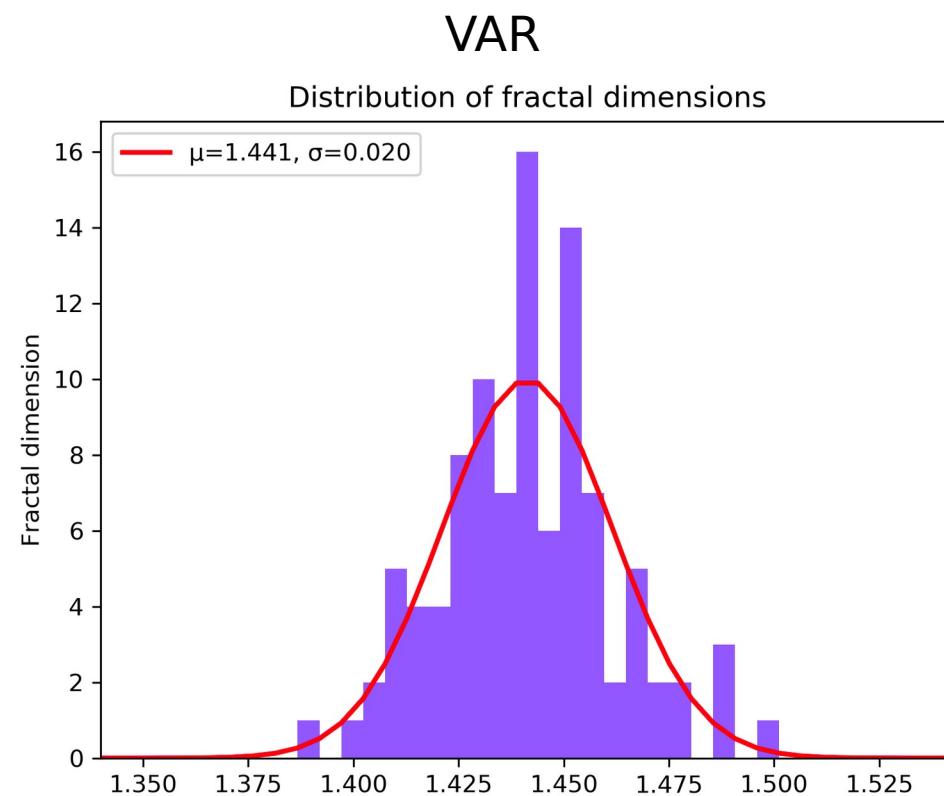
Tests on fake noise: White noise

Data length: 1 s, sampling rate: 4096 Hz. Decimate factor: 16/128 (VAR/ANAM)
Runs: 100. Theoretical fractal dimension: 2.000



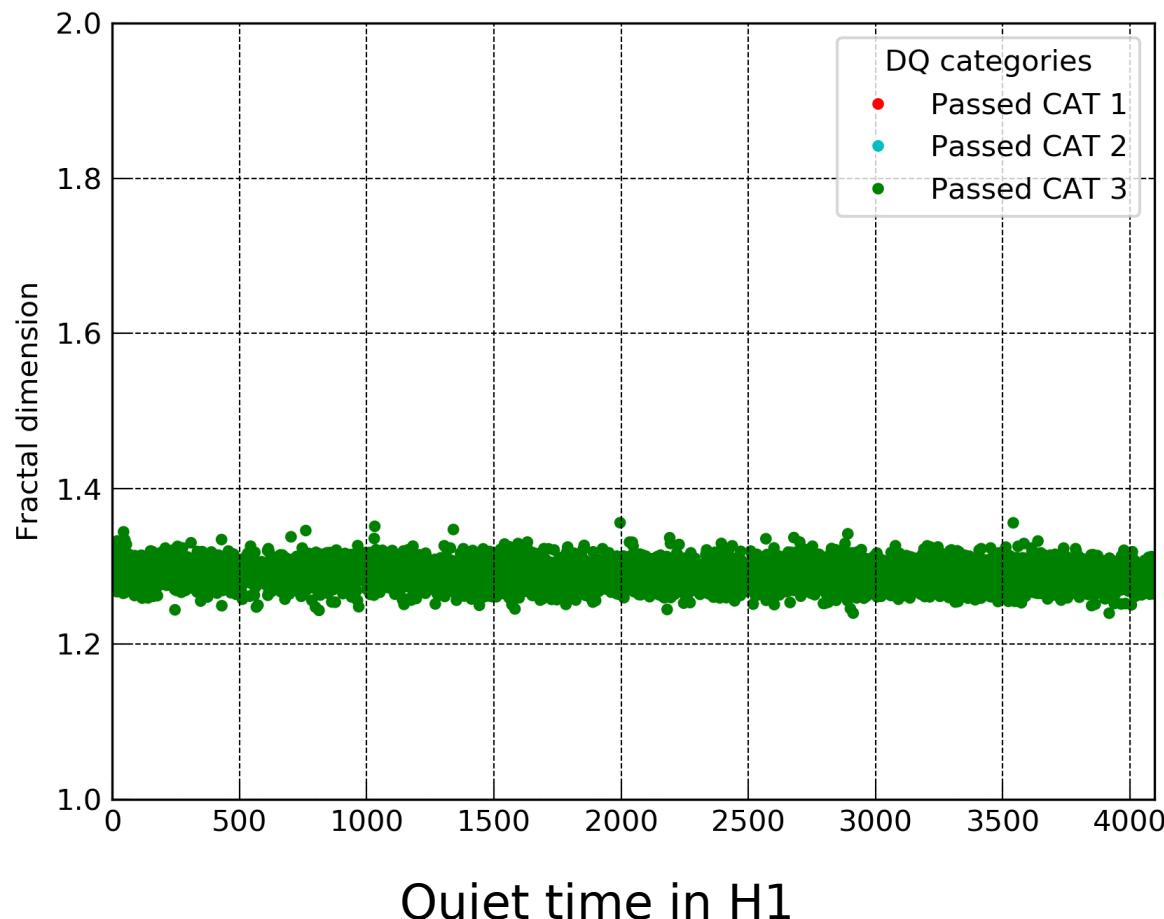
Tests on fake noise: Brownian noise

Data length: 1 s, sampling rate: 4096 Hz. Decimate factor: 16/128 (VAR/ANAM)
Brownian motion speed: 2. Runs: 100. Theoretical fractal dimension: 1.500



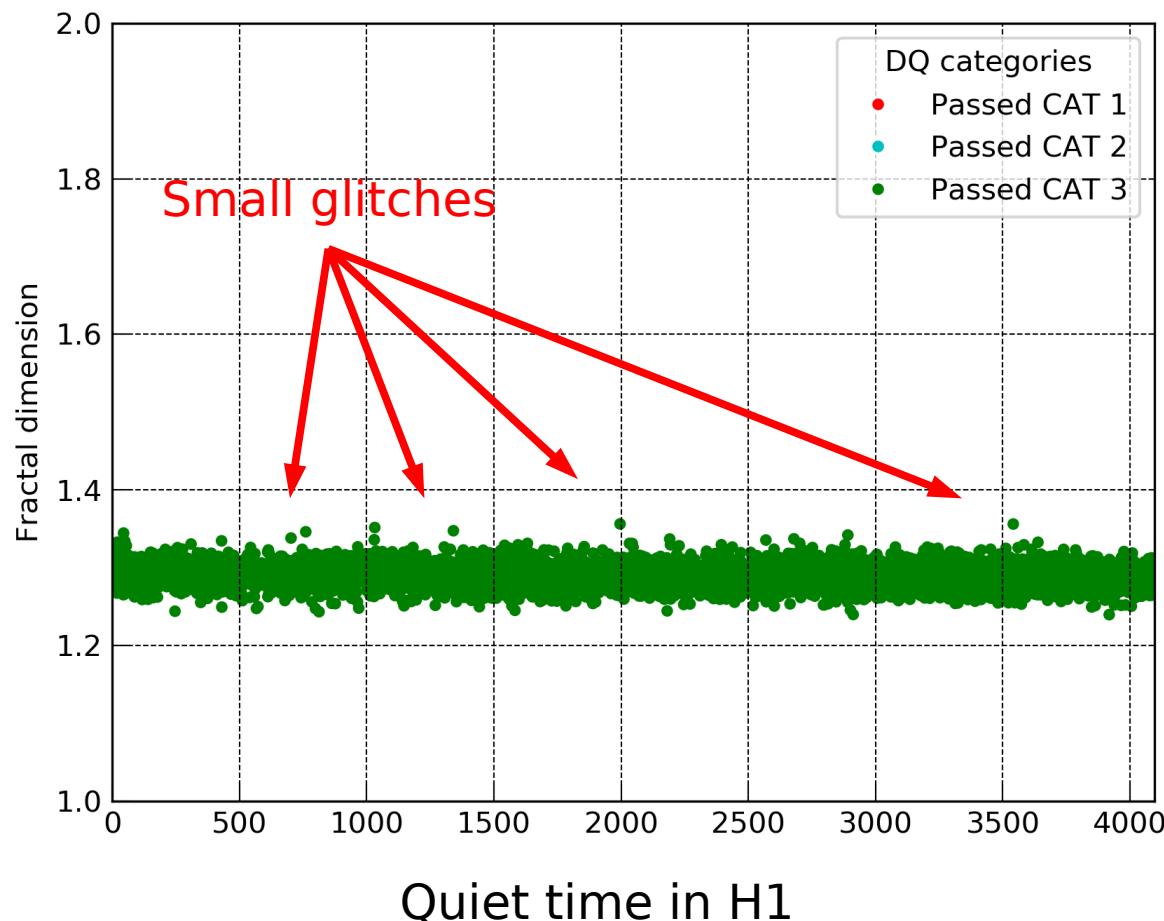
Tests on O2 LIGO (open) data

LIGO Hanford O2 data. Sampling rate: 4096 Hz. Fractal dimension over 1 sec.
Method: VAR. Decimate factor: 16



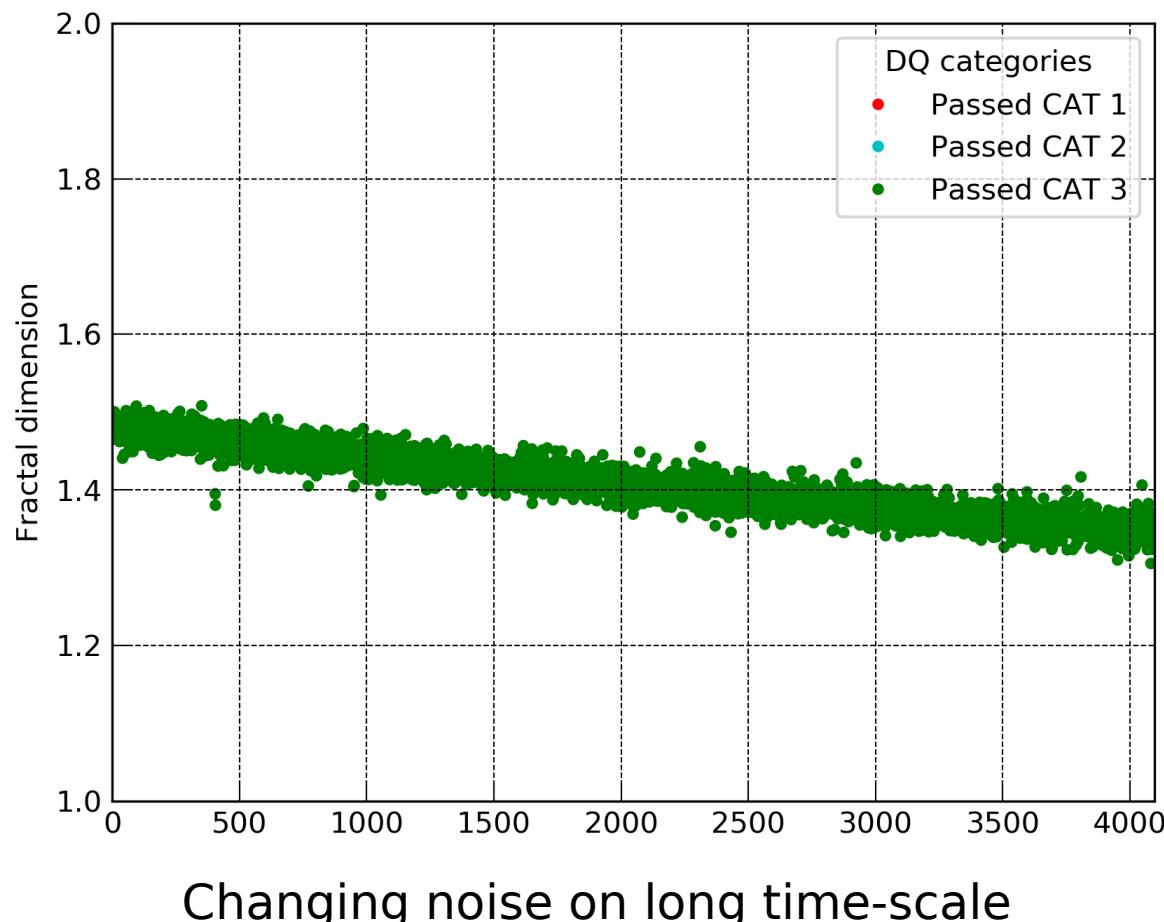
Tests on O2 LIGO (open) data

LIGO Hanford O2 data. Sampling rate: 4096 Hz. Fractal dimension over 1 sec.
Method: VAR. Decimate factor: 16



