

Resonance Ionization
Mass Spectrometry

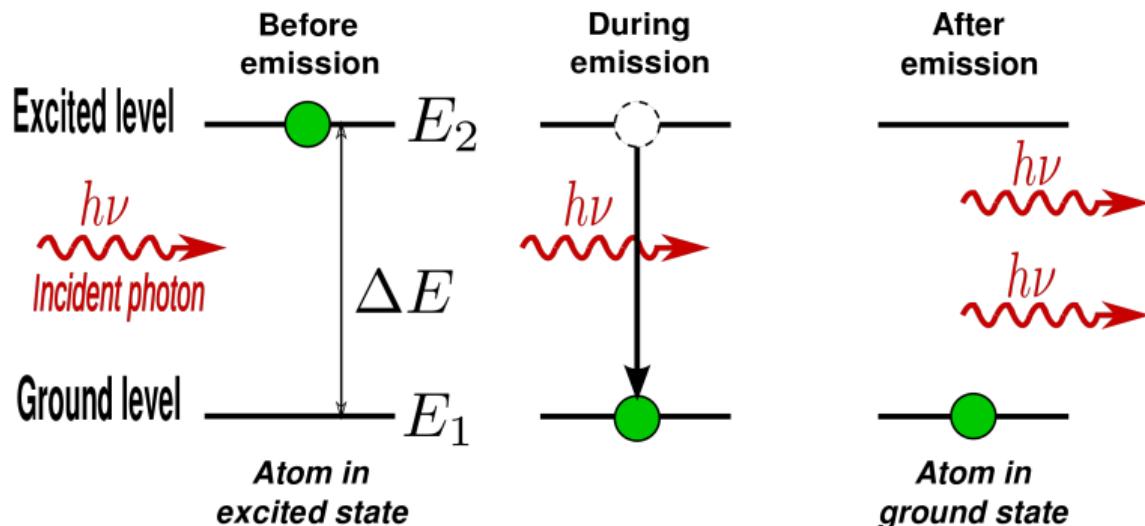
Reto Trappitsch
Laboratory for Biological Geochemistry

EPFL

January 17, 2023

An Idea as Old as the Laser

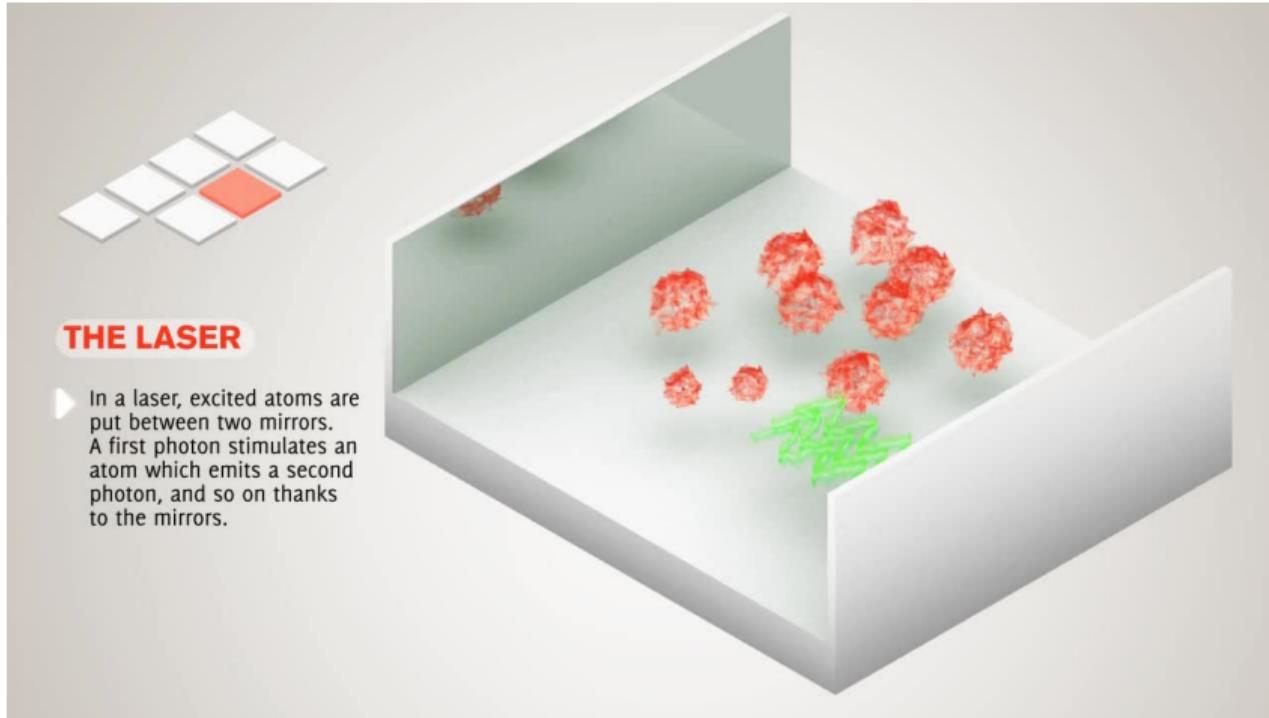
Laser: Light Amplification by Stimulated Emission of Radiation



$$E_2 - E_1 = \Delta E = h\nu$$

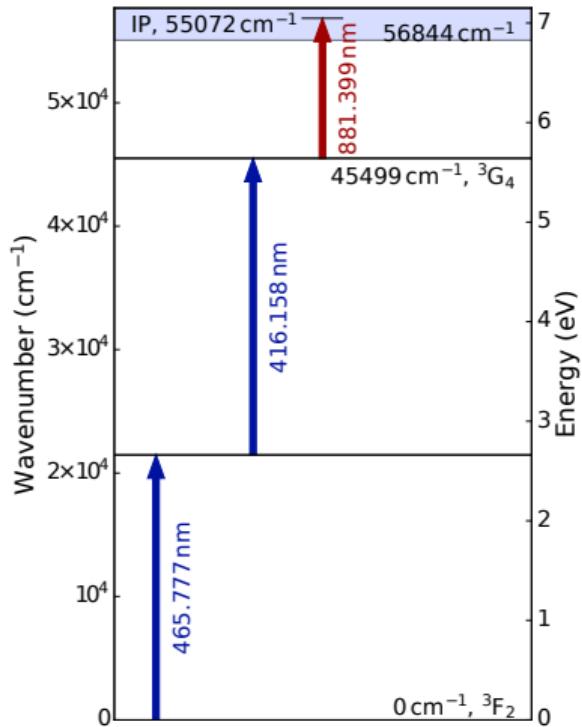
Credit: Wikipedia, Vladislav

Laser Principle



Credit: Wikipedia, Juboroff

The Process of Stimulated Emission is Reversible



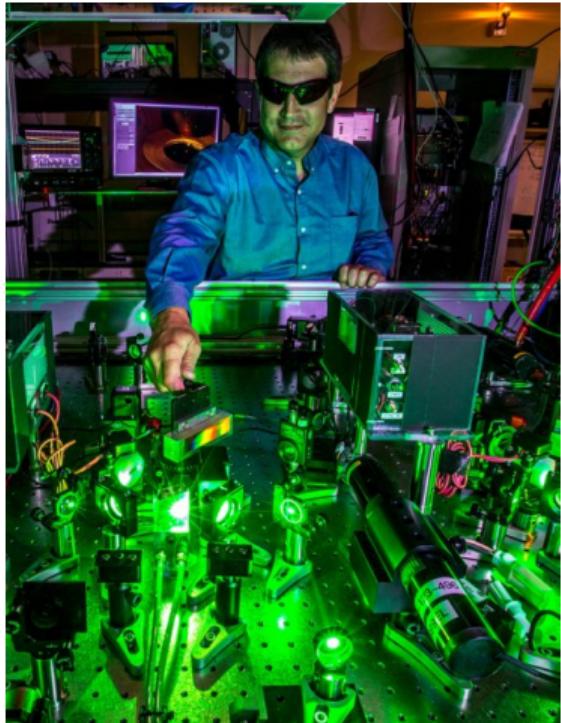
- Ionization of titanium (Trappitsch+, 2018)
- Absorption cross section σ is approximately equal to the square of wavelength λ

$$\sigma \approx \lambda^2$$

- Visible wavelength (400-800 nm): $\sigma \approx 10^{-9} \text{ cm}^2$
- Lifetime τ of a state is in the order of 10 ns
- Required photon flux to saturate transition:

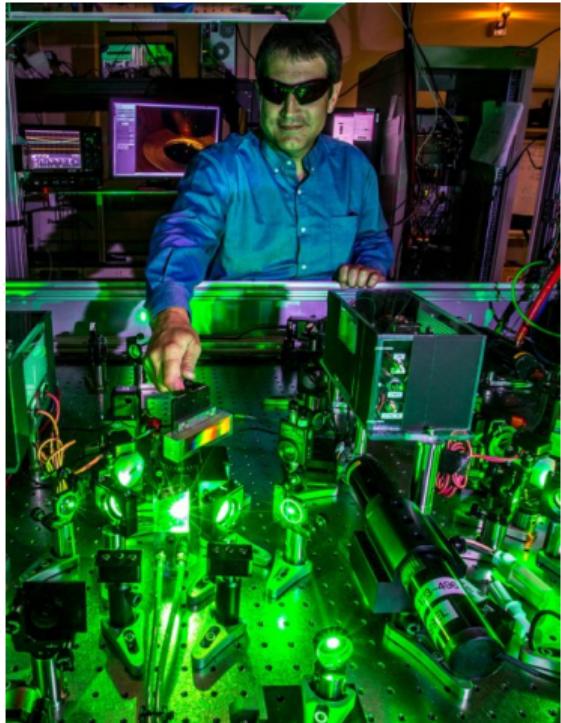
$$\phi = \frac{1}{\tau} \times \frac{1}{\sigma} \approx 10^{17} \text{ cm}^{-2} \text{ s}^{-1}$$

Pulsed Lasers can Achieve the Saturation Requirements



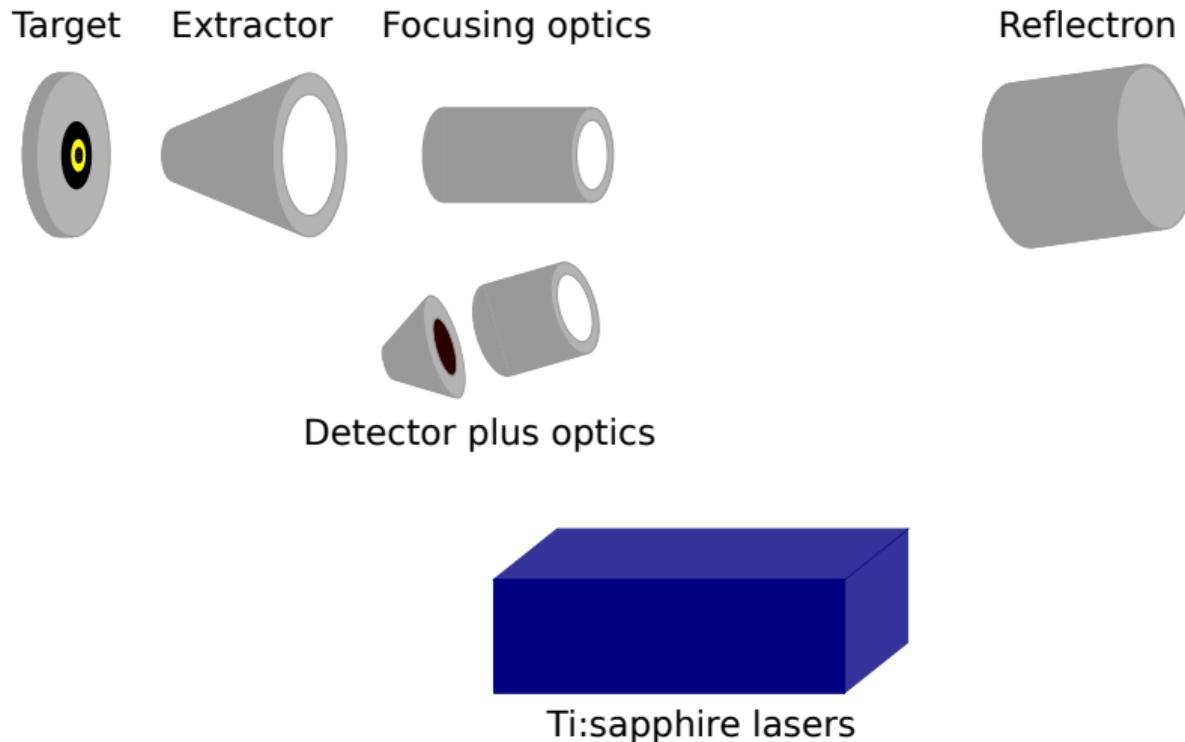
- For a pulsed laser at 1 kHz repetition rate, ~ 10 ns pulse width, this photon flux corresponds to
→ **A few mW average power**
- Requirement increased due to Doppler and power broadening and laser spectral bandwidth
- Pulsed lasers at 1 kHz, ~ 10 ns pulse width can achieve up to about 2 W mm^{-2}

