EDA2 Imager PART B: Gridding and Imaging

Start Date: 06/10/2022End Date: 25/10/2022

Test data used:

• EDA2 Data 2: 20200209

Part 1 of the code: imager_GalacticPlane

- · Calculation of uvw coordinates
- Creation of one uniform df, with all values of real, imaginary visibilities, corresponding to every baseline

Tasks performed:

- Step 1: Initialise a dataframe for gridding: df_uv_grid
- Step 2: Populate the uv grid dataframe with corresponding visibility values, from : df_uvw

Inputs (temporary):

- Excel: Dataframe with antenna positions, uvw coordinates, real and imaginary visibilities
- df_uvw_EDA2_20200209.xlsx

Output (temporary):

• Image of size: (180 x 180)

References:

ANITA Lectures on Imaging, link: https://www.youtube.com/watch?v=mRUZ9eckHZg&t=2315s (https://www.youtube.com/watch?v=mR

Importing required libraies

```
In [1]: # In order to save to .fits file for comparsion
        import astropy.io.fits as pyfits
        # Importing fits library for viewing the images
        from astropy.io import fits
        # For plotting
        import matplotlib.pyplot as plt
        # For plotting 3D plots
        from mpl toolkits.mplot3d import Axes3D
        # Using Pandas
        import pandas as pd
        # In order to use numpy
        import numpy as np
        # For performing math calculations
        import math
        from math import sin as sin
        from math import cos as cos
        from math import pi as pi
        from math import sqrt as sqrt
        # To check the system details
        import sys
        # To check the version of astropy
        import astropy.version
        import time
        from scipy.fft import fft, ifft
```

```
In [2]: # Defining start time
start = time.process_time()
```

Printing versions:

Calculated wavelength in (m): 1.881050716862745

```
In [3]: print('Versions Running on:')
        print(f'\tPython\t\t{sys.version[:31]})')
        print(f'\tAstropy\t\t{astropy. version }')
        print(f'\tMatplotlib\t{plt.matplotlib. version }')
        Versions Running on:
                                3.7.3 (default, Apr 24 2019, 15)
                Python
                Astropy
                                3.2.1
                Matplotlib
                                3.1.0
In [4]: # Defining and calculating all the constants
        # speed of Light : 3 x 10^8 m/s
        c = 299792458.0
        print('Speed of light(m/s):', c)
        # operating frequency = 159.375 MHz
        frequency = 159.375*(10**6)
        print('Operating frequency in (MHz):', frequency)
        # Calculating wavelength
        wavelength = c/frequency
        print('Calculated wavelength in (m):', wavelength)
        Speed of light(m/s): 299792458.0
        Operating frequency in (MHz): 159375000.0
```

Reading in the required inputs

```
In [5]: # Reading the values of real and imaginary visibilities
        vis real = fits.open("chan 204 20200209T034646 vis real.fits")
        vis imag = fits.open("chan 204 20200209T034646 vis imag.fits")
        vis real image = vis real[0].data
        vis imag image = vis imag[0].data
        # Creating a dataframe for all the real and imaginary visibilities
        df vis real = pd.DataFrame(vis real image)
        df vis imag = pd.DataFrame(vis imag image)
In [6]: df vis real.shape
Out[6]: (256, 256)
In [7]: df vis imag.shape
Out[7]: (256, 256)
In [8]: # Viewing the first five rows of the real visibilities
        df_vis_real.head()
Out[8]:
```

	0	1	2	3	4	5	6	7	8	9	 246	247	
0	328.867615	24.588898	21.415720	25.159893	23.272156	32.623756	67.544327	25.600283	-5.008373	56.130409	 20.584955	23.733694	25.835
1	24.588898	1001.324097	32.625900	30.562368	42.107533	67.392723	49.459095	42.223946	30.334806	45.333996	 21.941923	34.037449	69.689
2	21.415720	32.625900	1802.112183	-221.844772	5.765813	75.233185	28.925854	83.027077	53.337502	41.076969	 -11.819213	103.449448	56.257
3	25.159893	30.562368	-221.844772	1310.901611	119.502617	68.835815	46.054825	41.802647	53.148262	54.665218	 96.976486	-2.359920	38.590
4	23.272156	42.107533	5.765813	119.502617	1286.565674	65.288414	111.641373	72.028549	-37.939823	20.661905	 54.904049	21.105185	40.463