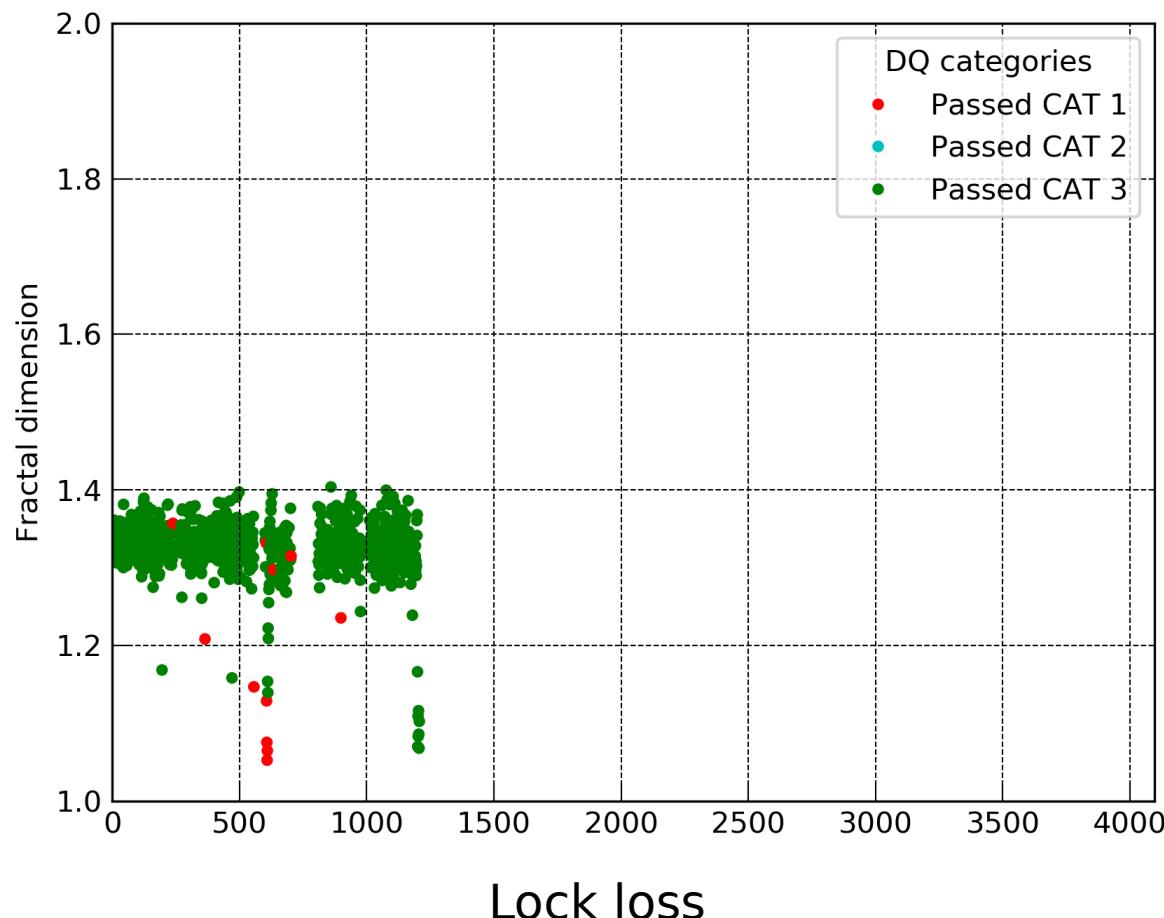
Tests on O2 LIGO (open) data

LIGO Hanford O2 data. Sampling rate: 4096 Hz. Fractal dimension over 1 sec.
Method: VAR. Decimate factor: 16



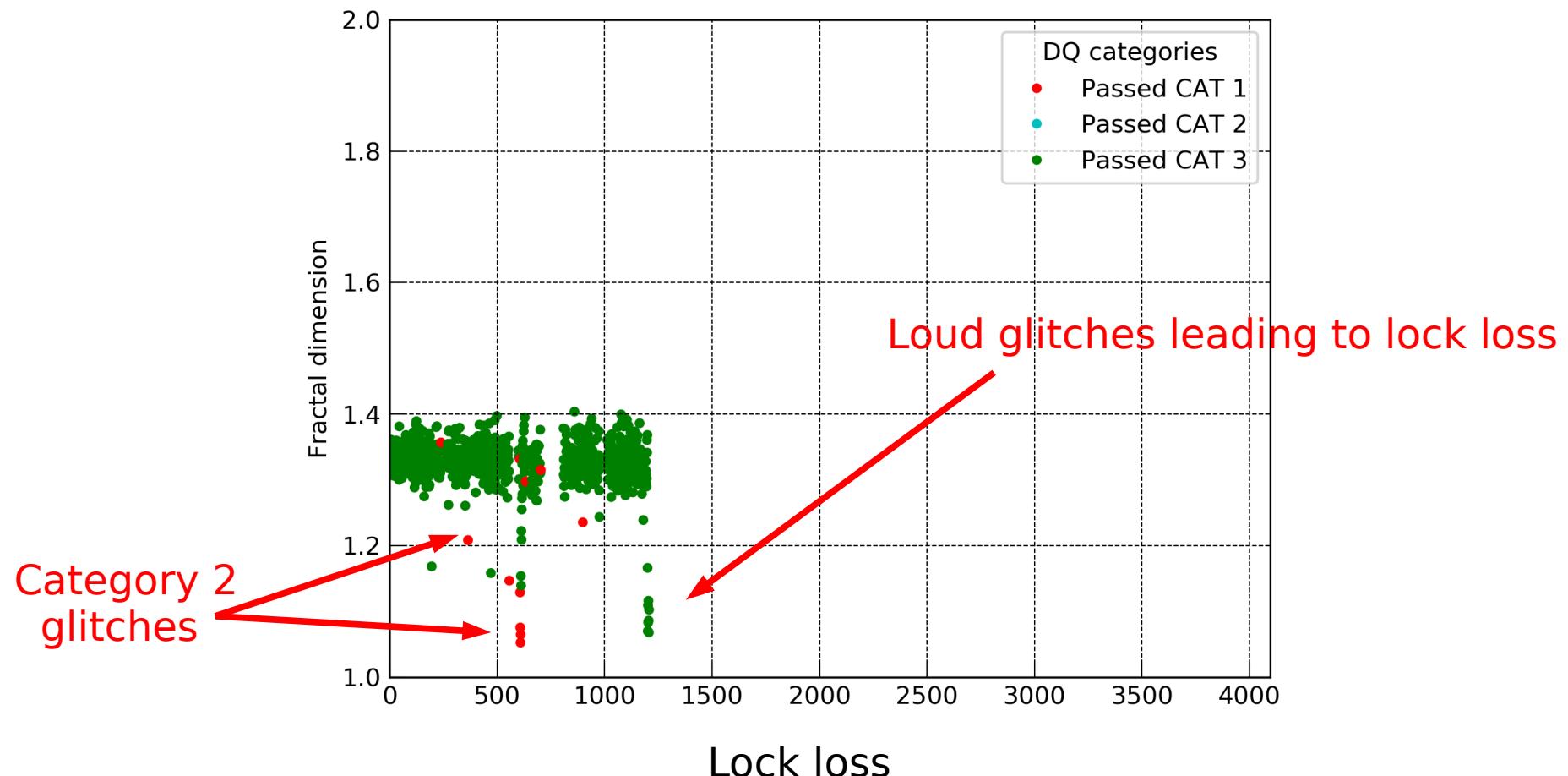
Tests on O2 LIGO (open) data

LIGO Hanford O2 data. Sampling rate: 4096 Hz. Fractal dimension over 1 sec.
Method: VAR. Decimate factor: 16



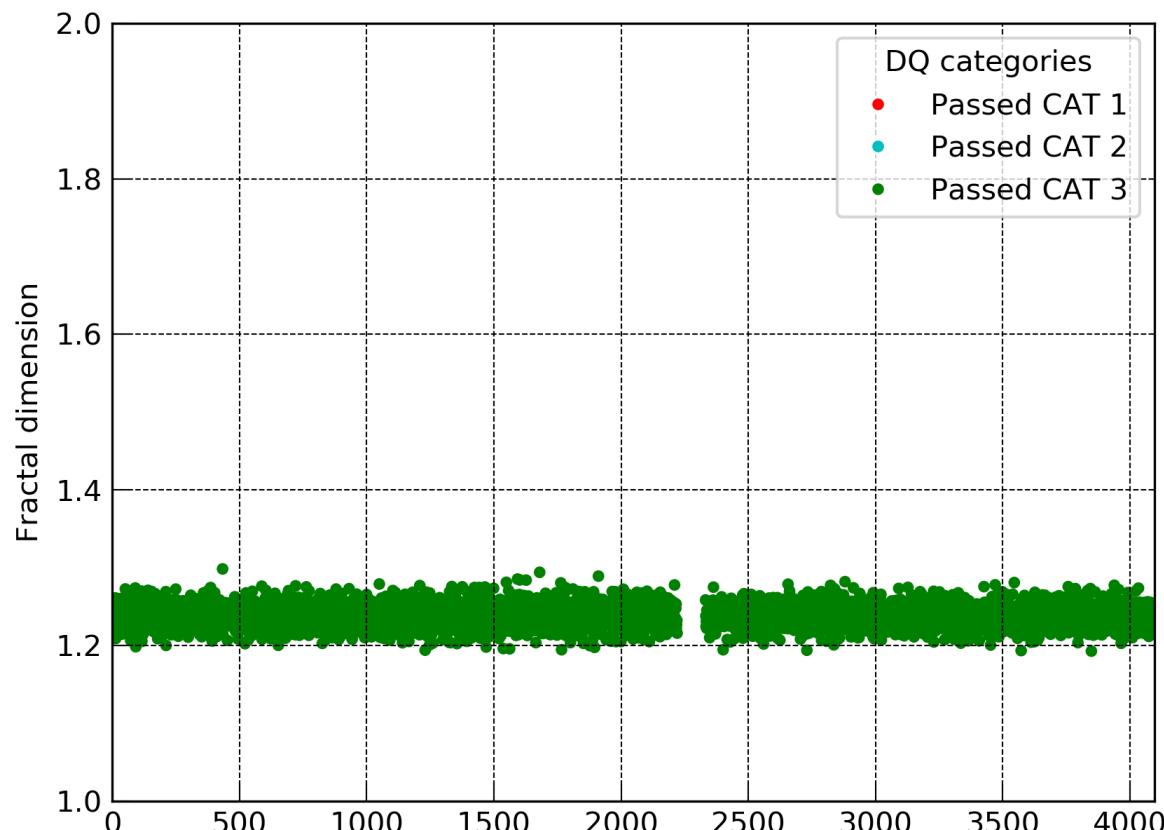
Tests on O2 LIGO (open) data

LIGO Hanford O2 data. Sampling rate: 4096 Hz. Fractal dimension over 1 sec.
Method: VAR. Decimate factor: 16



Tests on O2 LIGO (open) data

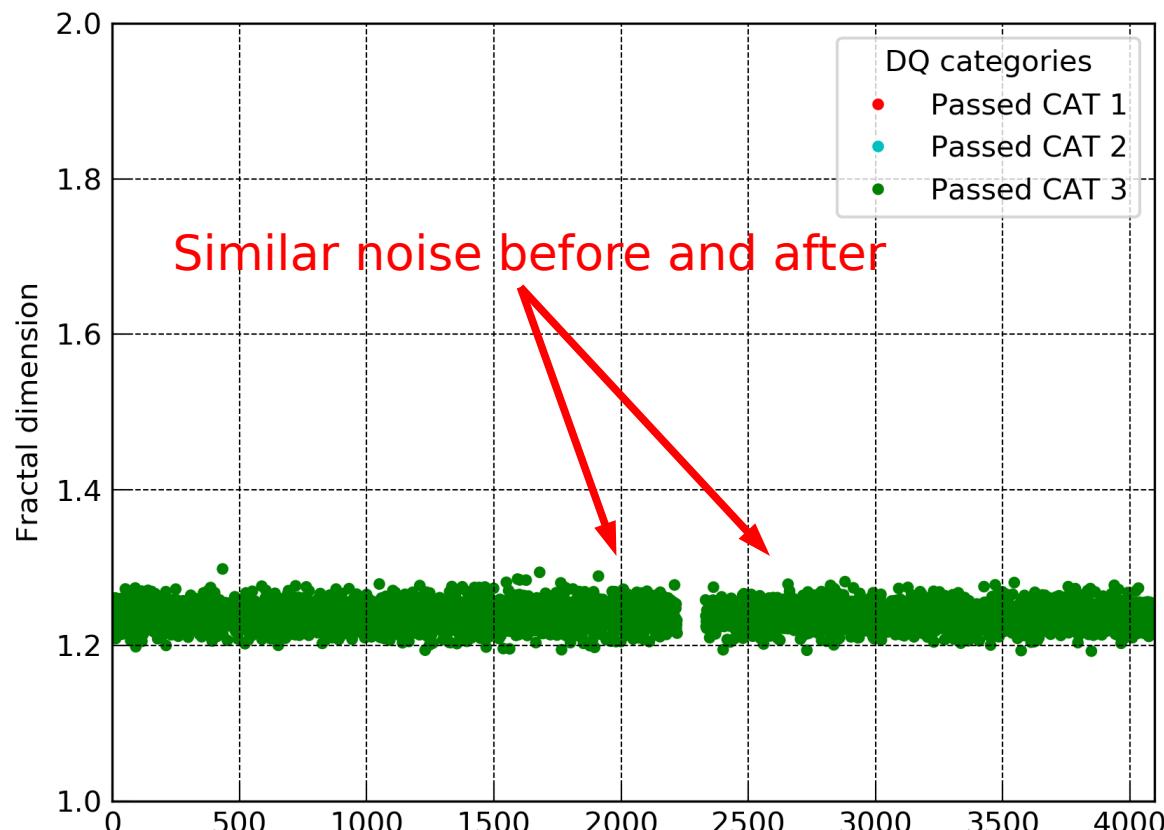
LIGO Hanford O2 data. Sampling rate: 4096 Hz. Fractal dimension over 1 sec.
Method: VAR. Decimate factor: 16



Category 1 data (interferometer in low noise, but locked)