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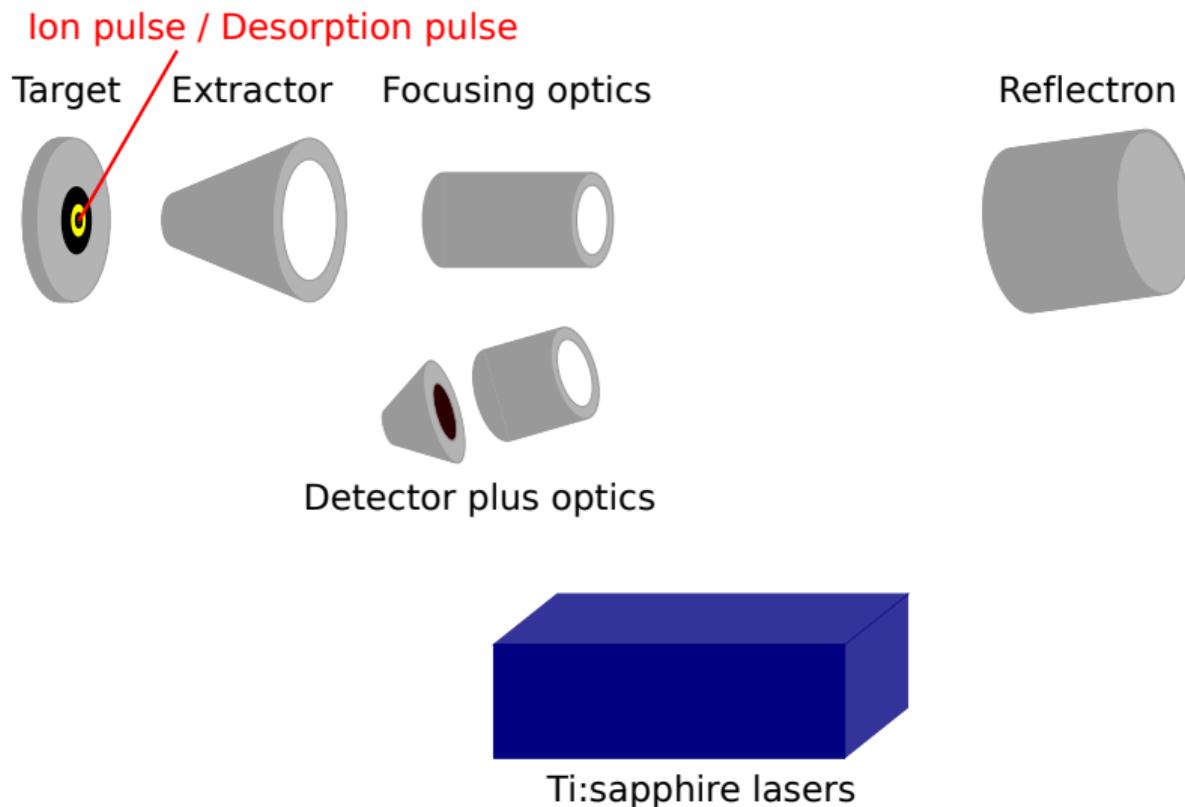


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- For a DC instrument (e.g., SIMS):
 - Pulsed laser-on time: $10^{-8}\text{ s} \times 1000\text{ Hz} = 10^{-5}$
 - Average power requirement would **hundreds of Watts** for a continuous-wave laser

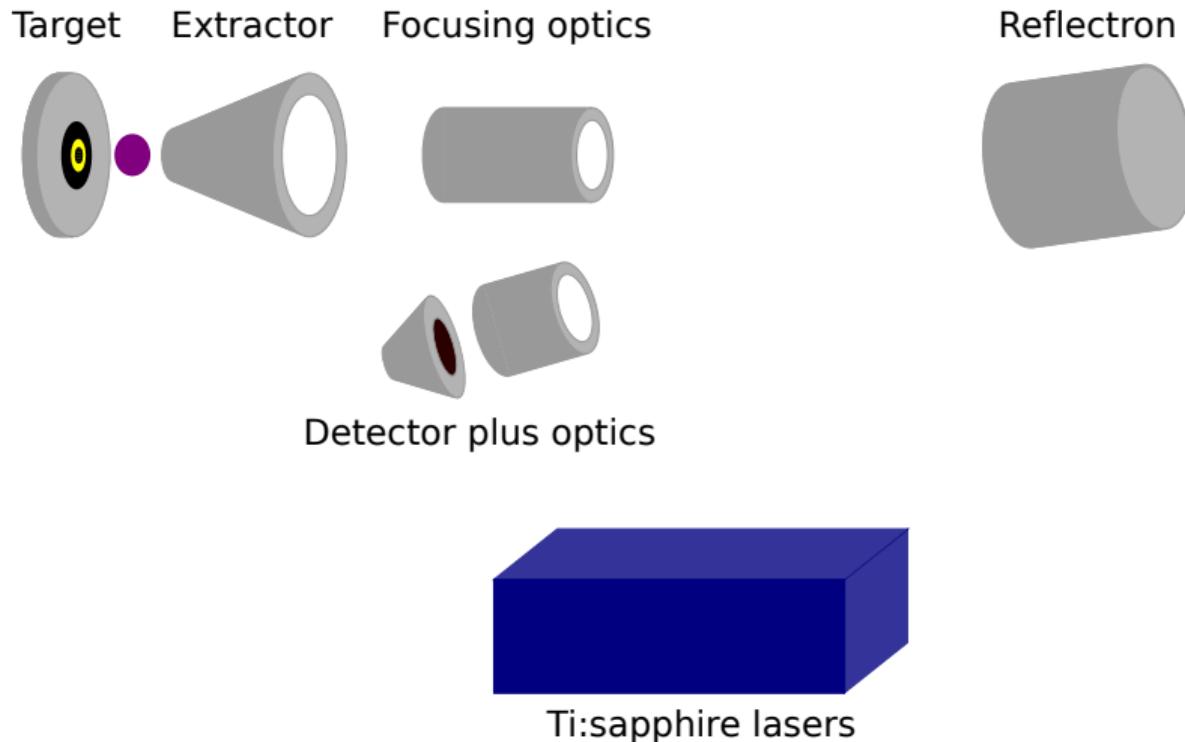
We Need a Pulsed Mass Spectrometer: Time of Flight



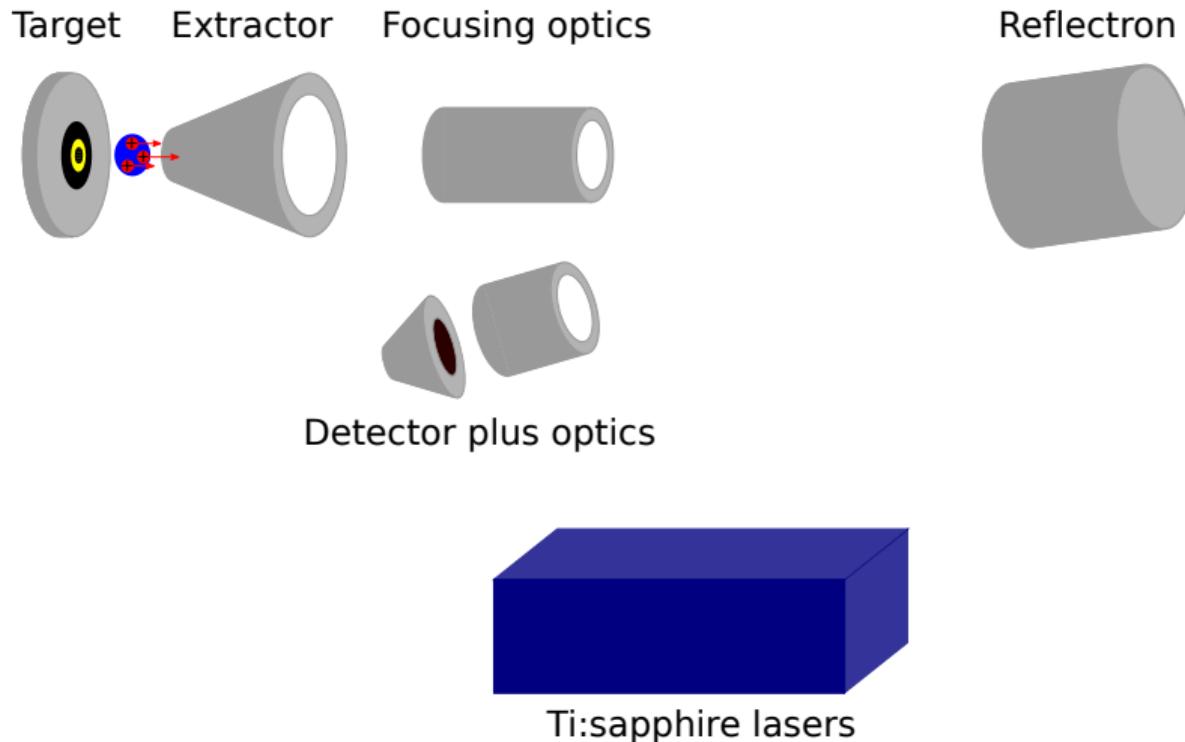
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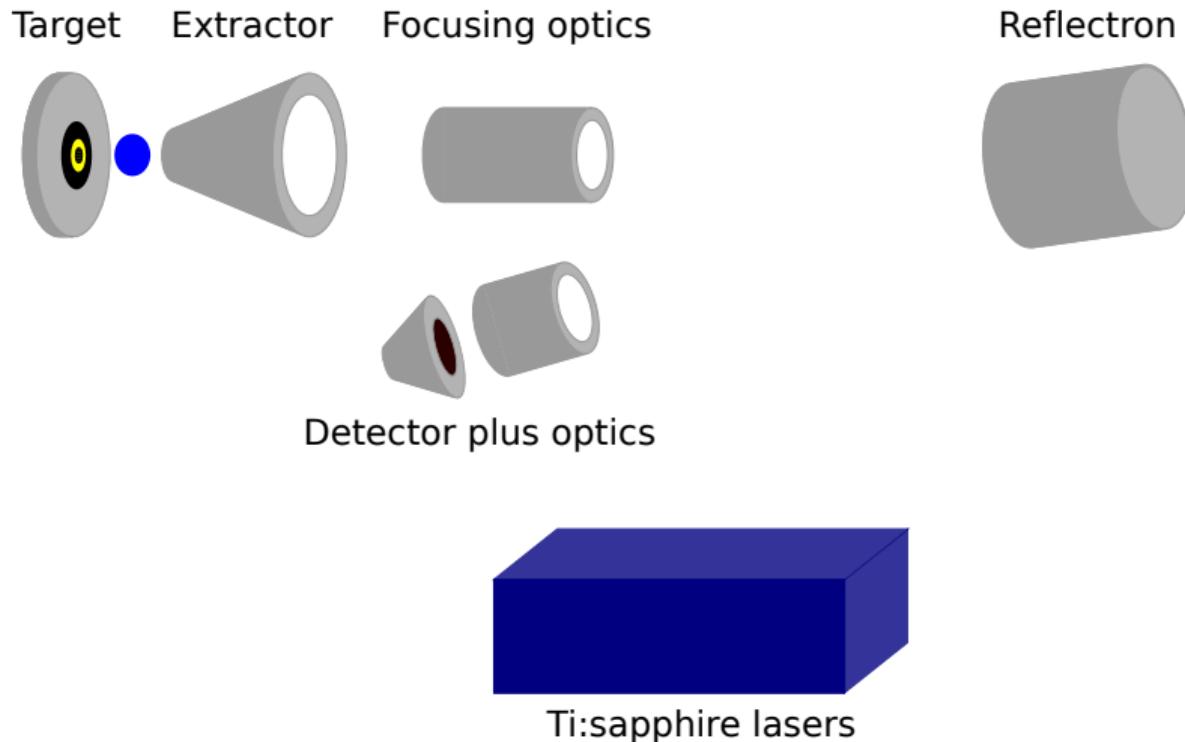
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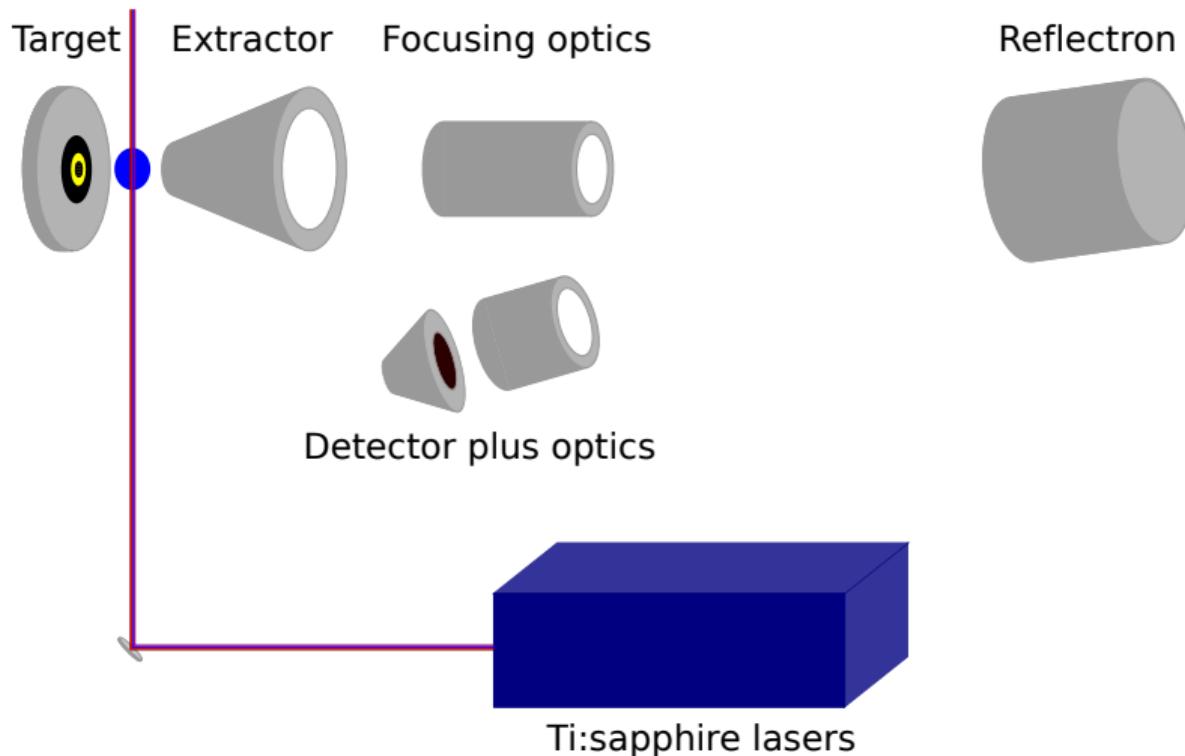
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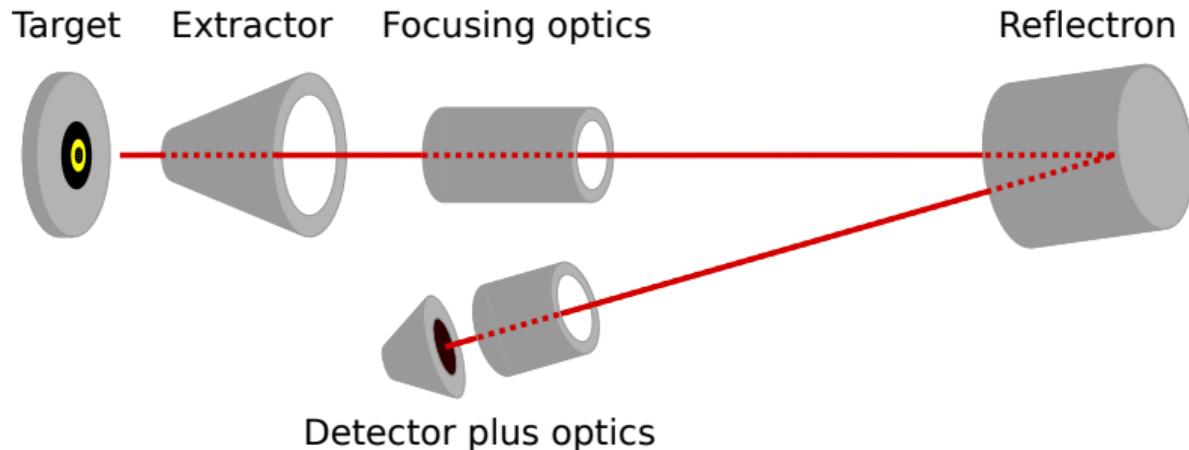
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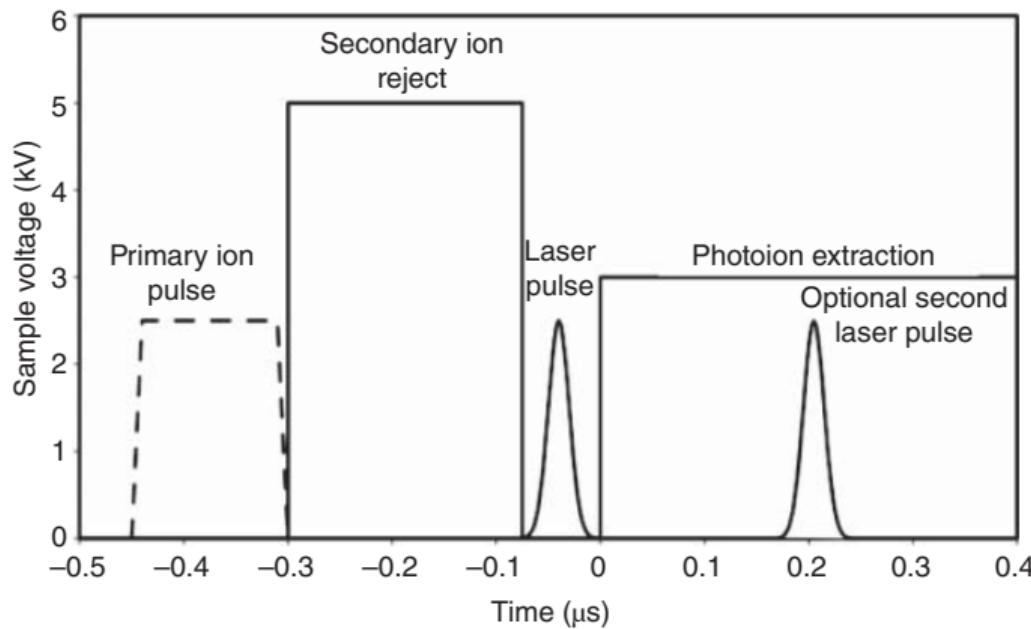