5 rows × 256 columns

4

<pre>In [9]: # Viewing the first five rows of the imag visibilities df_vis_imag.head()</pre>																
Out[9]:	_		•													
		0	1		2	3	4	5	6	7	8	9		246	247	248
	0	0.000000	20.347363	5.23	2766 <i>´</i>	17.098120	13.575706	40.088928	-4.494764	36.376900	4.382819	-5.101678		8.985415	0.704096	5.380109
	1	-20.347363	0.000000	-49.23	9357 2	23.557892	-10.475840	-7.355006	-58.698902	19.284822	27.034655	8.961044		-23.801003	-31.874340	-15.966197
	2	-5.232766	49.239357	0.00	0000 -15	58.680908	-65.660355	7.385582	-44.348717	24.389639	97.863960	54.119991		-25.575441	73.233078	62.957054
	3	-17.098120	-23.557892	158.68	0908	0.000000	-37.410690	-3.263887	-64.458542	-7.080525	10.403498	25.782703		-18.212944	-3.222548	-53.895336
	4	-13.575706	10.475840	65.66	0355	37.410690	0.000000	42.297253	-14.743159	37.618340	86.671875	59.724304		5.650125	19.793768	14.020086
	5 rc	ows × 256 c	olumns													
←																
In [10]: # Read Excel file, which was generated from the previous steps																
111 [10].	<pre>In [10]: # Read Excel file, which was generated from the previous steps df_uvw = pd.read_excel('df_uvw_EDA2_20200209.xlsx')</pre>															
In [11]:	dҒ	_uvw[[" <mark>Ant</mark>	enna 1"."	∆ntenn	na 2".")	(lambda"	."Y lambd	a"."7 lam	ıhda" . "ıı" .	"v"."w".'	'vis real"	."vis ima	g"]	 1		
Out[11]:	٠	_avn[[/me			<u></u>		,	u ,	, ,	. , ,	V15 cd1	, 125_2	6 1			
out[II].		Antenn	a_1 Antenr	na_2 X	_lambda	Y_lambda	Z_lambda	u	v	w	vis_real	vis_imag				
		0	0	1 3	3.576193	2.780893	-0.022328	3.576193	2.780893	-0.022328	24.588898	20.347363				
		1	0	2 -(0.195104	1.943063	-0.013290	-0.195104	1.943063	-0.013290	21.415720	5.232766				
		2	0	3 (0.587969	2.073309	-0.006911	0.587969	2.073309	-0.006911	25.159893	17.098120				
		3	0	4 3	3.615001	0.913851	-0.008506	3.615001	0.913851	-0.008506	23.272156	13.575706				
		4	0	5 2	2.433215	-2.140825	0.005848	2.433215	-2.140825	0.005848	32.623756	40.088928				
	65	275	255	250 -	1.115334	1.952632	0.007443	-1.115334	1.952632	0.007443	0.000000	0.000000				
	65	276	255	251 -(0.027644	1.032402	0.002126	-0.027644	1.032402	0.002126	0.000000	0.000000				

1.314159 -0.979772

0.024454 -0.611892 -1.671938

0.008506 -0.687382 -0.392334

0.011696

0.024454

0.008506

0.000000

0.000000

0.000000

0.000000

0.000000

0.000000

65280 rows × 10 columns

255

255

255

252

1.314159 -0.979772

-0.611892 -1.671938

-0.687382 -0.392334

0.011696

Notes:

65277

65278

65279

- · Conjugate values have been included.
- (Code has been written in EDA2 Imager Part A to include the conjugate values. The execution took almost 10-20 minutes.)

```
In [12]: # Defining numpy arrays for u, v and w coordinates
         u python = np.zeros((256, 256), dtype=float)
         v python = np.zeros((256, 256), dtype=float)
         w_python = np.zeros((256, 256), dtype=float)
In [13]: # Generating u, v and w coordinates for comparison with C++ imager
         # Iterate through all the rows of df_uvw
         for index,row in df uvw.iterrows():
             u = row['u m']
             v = row['v m']
             w = row['z m']
             u index = int(row['Antenna 1'])
             v index = int(row['Antenna 2'])
             # Assigning the u, v and w coordinates
             u python[u index, v index] = u
             v python[u index, v index] = v
             w python[u index, v index] = w
In [14]: # # Converting to .fits files
         # hdu = pyfits.PrimaryHDU()
         # hdu.data = u python
         # # Copying into .fits files
         # hdulist = pyfits.HDUList([hdu])
         # hdulist.writeto('u python.fits',clobber=True)
         # hdu = pyfits.PrimaryHDU()
         # hdu.data = v python
         # # Copying into .fits files
         # hdulist = pyfits.HDUList([hdu])
         # hdulist.writeto('v python.fits',clobber=True)
         # hdu = pyfits.PrimaryHDU()
         # hdu.data = w python
         # # Copying into .fits files
         # hdulist = pyfits.HDUList([hdu])
         # hdulist.writeto('w python.fits',clobber=True)
```

```
In [15]: # Reading the values of real and imaginary visibilities (Marcin's u.fits and v.fits)
u = fits.open("u_python.fits")
v = fits.open("v_python.fits")

u_image = u[0].data
v_image = v[0].data

# Creating a dataframe for all u, v .fits files
df_u = pd.DataFrame(u_image)
df_v = pd.DataFrame(v_image)

# These visibilities are in meters! Need to divide them by wavelengths
```