Tests on O2 LIGO (open) data

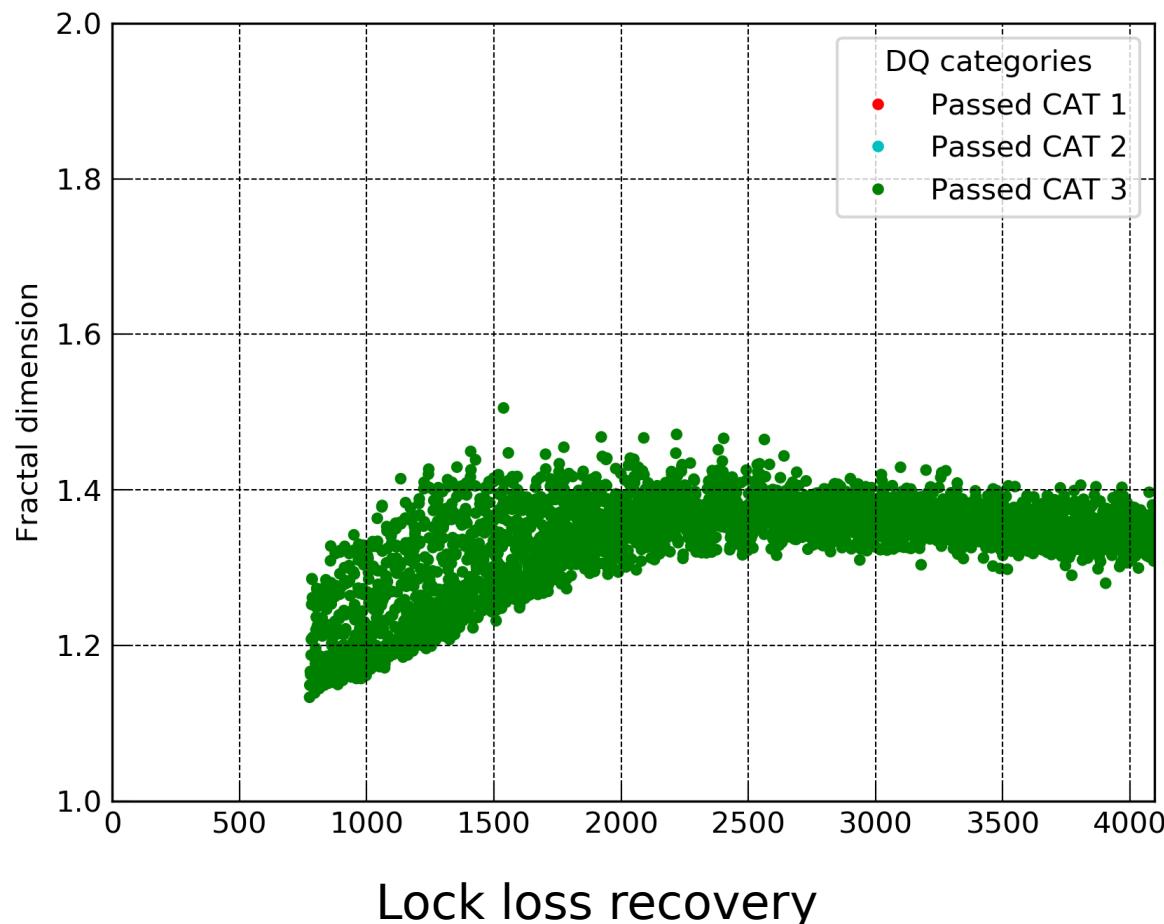
LIGO Hanford O2 data. Sampling rate: 4096 Hz. Fractal dimension over 1 sec.
Method: VAR. Decimate factor: 16



Category 1 data (interferometer in low noise, but locked)

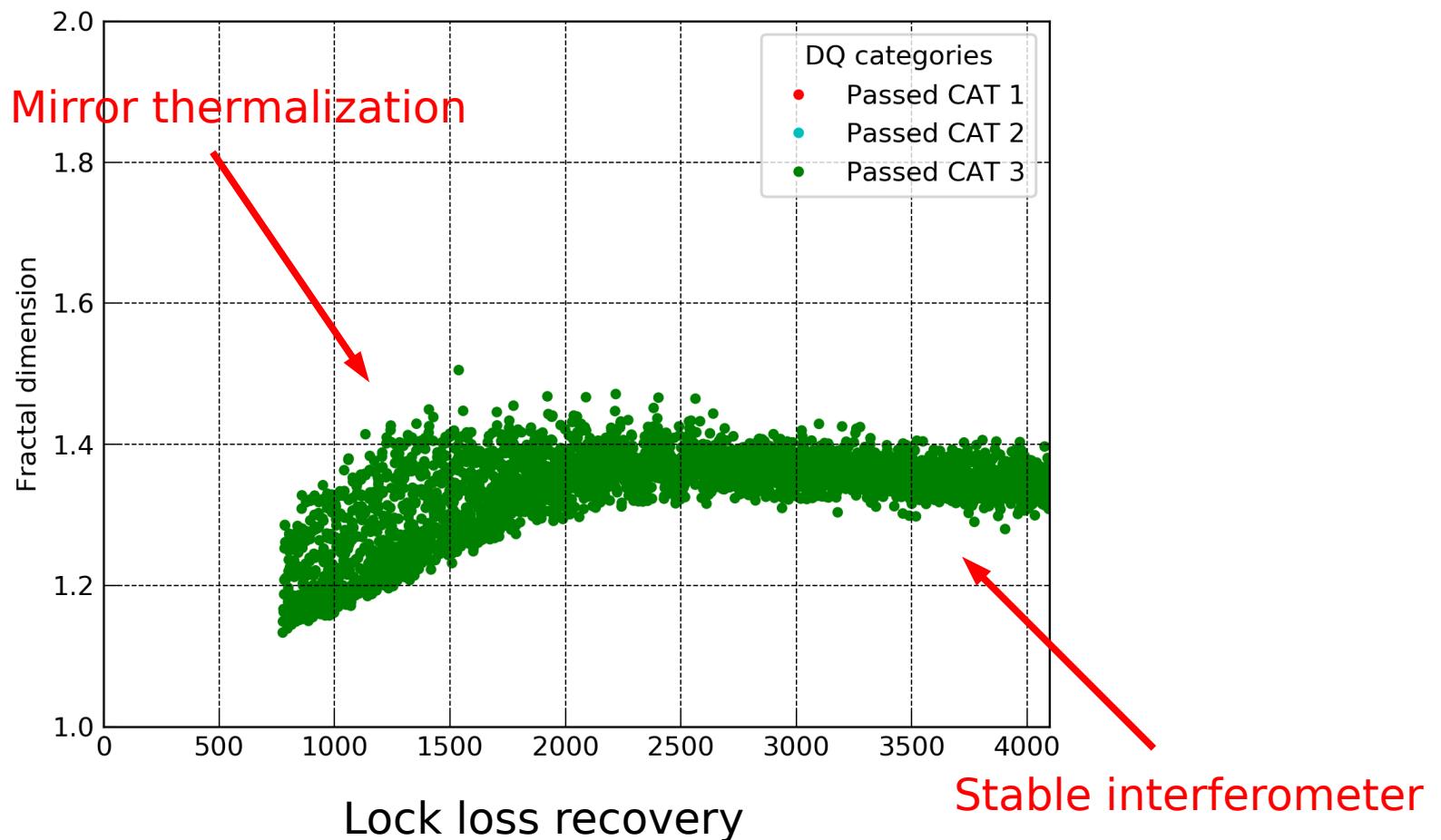
Tests on O2 LIGO (open) data

LIGO Hanford O2 data. Sampling rate: 4096 Hz. Fractal dimension over 1 sec.
Method: VAR. Decimate factor: 16

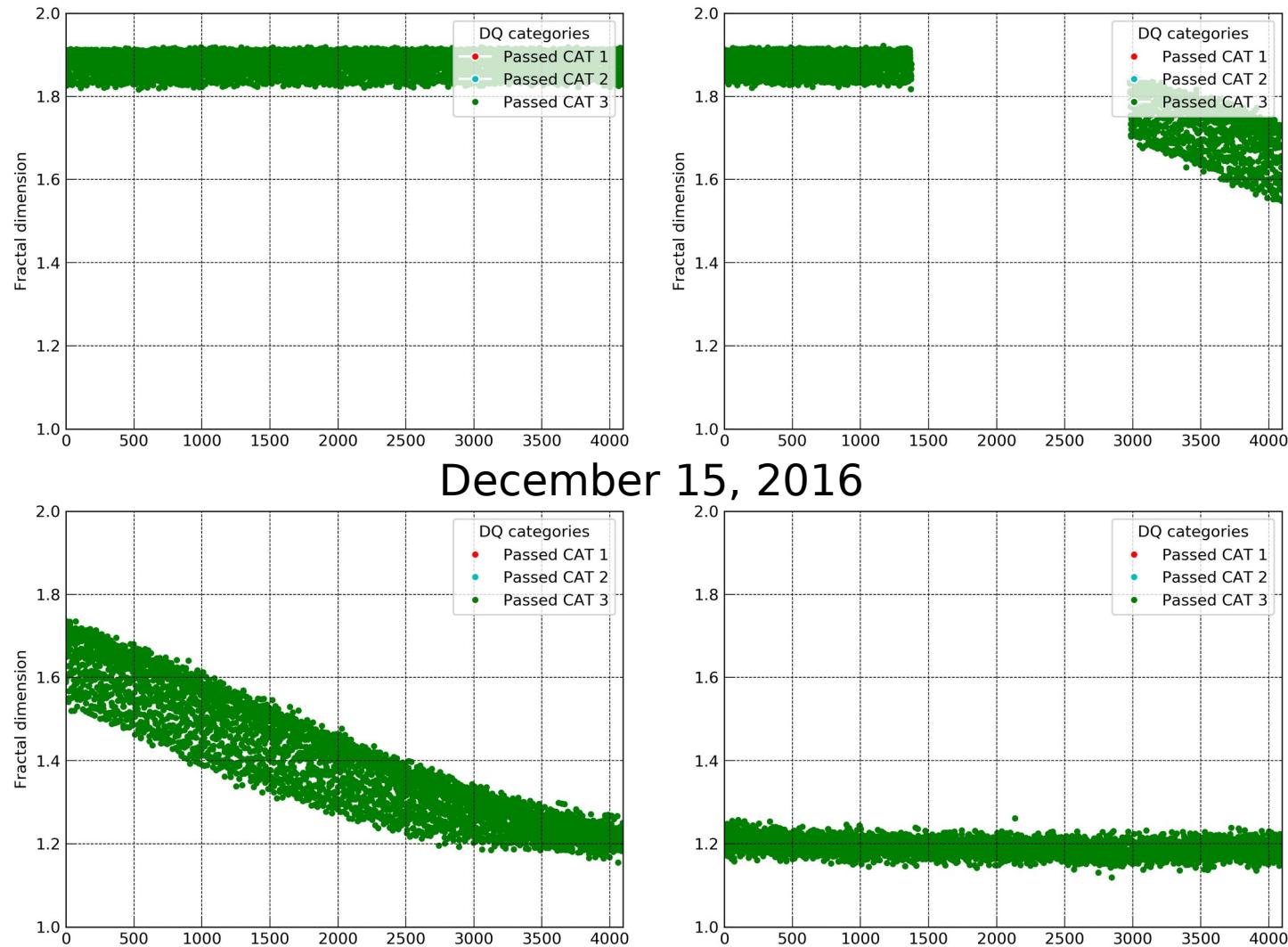


Tests on O2 LIGO (open) data

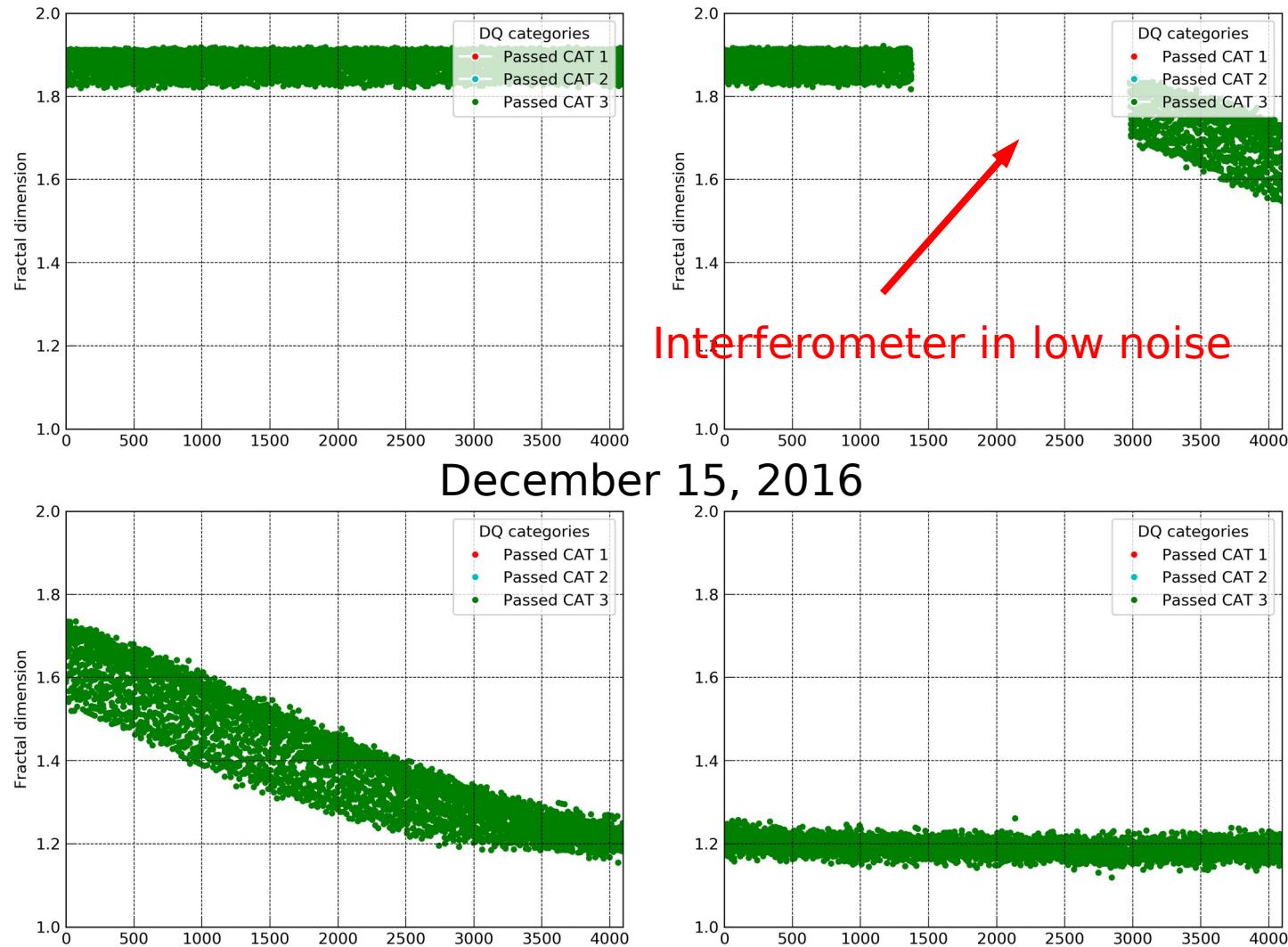
LIGO Hanford O2 data. Sampling rate: 4096 Hz. Fractal dimension over 1 sec.
Method: VAR. Decimate factor: 16