Ti:sapphire lasers

Measurement Cycles repeat at 1 kHz

- ① Desorption / Sputtering of sample
- ② Ejection of secondary ions
- ③ Resonance ionization of photoions
- ④ Extraction
- ⑤ Mass / Charge separation and detection

Optional second ionization laser pulse allows for separation of isobars



Savina and Trappitsch (2021)

Sample Removal: Sputtering vs. Laser Desorption

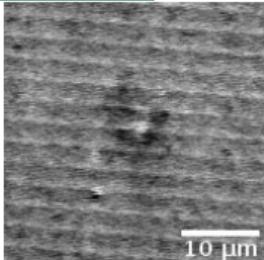


- Sputtering with Ga ion beam
 - < 100 nm spatial resolution
 - Motionless blanking required
 - Trade off between high current or high spatial resolution
 - Duty cycle compared to SIMS: $\sim 10^{-4}$
- Desorption laser
 - Various wavelength possible to couple with different materials
 - Spot-size down to around 1 μm
 - Very low secondary ion backgrounds can be achieved

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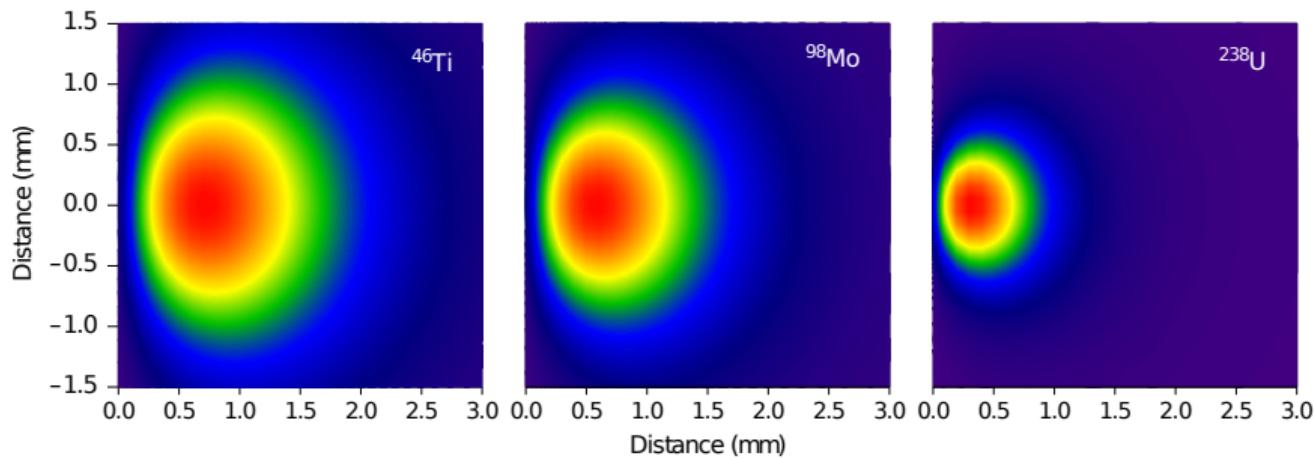


EKSPLA 1064 nm Desorption Laser



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Ionizing of Neutral Atoms: You only get One Chance!

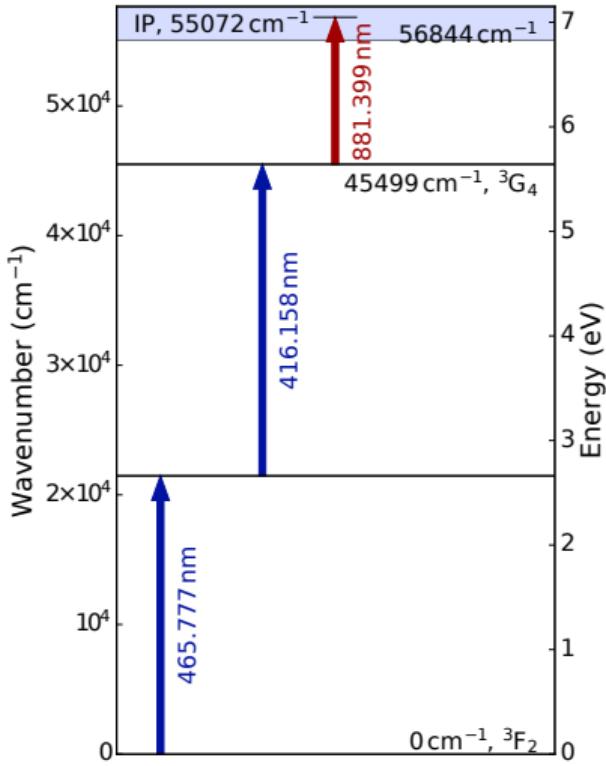


Savina and Trappitsch (2021)

- Ionization laser beam size: ~ 1.5 mm diameter cylinder
- Laser intercepts cloud of neutrals above sample surface
- Neutrals that do not get ionized in first shot will be lost due to cloud expansion

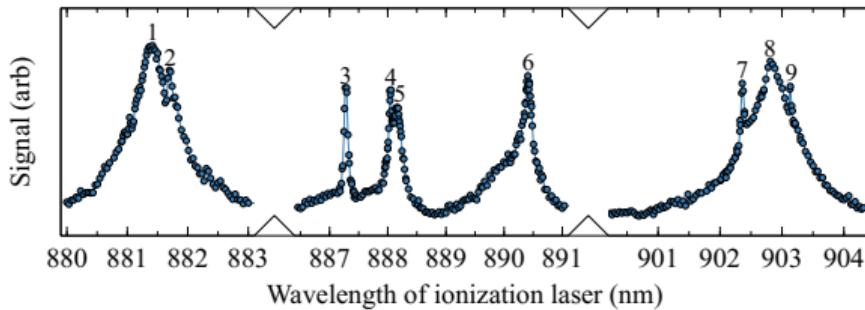
Multi-step Laser Ionization of, e.g., Titanium (Trappitsch et. al, 2018)

- Resonance Ionization of Titanium requires three lasers
- Each ionization step is highly selective
- Ionization schemes need to be tested:
 - Spectroscopy of states above ionization potential
 - Saturation: Irradiance counts!
- Ti has low lying states
 - Understand population of these states
 - Scheme specific
 - Here: majority after sputtering in ground state



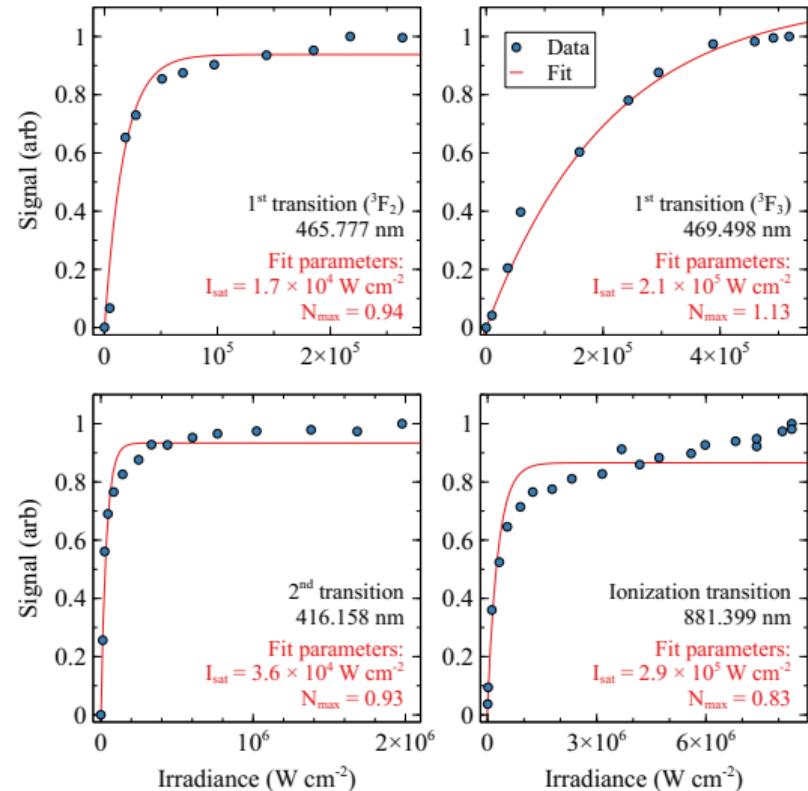
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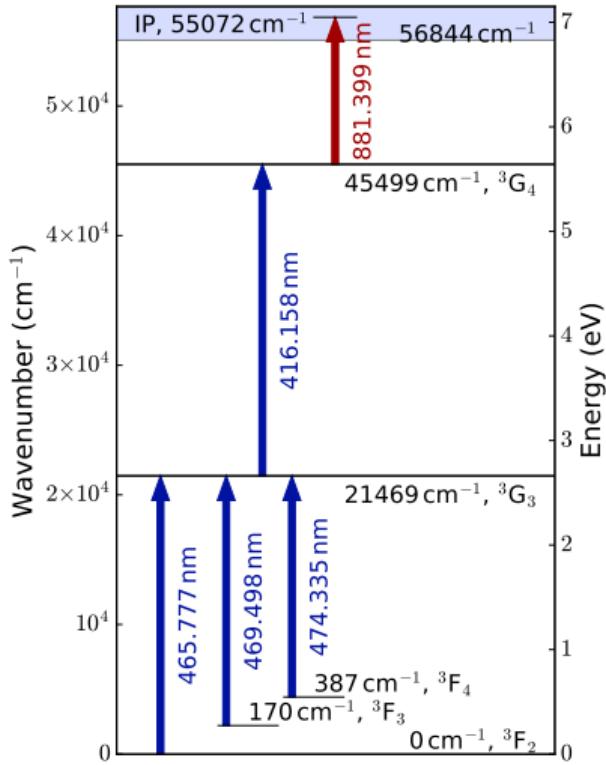
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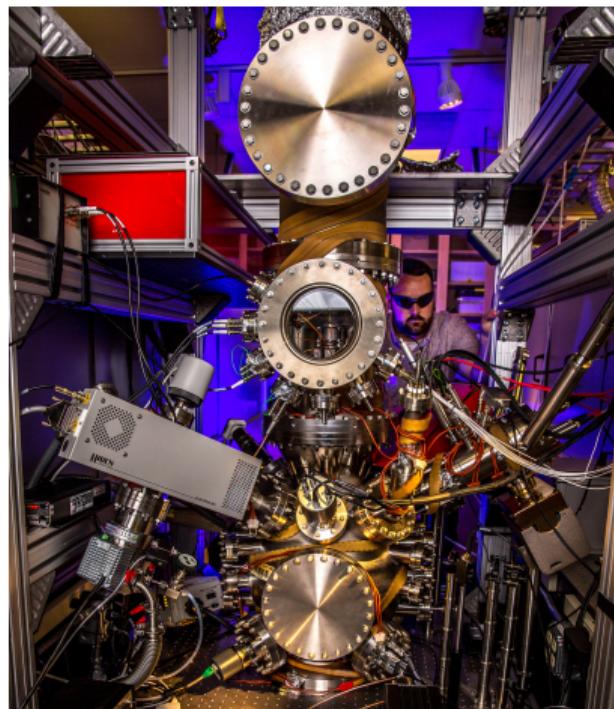
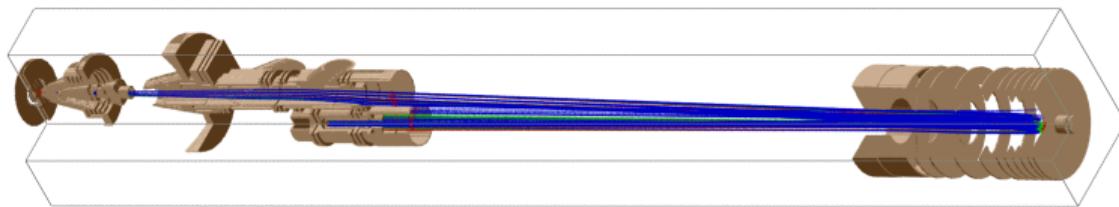
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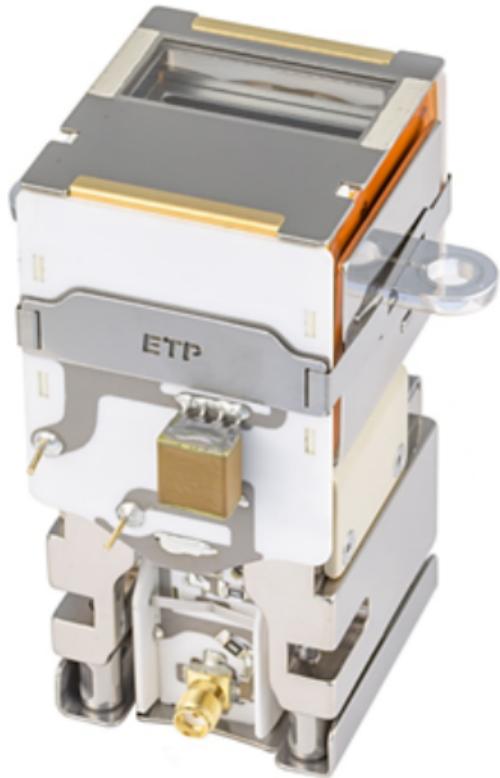
Separation of m/q in Time-of-Flight (TOF) Mass Analyzer