In [16]: df_u

Out[16]:

	0	1	2	3	4	5	6	7	8	9	 246	247	248	249	250	251	252	253	2!
0	0.000	6.727	-0.367	1.106	6.800	4.577	5.514	4.665	2.137	2.173	 -4.203	-5.070	-7.659	-8.616	-8.269	-6.223	-3.699	-7.322	-7.46
1	-6.727	0.000	-7.094	-5.621	0.073	-2.150	-1.213	-2.062	-4.590	-4.554	 -10.930	-11.797	-14.386	-15.343	-14.996	-12.950	-10.426	-14.049	-14.19
2	0.367	7.094	0.000	1.473	7.167	4.944	5.881	5.032	2.504	2.540	 -3.836	-4.703	-7.292	-8.249	-7.902	-5.856	-3.332	-6.955	-7.09
3	-1.106	5.621	-1.473	0.000	5.694	3.471	4.408	3.559	1.031	1.067	 -5.309	-6.176	-8.765	-9.722	-9.375	-7.329	-4.805	-8.428	-8.57
4	-6.800	-0.073	-7.167	-5.694	0.000	-2.223	-1.286	-2.135	-4.663	-4.627	 -11.003	-11.870	-14.459	-15.416	-15.069	-13.023	-10.499	-14.122	-14.26
251	6.223	12.950	5.856	7.329	13.023	10.800	11.737	10.888	8.360	8.396	 2.020	1.153	-1.436	-2.393	-2.046	0.000	2.524	-1.099	-1.24
252	3.699	10.426	3.332	4.805	10.499	8.276	9.213	8.364	5.836	5.872	 -0.504	-1.371	-3.960	-4.917	-4.570	-2.524	0.000	-3.623	-3.76
253	7.322	14.049	6.955	8.428	14.122	11.899	12.836	11.987	9.459	9.495	 3.119	2.252	-0.337	-1.294	-0.947	1.099	3.623	0.000	-0.14
254	7.464	14.191	7.097	8.570	14.264	12.041	12.978	12.129	9.601	9.637	 3.261	2.394	-0.195	-1.152	-0.805	1.241	3.765	0.142	0.00
255	6.171	12.898	5.804	7.277	12.971	10.748	11.685	10.836	8.308	8.344	 1.968	1.101	-1.488	-2.445	-2.098	-0.052	2.472	-1.151	-1.29

256 rows × 256 columns

4

In [17]: df_v

Out[17]:

	0	1	2	3	4	5	6	7	8	9	 246	247	248	249	250	251	252	253	254	255
0	0.000	5.231	3.655	3.900	1.719	-4.027	-1.129	0.726	1.693	-0.977	 5.754	3.951	5.618	6.812	8.551	6.820	3.035	1.733	4.140	4.878
1	-5.231	0.000	-1.576	-1.331	-3.512	-9.258	-6.360	-4.505	-3.538	-6.208	 0.523	-1.280	0.387	1.581	3.320	1.589	-2.196	-3.498	-1.091	-0.353
2	-3.655	1.576	0.000	0.245	-1.936	-7.682	-4.784	-2.929	-1.962	-4.632	 2.099	0.296	1.963	3.157	4.896	3.165	-0.620	-1.922	0.485	1.223
3	-3.900	1.331	-0.245	0.000	-2.181	-7.927	-5.029	-3.174	-2.207	-4.877	 1.854	0.051	1.718	2.912	4.651	2.920	-0.865	-2.167	0.240	0.978
4	-1.719	3.512	1.936	2.181	0.000	-5.746	-2.848	-0.993	-0.026	-2.696	 4.035	2.232	3.899	5.093	6.832	5.101	1.316	0.014	2.421	3.159
251	-6.820	-1.589	-3.165	-2.920	-5.101	-10.847	-7.949	-6.094	-5.127	-7.797	 -1.066	-2.869	-1.202	-0.008	1.731	0.000	-3.785	-5.087	-2.680	-1.942
252	-3.035	2.196	0.620	0.865	-1.316	-7.062	-4.164	-2.309	-1.342	-4.012	 2.719	0.916	2.583	3.777	5.516	3.785	0.000	-1.302	1.105	1.843
253	-1.733	3.498	1.922	2.167	-0.014	-5.760	-2.862	-1.007	-0.040	-2.710	 4.021	2.218	3.885	5.079	6.818	5.087	1.302	0.000	2.407	3.145
254	-4.140	1.091	-0.485	-0.240	-2.421	-8.167	-5.269	-3.414	-2.447	-5.117	 1.614	-0.189	1.478	2.672	4.411	2.680	-1.105	-2.407	0.000	0.738
255	-4.878	0.353	-1.223	-0.978	-3.159	-8.905	-6.007	-4.152	-3.185	-5.855	 0.876	-0.927	0.740	1.934	3.673	1.942	-1.843	-3.145	-0.738	0.000

256 rows × 256 columns

Gridding

Theory:

• Visibilities need to be "gridded" on a regularly spaced grid, in-order to perform a fourier transform on them.