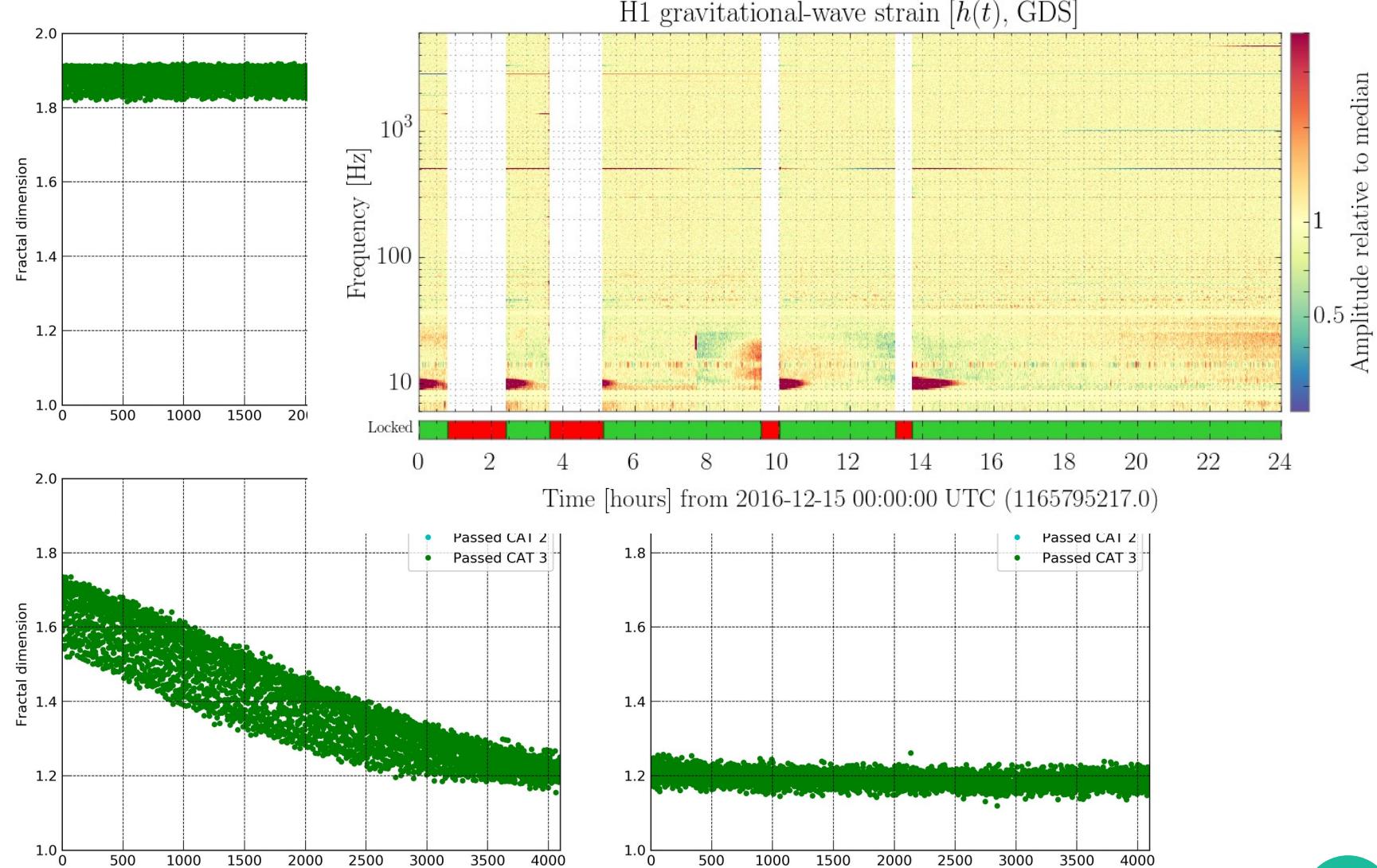
Tests on O2 LIGO (open) data



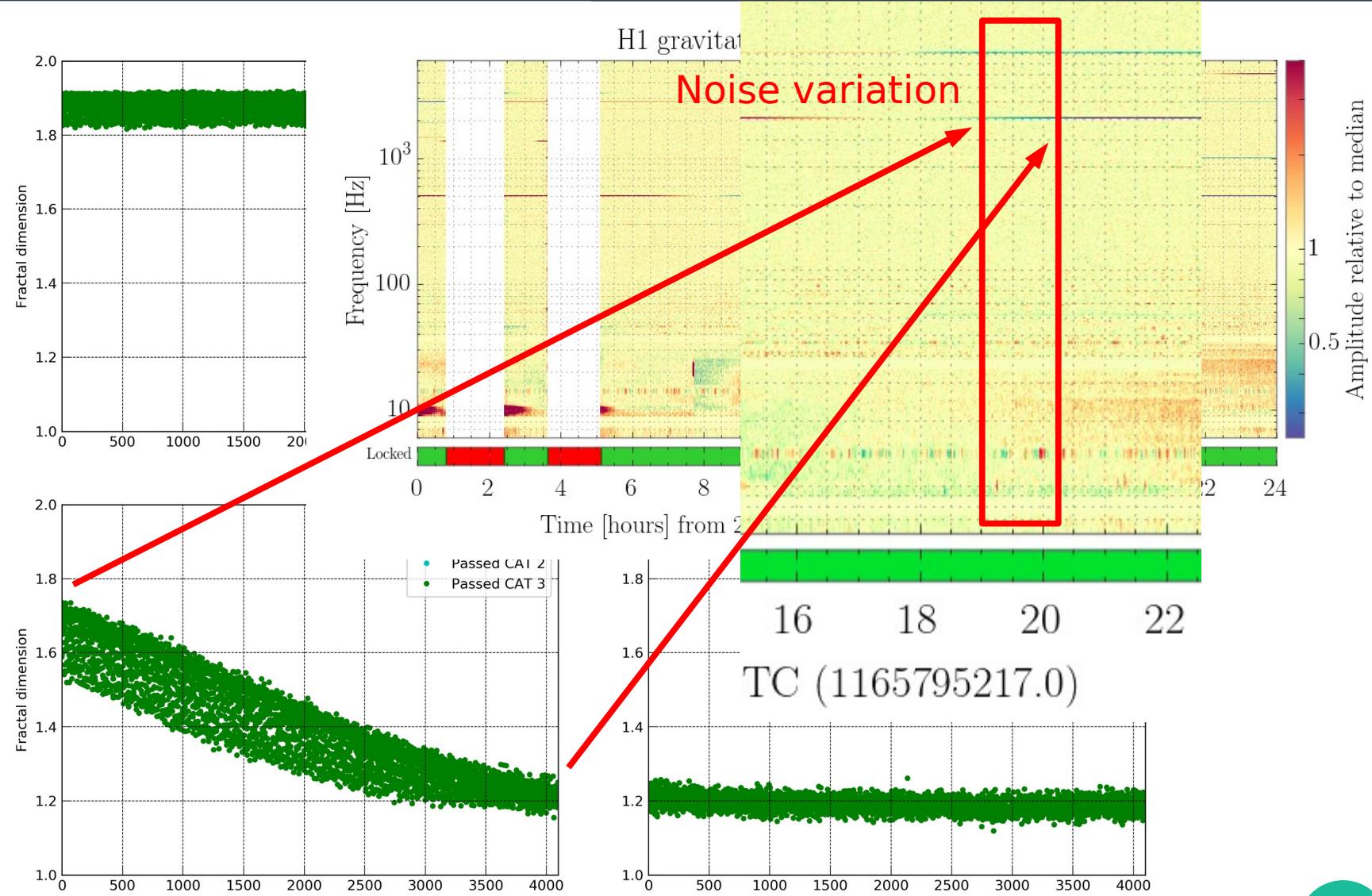
Tests on O2 LIGO (open) data



Tests on O2 LIGO (open) data

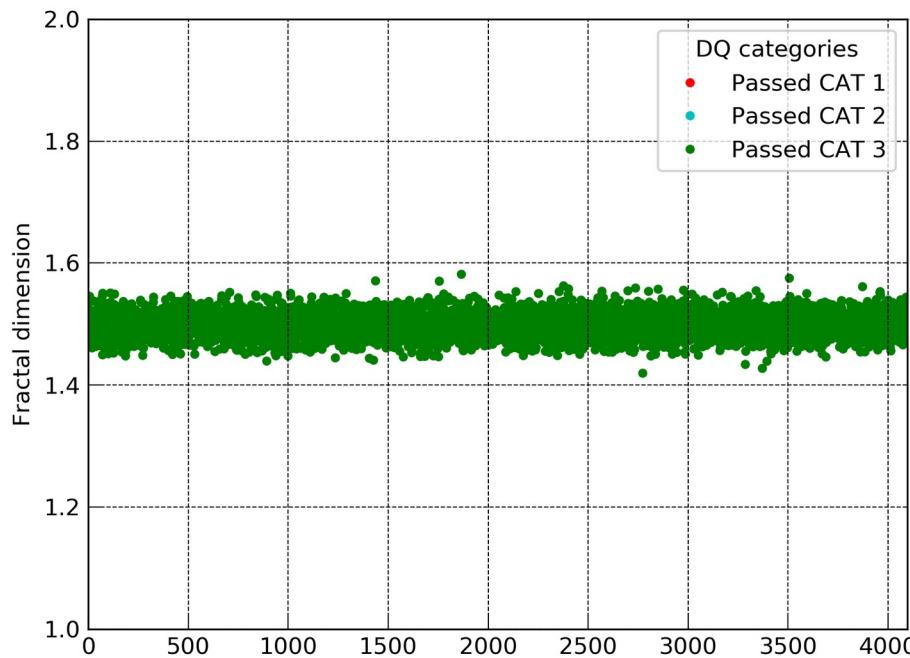


Tests on O2 LIGO (open) data

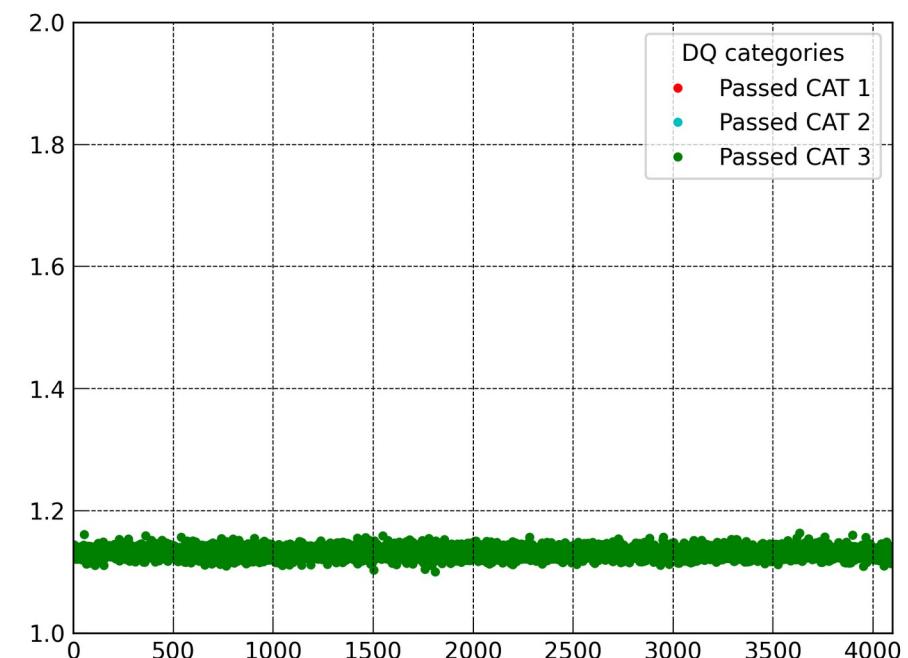


Tests on O2 LIGO (open) data

Background noise is not the same across the run!