- Time of Flight Mass Analyzer
 - ~ 3.5 m flight path
 - Grid-less reflectron to optimize transmission
 - Mass resolution $\frac{m}{\Delta m} > 1000$
- Difficulty: Map a photoion volume in time onto detector
- Lasers however take care of isobars



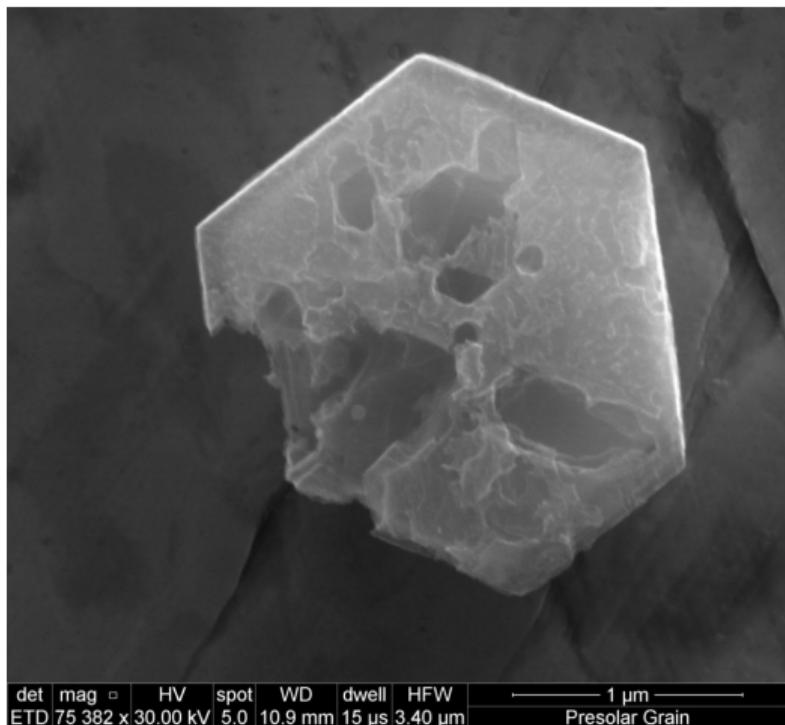
Ion Counting — Record every Arrival Time

- Ion counting detectors
 - Microchannel plate detectors (MCPs)
 - TOF Electron Multipliers
- Time-to-Digital Conversion: 80 ps time resolution
- Overall system dead-time: ~ 700 ps
- Reasonable count rates: $\sim 2,000$ cps



RIMS — A Versatile Technique for Trace Element Analyses

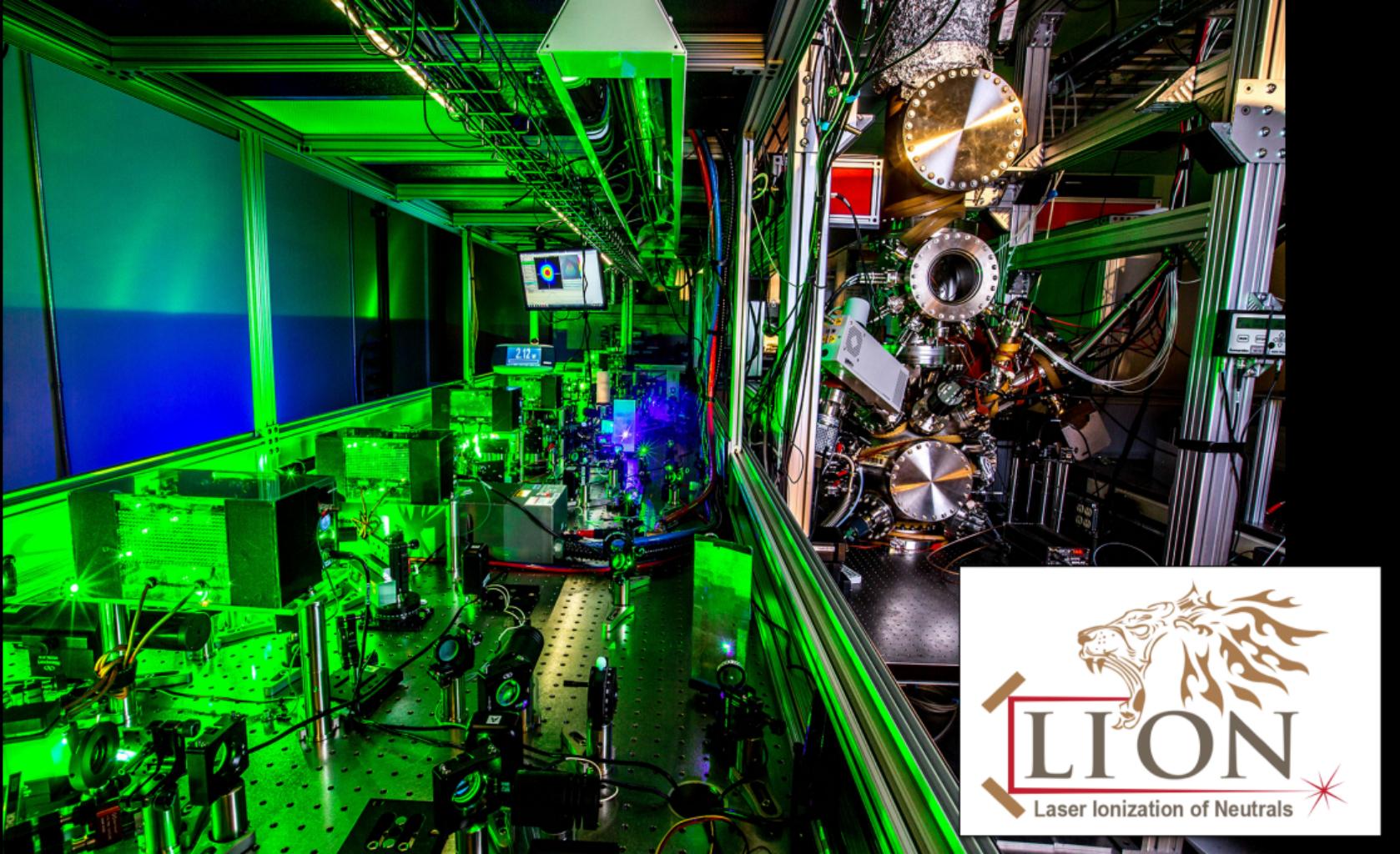
- High sensitivity for small, atom-limited samples
- Minimal sample preparation
- Resonance ionization with tunable Ti:Sapphire lasers
- High spatial resolution
 - $\sim 1 \mu\text{m}$ for laser desorption
 - $< 100 \text{ nm}$ for ion sputtering
- High useful yield
 - 38% for U analysis (Savina+ 2018)
 - $\sim 18\%$ for Ti analysis (Trappitsch+ 2018)
- Low backgrounds and high isobar suppression

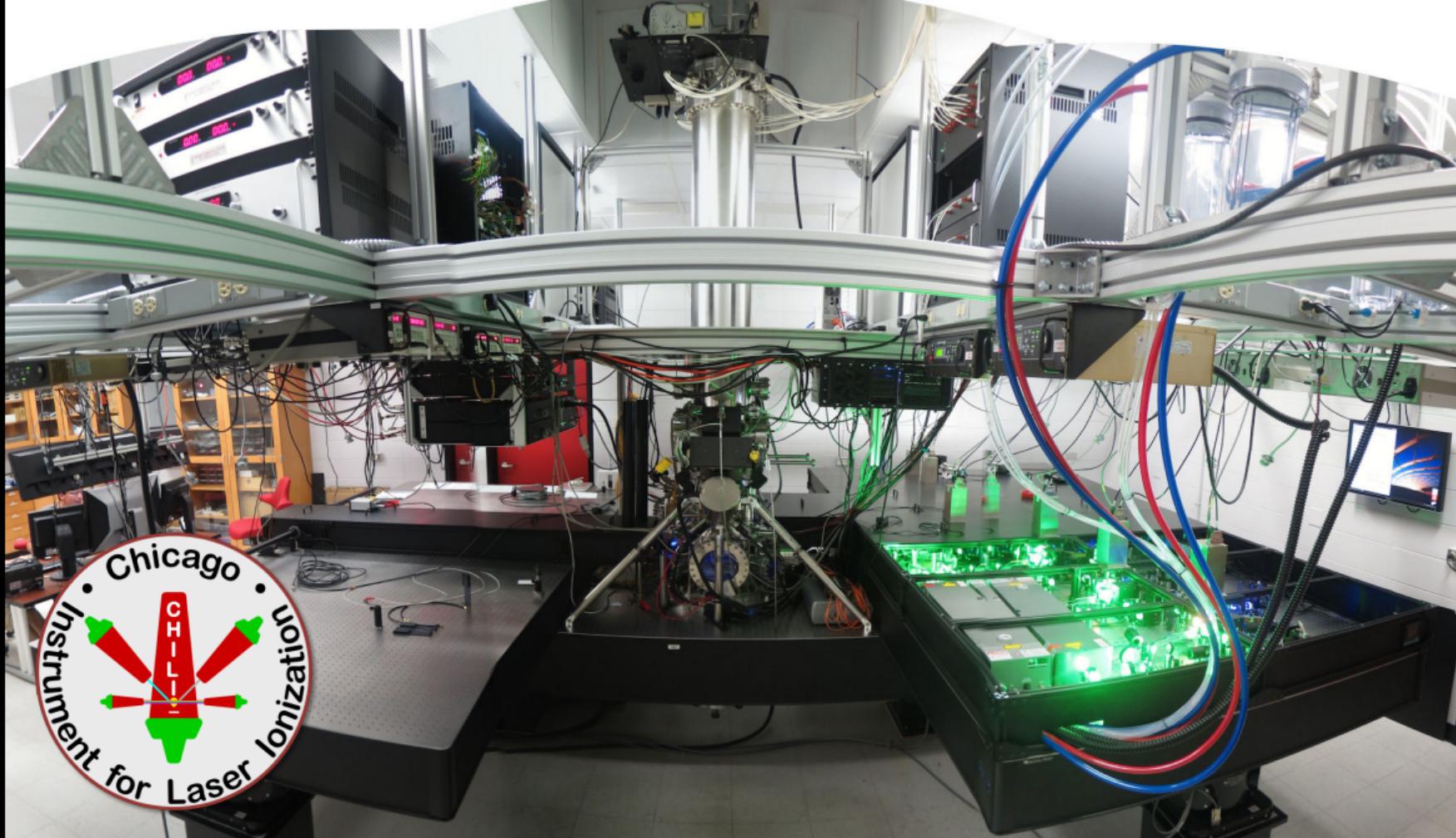


A RIMS Table of Elements

H																				He
Li	Be																			
Na	Mg																			
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr			
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe			
Cs	Ba	*	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn			
Fr	Ra	**																		
*	La	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu					
**	Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr					

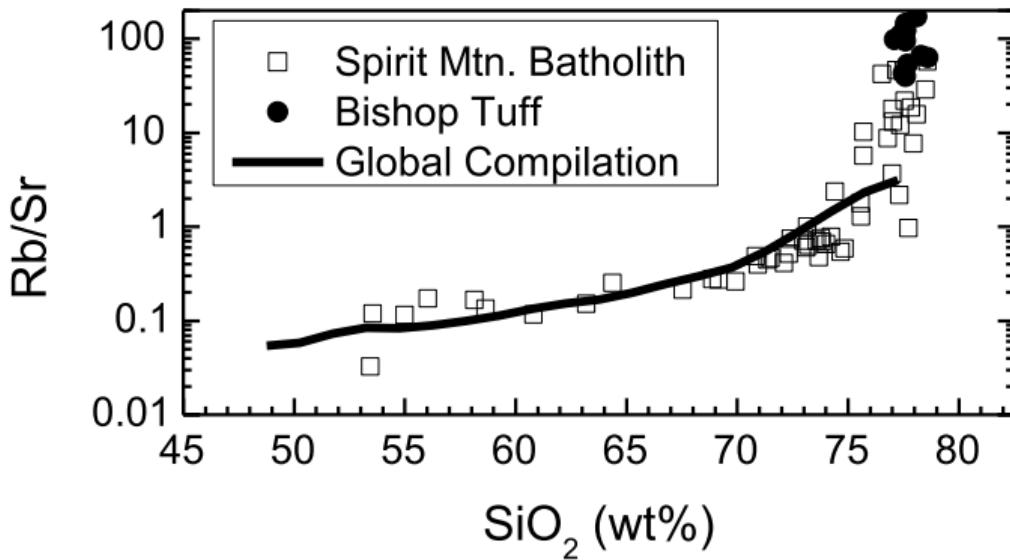
Savina and Trappitsch (2021)