Steps involved:

Step 1:

- Initialise two 2D grids, one for each real and imaginary visibilities with 0 values to begin with.
- The grid cell size depends on:
 - (Nx, Ny) : Image size
 - (delta_x, delta_y): pixel size of the image, which you get from the Full Width Half Max(FWHM)
 - delta_x = 1/3 x angular resolution(theta_degrees)
 - delta_y = 1/3 x angular resolution(theta_degrees)
 - lambda/D : angular resolution
 - lambda: which you get from operating frequency
 - **D**: Maximum baseline length

- delta_u : 1/(Nx*(delta_x)) : Dependent on Nx and delta_x
- delta_v : 1/(Ny*(delta_y)) : Dependent on Ny and delta y
- If Nx = Ny and delta_x = delta_y, then:
 - o delta_u = delta_v

Step 2:

- Populate the grid with visibility values, based on u and v coordinates
- Think about, how you are going to handle the weighting aspect, to start with keep it natural weighting!

References:

- Fundamentals of Radio Astronomy, Pg 266 onwards
- ANITA Lectures on Imaging and Deconvolution

Visualising gridded visibilities:

In [18]: # Viewing the first five rows of the
df_uvw

Out[18]:

	Antenna_1	Antenna_2	X_lambda	Y_lambda	Z_lambda	u	V	w	u_m	v_m	z_m	vis_real	vis_imag
0	0	1	3.576193	2.780893	-0.022328	3.576193	2.780893	-0.022328	6.727	5.231	-0.042	24.588898	20.347363
1	0	2	-0.195104	1.943063	-0.013290	-0.195104	1.943063	-0.013290	-0.367	3.655	-0.025	21.415720	5.232766
2	0	3	0.587969	2.073309	-0.006911	0.587969	2.073309	-0.006911	1.106	3.900	-0.013	25.159893	17.098120
3	0	4	3.615001	0.913851	-0.008506	3.615001	0.913851	-0.008506	6.800	1.719	-0.016	23.272156	13.575706
4	0	5	2.433215	-2.140825	0.005848	2.433215	-2.140825	0.005848	4.577	-4.027	0.011	32.623756	40.088928
65275	255	250	-1.115334	1.952632	0.007443	-1.115334	1.952632	0.007443	-2.098	3.673	0.014	0.000000	0.000000
65276	255	251	-0.027644	1.032402	0.002126	-0.027644	1.032402	0.002126	-0.052	1.942	0.004	0.000000	0.000000
65277	255	252	1.314159	-0.979772	0.011696	1.314159	-0.979772	0.011696	2.472	-1.843	0.022	0.000000	0.000000
65278	255	253	-0.611892	-1.671938	0.024454	-0.611892	-1.671938	0.024454	-1.151	-3.145	0.046	0.000000	0.000000
65279	255	254	-0.687382	-0.392334	0.008506	-0.687382	-0.392334	0.008506	-1.293	-0.738	0.016	0.000000	0.000000

65280 rows × 13 columns

```
In [19]: # Here, u v and w are in wavelengths
         u list = df uvw["u"].values.tolist()
         v list = df uvw["v"].values.tolist()
         w list = df uvw["w"].values.tolist()
In [20]: print('min(u_list) and max(u_list):', round(min(u_list),2),',', round(max(u_list),2))
         print('min(v list) and max(v list) :', round(min(v list),2),',', round(max(v list),2))
         print('min(w list) and max(w list) :', round(min(w_list),2),',', round(max(w_list),2))
         min(u list) and max(u list): -18.53, 18.53
         min(v list) and max(v list): -18.58, 18.58
         min(w list) and max(w list) : -0.11 , 0.11
In [21]: df uvw.shape
Out[21]: (65280, 13)
In [22]: df uvw.count()
Out[22]: Antenna 1
                      65280
         Antenna 2
                      65280
         X lambda
                      65280
         Y lambda
                      65280
         Z lambda
                      65280
                      65280
         u
                      65280
                      65280
                      65280
         u m
                      65280
         v_m
                      65280
         z m
         vis real
                      65280
                      65280
         vis imag
         dtype: int64
```

```
In [23]: df uvw.info()
         <class 'pandas.core.frame.DataFrame'>
        RangeIndex: 65280 entries, 0 to 65279
        Data columns (total 13 columns):
         # Column
                        Non-Null Count Dtype
             -----
         0 Antenna 1 65280 non-null int64
             Antenna 2 65280 non-null int64
             X lambda
                        65280 non-null float64
         3
             Y lambda
                        65280 non-null float64
             Z lambda
                        65280 non-null float64
         5
                        65280 non-null float64
             u
                        65280 non-null float64
         6
```