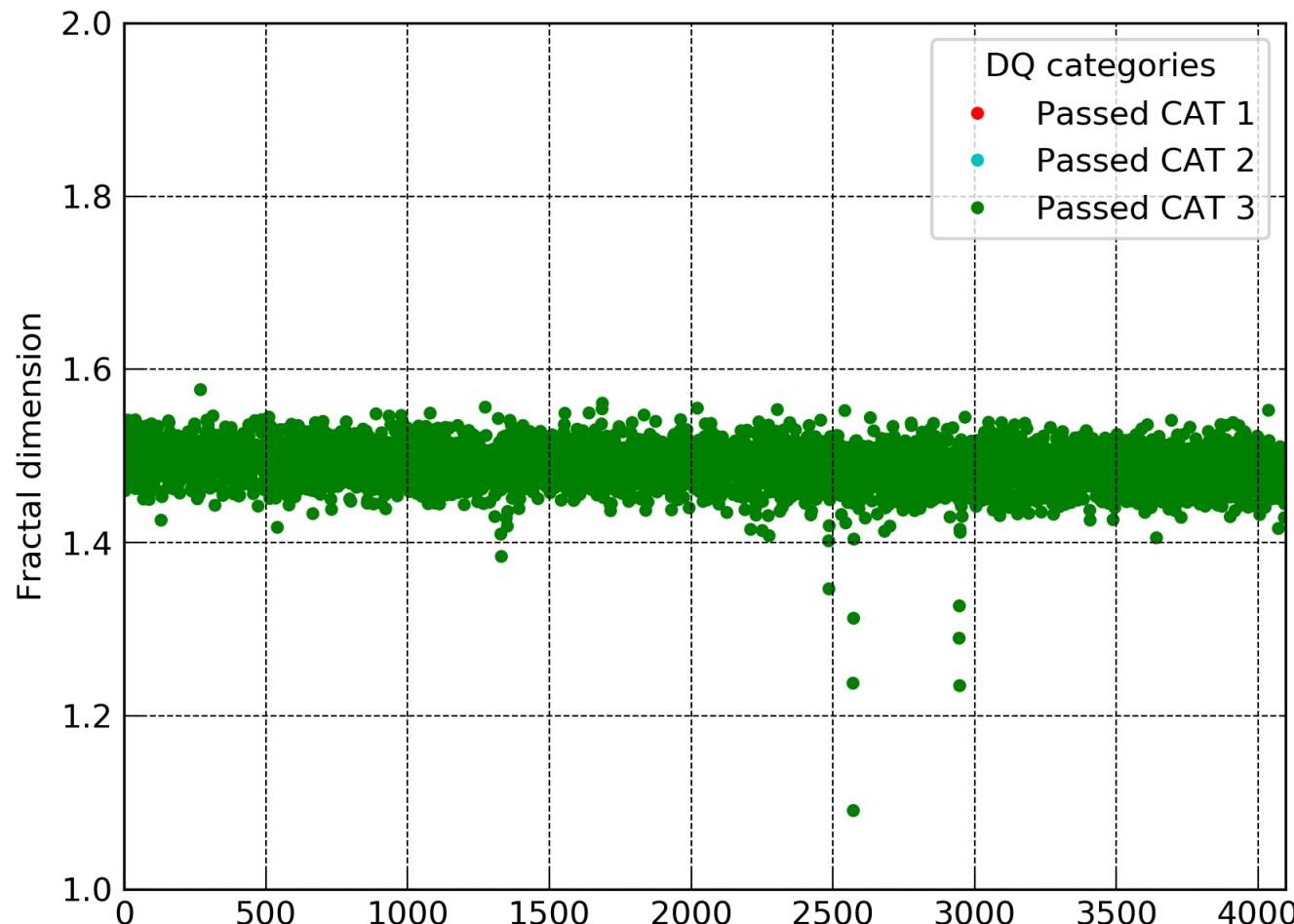


December 8, 2016

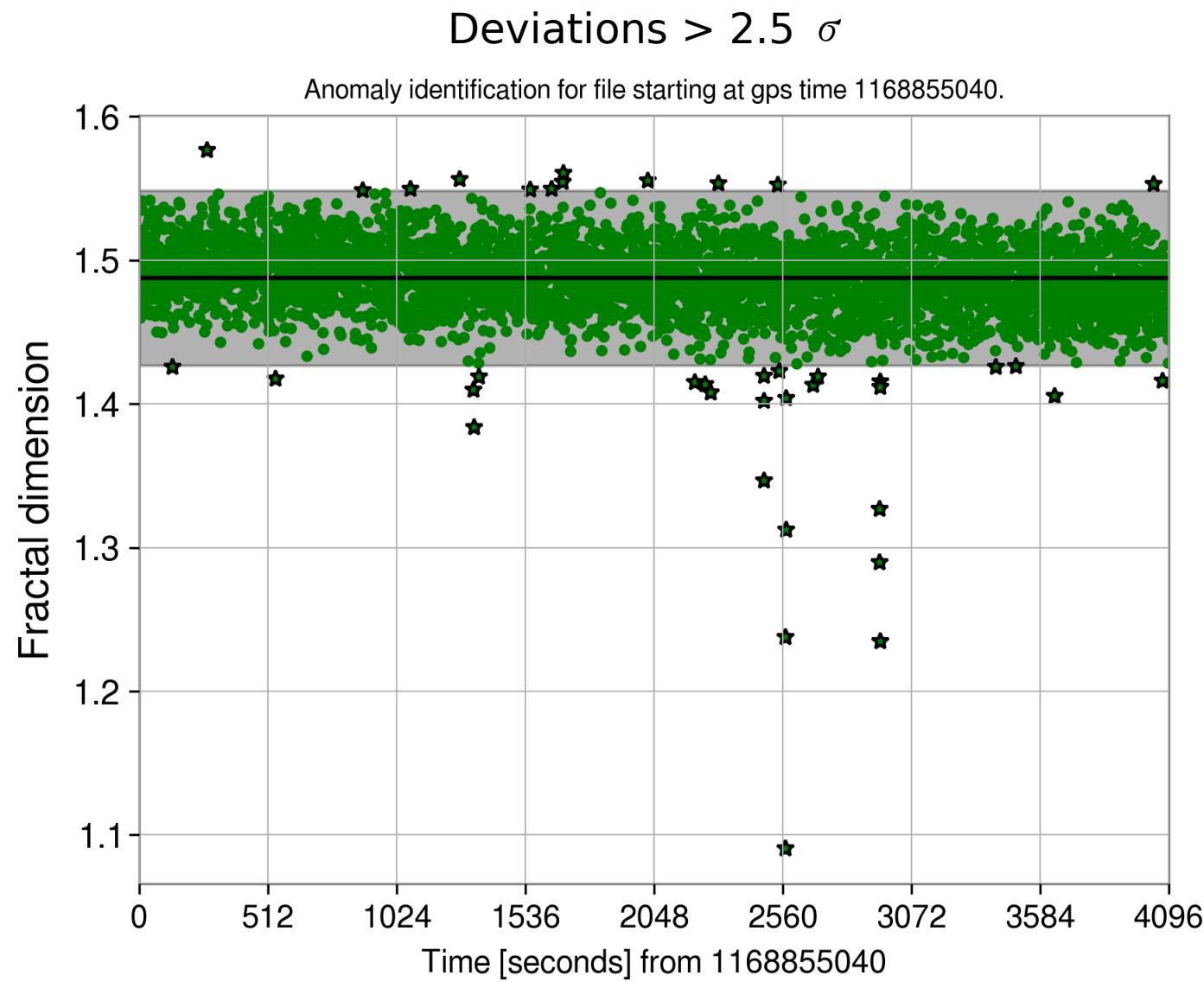


February 24, 2017

Glitch identification



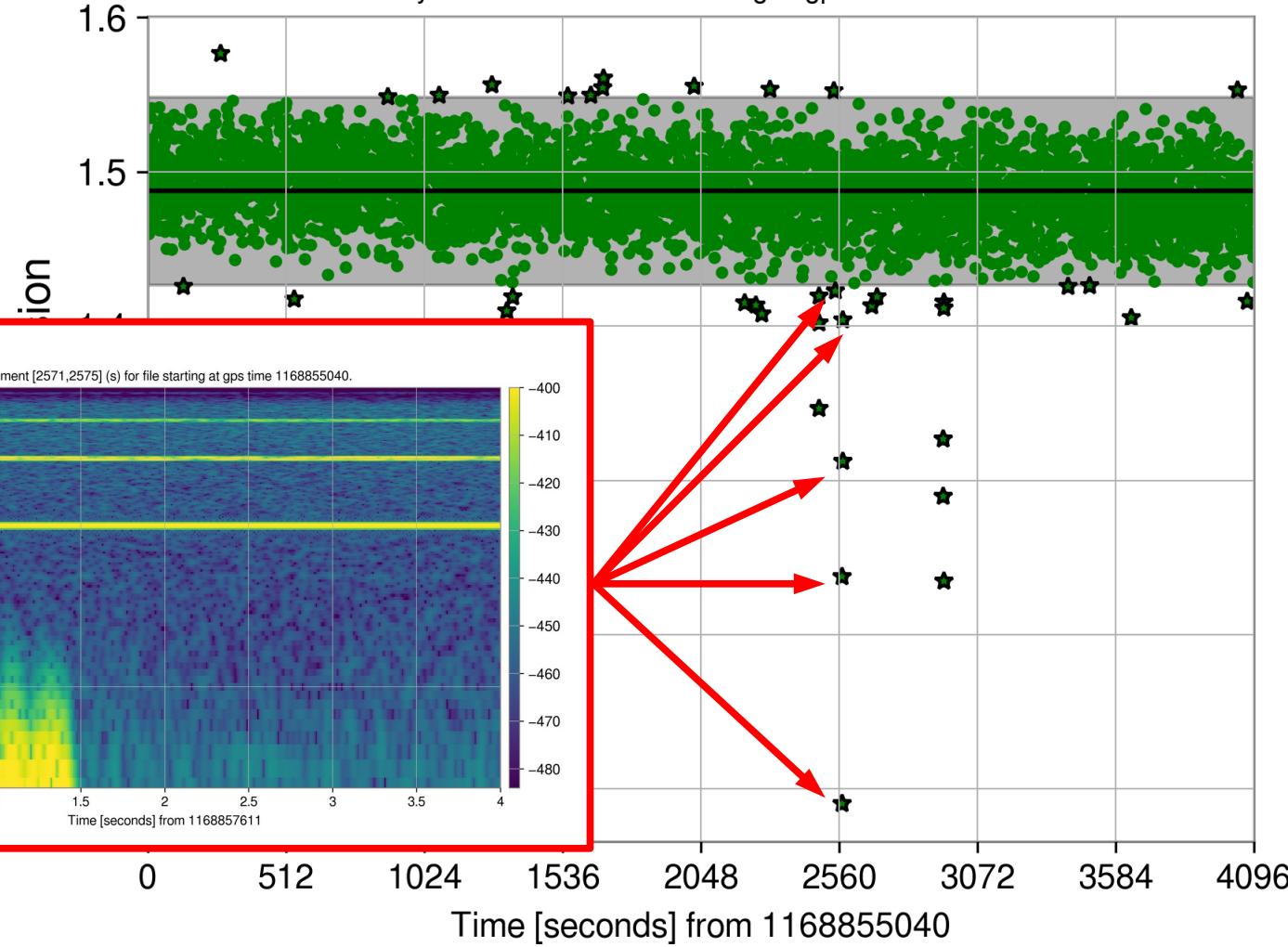
Glitch identification



Glitch identification

Deviations $> 2.5 \sigma$

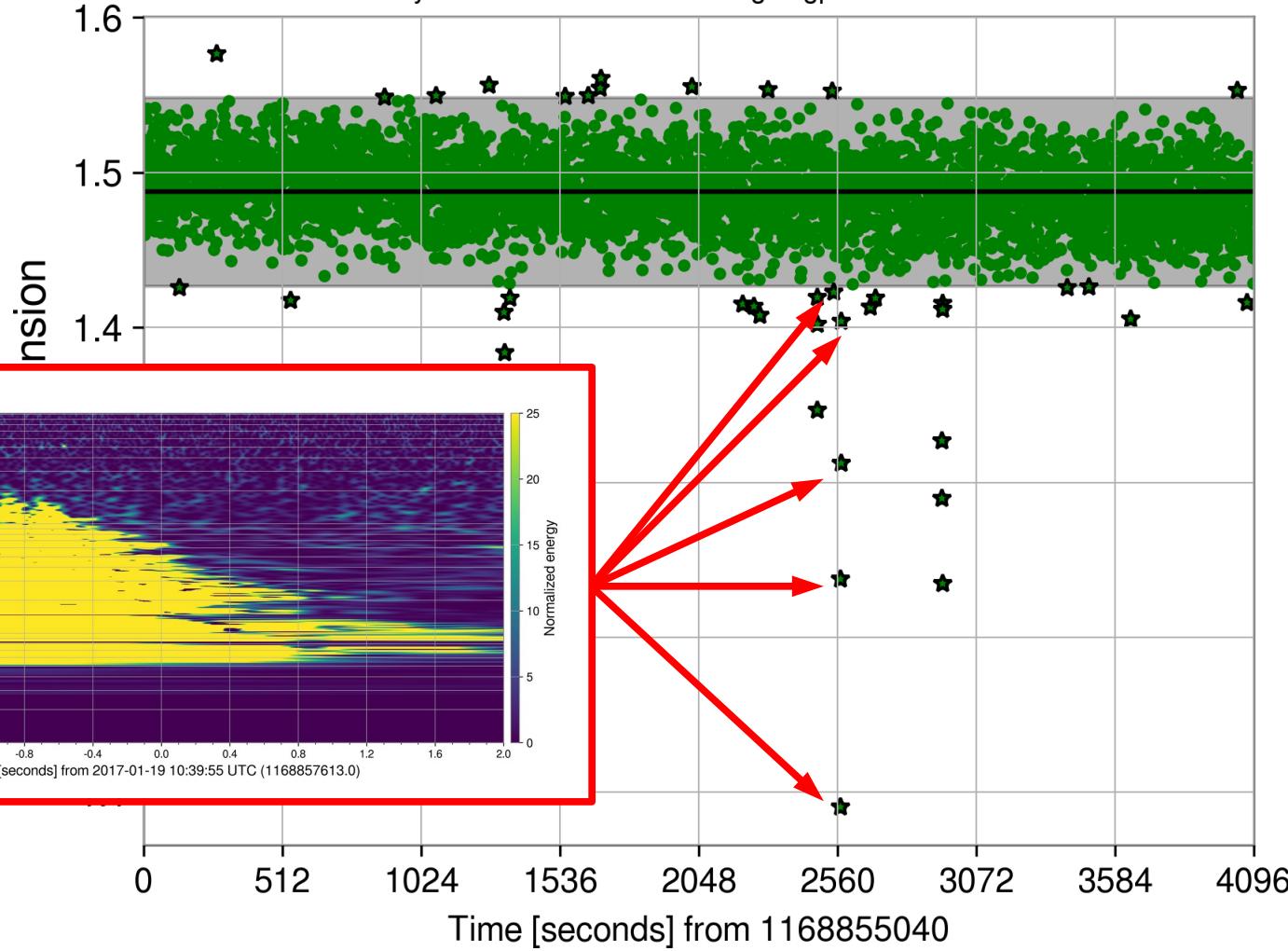
Anomaly identification for file starting at gps time 1168855040.



Glitch identification

Deviations $> 2.5 \sigma$

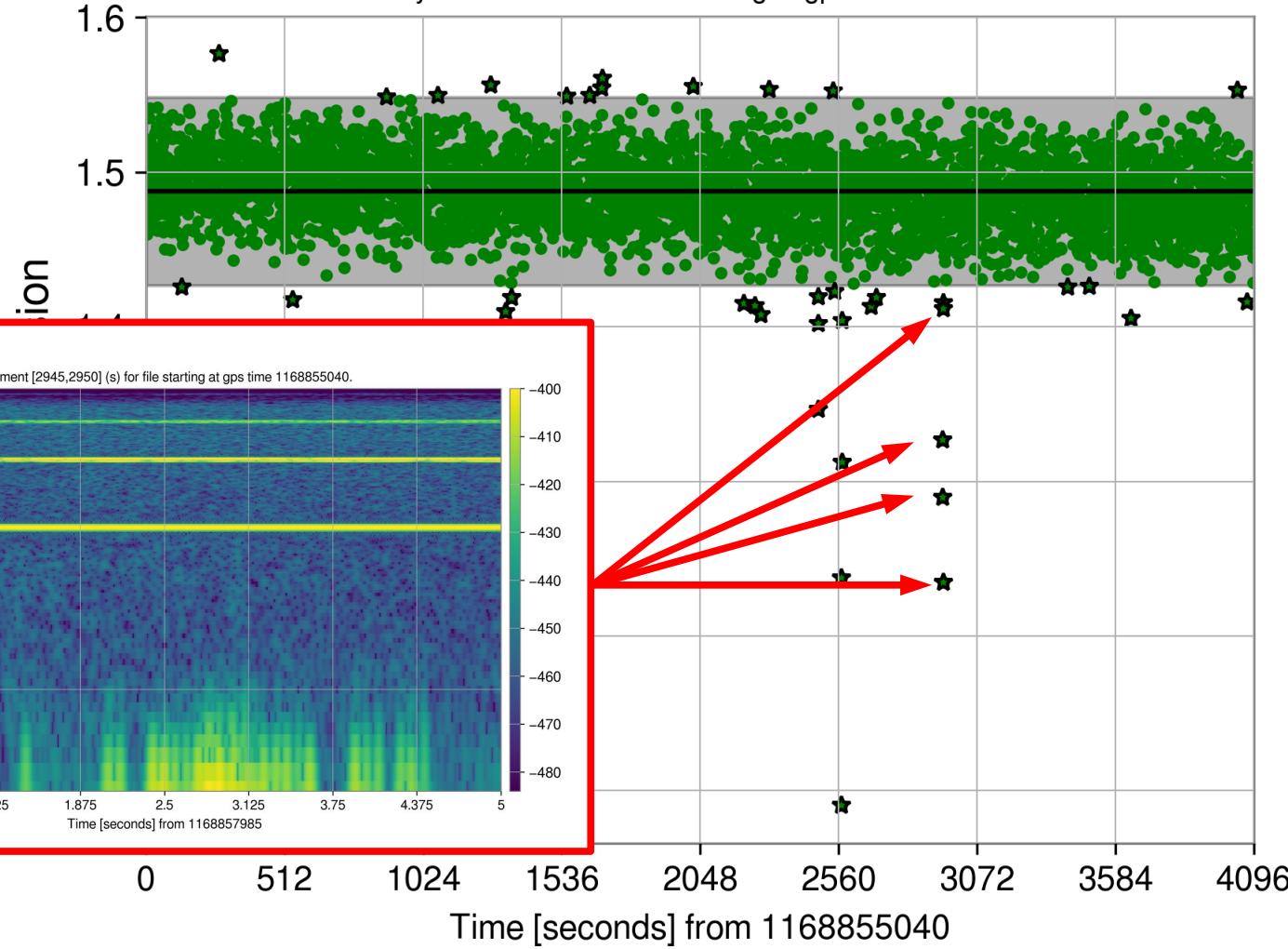
Anomaly identification for file starting at gps time 1168855040.