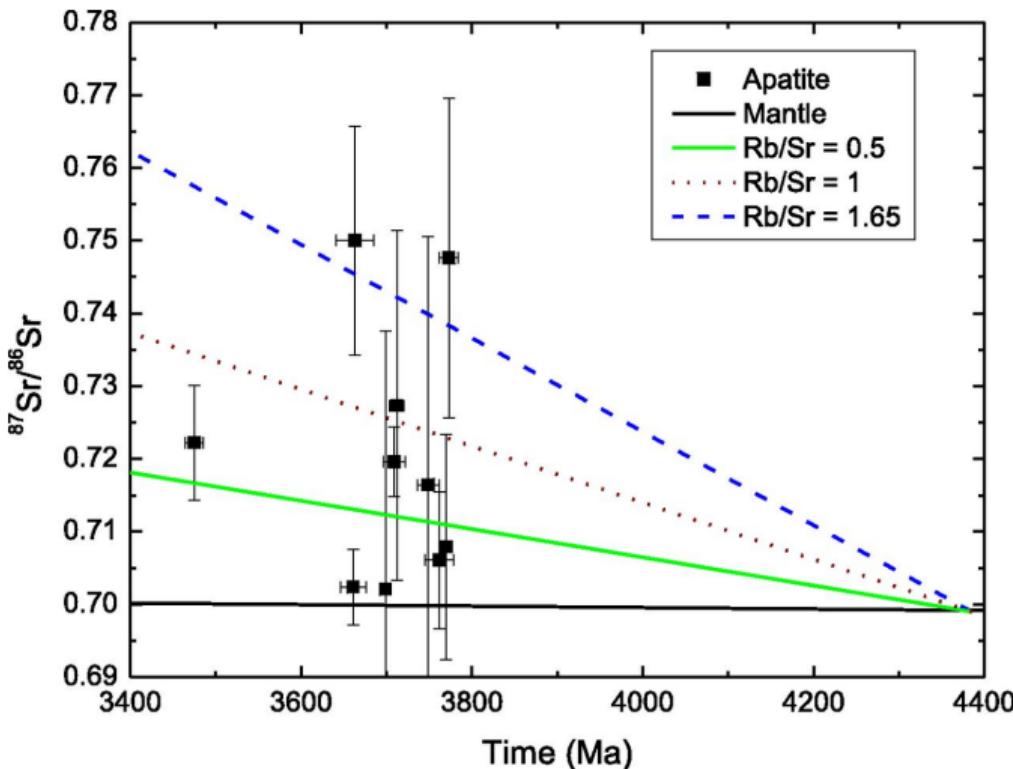
Potassic, High-Silica Hadean Crust (Boehnke+, 2018)

- Composition of the Hadean crust poorly understood
- Rb/Sr ratio correlates with SiO₂
- Analyzed 10 apatite inclusions in Archean zircons
- Sr isotope ratios indicate high Rb/Sr ratio
- Suggest a felsic crust formed by ~ 4.4 Ga



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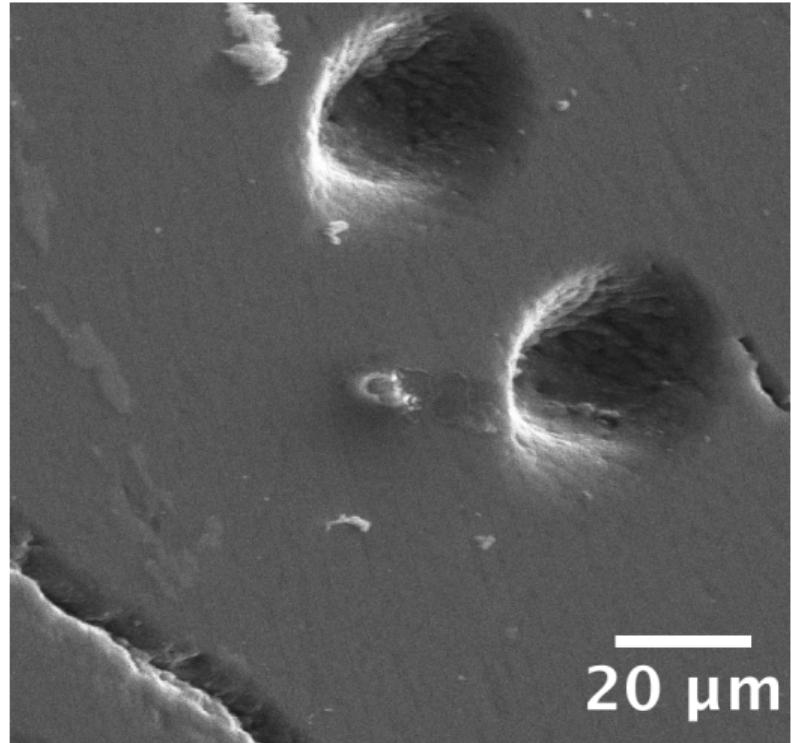
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Did a Supernova Contribute ^{60}Fe to the Early Solar System?

- Long-standing controversy:
 - In-situ SIMS measurements: “High” ^{60}Fe
→ Implies supernova injection
 - Bulk ICPMS measurements: “Low” ^{60}Fe
→ Can be explained as “galactic background”
- Re-analysis by RIMS
- Correlated effects minimized thanks to normalization to $^{62,58}\text{Ni}$
- RIMS measurements do not detect the previously reported “high” ^{60}Fe content

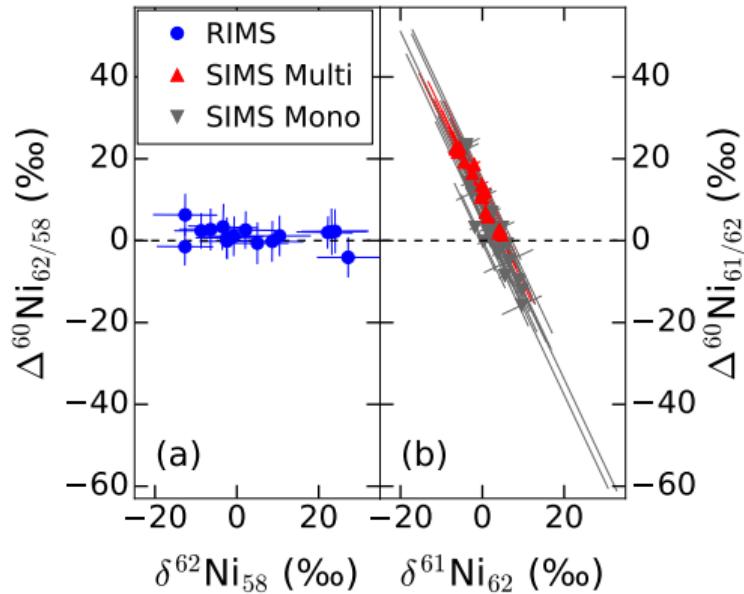
Beware your measurement uncertainties!



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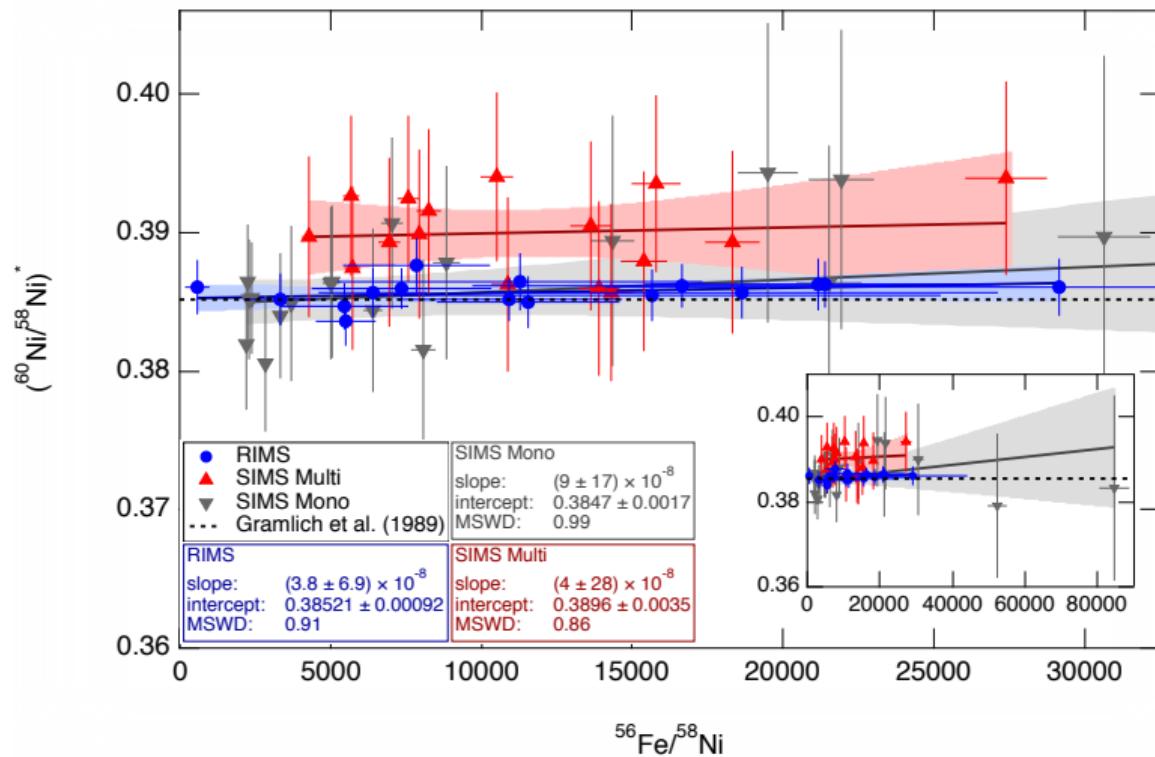
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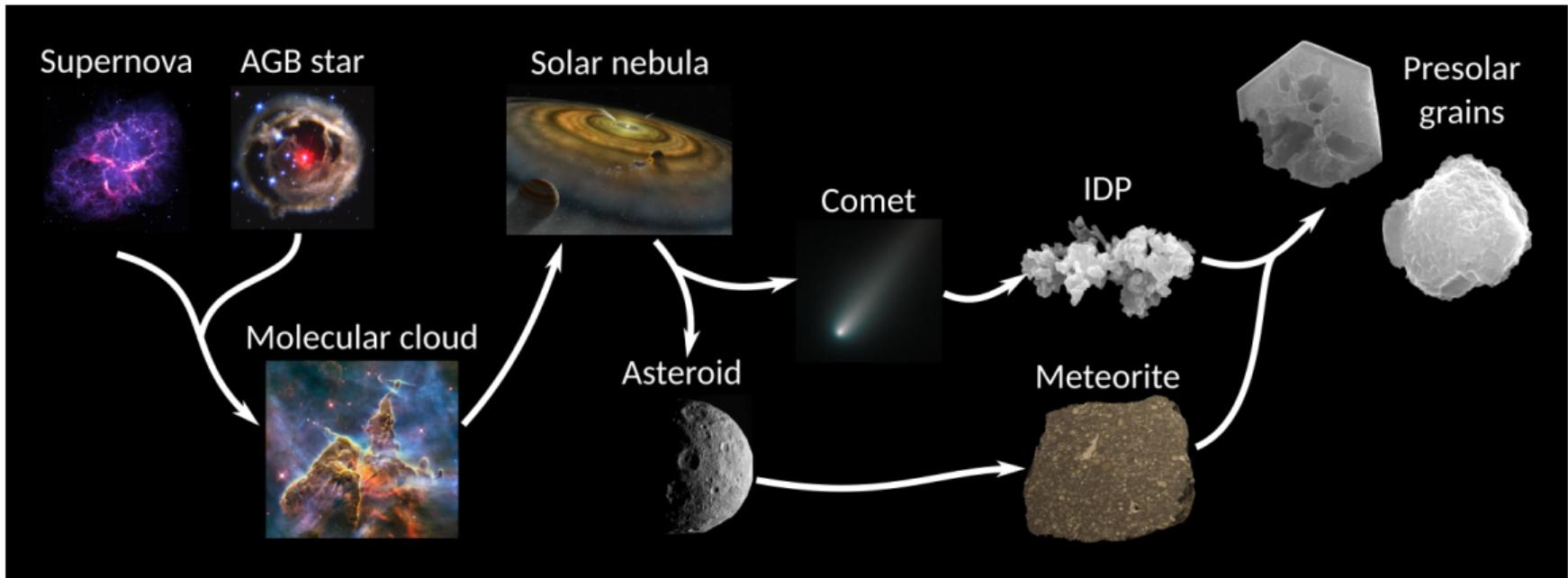
Trappitsch+ (2018)