65280 non-null float64

dtypes: float64(11), int64(2)

memory usage: 6.5 MB

7

8

9

W

u m

v m 10 z m

11 vis real

12 vis imag

```
In [24]: # Defining maximum baseline in meters, based on EDA2
         # Can calculate this value later based on x and y coordinates
         max baseline = 35
         print('Maximum baseline (m):', max baseline)
         # Calculating maximum angular resolution in radians
         theta radians = wavelength/max baseline
         print('theta radians:', round(theta radians, 4))
         # Calculating delta x : angular width
         delta x = (1/3)*(theta radians)
         print('delta x:', round(delta x,4))
         # Calculating delta y : angular height
         delta y = (1/3)*(theta radians)
         print('delta y:', round(delta y,4))
         # Assigning Nx
         Nx = 180
         print('Nx:', Nx)
         # Assigning Ny
         Ny = 180
         print('Ny:', Nx)
         # Source: Taken from Marcin's C++ code for gridding()
         # Calculating centre positions of the grid
         centre x = int(Nx/2)
         centre y = int(Ny/2)
         # For images, with even dimensions, default values = 0
         is odd x = 0
         is odd y = 0
         # For images, with odd dimensions, setting values to 1
         if(Nx\%2 ==1):
             is odd x = 1
         if(Ny\%2 ==1):
             is odd y = 1
         # Printing out the values
         print("centre_x =", centre_x)
         print("centre_x =", centre_x)
         print("is odd x =", is odd x)
         print("is odd y =", is odd y)
```

```
# Calculating delta u
delta u = 1/(Nx*(delta x))
print('delta u:', round(delta u,2))
# Calculate delta v
delta v = 1/(Nv*(delta v))
print('delta v:', round(delta v, 2))
# u max, u min values
u max = round(max(u list),2)
u min = round(min(u list),2)
print("u(min, max):", u min,",", u max)
# v max, v min values
v max = round(max(v list), 2)
v min = round(min(v list),2)
print("v(min, max):", v min,",", v max)
Maximum baseline (m): 35
theta radians: 0.0537
delta x: 0.0179
delta v: 0.0179
Nx: 180
Nv: 180
```

Gridding:

centre_x = 90
centre_x = 90
is_odd_x = 0
is_odd_y = 0
delta_u: 0.31
delta v: 0.31

u(min, max): -18.53 , 18.53 v(min, max): -18.58 , 18.58

```
In [25]: # Initialising an uv_grid_counter, to count the number of visibilities stored in every grid
uv_grid_counter = np.zeros((Nx, Ny), dtype=int)

# Initialising an uv_grids for real and imaginary visibilities, as per the image size
uv_grid_real = np.zeros((Nx, Ny), dtype=float)
uv_grid_imag = np.zeros((Nx, Ny), dtype=float)
uv_grid = np.zeros((Nx, Ny), dtype=complex)
```