Glitch identification

Deviations $> 2.5 \sigma$

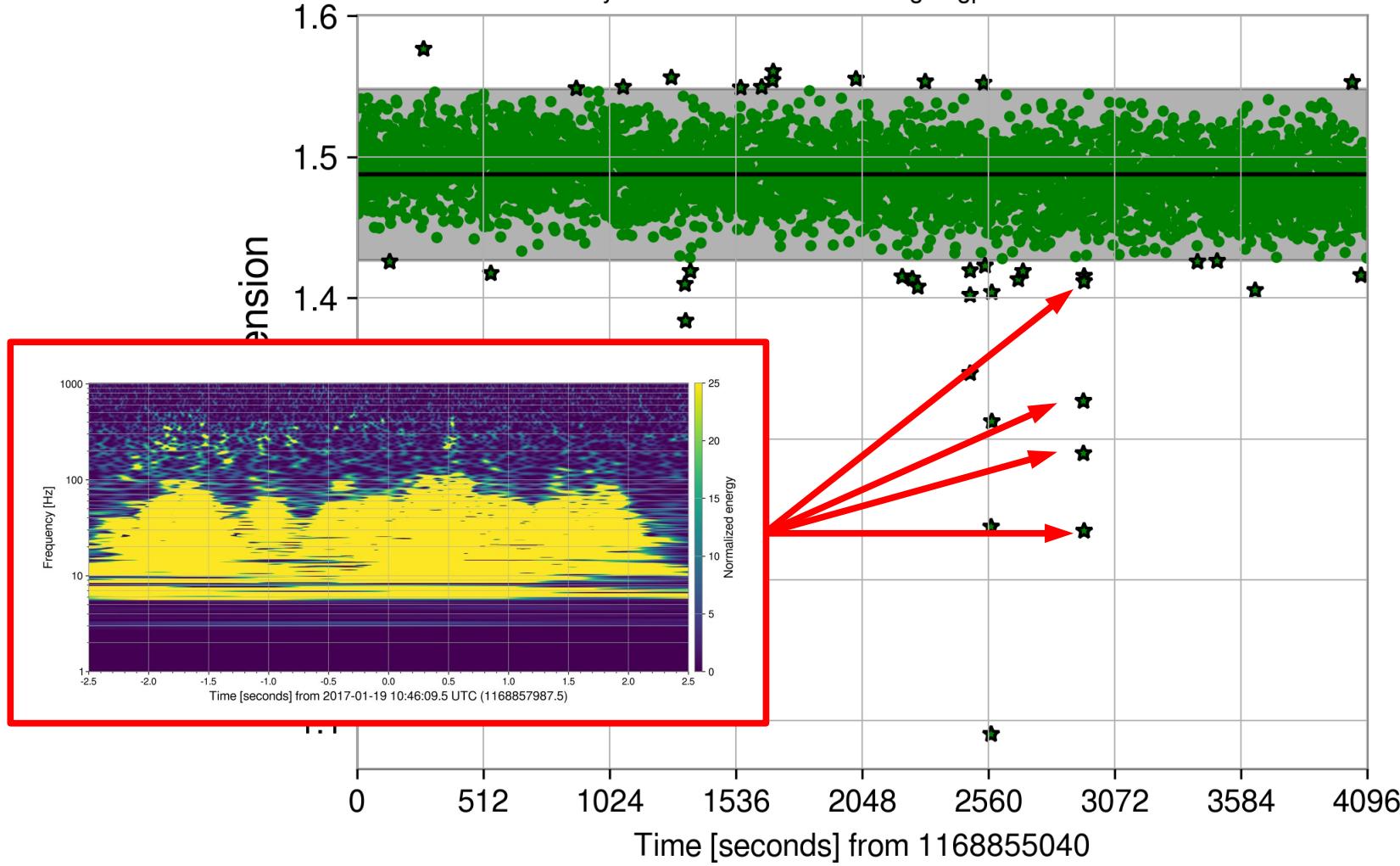
Anomaly identification for file starting at gps time 1168855040.



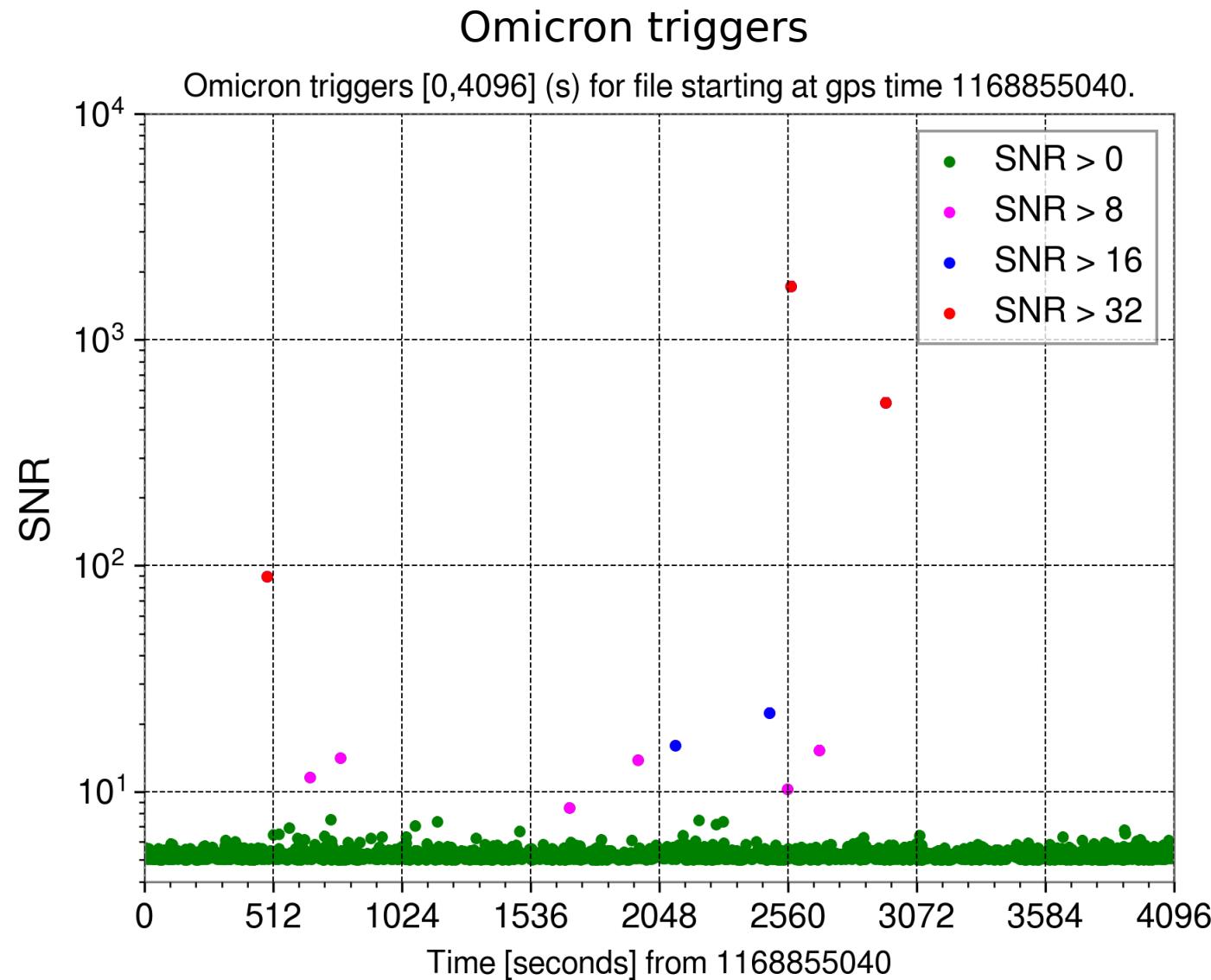
Glitch identification

Deviations $> 2.5 \sigma$

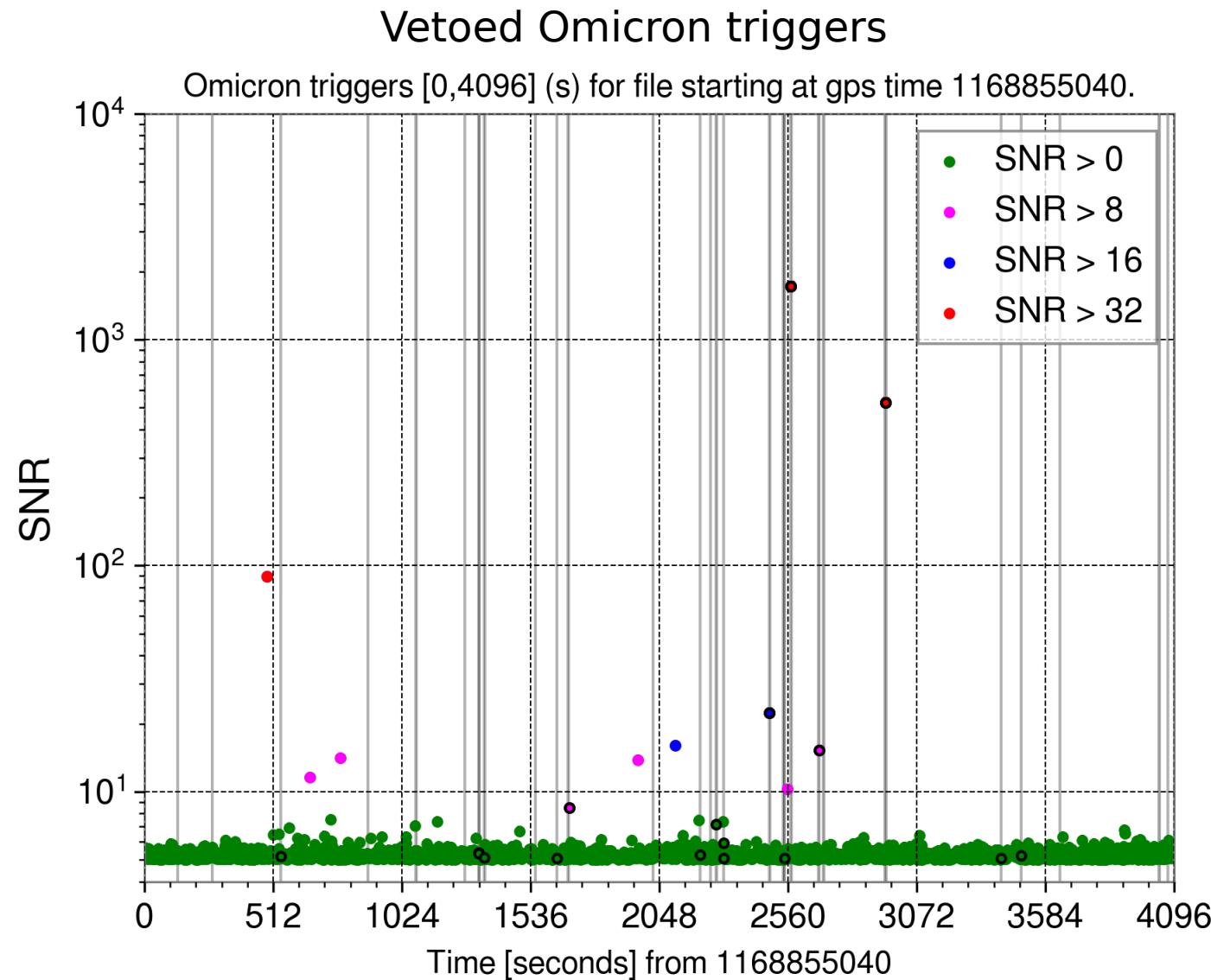
Anomaly identification for file starting at gps time 1168855040.