^{60}Fe in Early Solar System can be Explained as Galactic Background

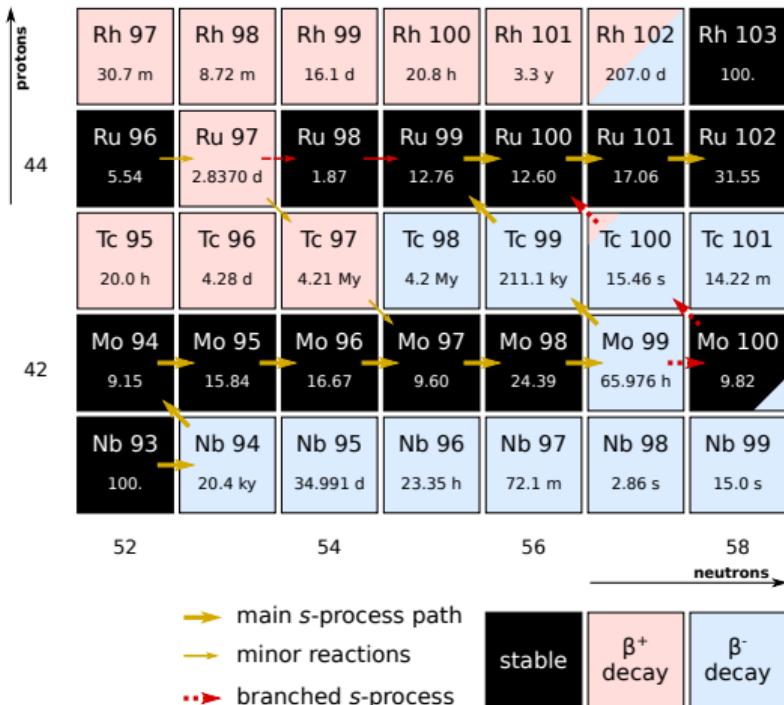


Trappitsch+ (2018)

Stardust grains: Tracing Stellar Nucleosynthesis



Observations of Live ^{99}Tc in AGB Stars



SPECTROSCOPIC OBSERVATIONS OF STARS OF CLASS S

PAUL W. MERRILL

MOUNT WILSON AND PALOMAR OBSERVATORIES
CARNEGIE INSTITUTION OF WASHINGTON
CALIFORNIA INSTITUTE OF TECHNOLOGY

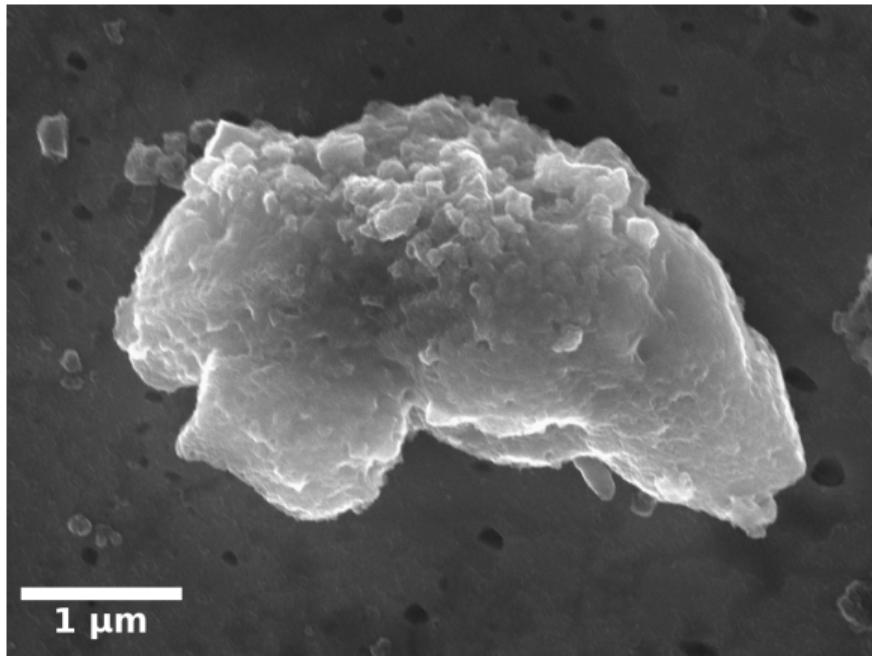
Received February 27, 1952

TABLE 2
INTENSITIES OF LINES AND BANDS

STAR	PLATE	ABSORPTION					EMISSION					
		ZrO	TiO	Ba II	Low-Temp.	Tc I	H	Fe II	Mg I	Si I	In I	Co I
R And...	Ce 3522	8	3	5	8	4	10	3	2	3	3	2
U Cas...	Pc 127	7	7	5	6	3	10	3	1	3	1	2
HD 22649	Pc 192	2	2	5	6	1	0	0	0	0	0	0
R Gem...	Pc 68	5	0	10	7	5	10	3	2	2	3	3
S UMa...	Pc 110	1	0	7	4	1	10	3	1	2	1	1
T Sgr...	Pc 124	7	0	7	5	3	10	3	2	3	4	3
R Cyg...	Pc 137	10	0	10	5	3	10	2	2	2	2	3
AA Cyg...	Pc 115	8	7	7	8	4	0	0	0	0	0	0
Z Del...	Pc 112	2	7	3	3	1	10	3	1	2	0	2
x Cyg...	Ce 3762	5	20	3	10	3	10	3	2	5	4	2
o Cet...	(Ce 4109	1	15	1	7	2	5	1	0	2	1	0
R Hya...	(Ce 5925	1	10	2	6	1	10	3	1	2	0	1
R Leo...	Pc 40	0	20	1	10	0	10	4	4	6	3	0

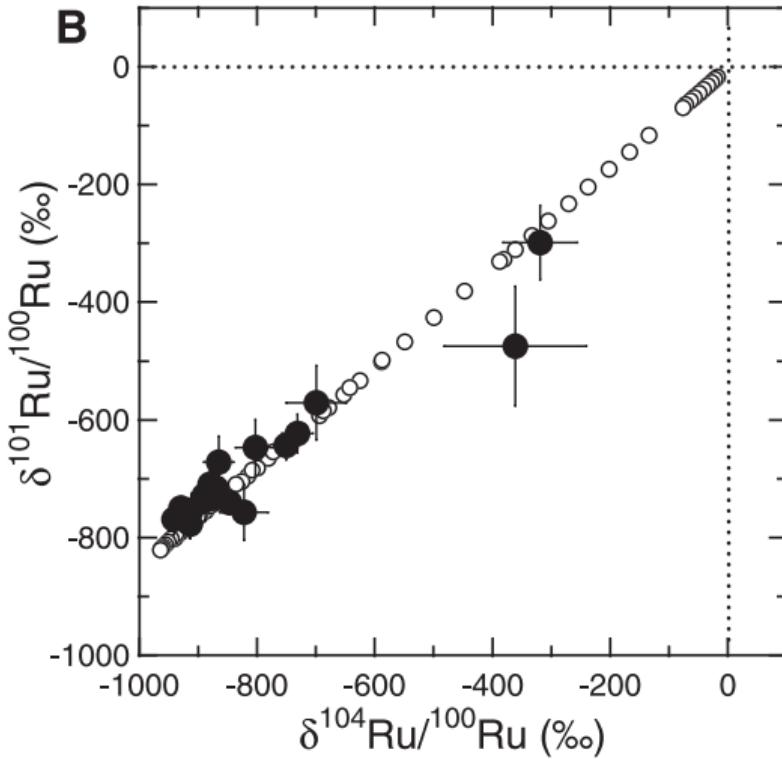
Enhancements in ^{99}Ru in Presolar Stardust (Savina+, 2004)

- Ruthenium isotopic composition measured in μm -sized SiC grains by RIMS
- Comparison with slow neutron capture process models
 - $^{101}\text{Ru}/^{100}\text{Ru}$ agrees with models
 - $^{99}\text{Ru}/^{100}\text{Ru}$ elevated due to in-situ decay of ^{99}Tc
- Measurements require in-situ decay of ^{99}Tc
- Proof that these grains come from AGB stars (stars of class S)
- Many further measurements since
 - Stellar nucleosynthesis
 - Galactic chemical evolution



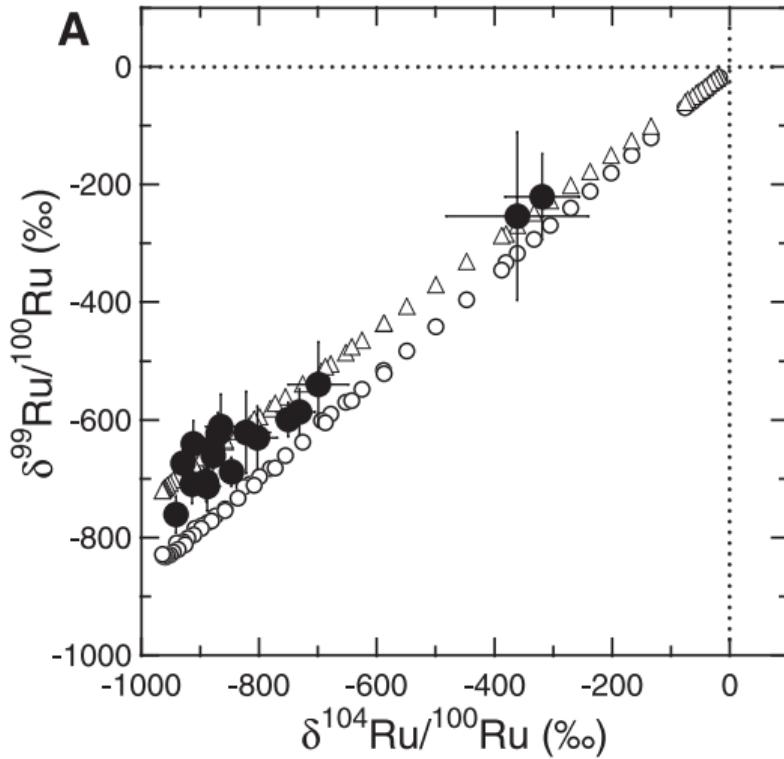
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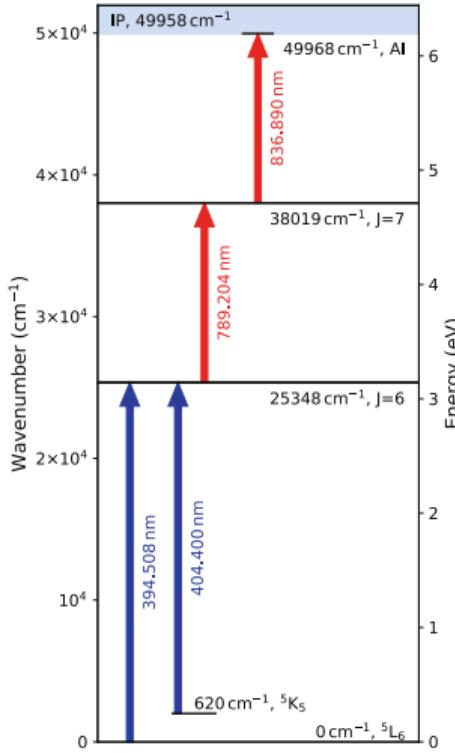
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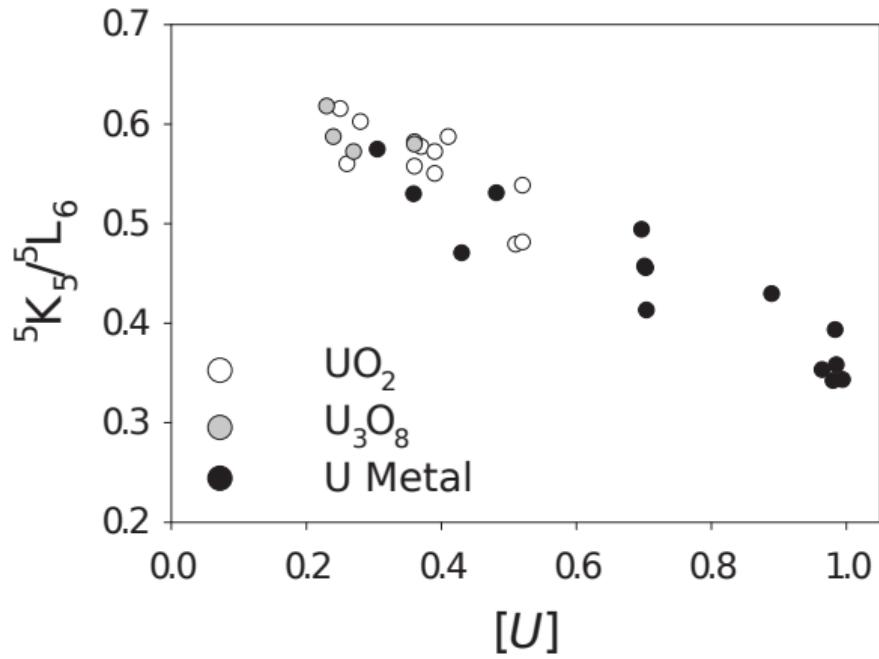
And there are Many More Uses for RIMS!

- Surface Chemistry
 - Oxidation state of U
 - Enabled by low-lying state
- Nuclear forensics
 - Verification
 - Attribution
- Radionuclides in the environment
 - Tracing nuclear accidents
 - Study remediation measures
- Medical applications
 - Ca metabolism
 - Drug uptake
- ...



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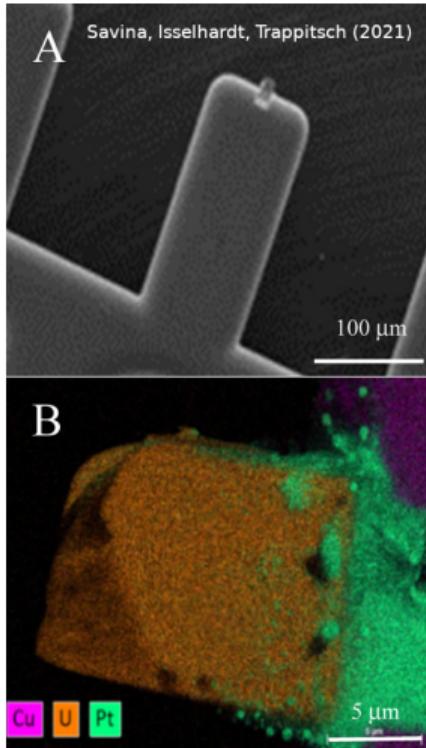
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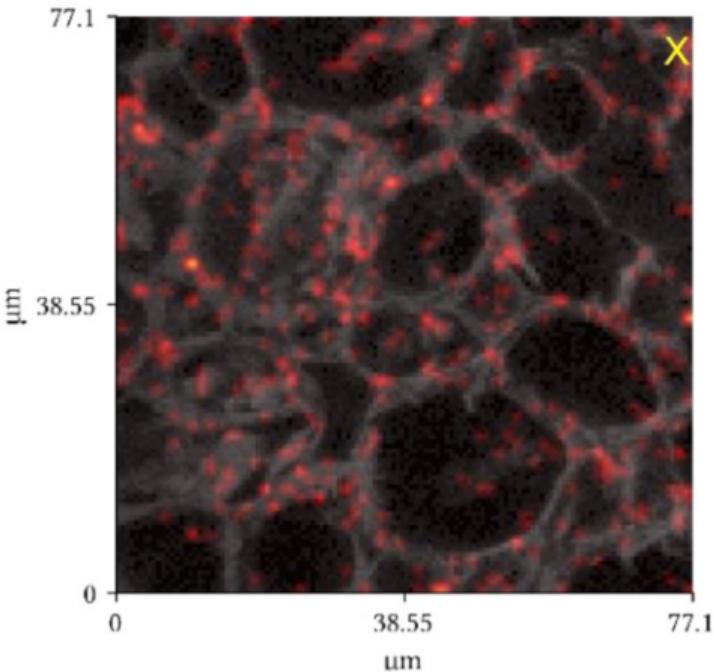
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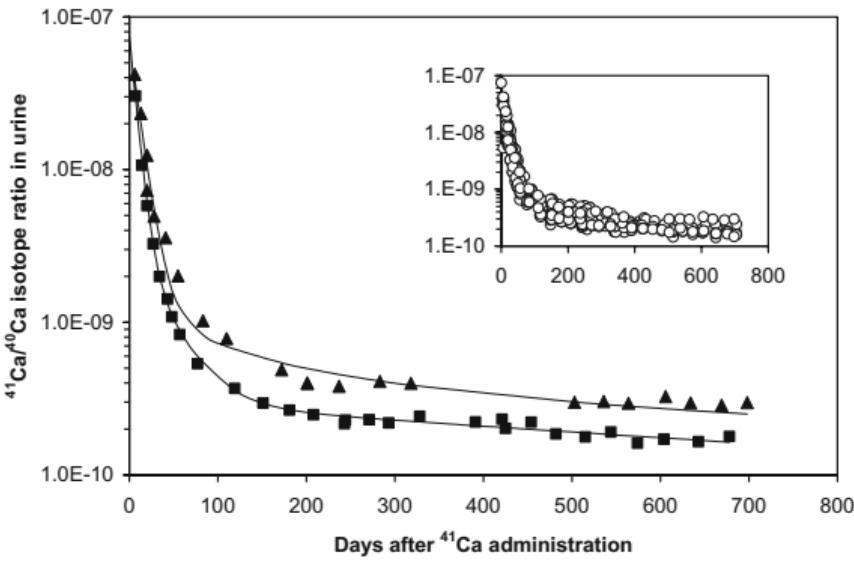
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^{99}Tc in *pisum sativum* (Mandel+, 2022)

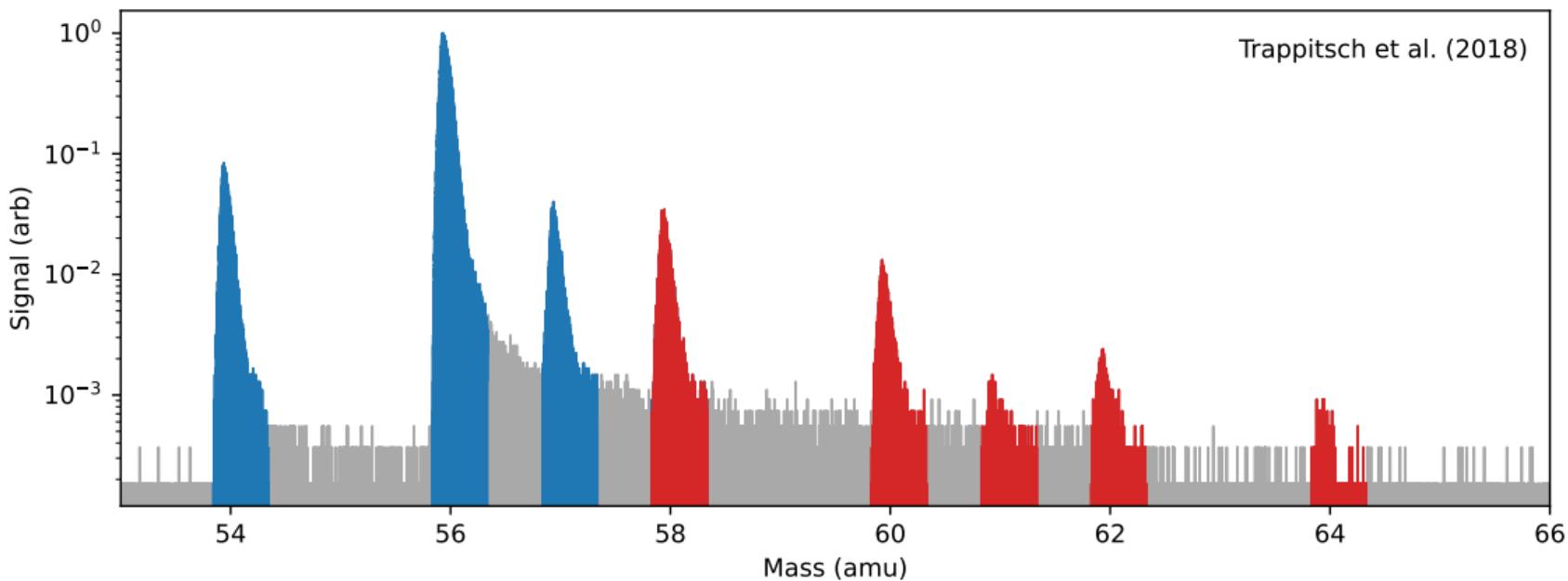
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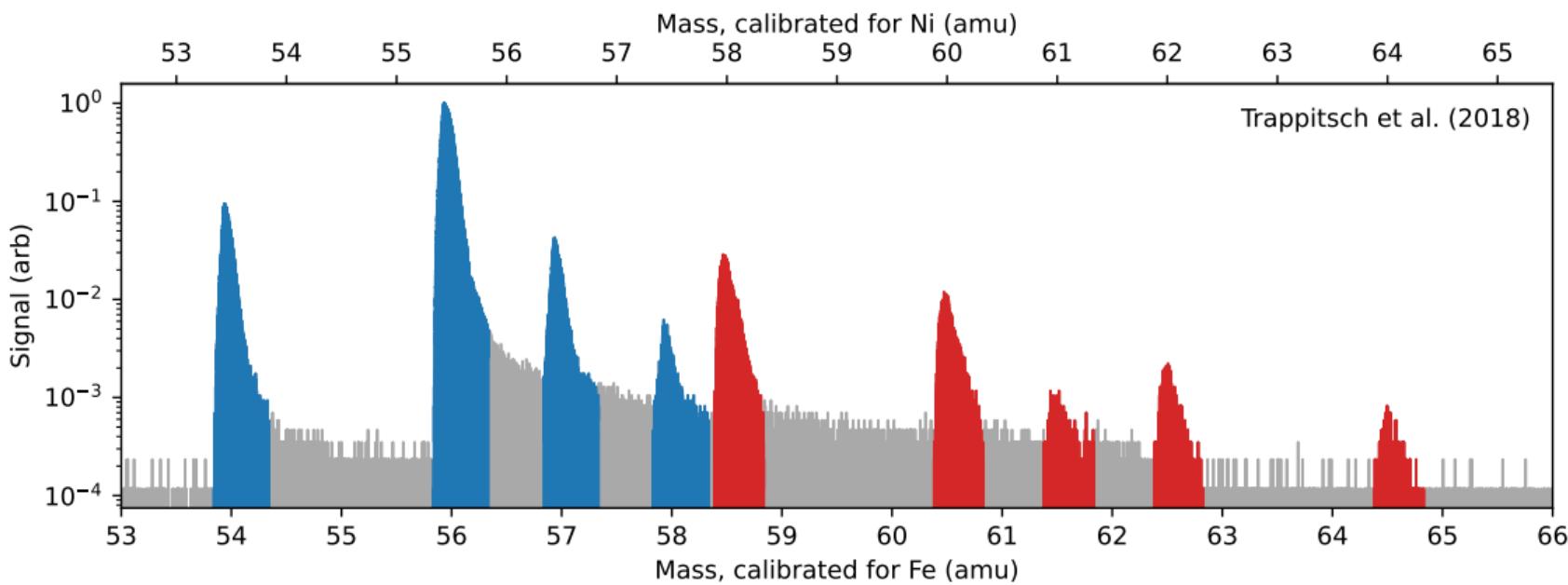


Denk+ (2006)

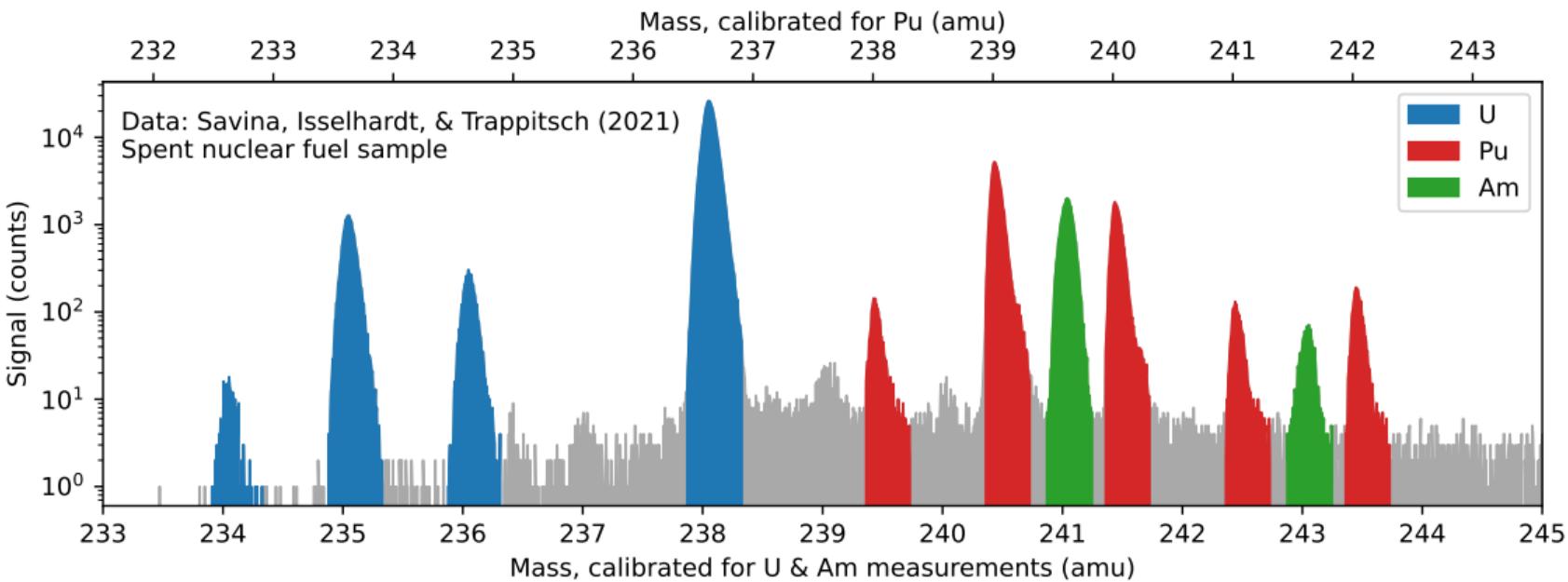
Simultaneous Measurements of Iron and Nickel



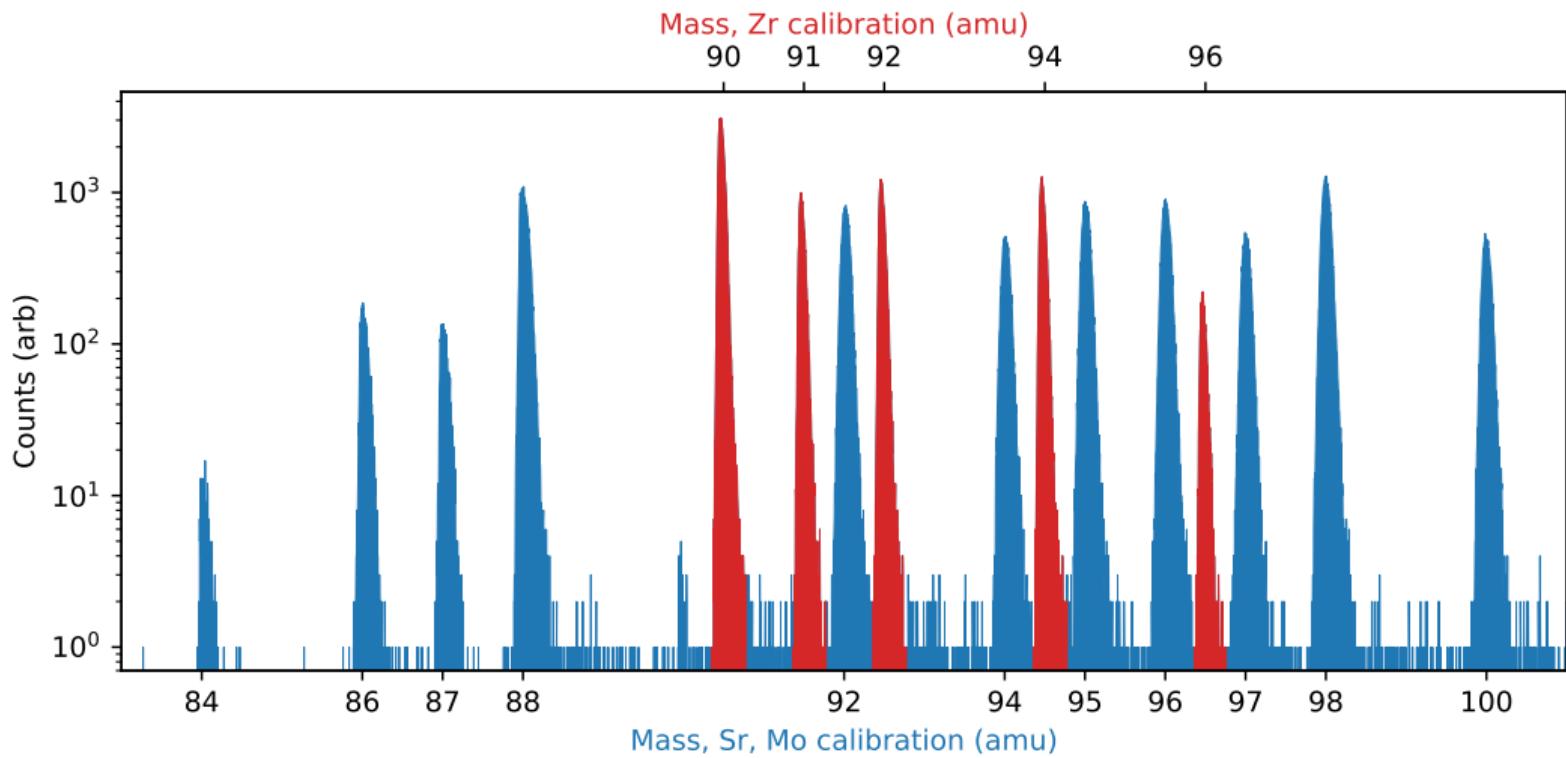
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Multi-Element Analysis Avoiding Isobaric Overlap

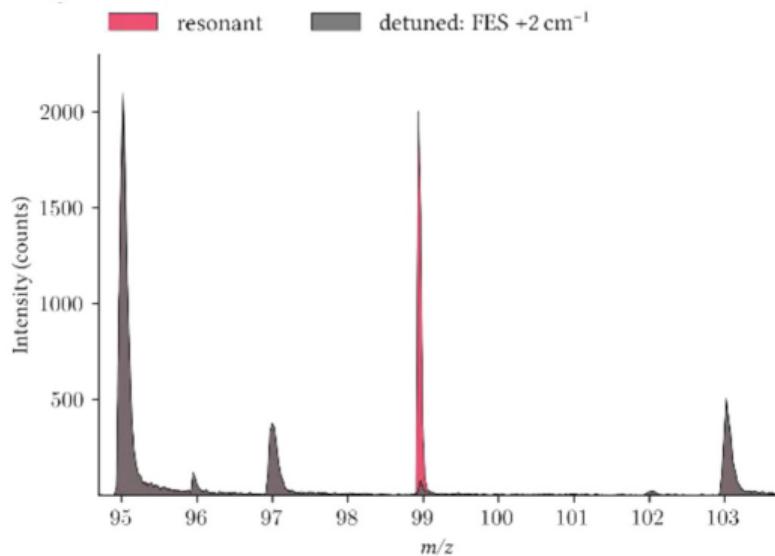


Simultaneous Sr, Zr, and Mo analysis (Shulaker+, 2022)



Signal and Noise: Quasi-Simultaneous Detection and Correction

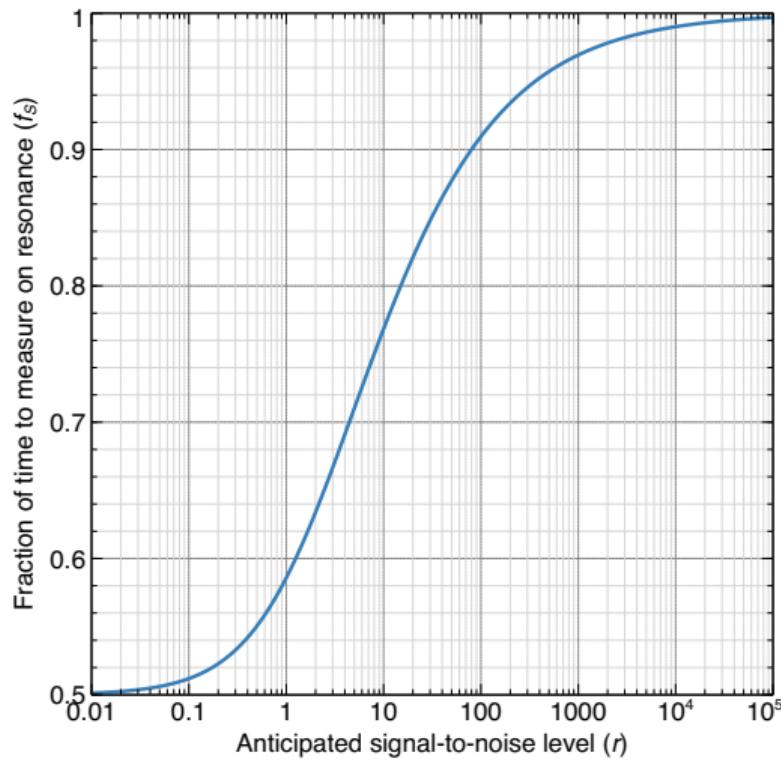
- Quantitative analysis of noise possible by RIMS
- Signal: Lasers are on-resonance
- Noise: Detune laser to be off resonance
- Quasi-simultaneous
 - Tune additional laser and blink on-/off-resonance
 - Use single laser and electro-optic deflectors (in development)
- Optimal “blinking” rate depends on signal/noise



Mandel+ (2022)

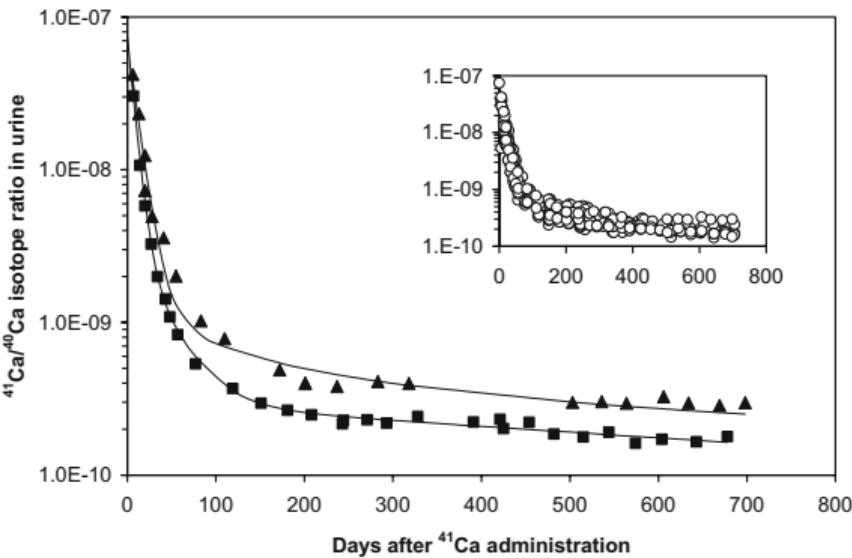
Signal and Noise: Quasi-Simultaneous Detection and Correction

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Full Disclosure — Limitations of RIMS

- Count rate limitations significantly limit the dynamic range
 - Narrowband cw lasers can be used in special cases to increase dynamic range
 - Example: $^{41}\text{Ca}/^{40}\text{Ca}$ analysis
- Duty cycle compared to SIMS: $\sim 10^{-4}$
- Desorption laser coupling depends on material and wavelength
 - Choose the right wavelength and pulse width
- Sample material might come off as molecules
 - In-vacuo surface chemistry



Denk+ (2006)

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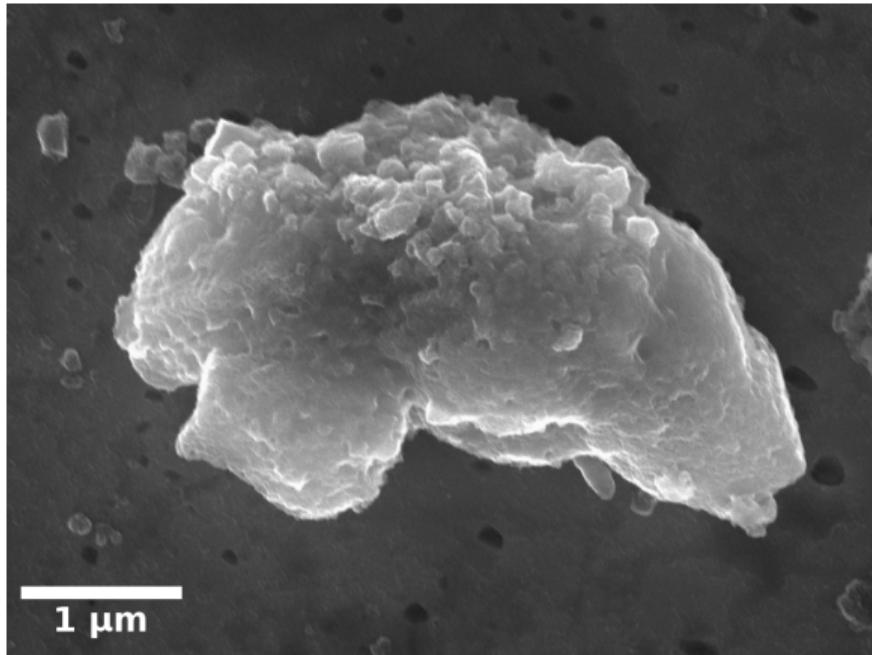
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RIMS — A Tool for Your Analysis?

- Small, atom-limited samples
- Spot size:
 - < 100 nm ion gun
 - $\sim 1 \mu\text{m}$ desorption laser
- High useful yield: Up to $\sim 40\%$ (U)
- Isotope ratio uncertainties $\gtrsim 2\%$
- Isobar suppression and separation
- Quantitative background measurement

RIMS: An ultra-sensitive technique that is complementary to, e.g., NanoSIMS

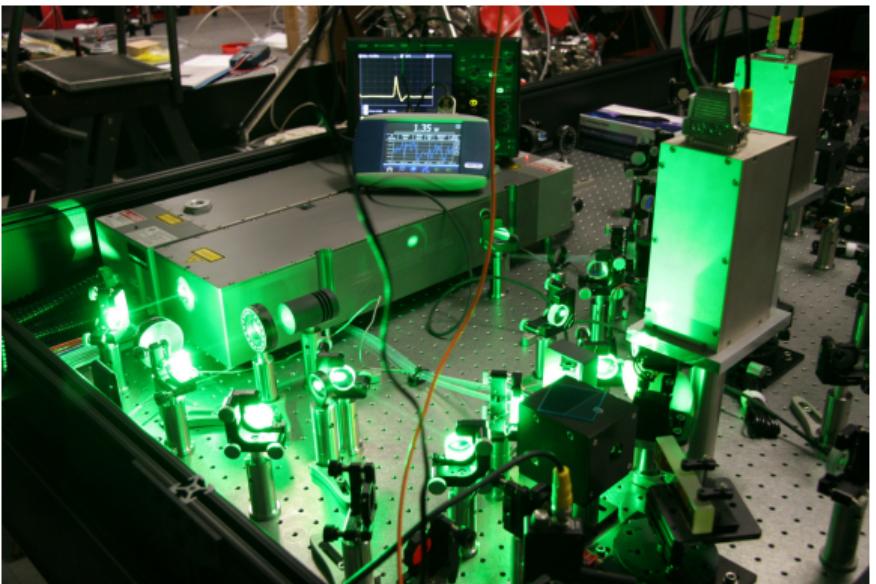


Presolar SiC grain

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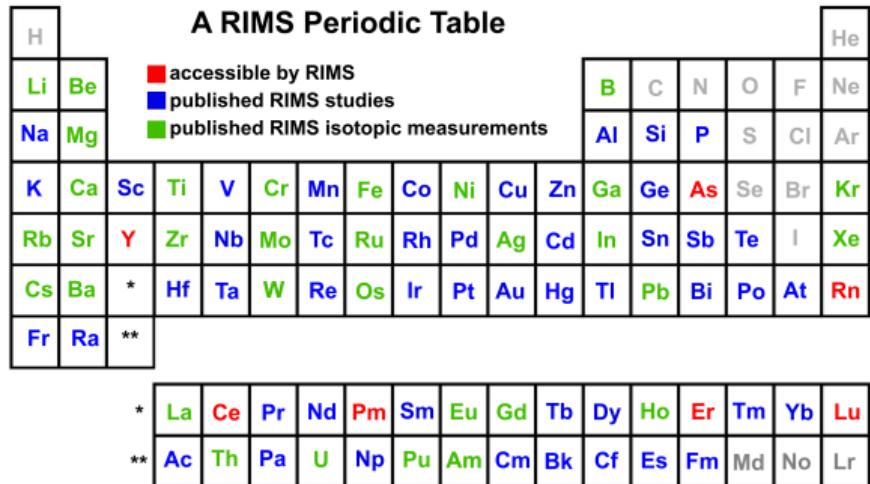
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Laser cavity on CHILI

RIMS — A Tool for Your Analysis?

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 - ~ 1 μm desorption laser
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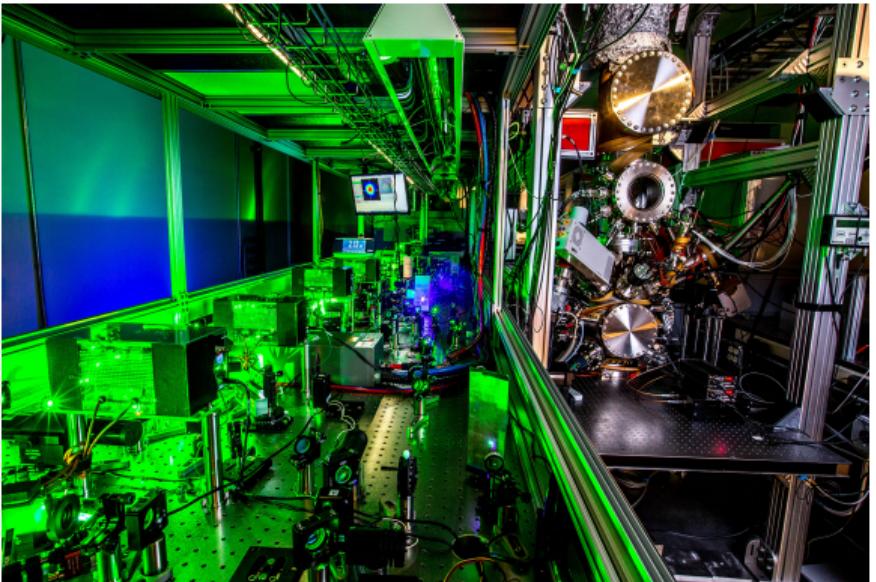
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Savina and Trappitsch (2021)

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LION at LLNL

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- Ziva Shulaker
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