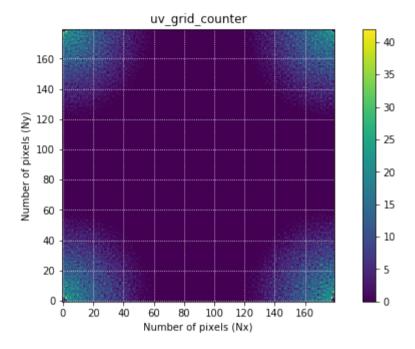
```
In [30]: count = 0
         grid count = 0
         # Defining start time
         start gridding = time.process time()
         # Iterate through all the rows of df uvw, total: 256x255 iterations: 65280
         for index,row in df uvw.iterrows():
             count += 1
             # Get the u,v coordinates in wavelengths only
             u = row['u']
             v = row['v']
             # Get the antenna numbers
             a1 = row['Antenna 1']
             a2 = row['Antenna 2']
             vis real = row['vis real']
             vis_imag = row['vis_imag']
             # Fix your u and v coordinates
             u pix = round(u/delta u)
             v pix = round(v/delta v)
             # Calculating u_index, v_index for positive values,
             u index = int(u pix + Nx/2)
             v index = int(v pix + Ny/2)
             # Setting initial values of x grid, y grid
             u grid = 0
             v grid = 0
             # Doing an FFT-UNSHIFT, moving from centre to corners
             # u index < centre x, the index is pushed away further towards the corners, in the right half of the grid
             # u index > center x, the index is pushed away further towards the corners, in the left half of the grid
             if(u index < centre x):</pre>
                 u_grid = u_index + centre_x
             else:
                 u grid = u index - centre x
             if(v_index < centre_y):</pre>
                 v grid = v index + centre y
             else:
```

```
v grid = v index - centre x
    # Populate the uv grid, with Non NaN visibilities
    # If grid is empty, assign the visibility, else assign by adding it with the existing value
    if (math.isnan(vis real)==False) and (math.isnan(vis imag)==False):
        # Assigning real visibility
        uv grid real[u grid, v grid] += vis real
        # Assigning imag visibility
        uv grid imag[u grid, v grid] += vis imag
        uv grid[u grid,v grid] += vis real + 1j*vis imag
        # Incrementing the value in uv grid counter
        uv grid counter[u grid, v grid] += 1
        grid count += 1
        print("Baseline",count,"[ grid count:", grid count,"A1:", a1,"A2:",a2,"u:",round(u,2),"v:",round(v,2),"real:",round(vis
# Overall time taken in seconds
print("gridding A took:",time.process time() - start gridding)
Baseline 1 [ grid count: 1 A1: 0.0 A2: 1.0 u: 3.58 v: 2.78 real: 24.59 imag: 20.35 u index: 12 v index: 9
Baseline 2 [ grid count: 2 A1: 0.0 A2: 2.0 u: -0.2 v: 1.94 real: 21.42 imag: 5.23 u index: 179 v index: 6
Baseline 3 [ grid count: 3 A1: 0.0 A2: 3.0 u: 0.59 v: 2.07 real: 25.16 imag: 17.1 u index: 2 v index: 7
Baseline 4 [ grid count: 4 A1: 0.0 A2: 4.0 u: 3.62 v: 0.91 real: 23.27 imag: 13.58 u index: 12 v index: 3
Baseline 5 [ grid count: 5 A1: 0.0 A2: 5.0 u: 2.43 v: -2.14 real: 32.62 imag: 40.09 u index: 8 v index: 173
Baseline 6 [ grid count: 6 A1: 0.0 A2: 6.0 u: 2.93 v: -0.6 real: 67.54 imag: -4.49 u index: 9 v index: 178
Baseline 7 [ grid count: 7 A1: 0.0 A2: 7.0 u: 2.48 v: 0.39 real: 25.6 imag: 36.38 u index: 8 v index: 1
Baseline 8 [ grid count: 8 A1: 0.0 A2: 8.0 u: 1.14 v: 0.9 real: -5.01 imag: 4.38 u index: 4 v index: 3
Baseline 9 [ grid count: 9 A1: 0.0 A2: 9.0 u: 1.16 v: -0.52 real: 56.13 imag: -5.1 u index: 4 v index: 178
Baseline 10 [ grid count: 10 A1: 0.0 A2: 10.0 u: 0.41 v: -0.75 real: -17.87 imag: 13.12 u index: 1 v index: 178
Baseline 11 [ grid count: 11 A1: 0.0 A2: 11.0 u: 1.92 v: -0.75 real: 32.19 imag: 9.37 u index: 6 v index: 178
Baseline 12 [ grid count: 12 A1: 0.0 A2: 12.0 u: 2.83 v: 1.07 real: 25.65 imag: 10.83 u index: 9 v index: 3
Baseline 13 [ grid count: 13 A1: 0.0 A2: 13.0 u: 3.17 v: 1.74 real: 20.28 imag: -12.11 u index: 10 v index: 6
Baseline 14 [ grid count: 14 A1: 0.0 A2: 14.0 u: 2.07 v: 1.77 real: 32.97 imag: 10.98 u index: 7 v index: 6
Baseline 15 [ grid count: 15 A1: 0.0 A2: 15.0 u: 1.47 v: 0.16 real: 28.74 imag: 54.31 u index: 5 v index: 1
Baseline 16 [ grid count: 16 A1: 0.0 A2: 16.0 u: 3.7 v: -1.72 real: 10.1 imag: 4.52 u index: 12 v index: 174
Baseline 17 [ grid count: 17 A1: 0.0 A2: 17.0 u: 4.01 v: -0.79 real: 37.74 imag: 15.99 u index: 13 v index: 177
Baseline 18 [ grid count: 18 A1: 0.0 A2: 18.0 u: 4.91 v: 2.11 real: 0.95 imag: 10.05 u index: 16 v index: 7
Baseline 19 [ grid count: 19 A1: 0.0 A2: 19.0 u: 4.93 v: 0.7 real: 34.49 imag: 8.6 u index: 16 v index: 2
Daralina 20 [ amid accept. 20 11. 0 0 12. 20 0 ... 7 00 ... 1 02 marl. 10 16 imag. 7 17 ...
```

```
In [31]: # To check the number of visibilities getting gridded:
         uv grid counter.sum()
Out[31]: 65280
In [32]: hdu = pyfits.PrimaryHDU()
         hdu.data = uv grid counter
         # Copying into .fits files
         hdulist = pyfits.HDUList([hdu])
         hdulist.writeto('uv grid counter python.fits',overwrite=True)
In [33]: hdu = pyfits.PrimaryHDU()
         hdu.data = uv grid real
         # Copying into .fits files
         hdulist = pyfits.HDUList([hdu])
         hdulist.writeto('uv grid real python.fits',overwrite=True)
In [34]: hdu = pyfits.PrimaryHDU()
         hdu.data = uv grid imag
         # Copying into .fits files
         hdulist = pyfits.HDUList([hdu])
         hdulist.writeto('uv grid imag python.fits',overwrite=True)
In [35]: # hdu = pyfits.PrimaryHDU()
         # hdu.data = uv grid
         # # Copying into .fits files
         # hdulist = pyfits.HDUList([hdu])
         # hdulist.writeto('uv grid python.fits',overwrite=True)
```