Glitch identification

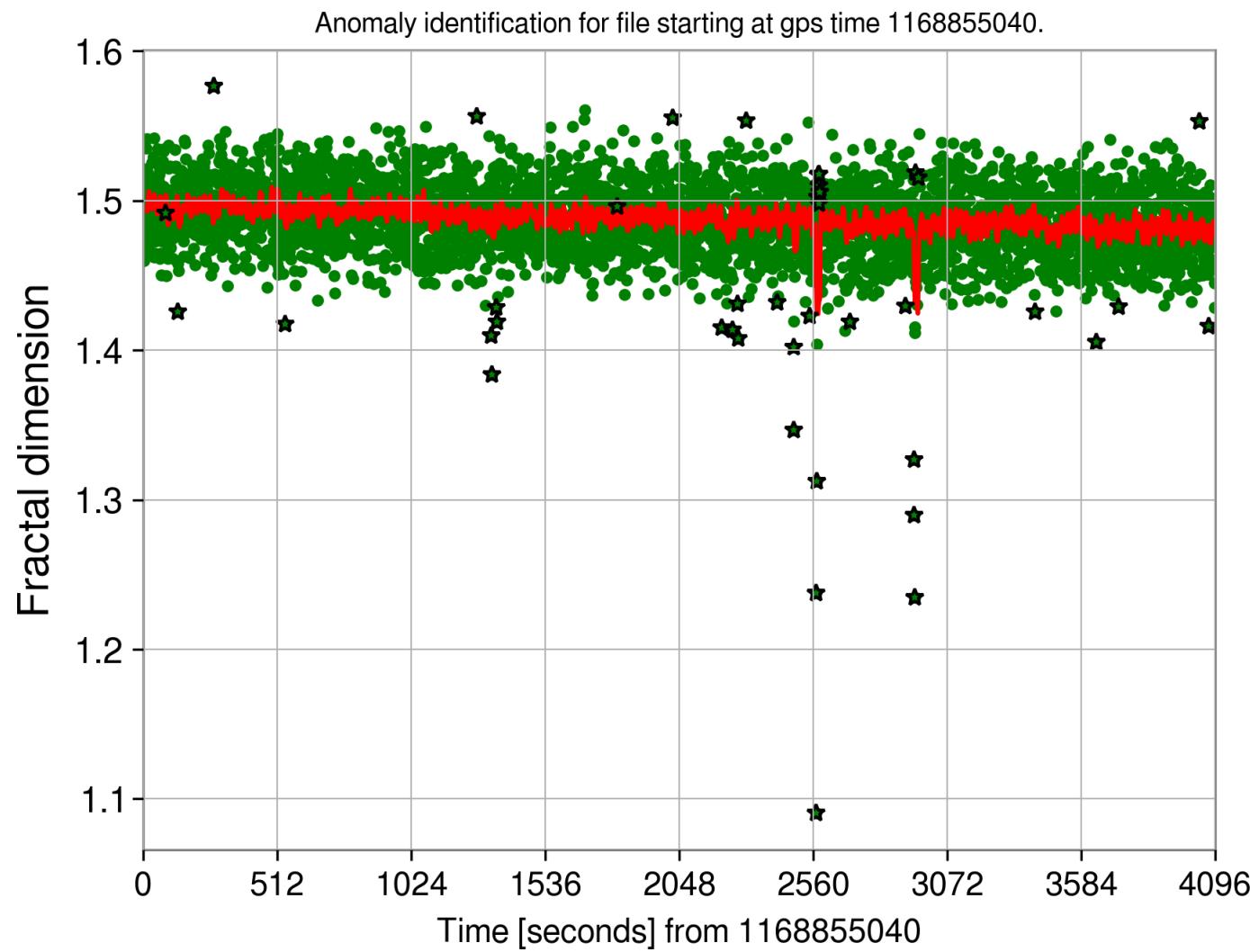


Glitch identification

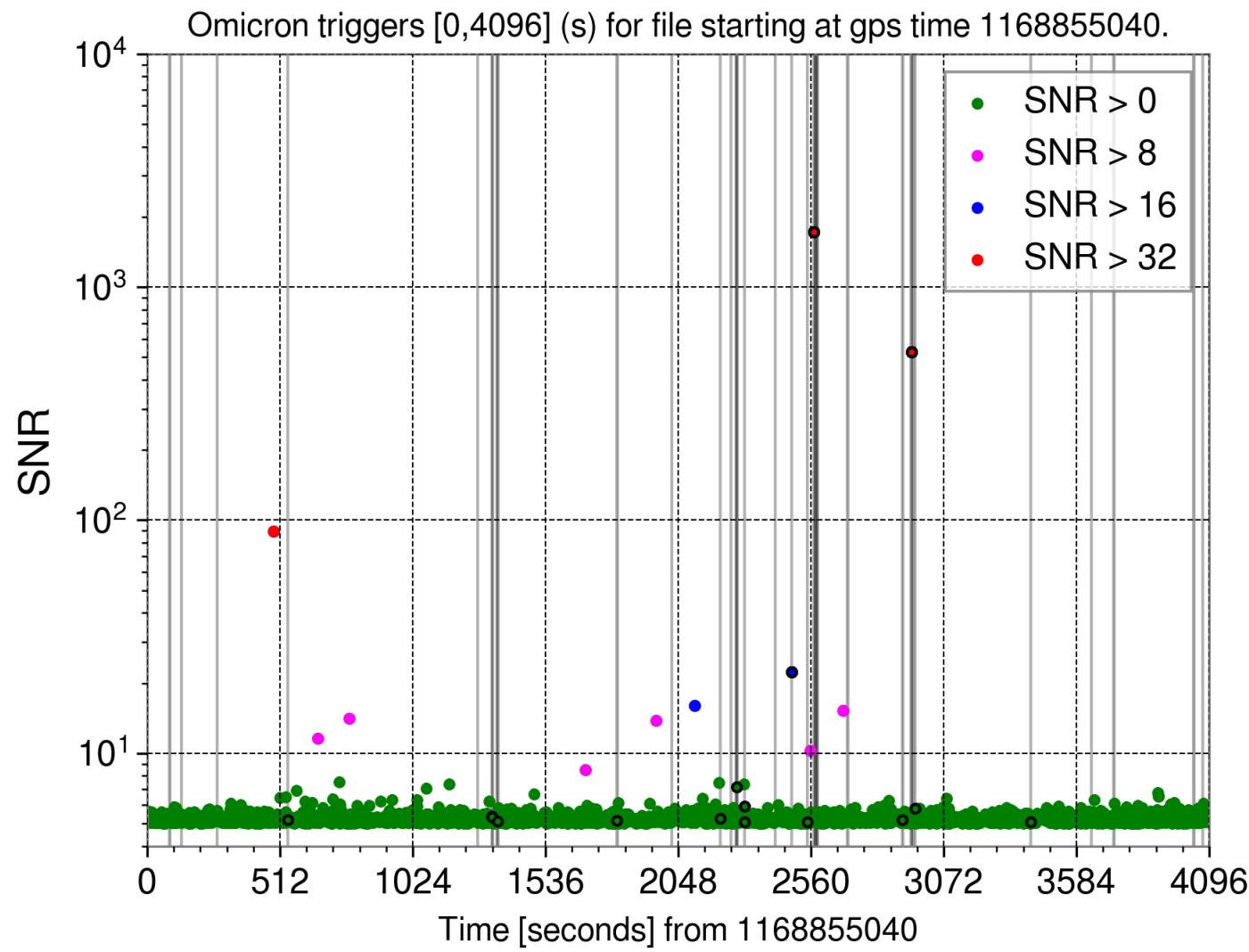


Glitch identification

Rolling mean + anomaly detection with Local Outlier Factor (LOF)

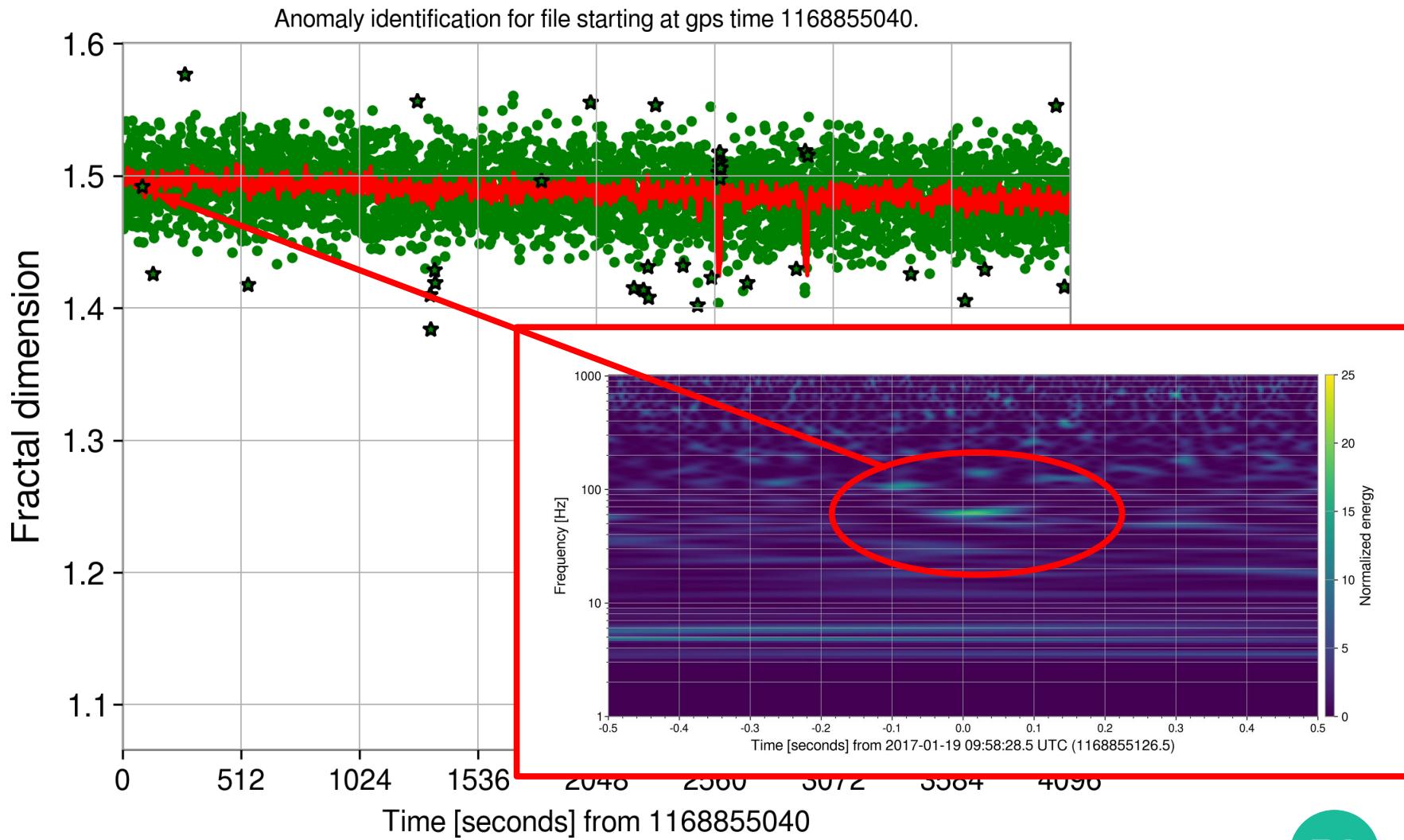


Glitch identification



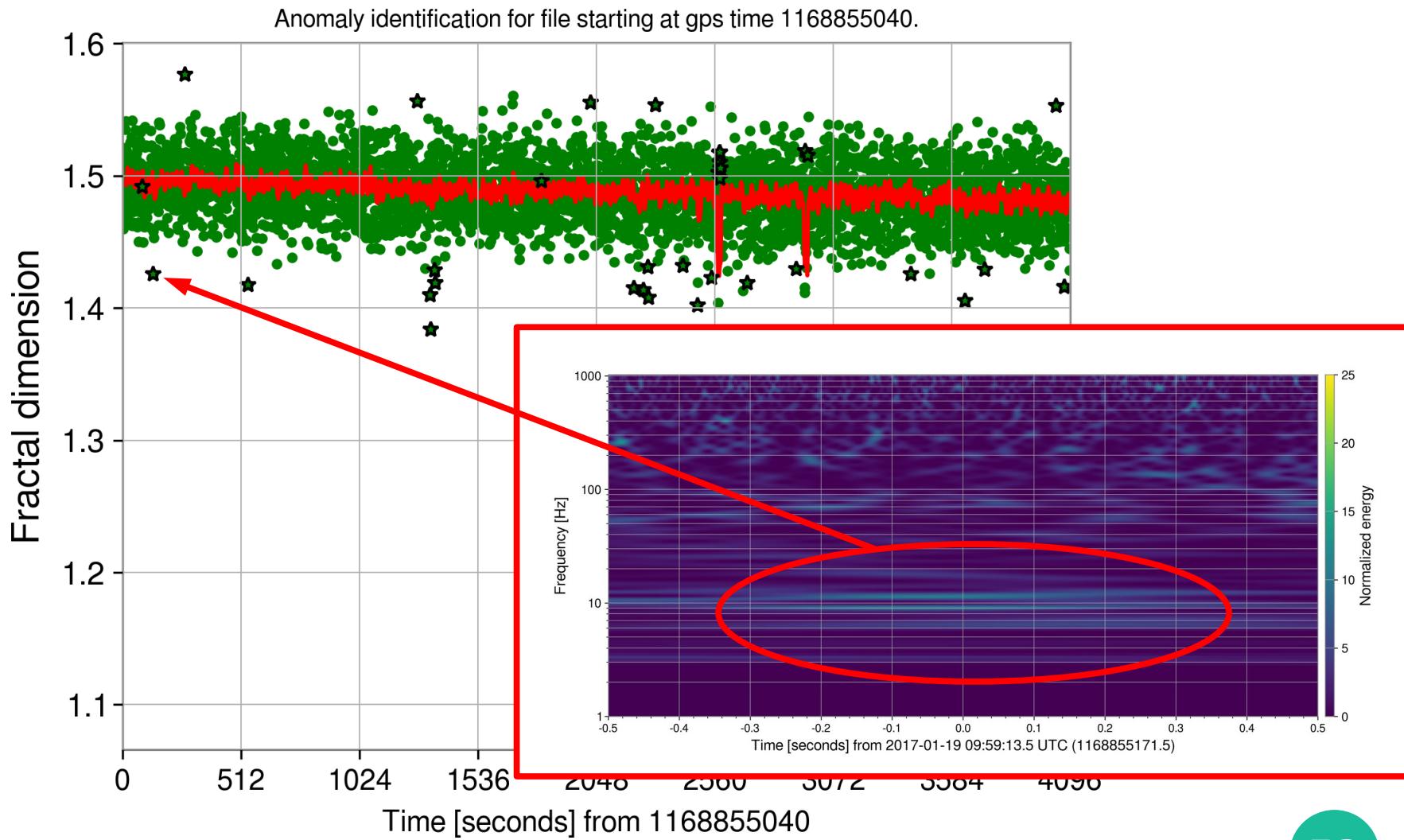
Glitch identification

Rolling mean + anomaly detection with Local Outlier Factor (LOF)

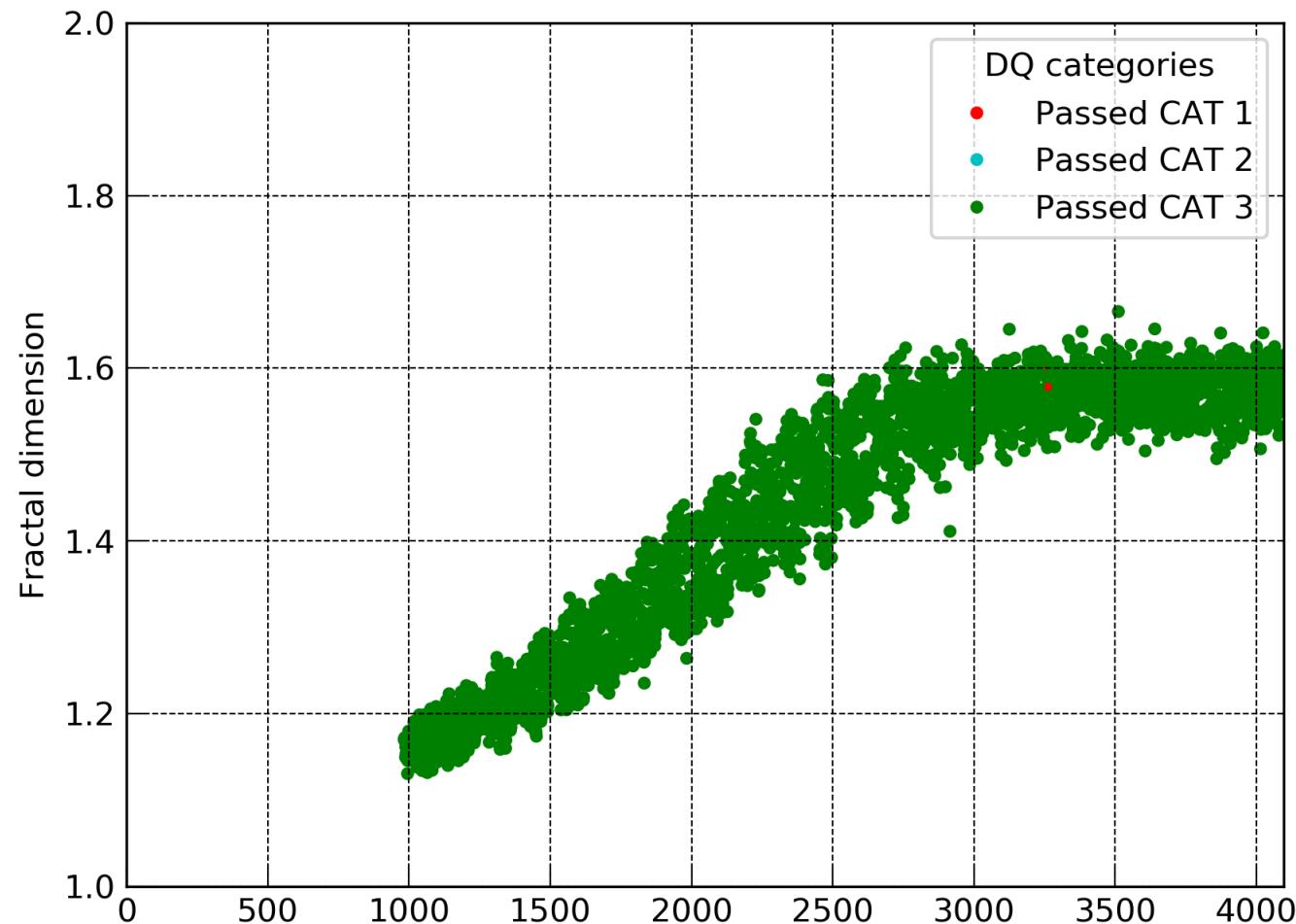


Glitch identification

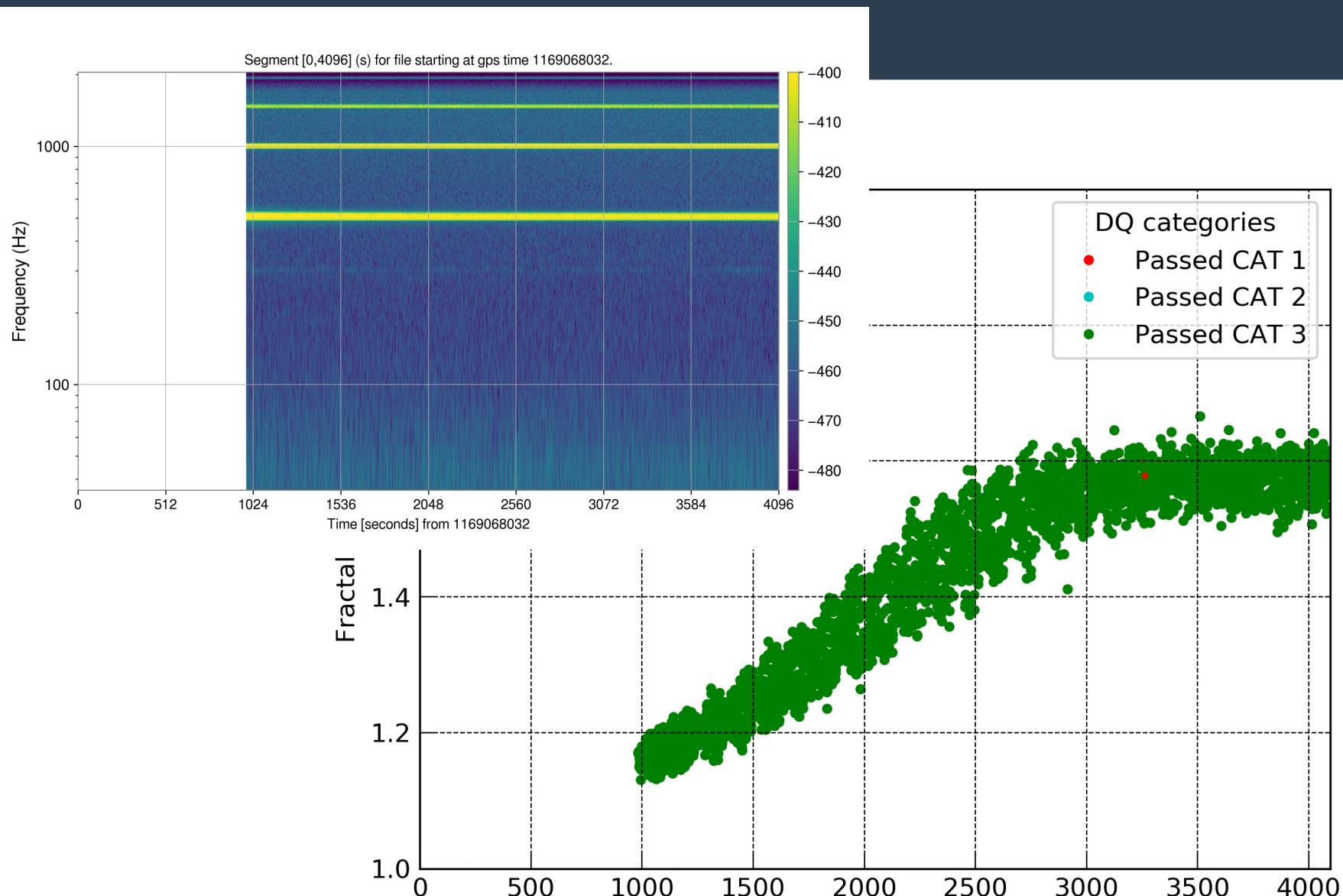
Rolling mean + anomaly detection with Local Outlier Factor (LOF)



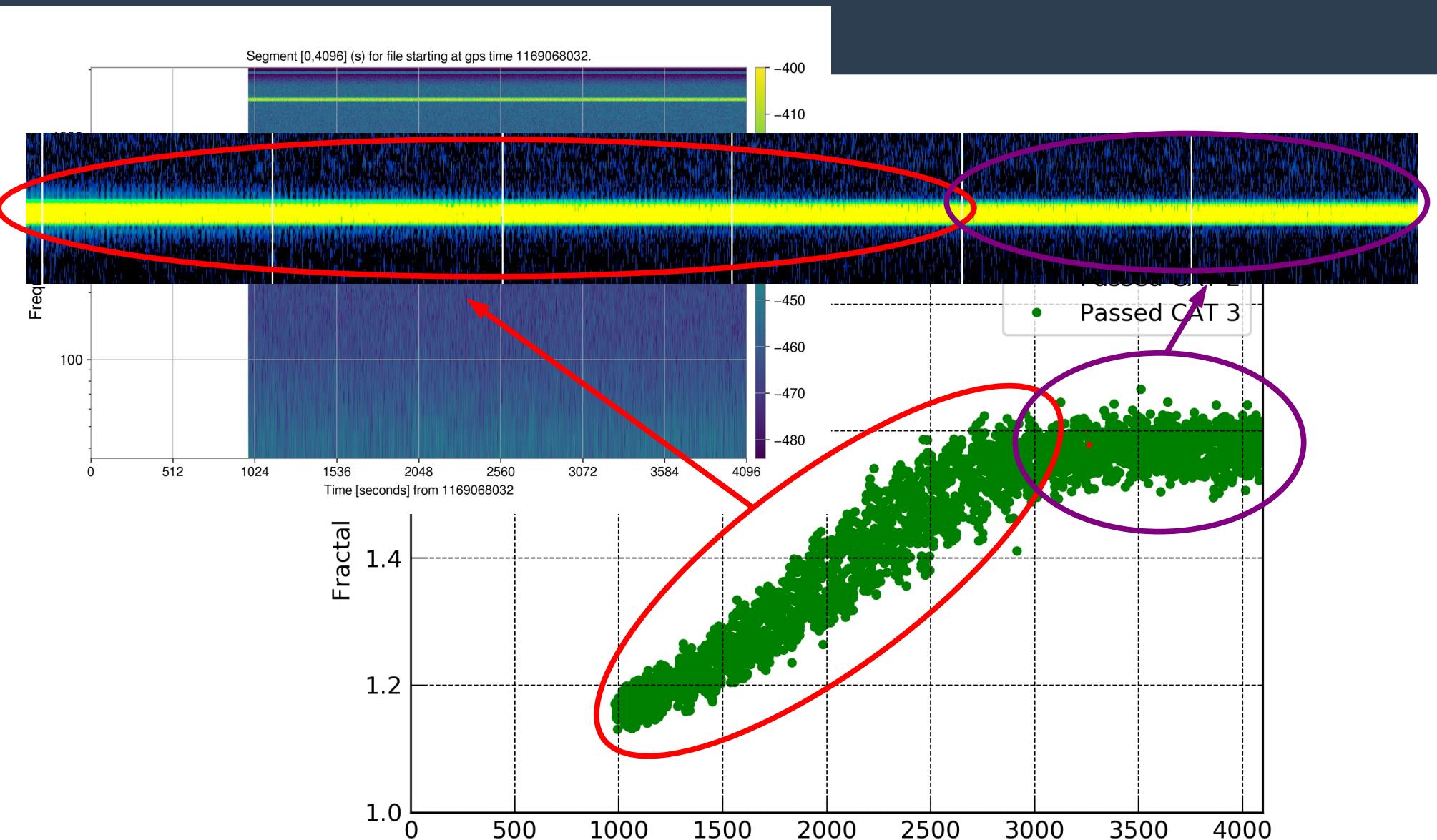
Long-term noise variation



Long-term noise variation



Long-term noise variation



Conclusions

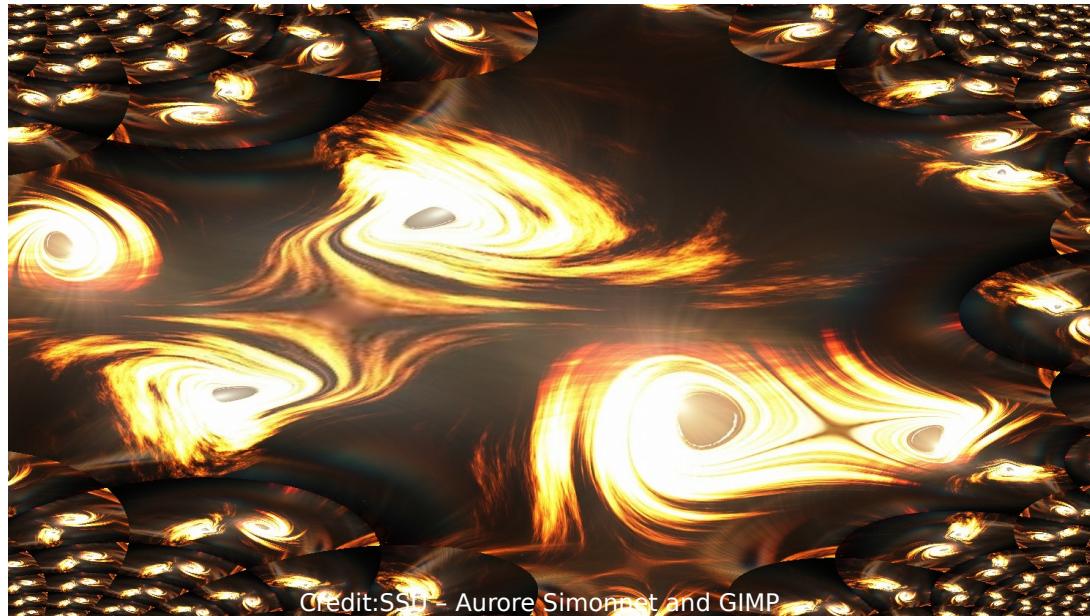
- Simple algorithms that can run on GPUs with numba → speed × 100
- One second of O2 open data at 4kHz decimated by a factor 16 is processed with the VAR algorithm on Caltech's GPU's pcdev11 in less than 0.6 s (including I/O)
- ~ 1/3 of O2 open data processed in a few days:
<https://ldas-jobs.ligo.caltech.edu/~marco.cavaglia/Fractals/>
- Fractal dimension can characterize the noise and can be processed in real time!

Useful references:

- M. Bigerelle, I. Alain, *Fractal dimension and classification of music*, Chaos Solitons & Fractals 11(14):2179-2192 (November 2000), DOI:10.1016/S0960-0779(99)00137-X
- P. Maragos, A. Potamianos, *Fractal dimensions of speech sounds: Computation and application to automatic speech recognition*, The Journal of the Acoustical Society of America 105(3):1925-32 (April 1999) DOI:10.1121/1.426738.
- M. Yamaguti et al. "Mathematics of Fractal" Mathematical monographs, vol.167 AMS 1997.

Thank you! Questions^(†) ?

(†) “If you ask me a question I do not know, I’m not going to answer it”
– Yogi Berra



The author thankfully acknowledges the human and material resources of the LIGO Scientific Collaboration and the Virgo Collaboration that have made possible the results presented in this talk, and the National Science Foundation for its continuous support of LIGO science and basic and applied research in the United States. This work has been partially supported by NSF grant PHY-2011334.