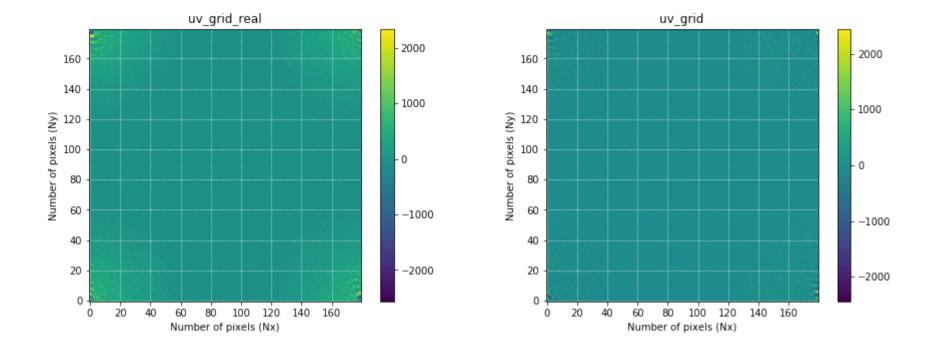
In [36]: # create figure fig = plt.figure(figsize=(10,5)) plt.imshow(uv_grid_counter, origin ='lower') plt.grid(color='white', ls='dotted') plt.xlabel("Number of pixels (Nx)") plt.ylabel("Number of pixels (Ny)") plt.title("uv_grid_counter") plt.colorbar()

Out[36]: <matplotlib.colorbar.Colorbar at 0x231ce454ef0>



```
In [37]: # create figure
         fig = plt.figure(figsize=(15, 5))
         # setting values to rows and column variables
         rows = 1
         columns = 2
         # create figure
         fig.add_subplot(rows, columns, 1)
         plt.imshow(uv grid real, origin ='lower')
         plt.grid(color='white', ls='dotted')
         plt.xlabel("Number of pixels (Nx)")
         plt.ylabel("Number of pixels (Ny)")
         plt.title("uv grid real")
         plt.colorbar()
         # create figure
         fig.add subplot(rows, columns, 2)
         plt.imshow(uv grid imag, origin ='lower')
         plt.grid(color='white', ls='dotted')
         plt.xlabel("Number of pixels (Nx)")
         plt.ylabel("Number of pixels (Ny)")
         plt.title("uv grid")
         plt.colorbar()
```

Out[37]: <matplotlib.colorbar.Colorbar at 0x231cf8512e8>



```
In [38]: hdu = pyfits.PrimaryHDU()
hdu.data = uv_grid_real
# Copying into .fits files
hdulist = pyfits.HDUList([hdu])
hdulist.writeto('uv_grid_real_python.fits',overwrite=True)
In [39]: hdu = pyfits.PrimaryHDU()
```

```
In [39]: hdu = pyfits.PrimaryHDU()
    hdu.data = uv_grid_imag
    # Copying into .fits files
    hdulist = pyfits.HDUList([hdu])
    hdulist.writeto('uv_grid_imag_python.fits',overwrite=True)
```

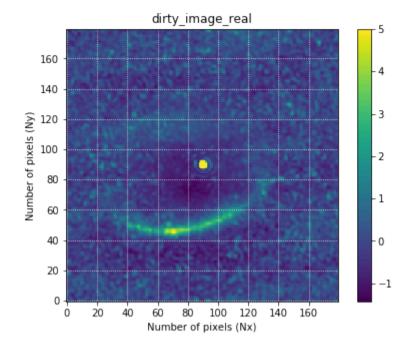
dirty_image (after FFT-SHIFT)

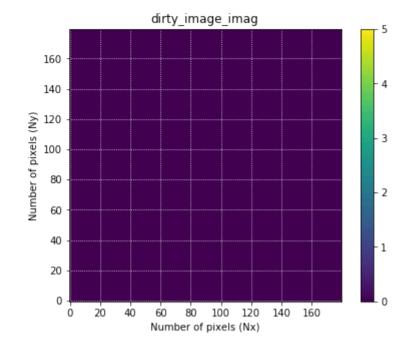
- dirty_image_real
- dirty_image_imag
- dirty_image

```
In [40]: # uv_grid, with complex data-type?
dirty_image = np.fft.ifftshift(np.fft.ifft2(uv_grid))
```

```
In [44]: # create figure
         fig = plt.figure(figsize=(15, 5))
         # setting values to rows and column variables
         rows = 1
         columns = 2
         # create figure
         fig.add_subplot(rows, columns, 1)
         plt.imshow(dirty image real, origin ='lower', vmax=5)
         plt.grid(color='white', ls='dotted')
         plt.xlabel("Number of pixels (Nx)")
         plt.ylabel("Number of pixels (Ny)")
         plt.title("dirty image real")
         plt.colorbar()
         # create figure
         fig.add subplot(rows, columns, 2)
         plt.imshow(dirty_image_imag, origin ='lower', vmax=5)
         plt.grid(color='white', ls='dotted')
         plt.xlabel("Number of pixels (Nx)")
         plt.ylabel("Number of pixels (Ny)")
         plt.title("dirty image imag")
         plt.colorbar()
```

Out[44]: <matplotlib.colorbar.Colorbar at 0x231cf65ac88>





```
In [45]: hdu = pyfits.PrimaryHDU()
hdu.data = dirty_image_real
# Copying into .fits files
hdulist = pyfits.HDUList([hdu])
hdulist.writeto('dirty_image_real_python.fits',overwrite=True)
```

```
In [46]: hdu = pyfits.PrimaryHDU()
    hdu.data = dirty_image_imag
    # Copying into .fits files
    hdulist = pyfits.HDUList([hdu])
    hdulist.writeto('dirty_image_imag_python.fits',overwrite=True)
```

```
In [47]: # Overall time taken in seconds
print(time.process_time() - start)
196.03125
```

In [48]: # Extra code

```
In [49]: # # Defining start time
          # start gridding = time.process time()
          \# count = 0
          # grid count = 0
         # # Outer loop, i for Antenna 1, i.e rows
          # for a1 in antenna list:
                # Inner Loop, j for Antenna 2, i.e columns
          #
                for a2 in antenna list:
                    if(a1!=a2):
          #
                             count += 1
                            # Getting the u,v coordinates for a1,a2
                            u = df u.iloc[a1][a2]
                            v = df u.iloc[a1][a2]
                             if(math.isnan(u)==False) and (math.isnan(v)==False):
                                 # Calculating the u, v coordinates in lambda/wavelengths
                                 u lambda = u/wavelength
                                 ν lambda = ν/wavelength
                                 # Get the corresponding real and imag visibility value for that antenna pair
                                 vis real = df vis real.iloc[a1][a2]
                                 vis imag = df vis imag.iloc[a1][a2]
                                 # Fix your u and v coordinates
                                 u pix = round(u lambda/delta u)
                                 v pix = round(v lambda/delta v)
                                 u_index = int(u_pix + Nx/2)
                                 v index = int(v pix + Ny/2)
                                 u grid = 0
                                 v \text{ arid} = 0
                                 # Doing an FFT-UNSHIFT, moving from centre to corners
                                 if(u index < centre x):</pre>
                                     u \text{ grid} = u \text{ index} + \text{centre } x
                                 else:
                                     u grid = u index - centre x
                                 if(v index < centre y):</pre>
                                     v_grid = v_index + centre_y
                                 else:
```

```
v grid = v index - centre x
#
                       # Populate the uv grid
#
                       if (math.isnan(vis real)==False) and (math.isnan(vis imag)==False):
#
                            uv grid real[u grid, v grid] += vis real
                           uv grid imag[u grid, v grid] += vis imag
                           uv grid counter[u grid, v grid] += 1
                           grid count += 1
                         # For conjugate values?
# #
                         u index = int(-u pix + Nx/2)
# #
                         v index = int(-v pix + Nv/2)
# #
# #
                          u \text{ arid} = 0
                         v \text{ arid} = 0
# #
                         # Doing an FFT-UNSHIFT, moving from centre to corners
# #
                         if(u index < centre x):</pre>
# #
                              u \text{ grid} = u \text{ index} + \text{centre } x
# #
# #
                          else:
                              u \text{ grid} = u \text{ index - centre } x
# #
# #
                         if(v index < centre y):</pre>
# #
                              v grid = v index + centre y
# #
                          else:
# #
                              v grid = v index - centre x
                         # Populate the uv grid
# #
# #
                         if (math.isnan(vis real)==False) and (math.isnan(vis imag)==False):
                              uv grid real[u grid, v grid] += vis real
# #
                              uv grid imag[u grid, v grid] += vis imag
# #
                              uv grid counter[u grid, v grid] += 1
# #
                              grid count += 1
# #
                       print("Baseline",count,"[ grid count:", grid count,"A1:", a1,"A2:",a2,"u:",round(u,2),"v:",round(v,2),"red
# # Overall time taken in seconds
# print("gridding A took(s):",time.process time() - start gridding)
```

In []: