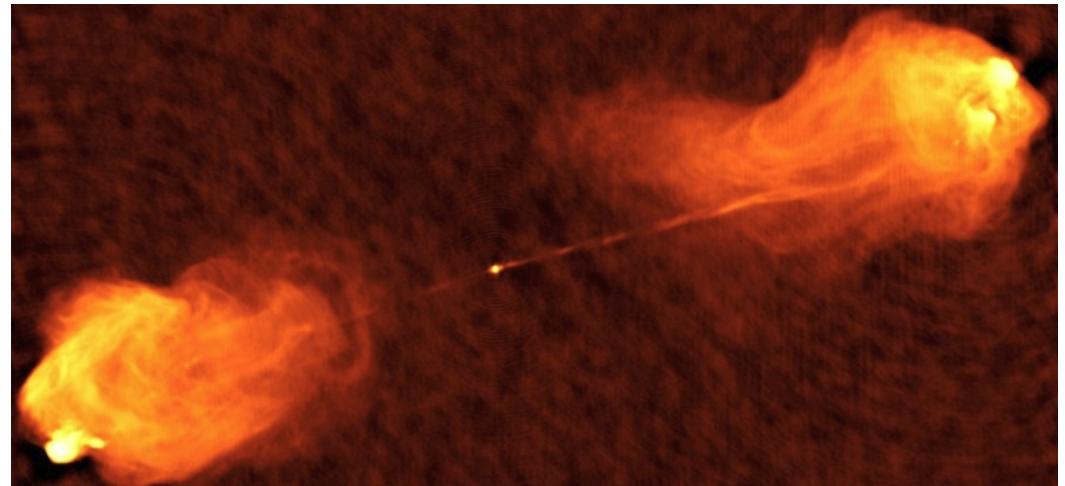


# Radio Interferometers and Coherence

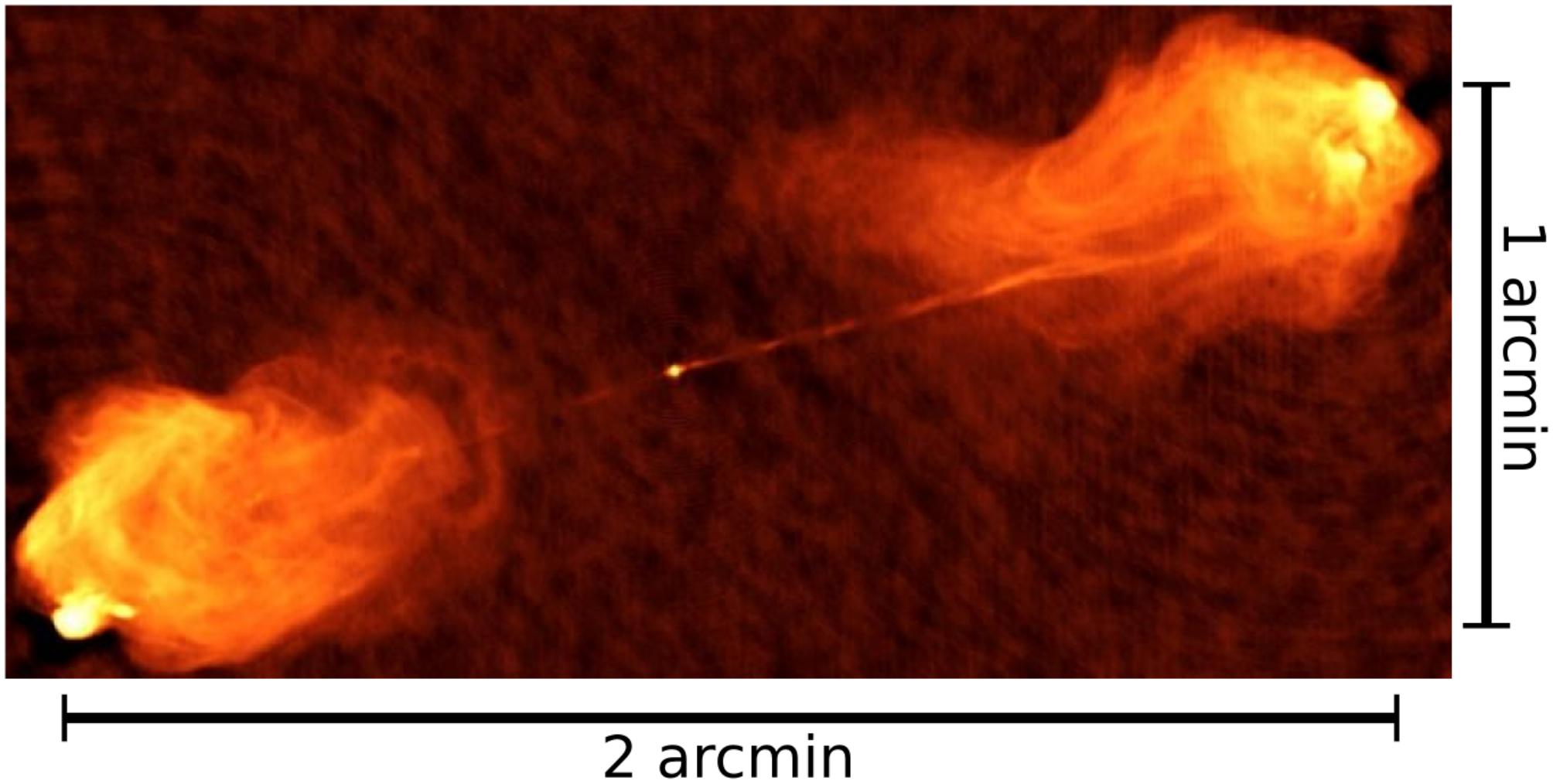


Jason Gallicchio  
Harvey Mudd College  
3 March 2015

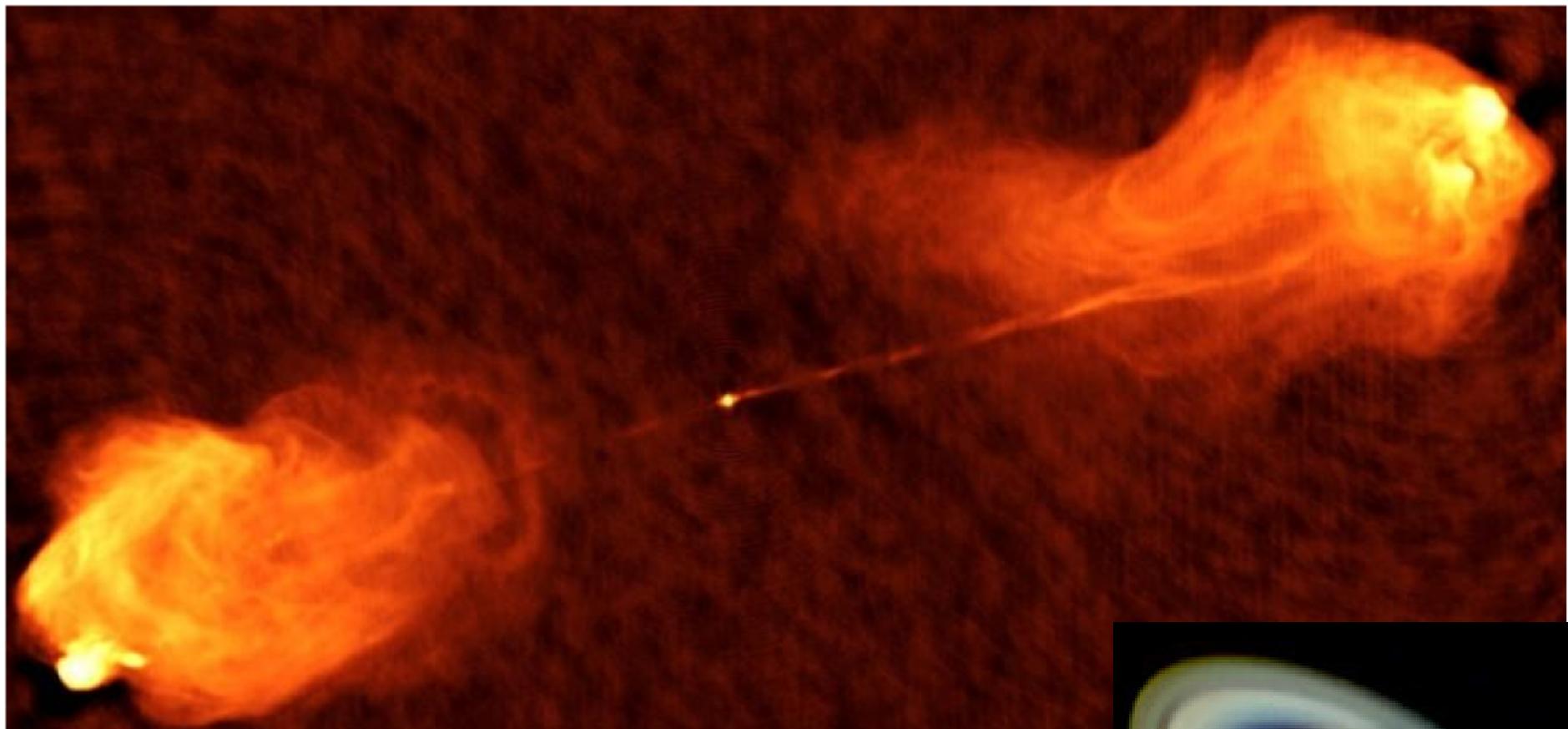
# Outline

- It's completely impossible (with one dish)
  - Telescope resolution
- It's possible, but subtle with two
  - Interferometry
- It relies on coherence and incoherence

At 5 GHz (6cm) with resolution (pixel size): 0.5 arcsec



At 5 GHz (6cm) with resolution (pixel size): 0.5 arcsec

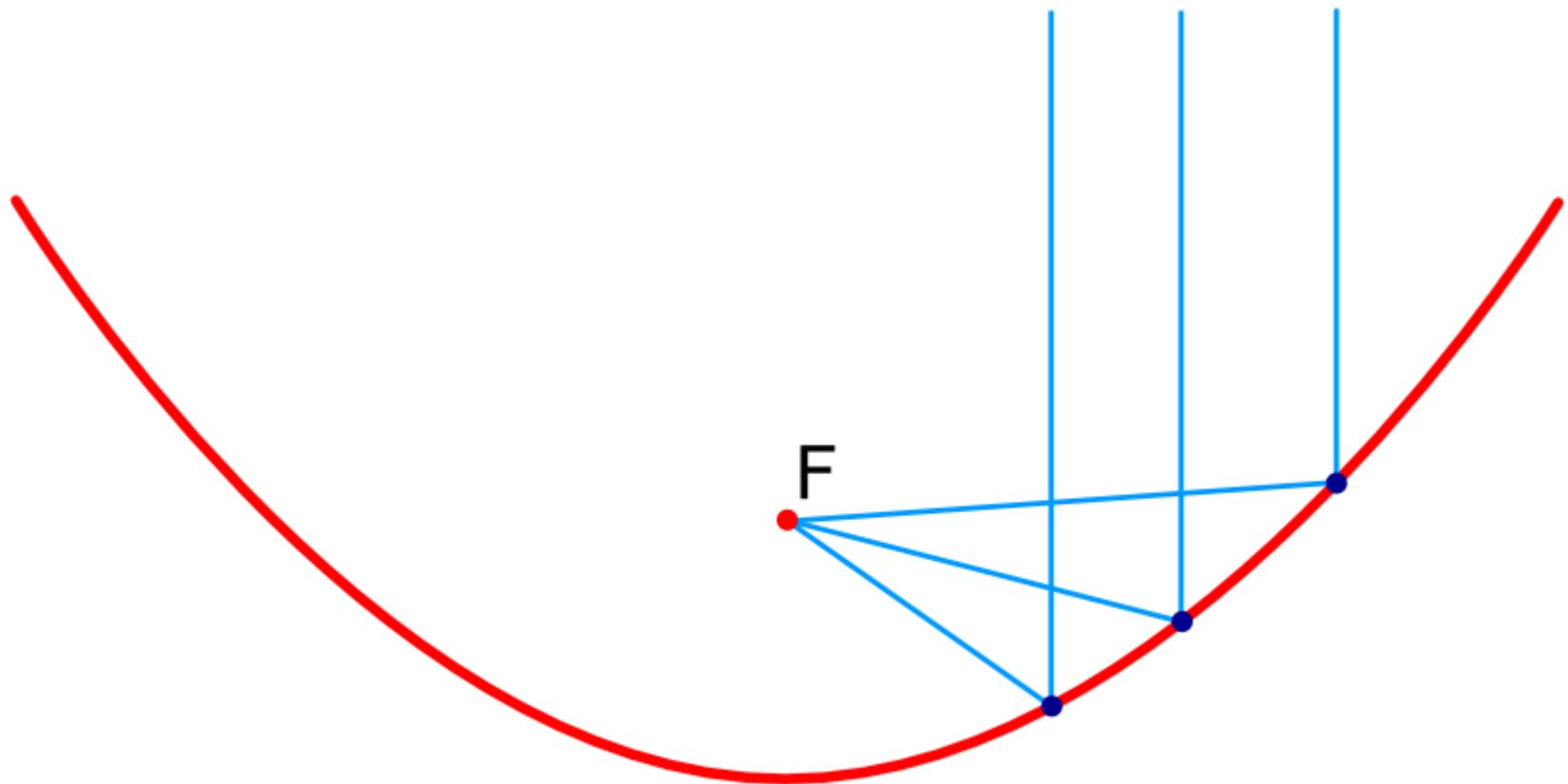


2 arcmin

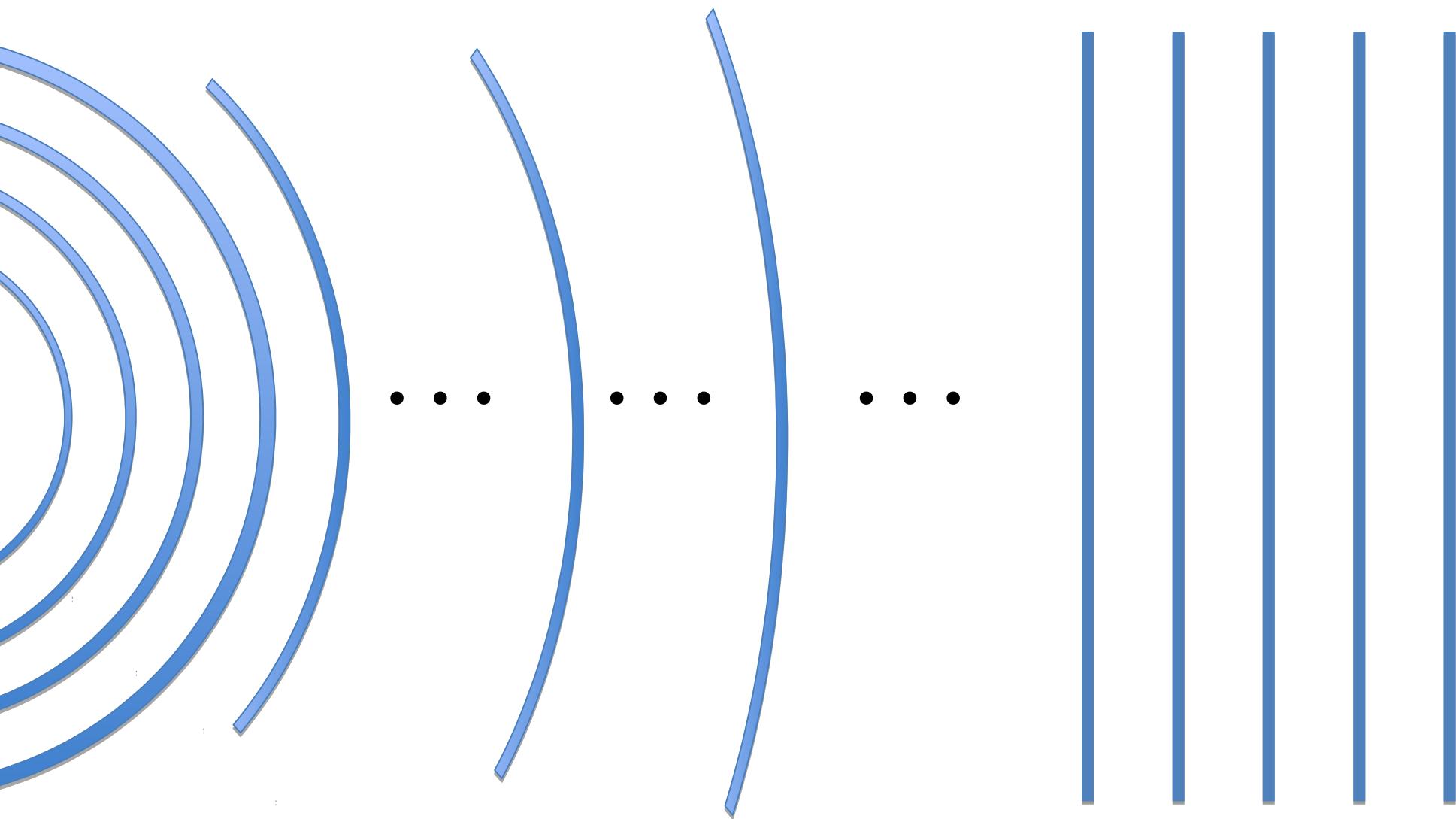


1 arcmin

# Telescope Resolution



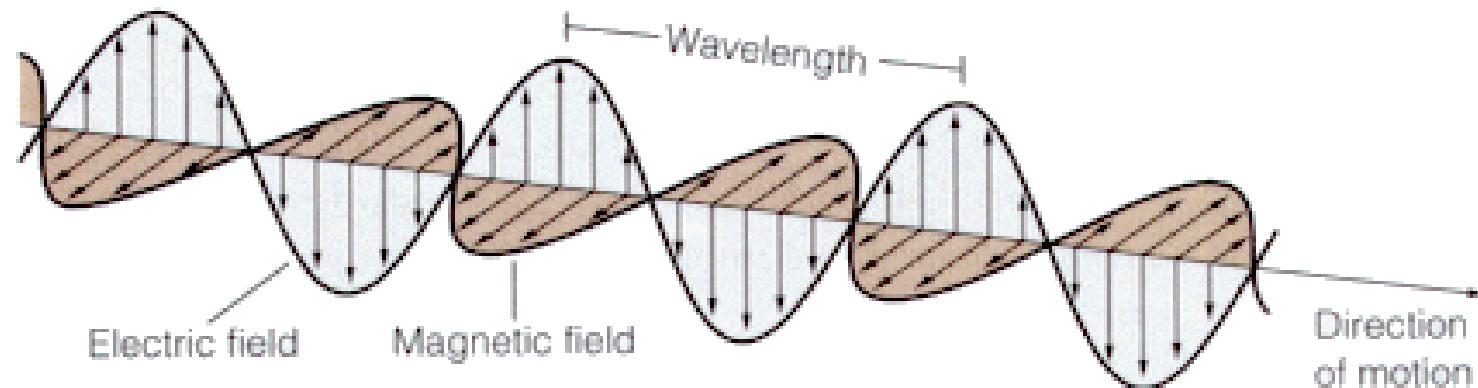
# EM Waves



# Energy, Intensity, Power

$$\text{Energy Density } [\text{J/m}^3] = \frac{1}{2}\epsilon_0 E^2$$

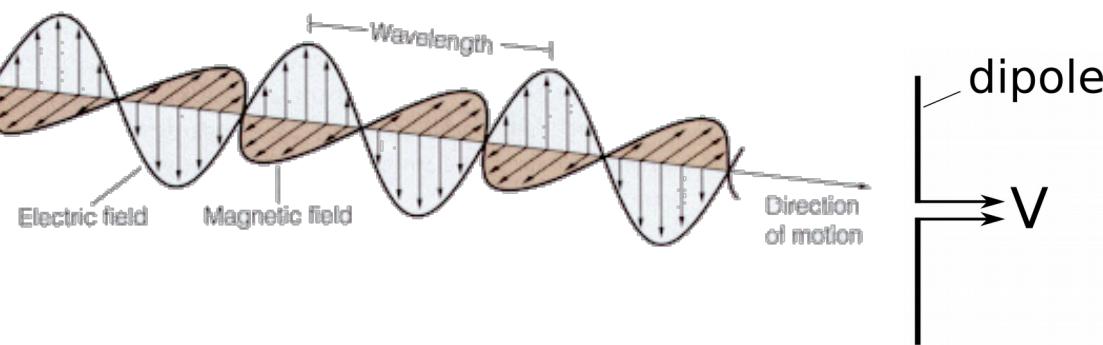
$$\text{Intensity of EM wave } [\text{W/m}^2] = \frac{1}{2}c\epsilon_0 E^2$$



# Energy, Intensity, Power

$$\text{Energy Density } [\text{J/m}^3] = \frac{1}{2}\epsilon_0 E^2$$

$$\text{Intensity of EM wave } [\text{W/m}^2] = \frac{1}{2}c\epsilon_0 E^2$$

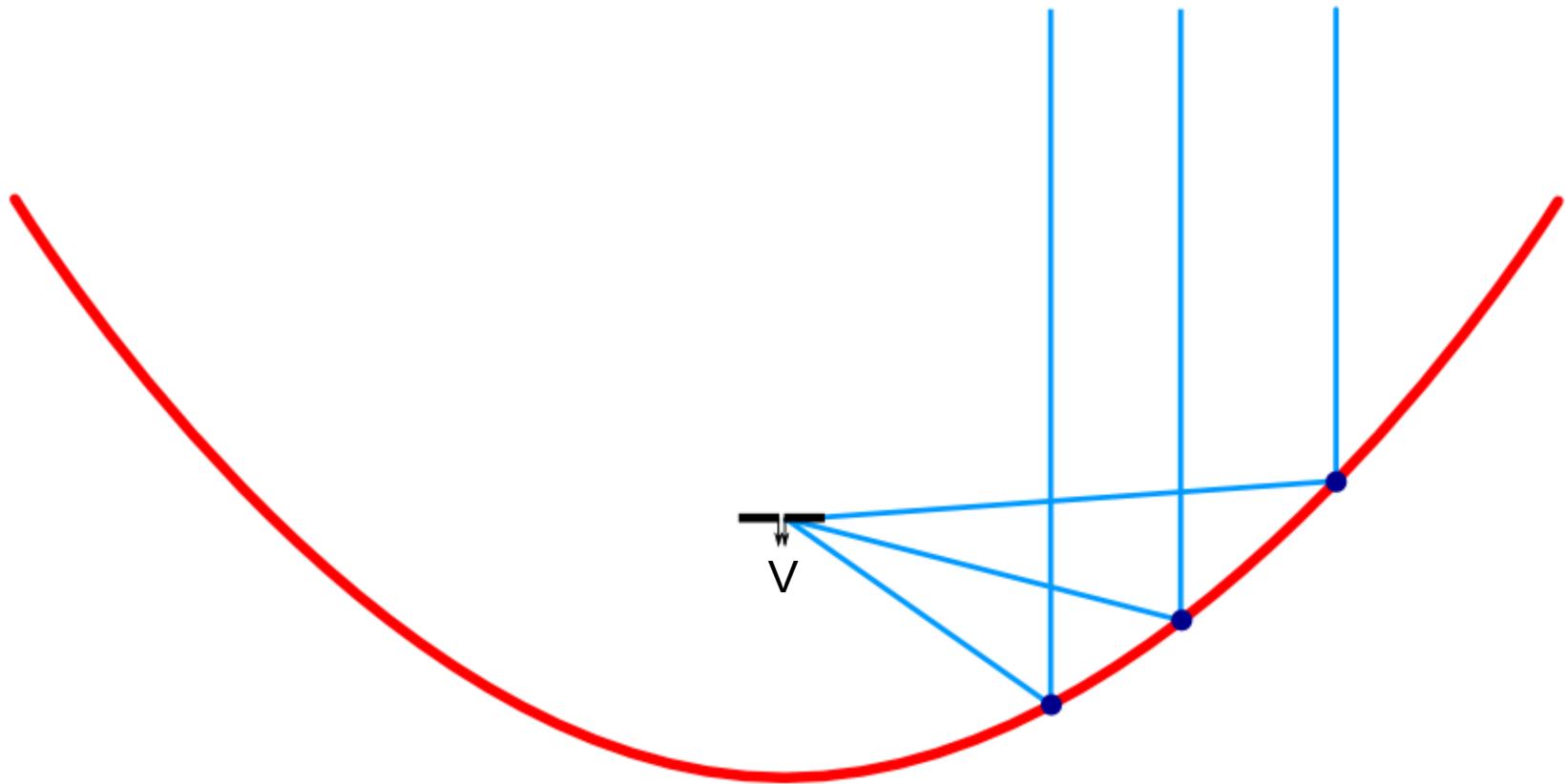


For Antenna:  $V \propto E$

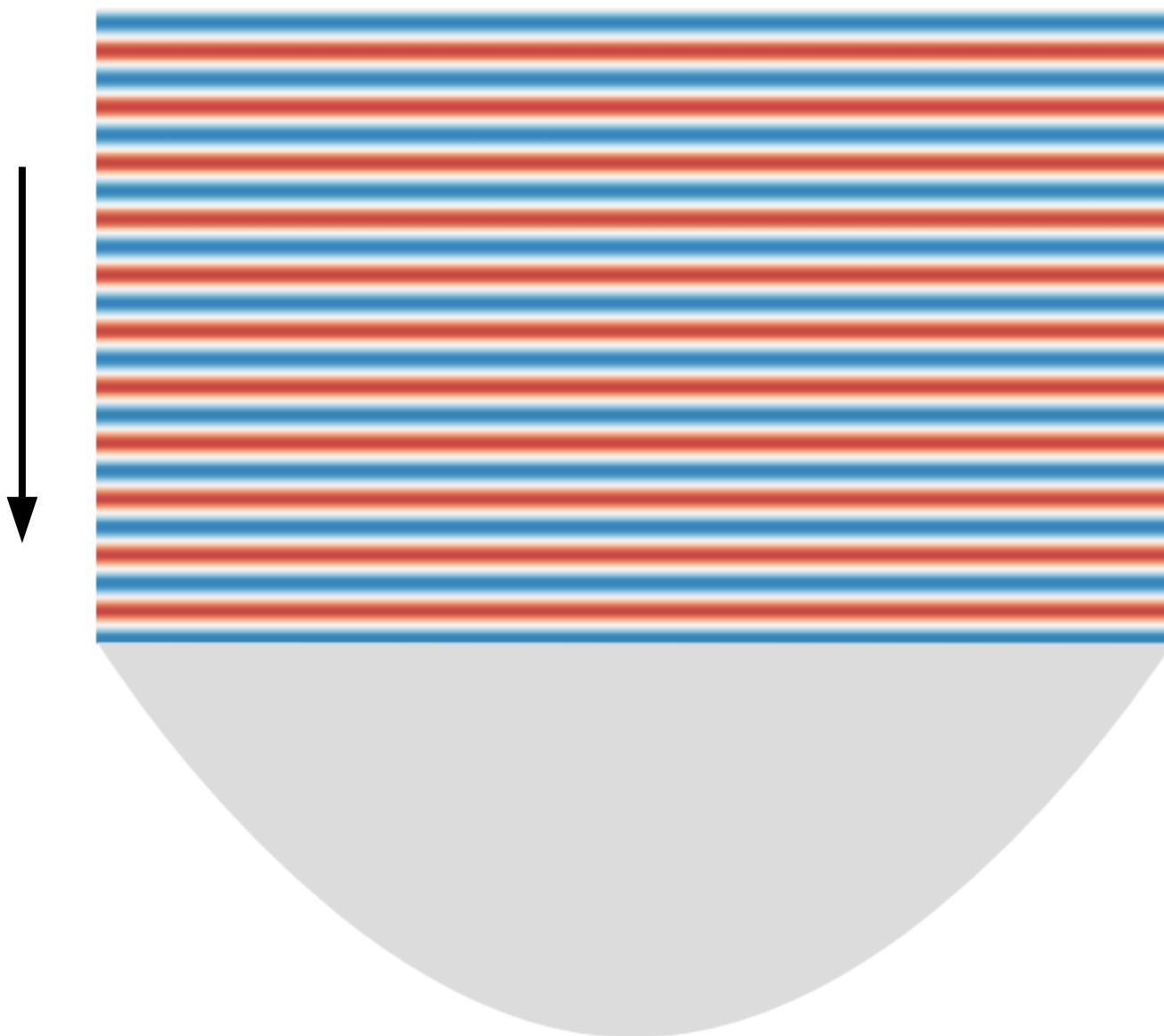
$$\text{Power} = \frac{V^2}{R} \propto E^2$$

Simplify: Treat Electric field as a scalar (i.e. consider just one polarization)

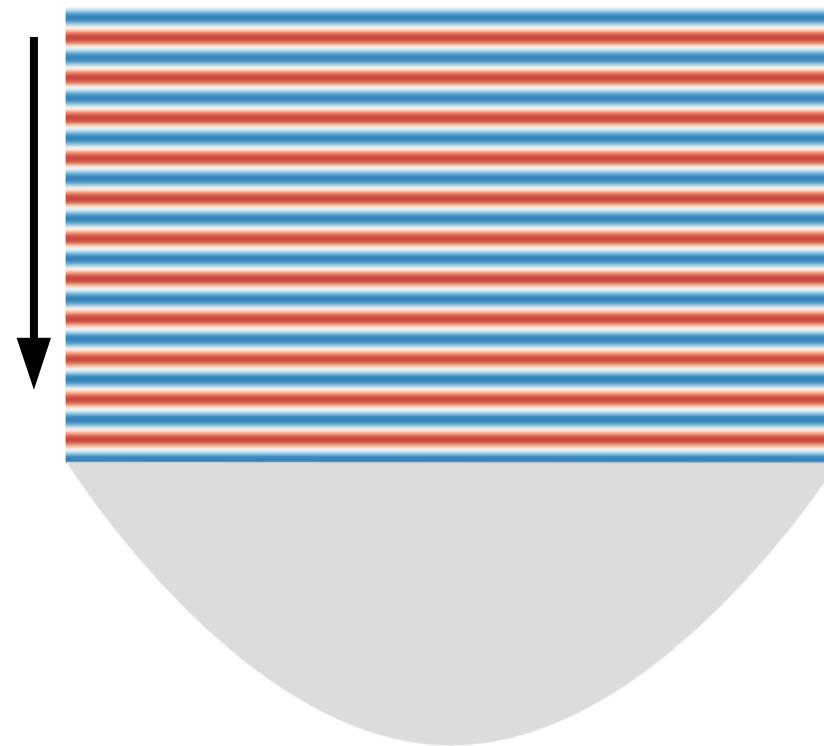
# Parabolic Mirror



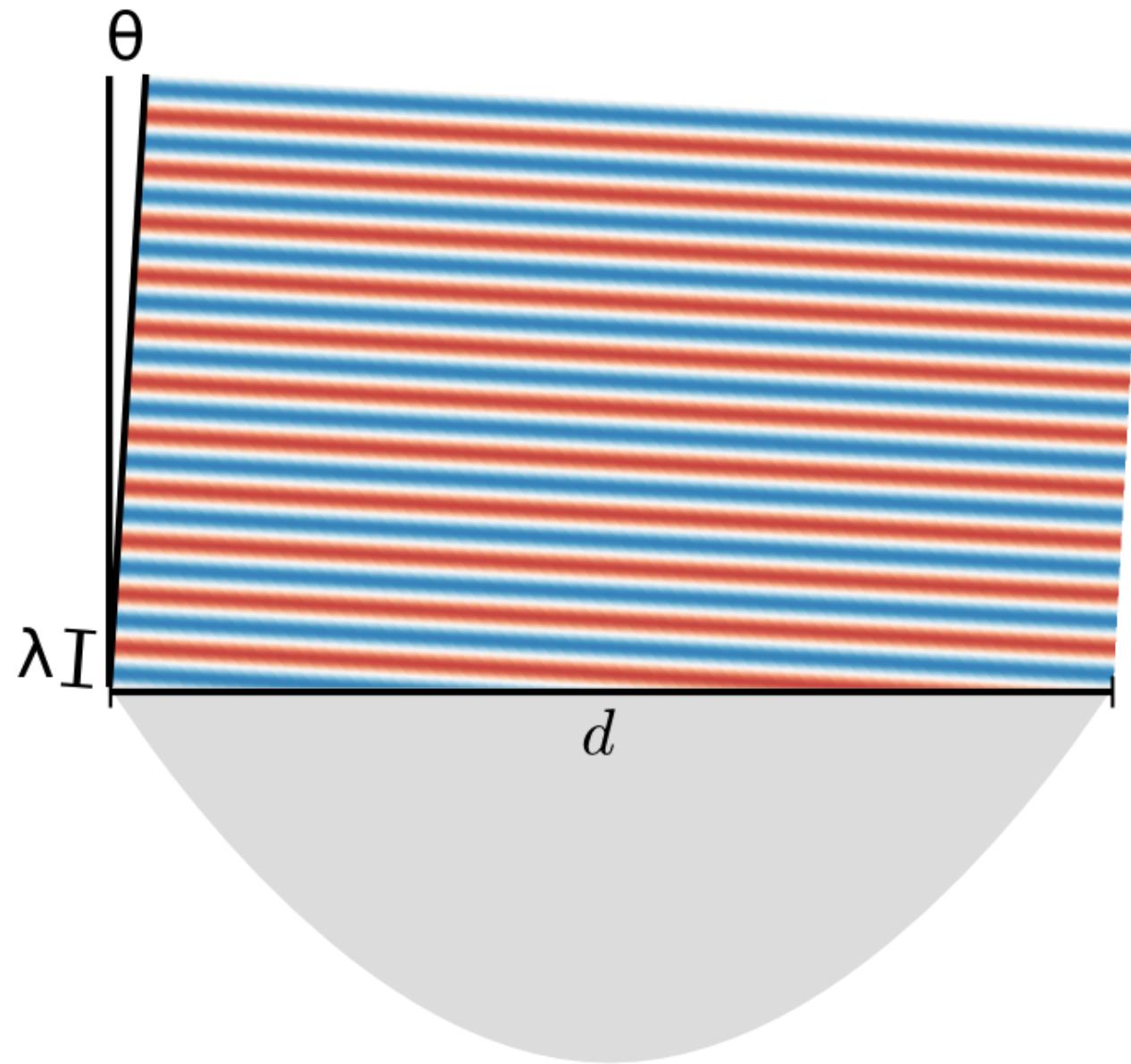
# Incoming Plane Wave



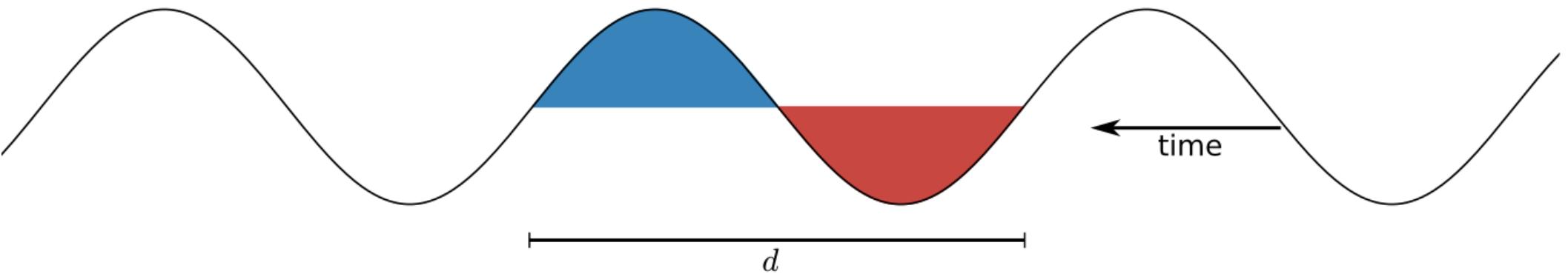
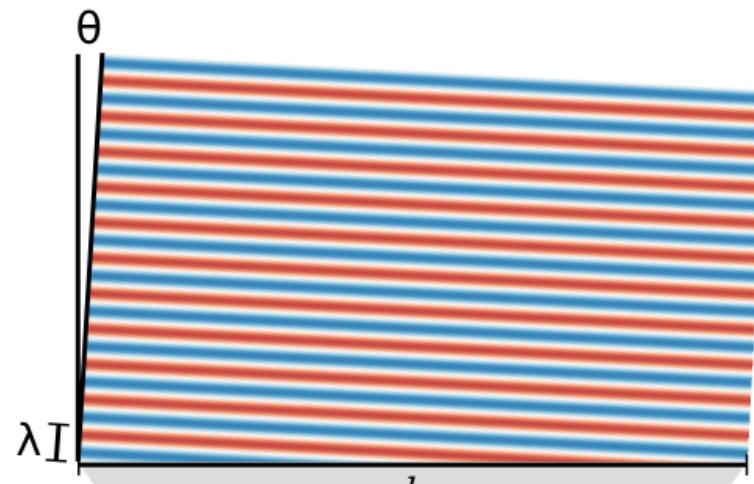
# At the Entrance Aperture



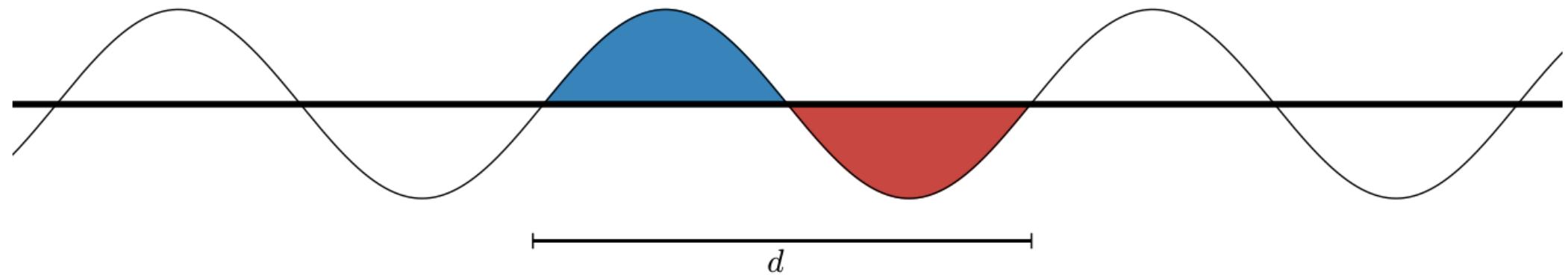
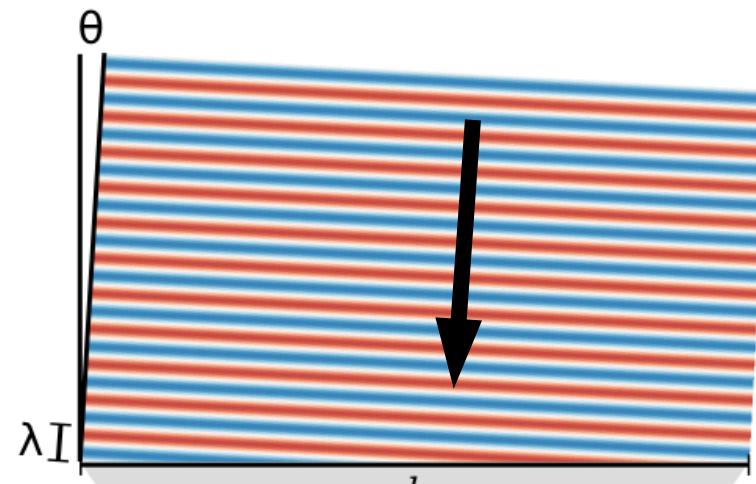
# Telescope Resolution

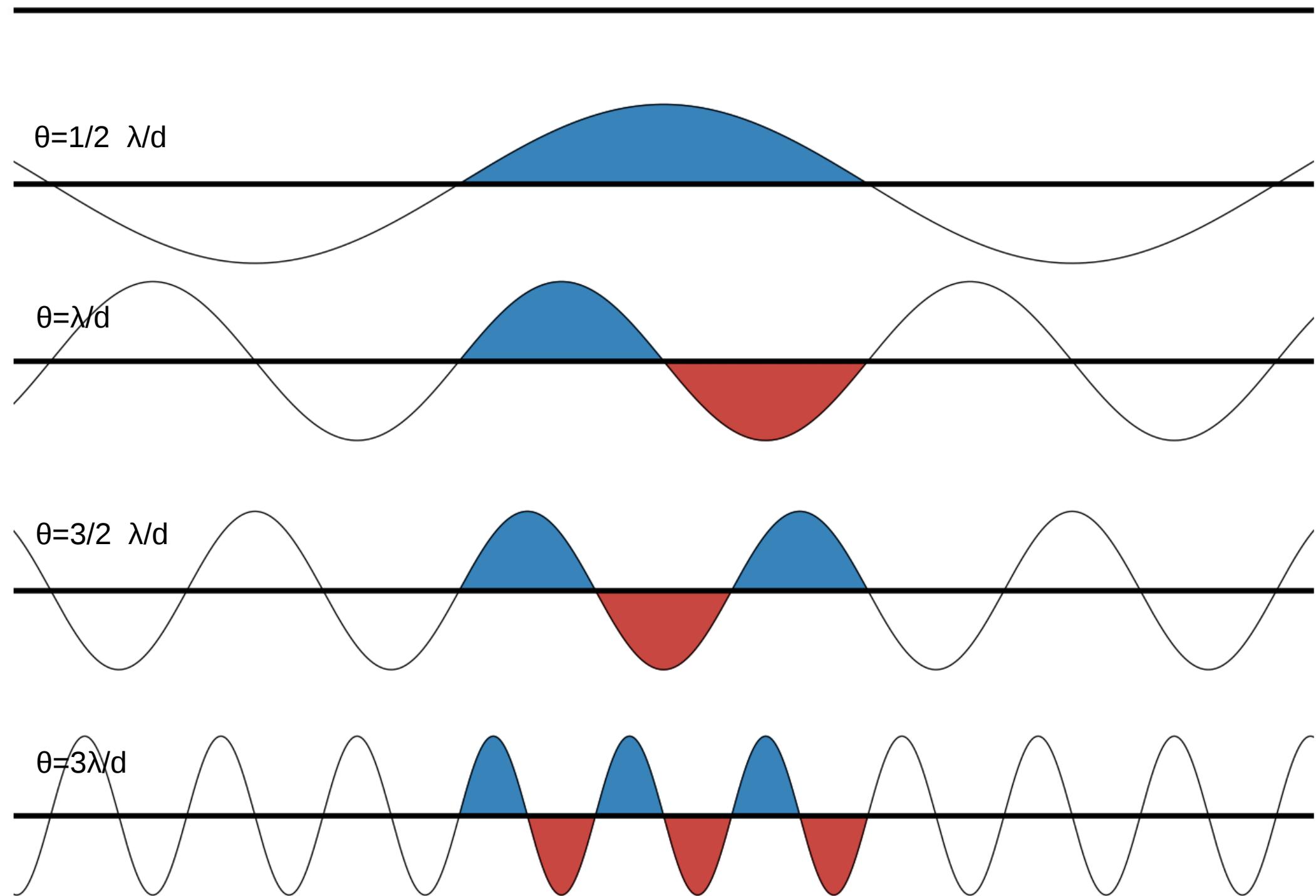


# Telescope Resolution

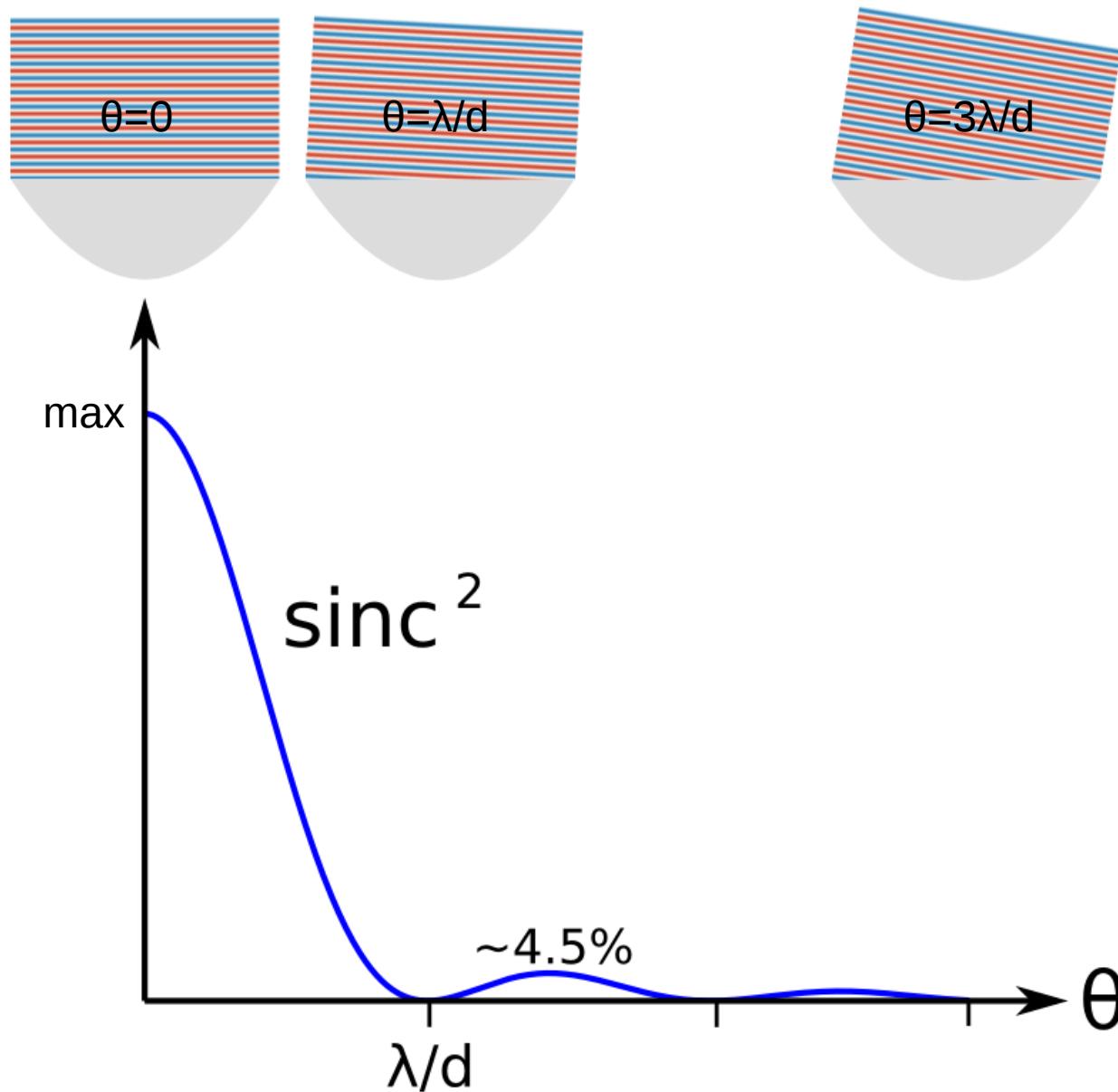


# Telescope Resolution

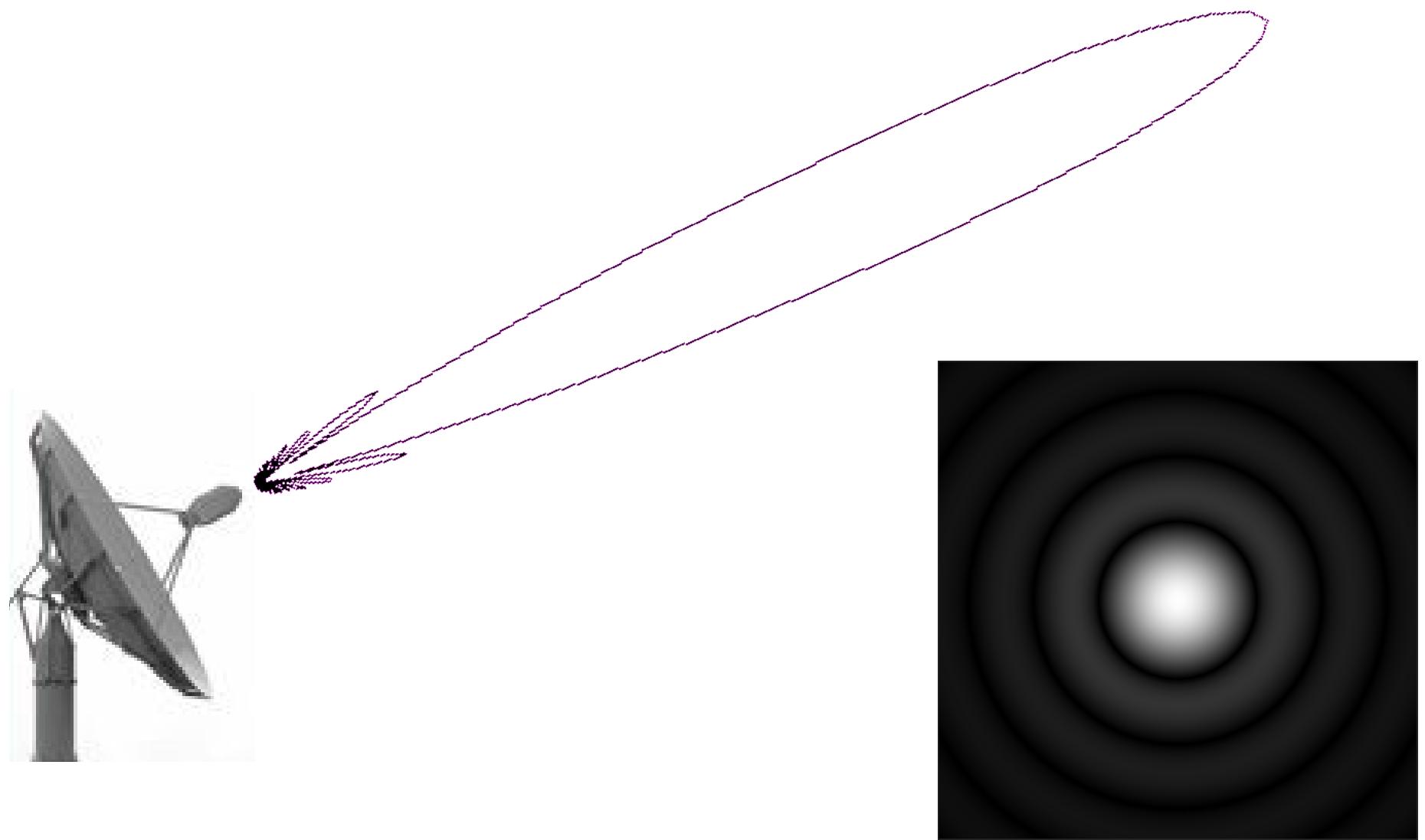


$\theta=0$ 

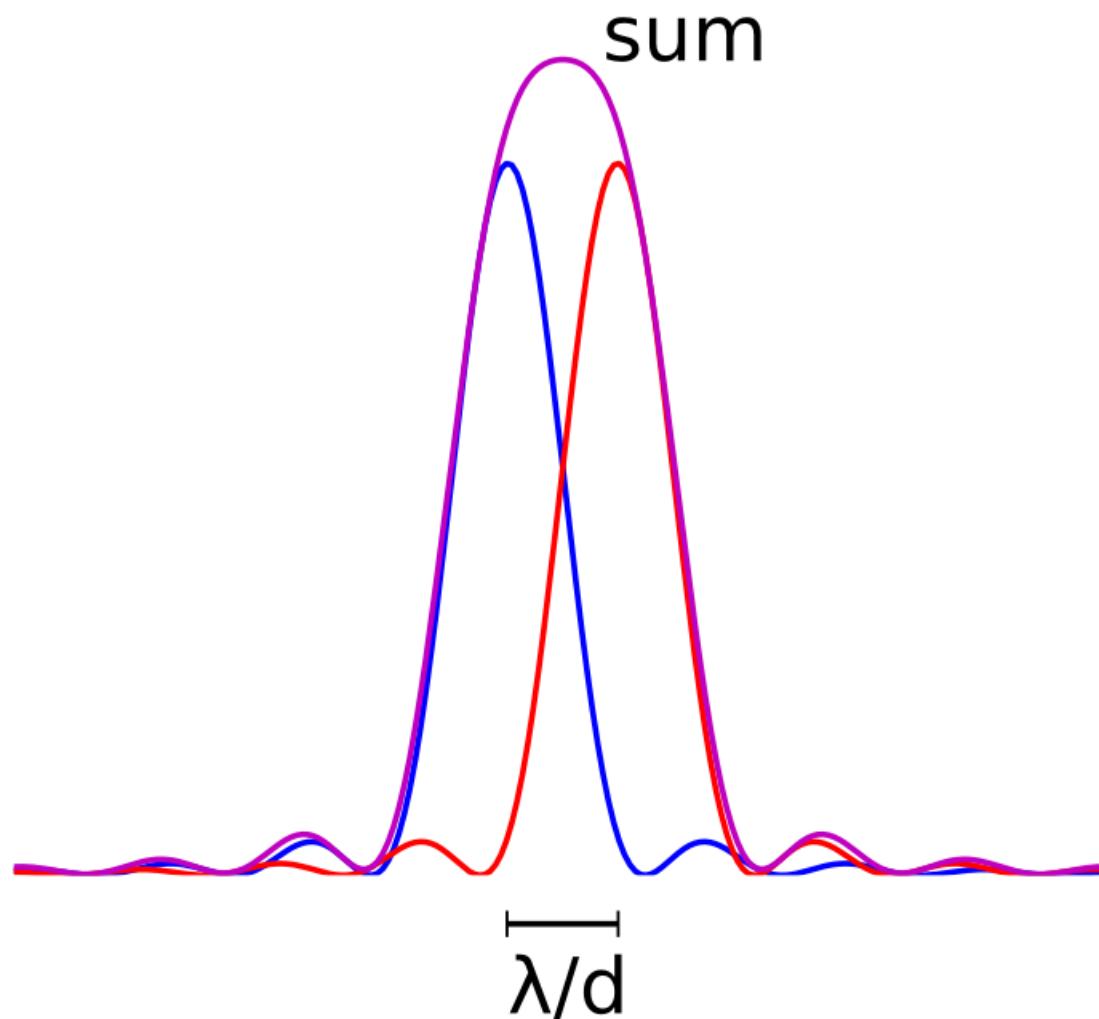
# Telescope Response



# Telescope Resolution

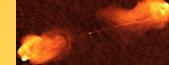


# Rayleigh Criterion

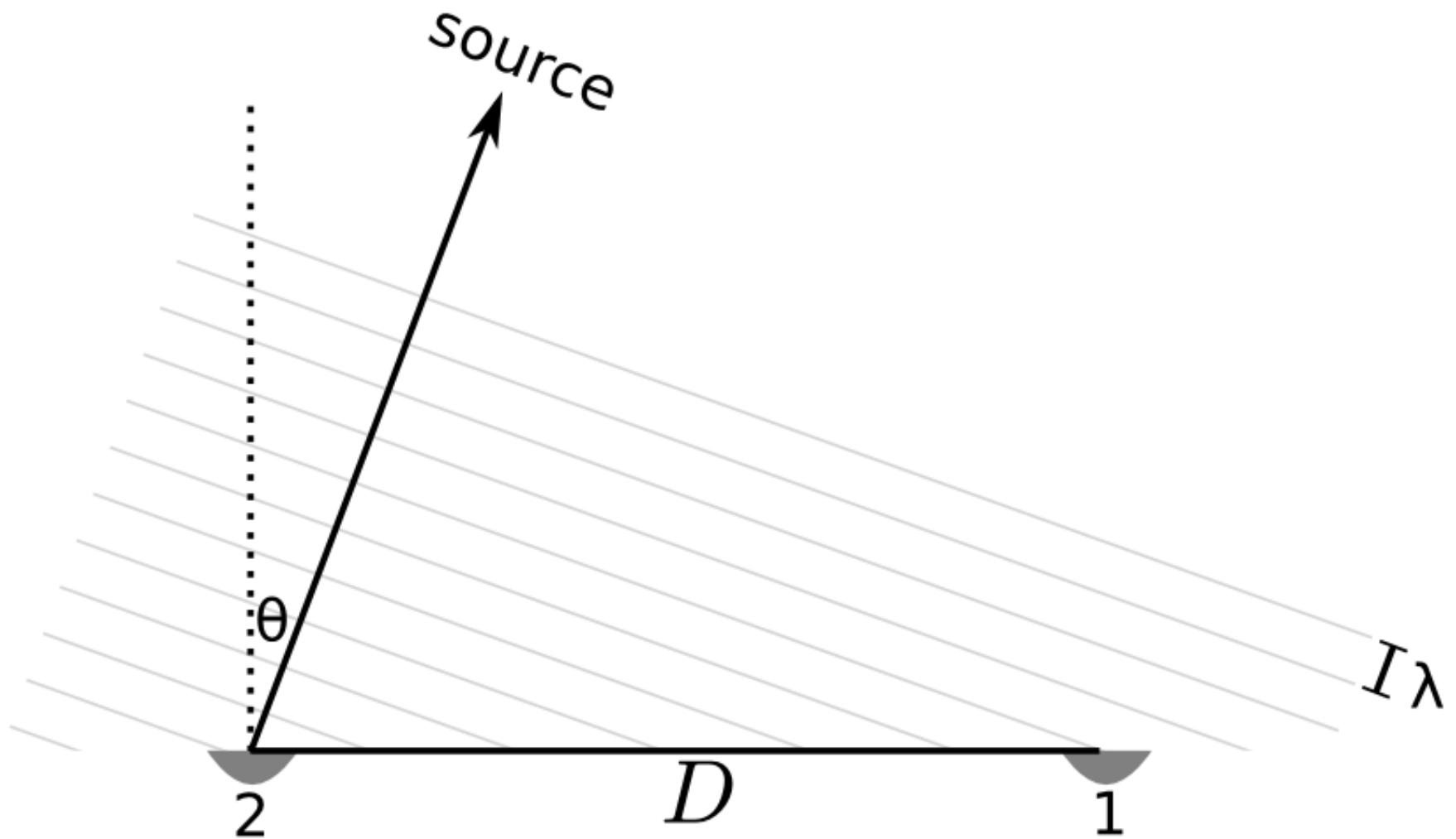


# One Dish Sees

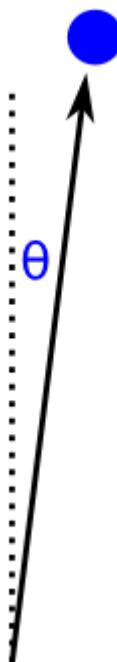
not this



# Two Telescopes

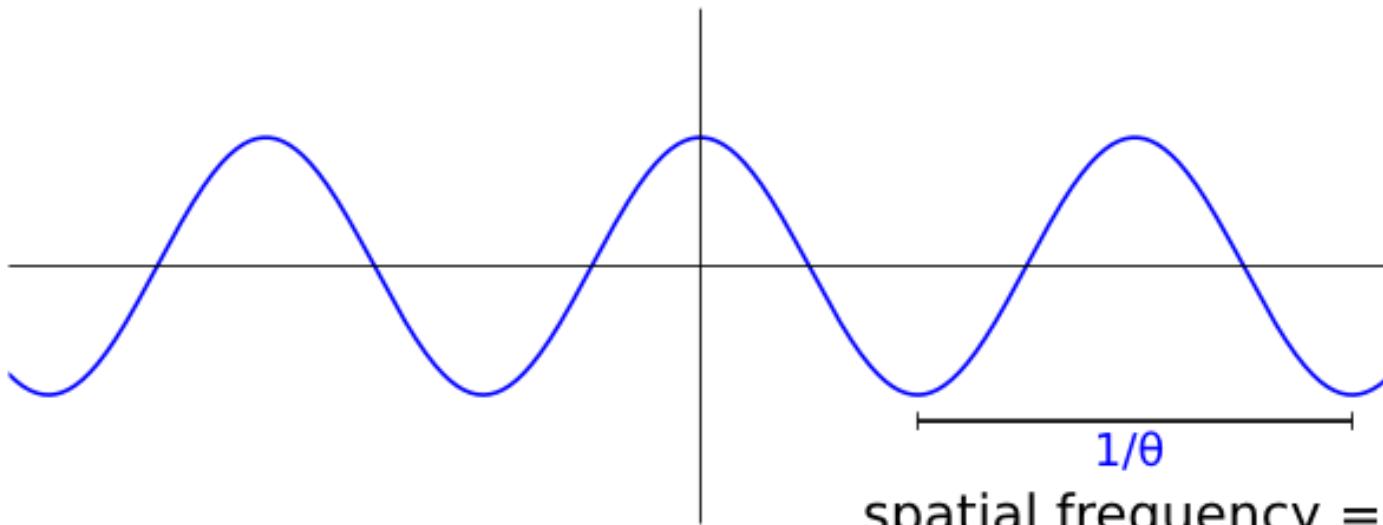


# Interference Fringes



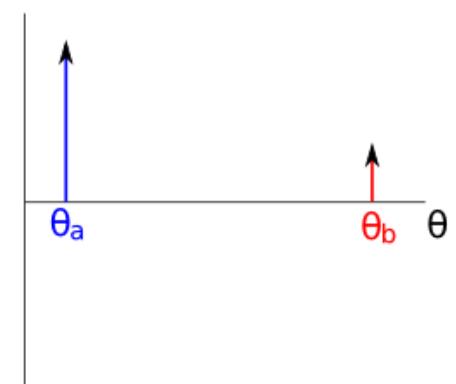
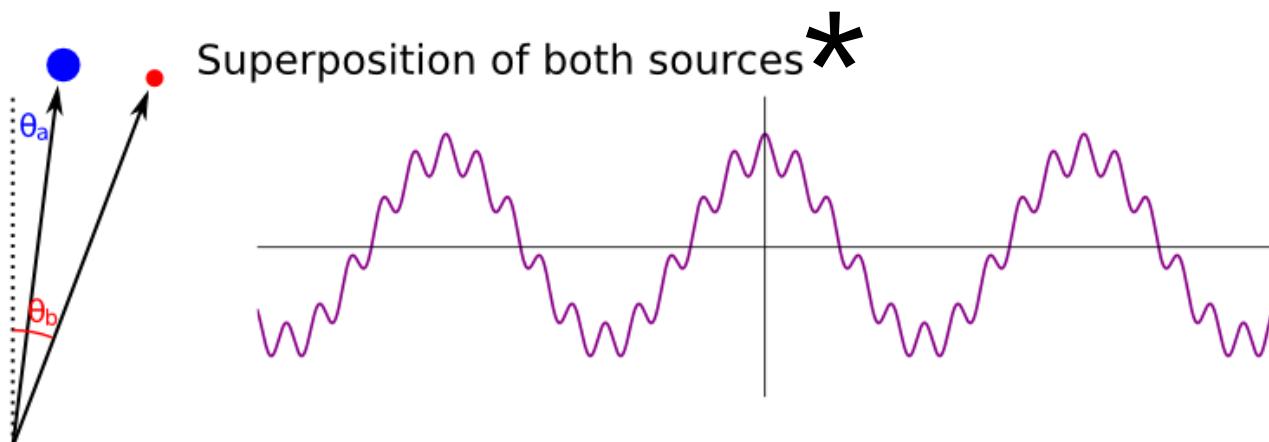
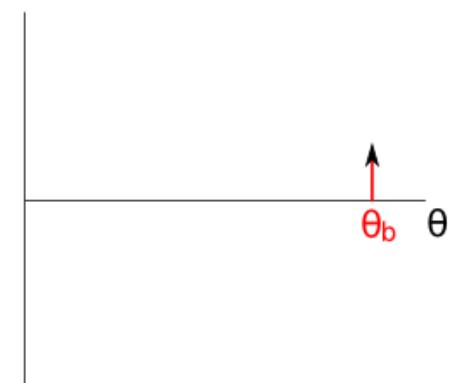
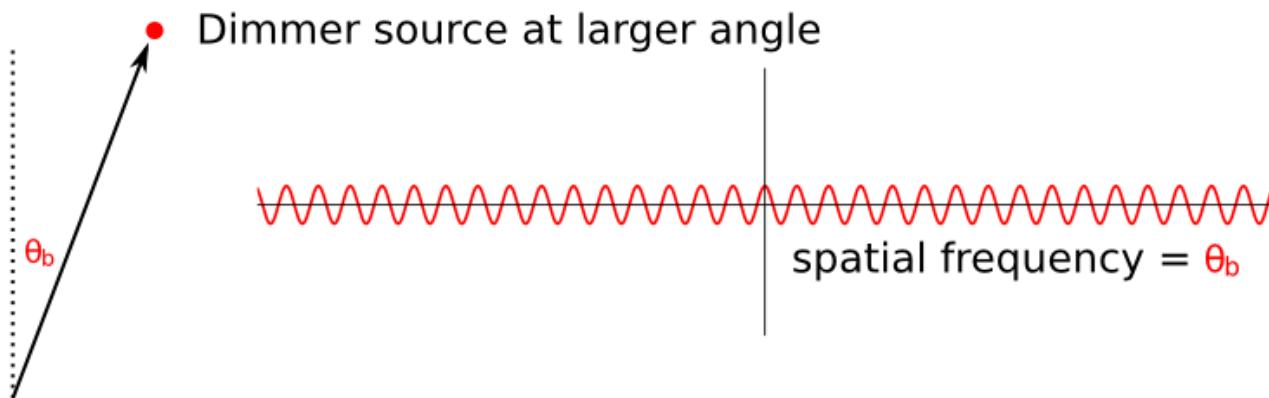
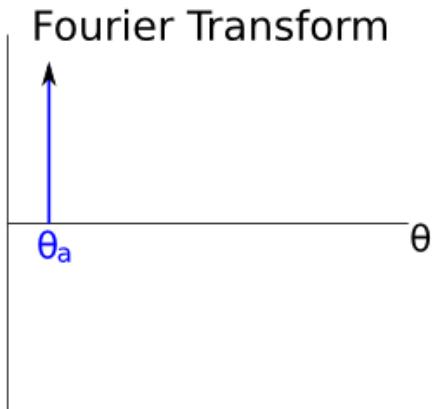
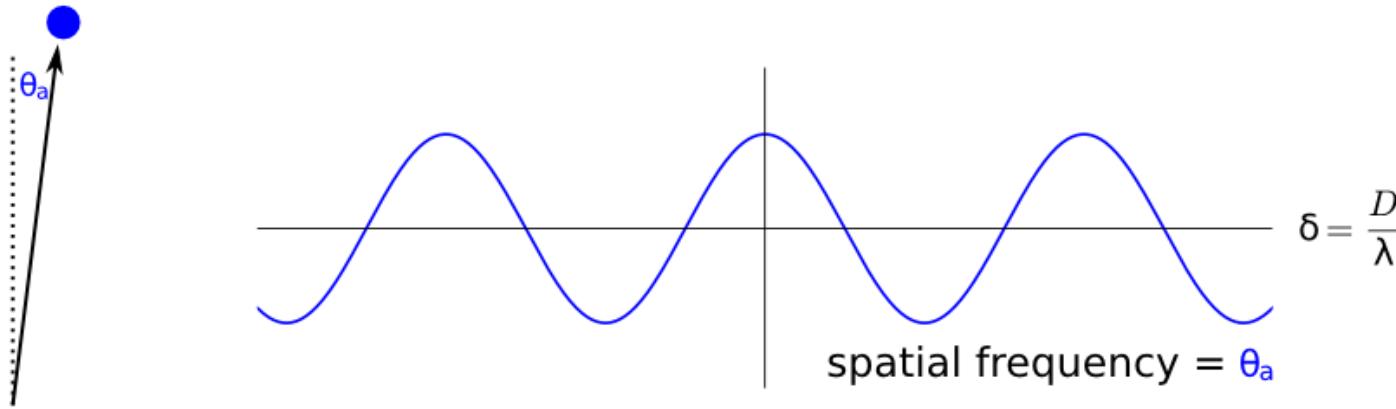
Correlation

$$\langle v_1(t)v_2(t) \rangle = I \cos(2\pi\theta\delta)$$



Telescope Separation

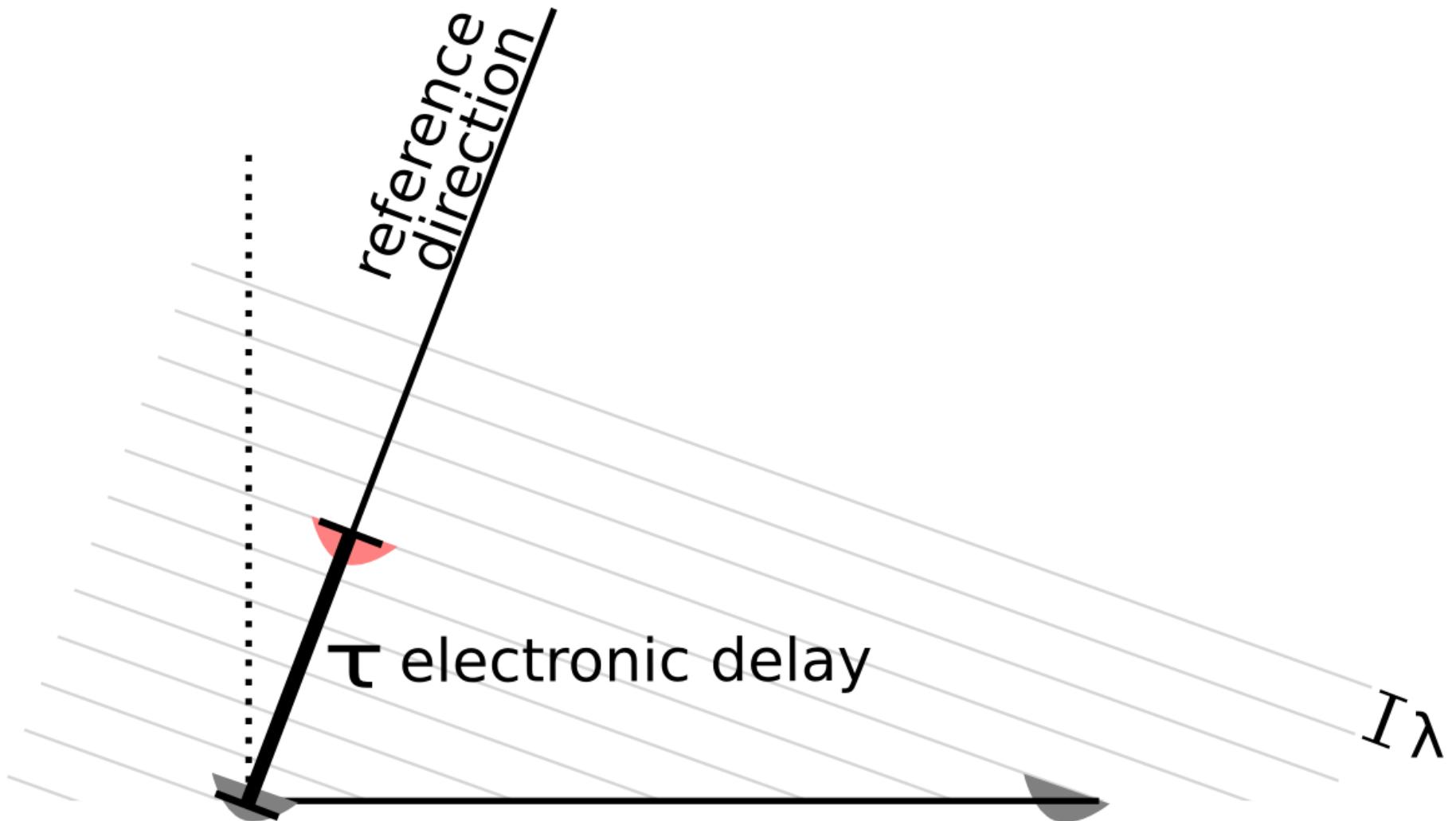
$$\delta = \frac{D}{\lambda}$$



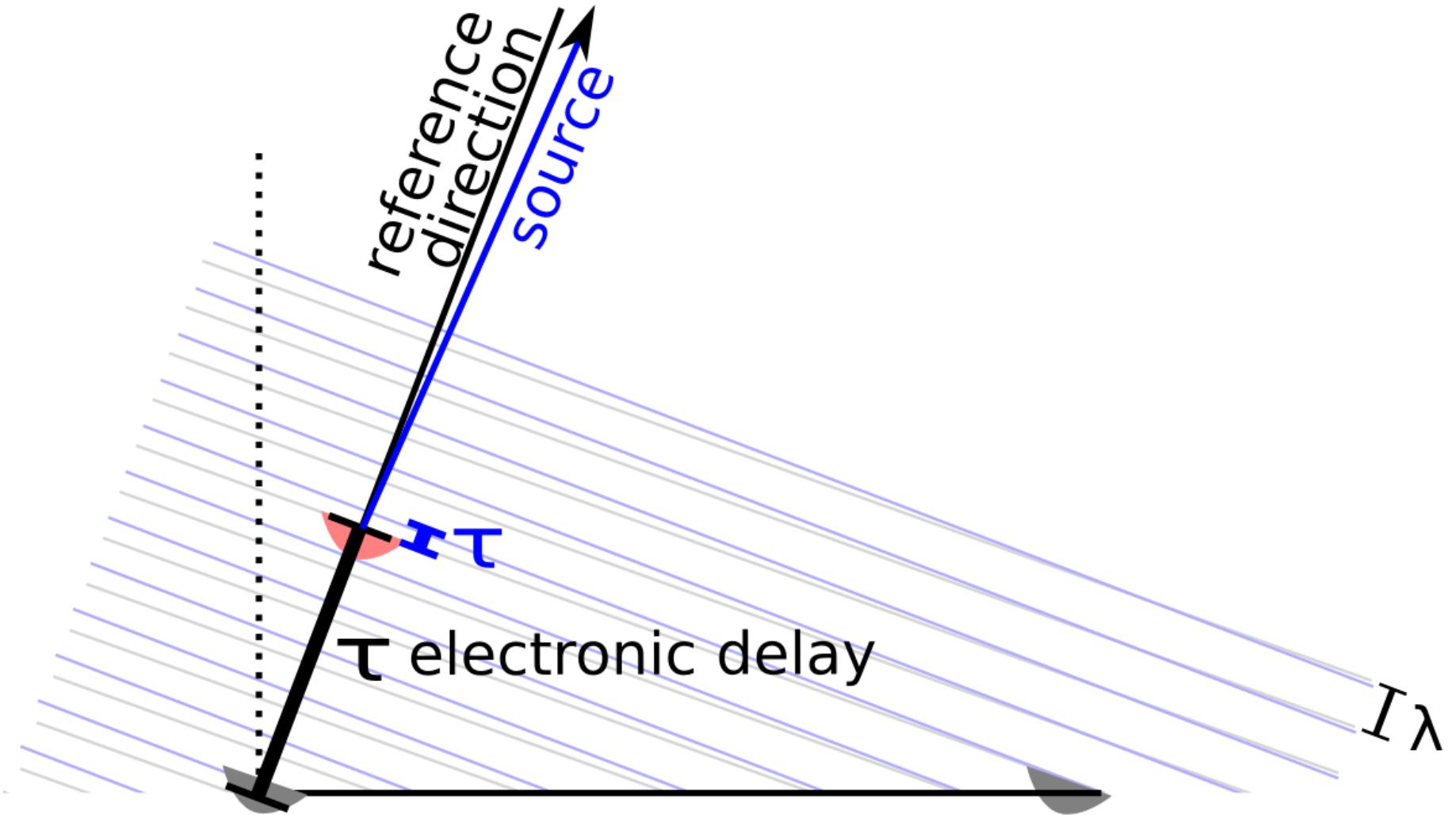
# Two Related Questions

- Sources that aren't almost straight up?
- How do you tell left vs right?

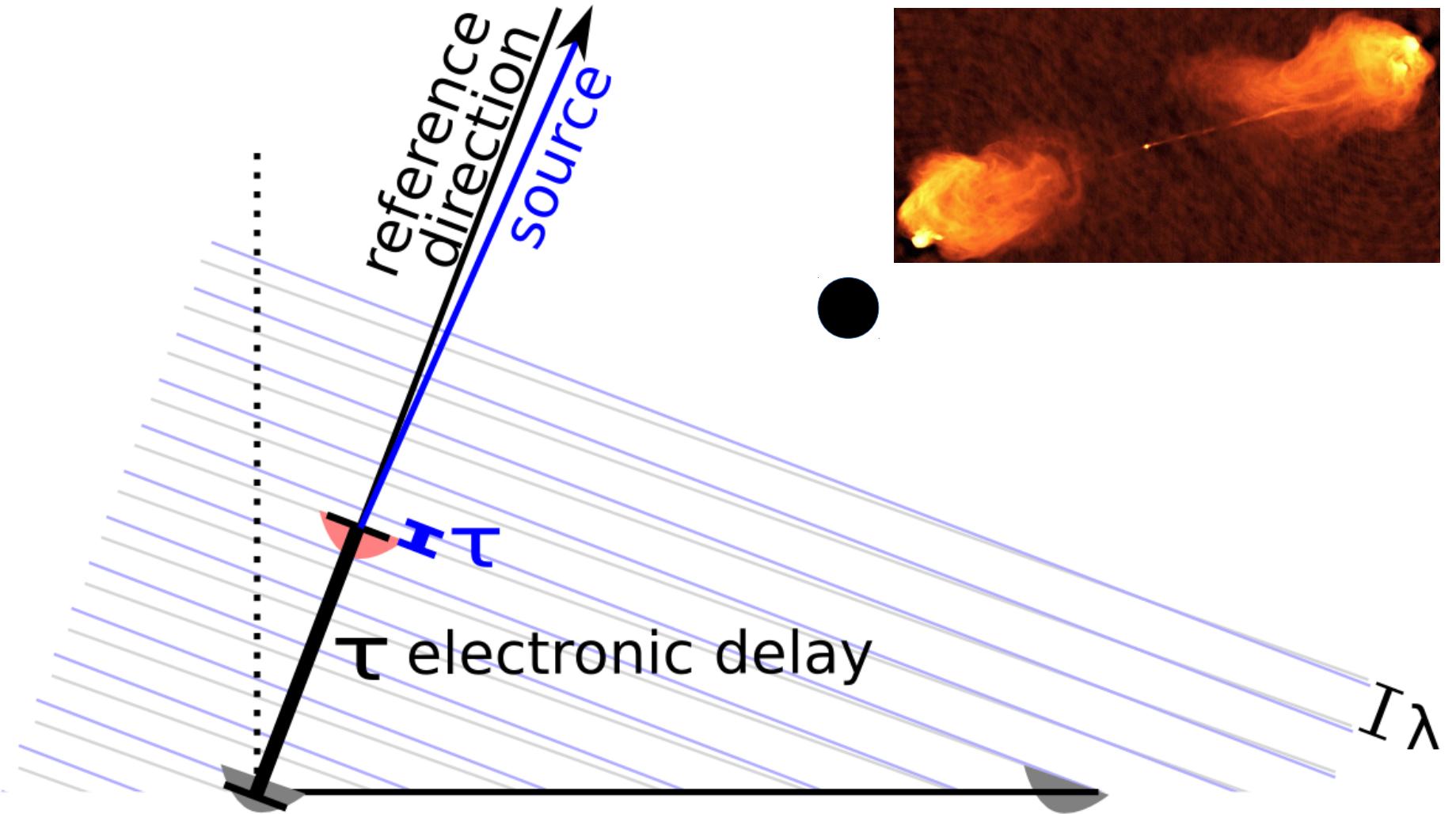
# Sources at Realistic Angles



# Sources at Realistic Angles

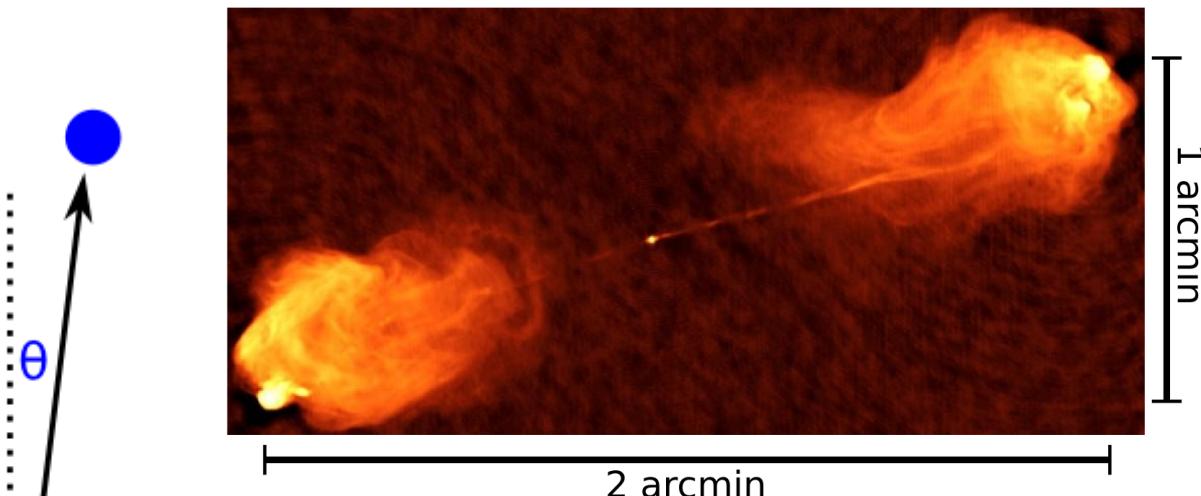


# Sources at Realistic Angles



# 1D version of our example...

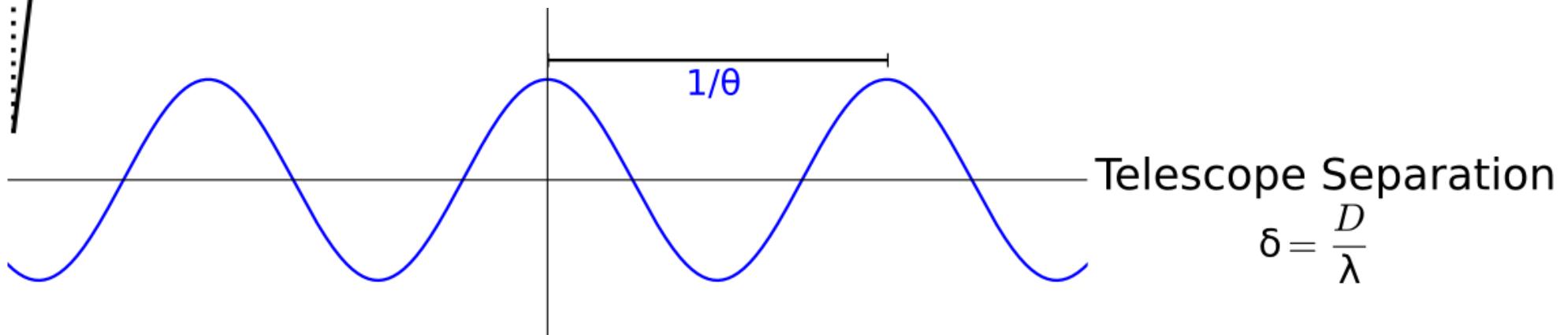
At 5 GHz (6cm) with resolution (pixel size): 0.5 arcsec



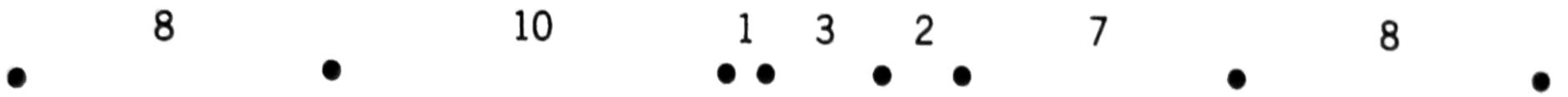
## Questions:

For resolution of 0.5 arcsec,  
what is smallest D required  
to see **one** fringe?

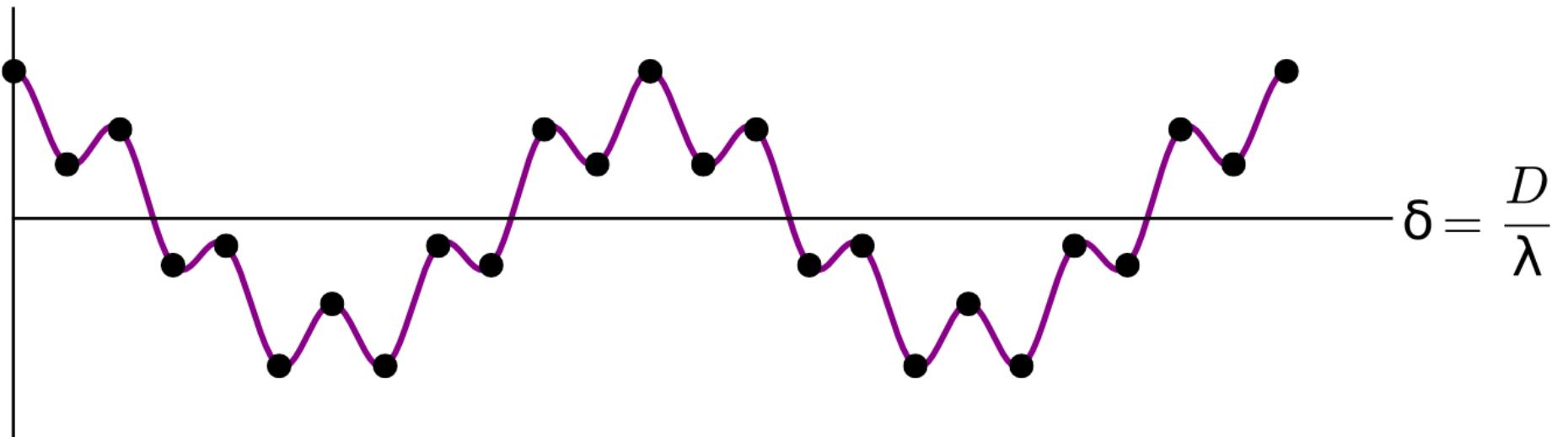
For separation of 2 arcmin,  
how many fringes at this D?



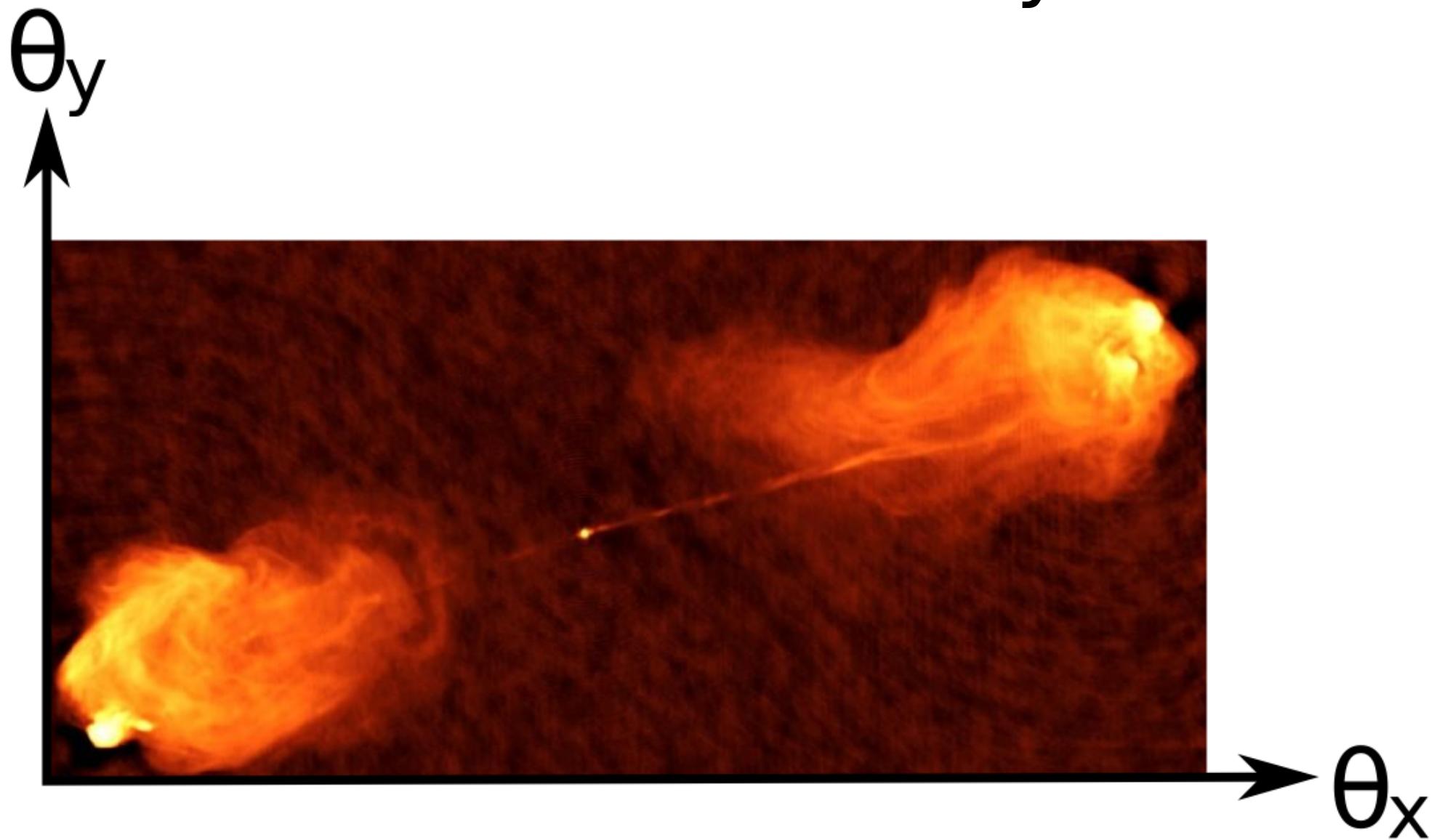
# Few Antennas, Many Baselines



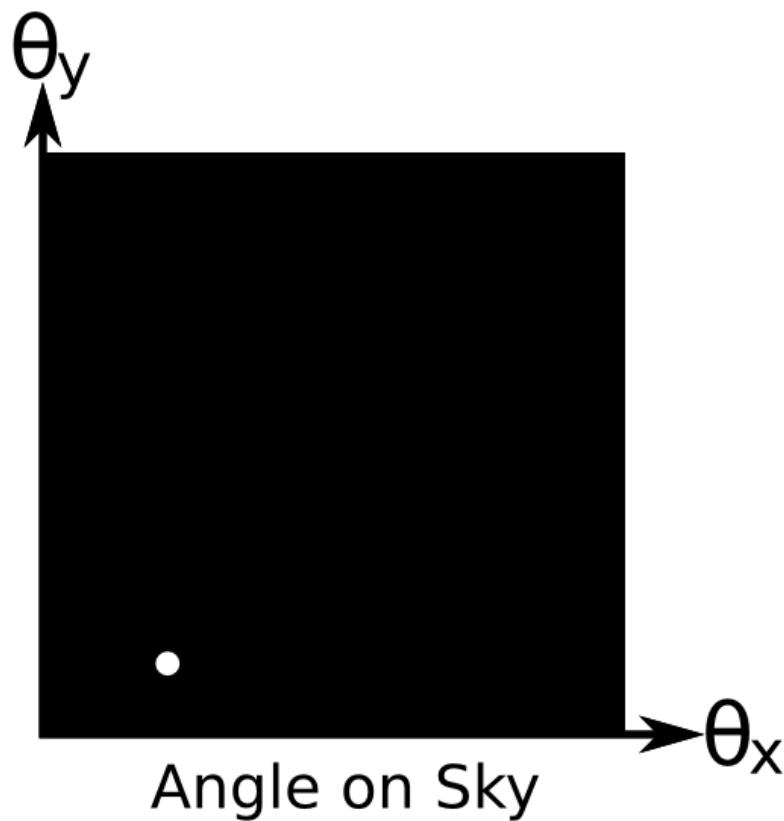
All spacings between 1 and 24 need only 8 antennas!



# 2D Source on Sky

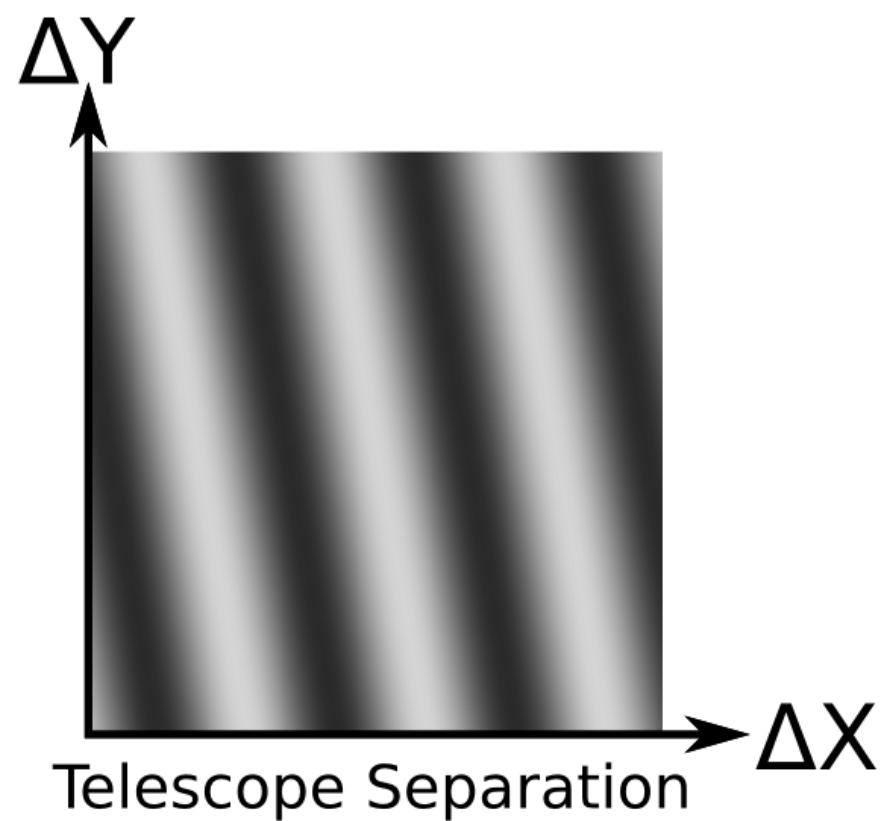


# Source

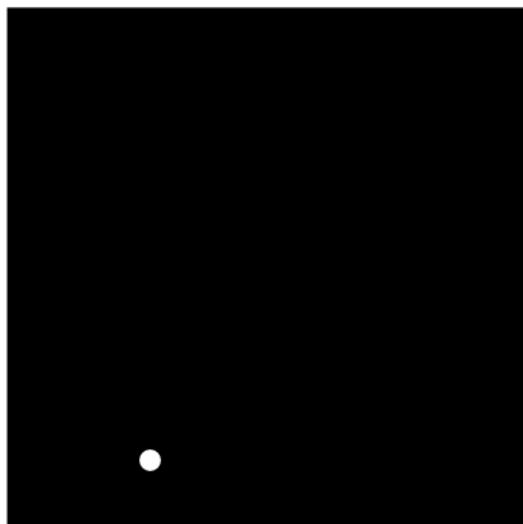


Angle on Sky

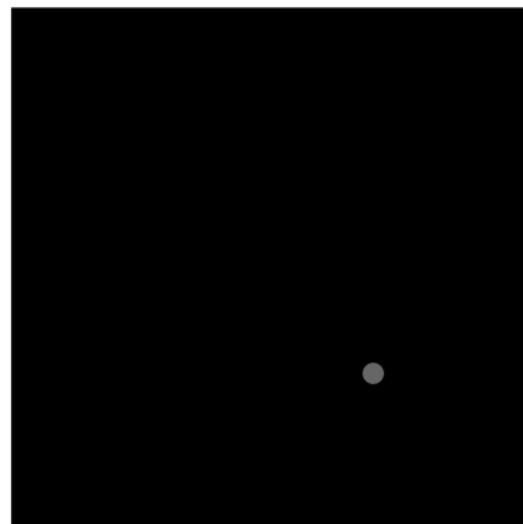
# Correlation



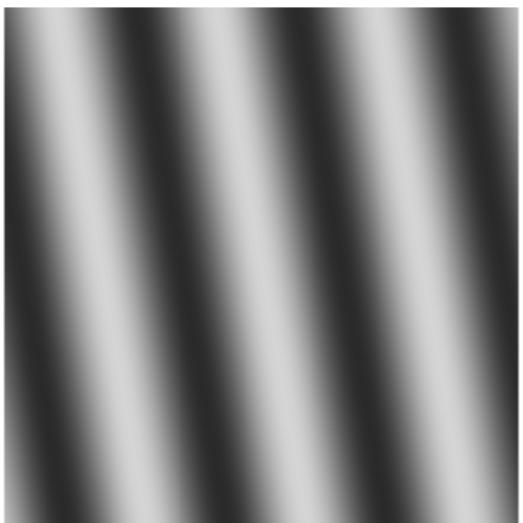
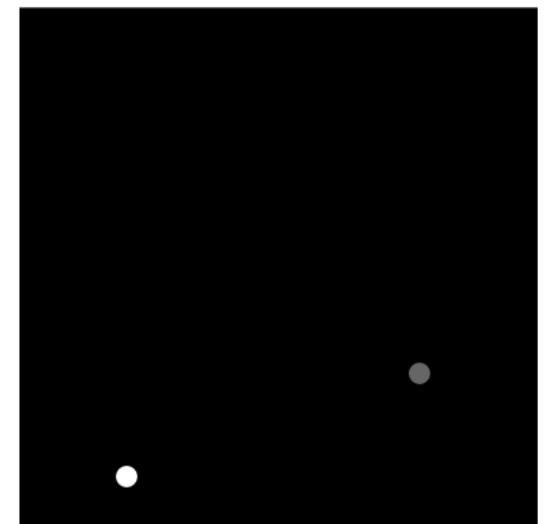
Telescope Separation



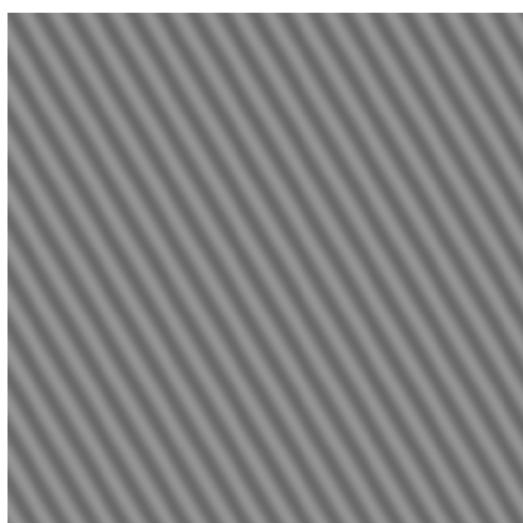
+



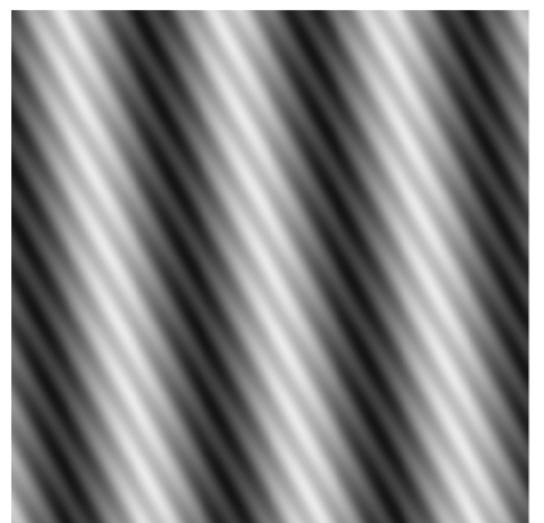
=



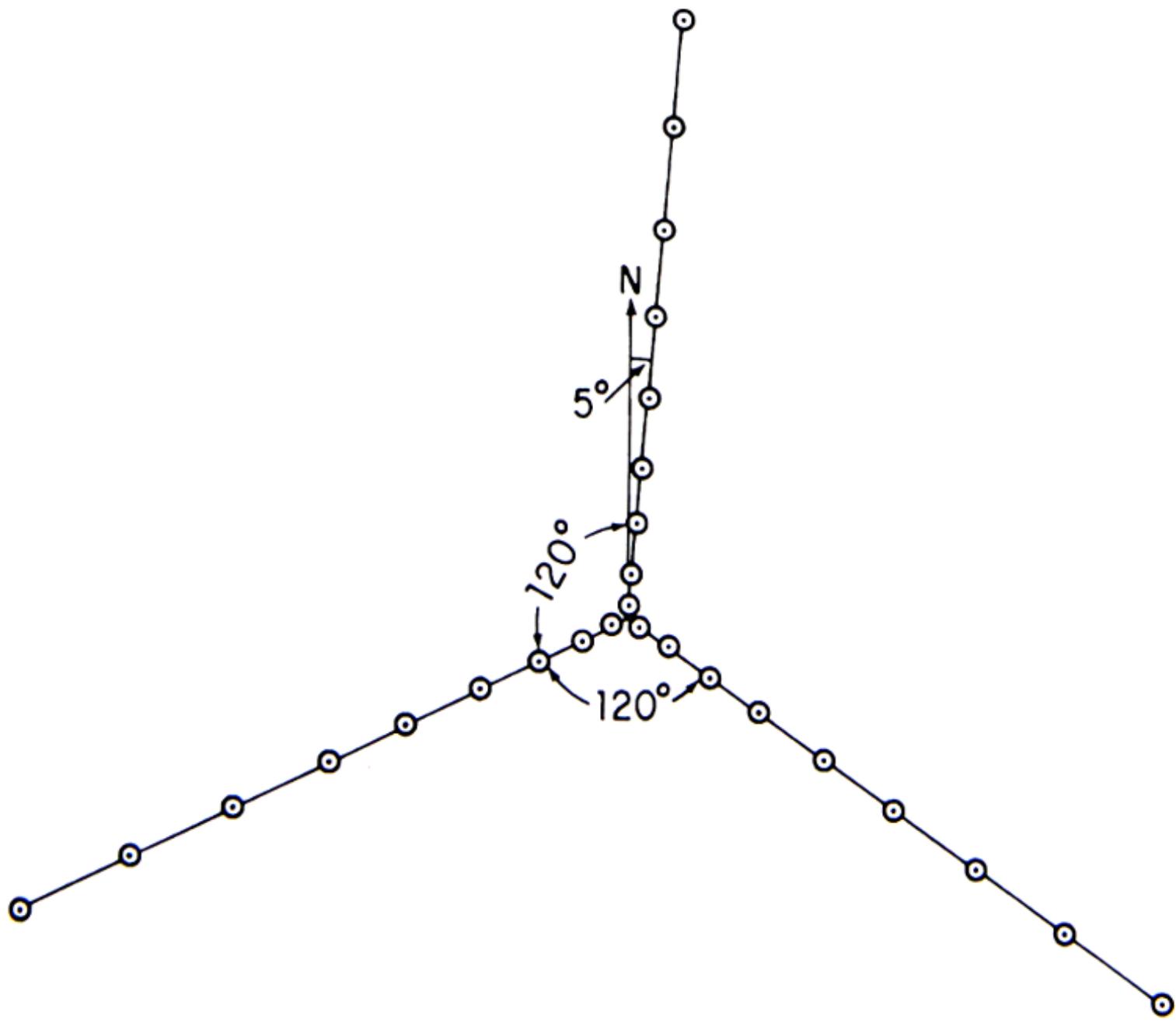
+



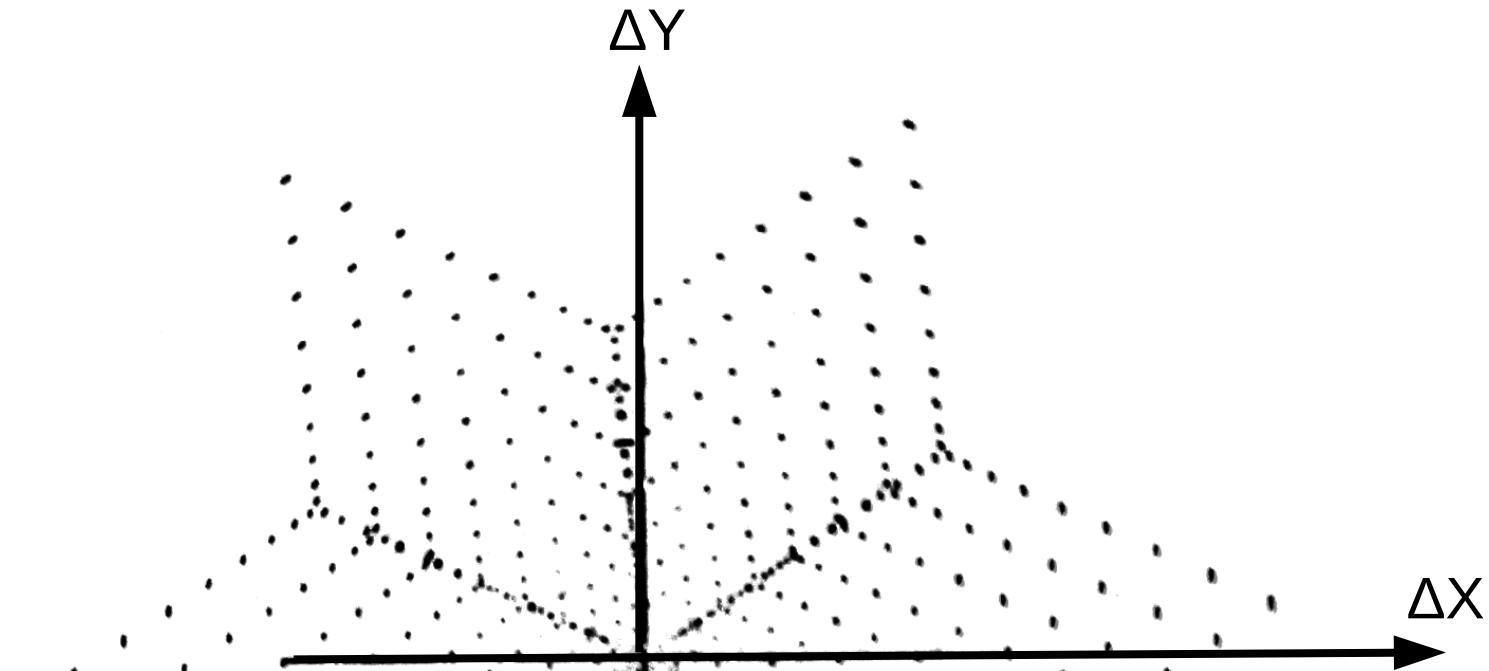
=





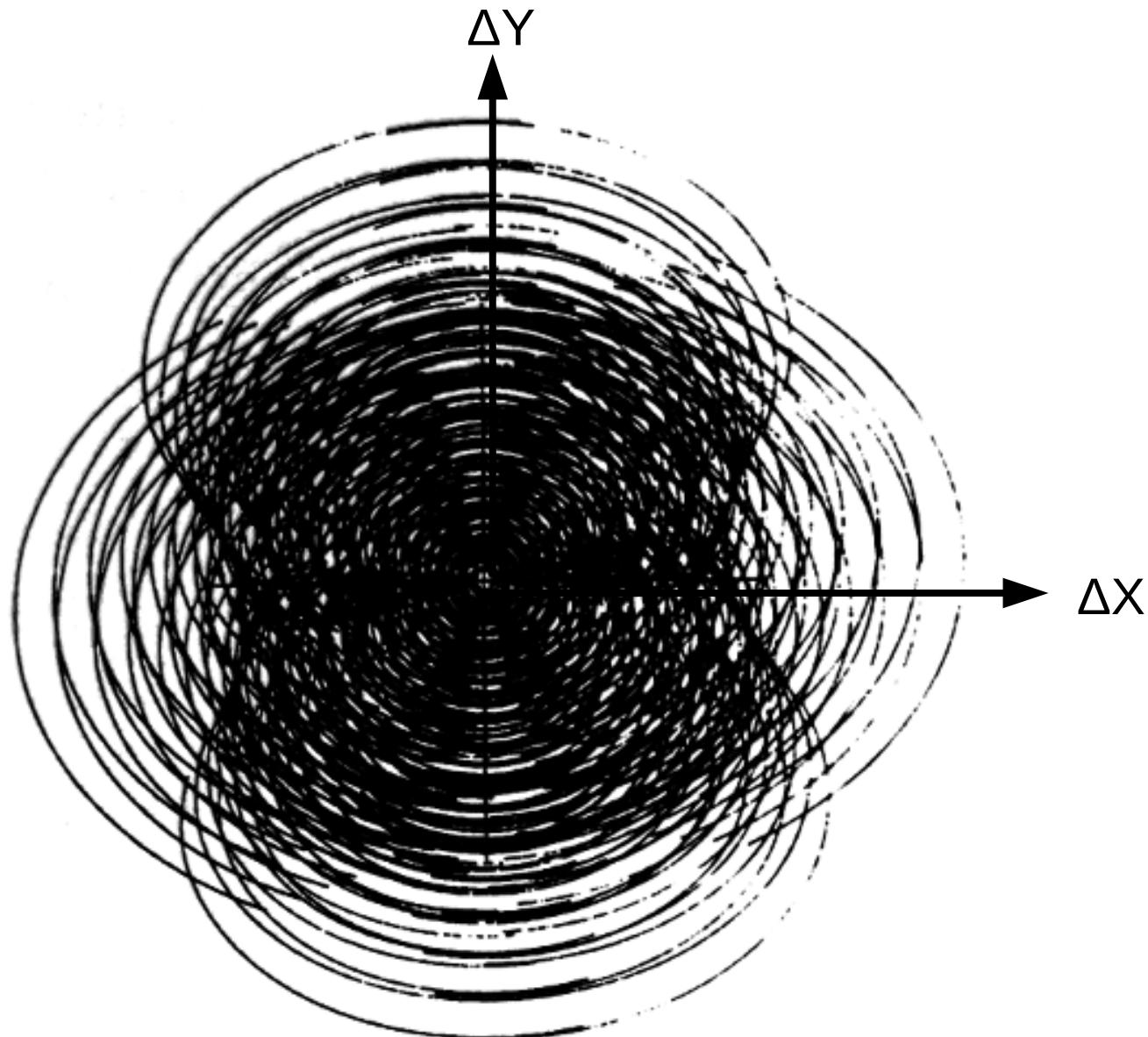


# Few Antennas, Many Baselines



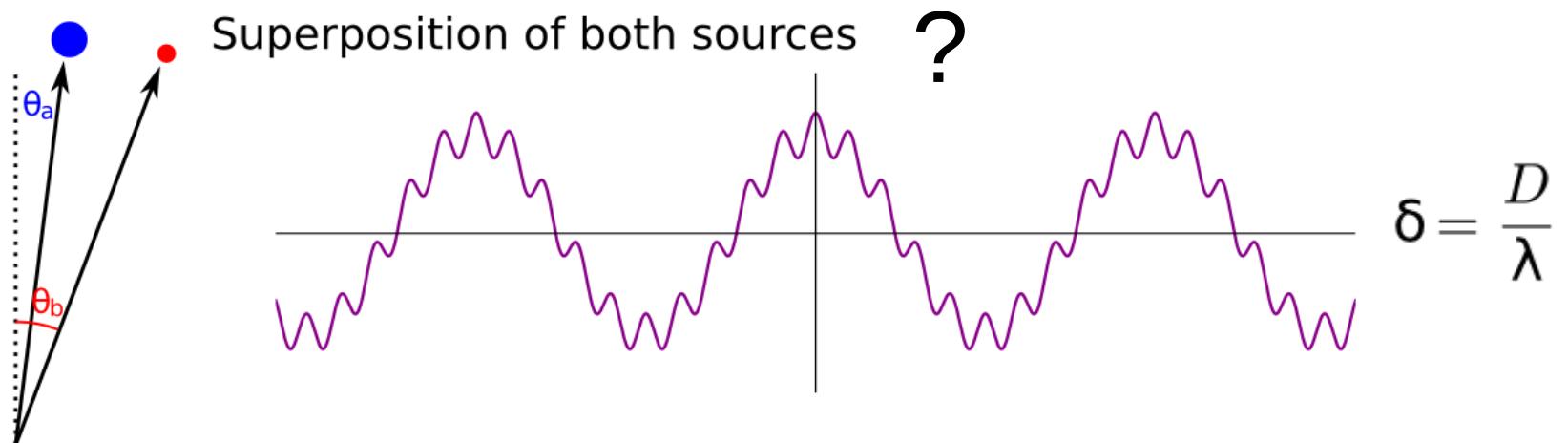
$$\frac{n(n - 1)}{2} \sim \mathcal{O}(n^2)$$

# Earth Rotation

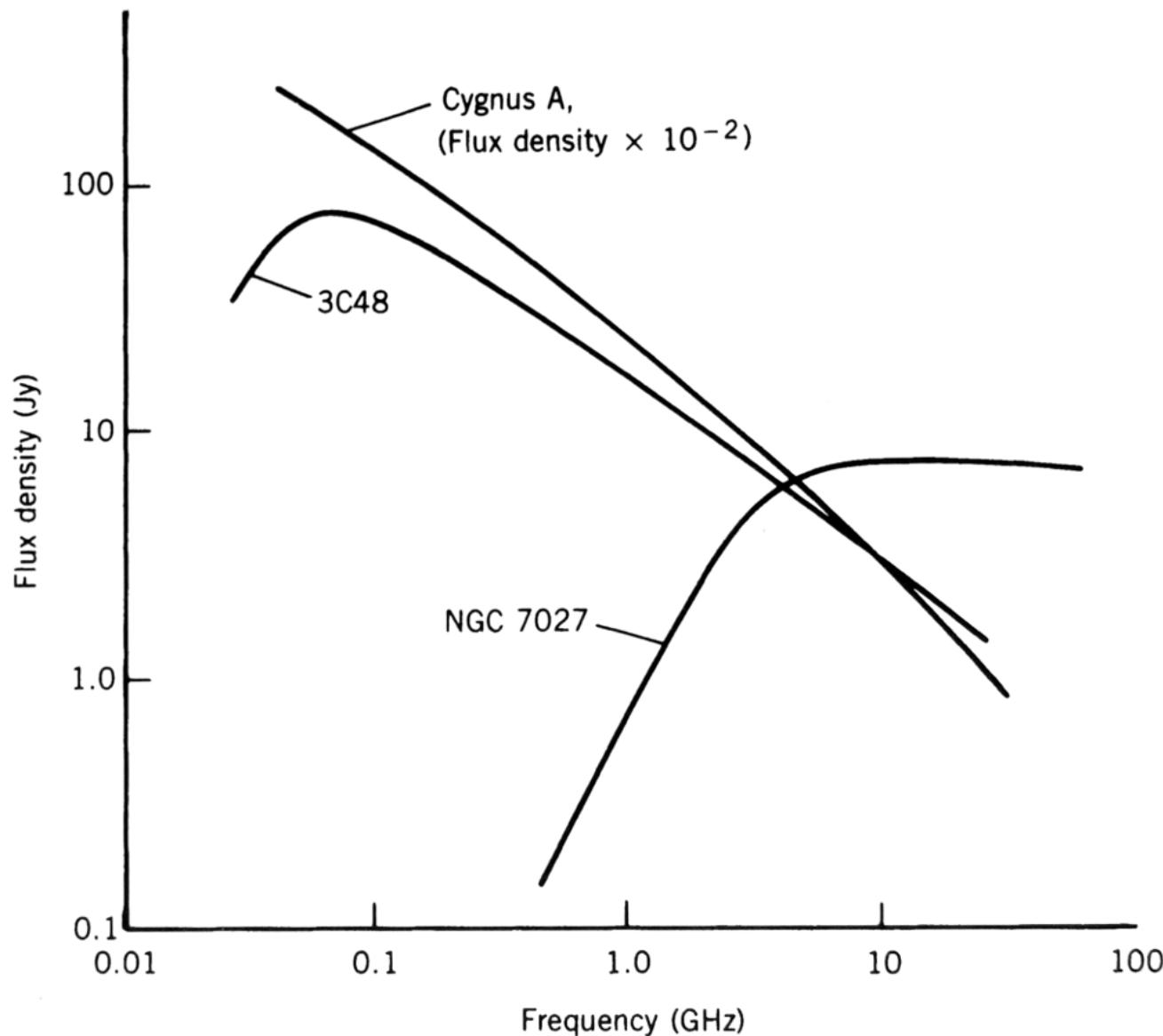


# \*Superposition

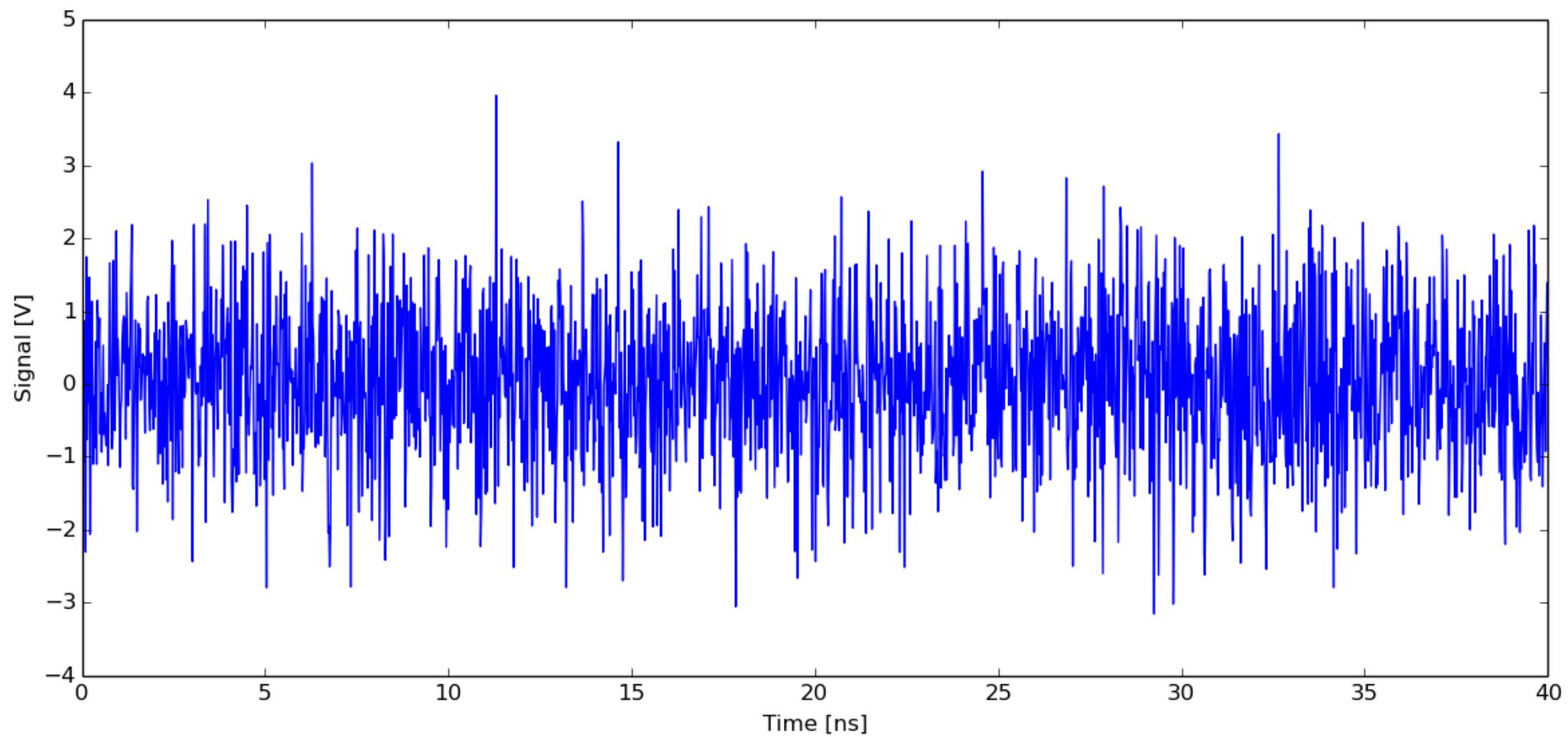
- E-fields and voltages add.
  - But not the “square-and-average” correlation
  - Right? (To the board!)
- So far we considered pure plane waves:
  - sines at a definite frequency (zero bandwidth)



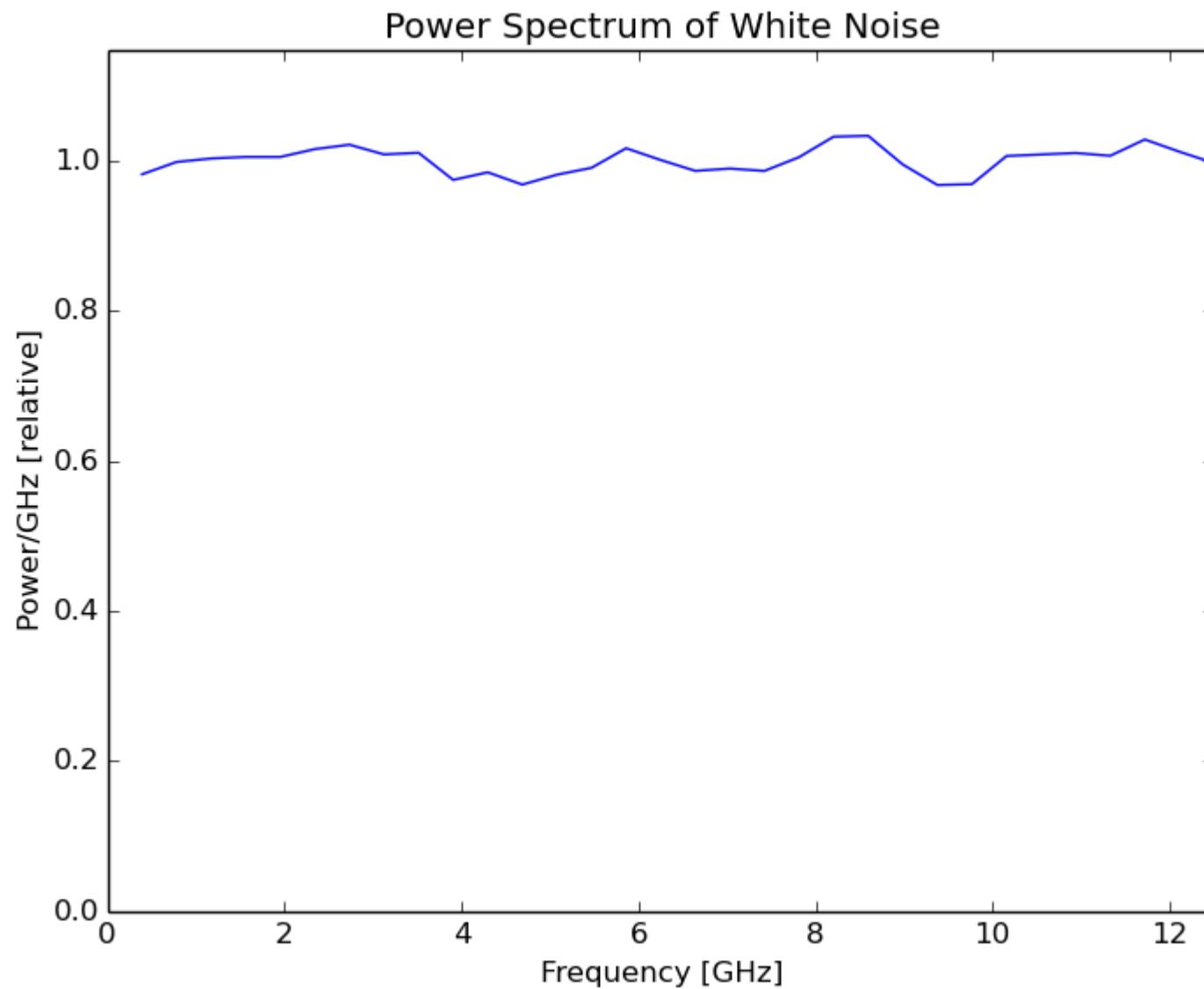
# Spectra of Sources



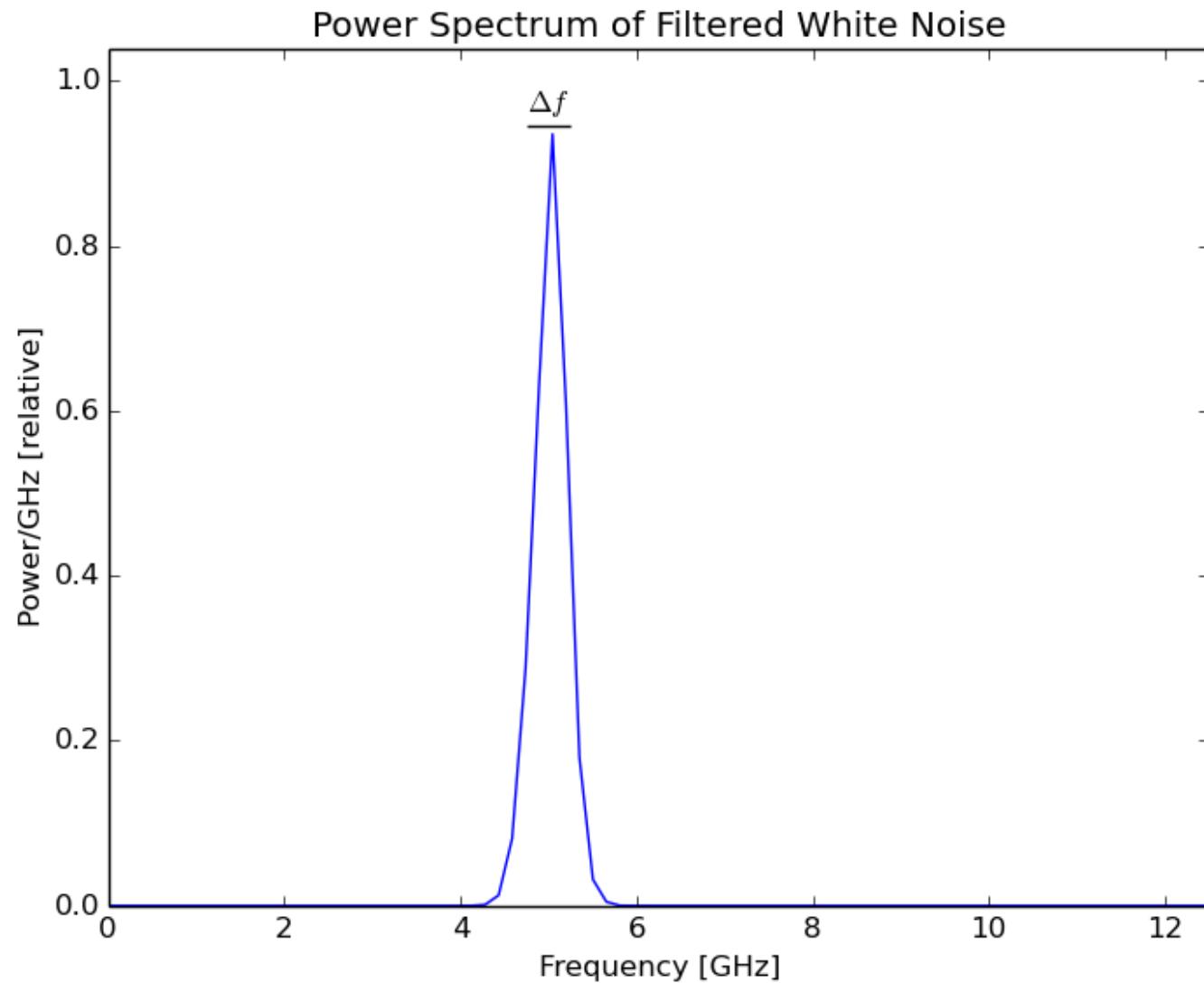
# White Noise



# Flat Spectrum

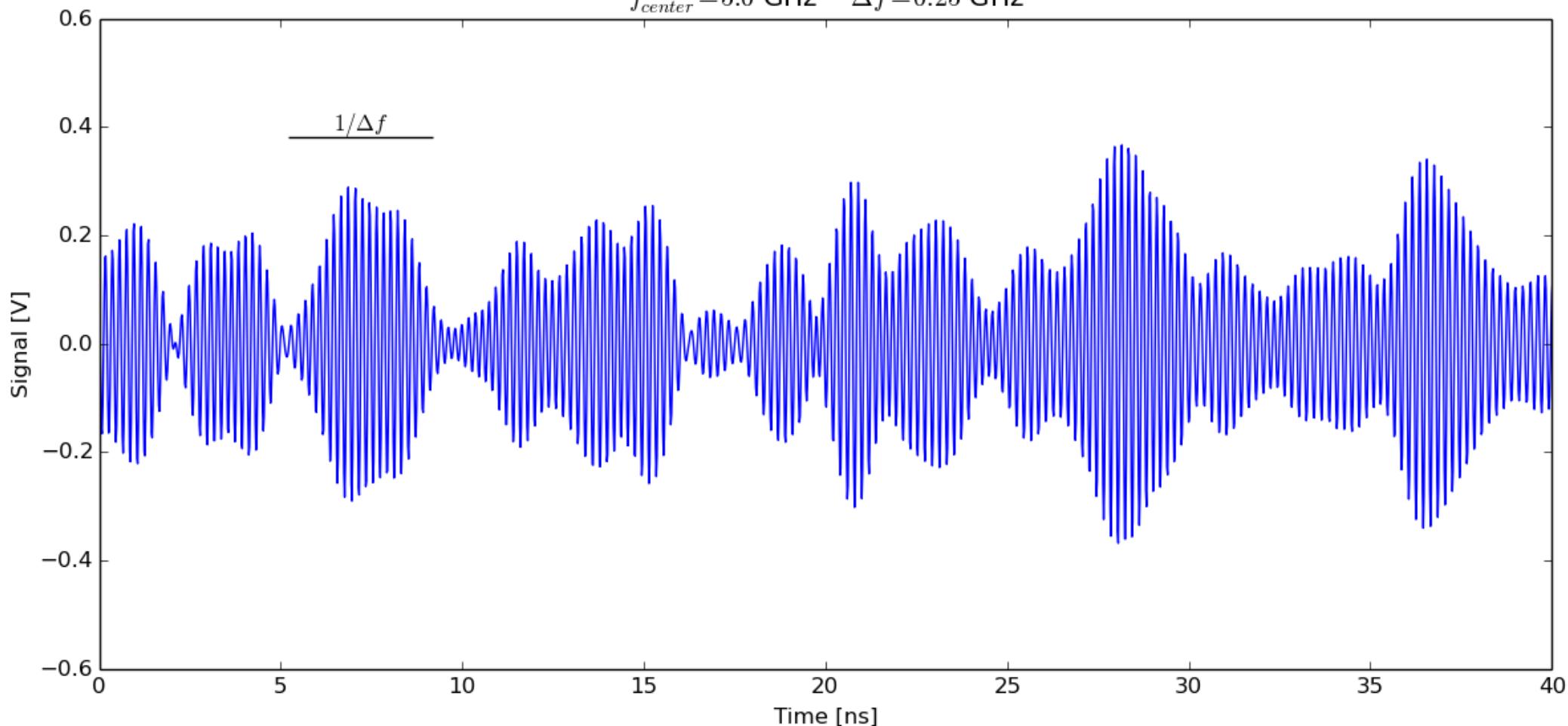


# Bandpass Spectrum (Quasi-monochromatic)



# Narrow-Band Noise (Quasi-monochromatic Light)

White noise filtered with bandpass  
 $f_{center} = 5.0 \text{ GHz}$     $\Delta f = 0.25 \text{ GHz}$

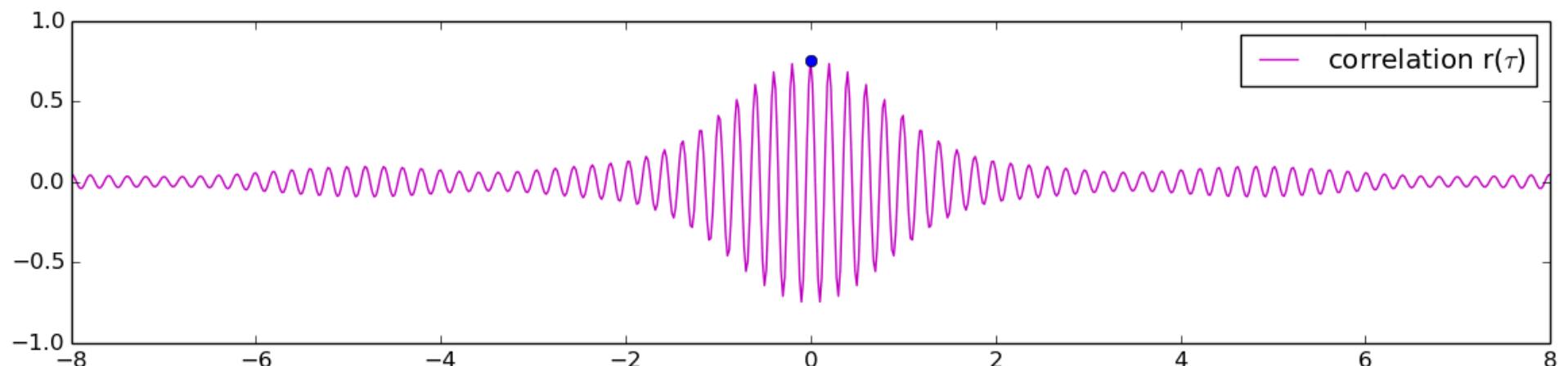
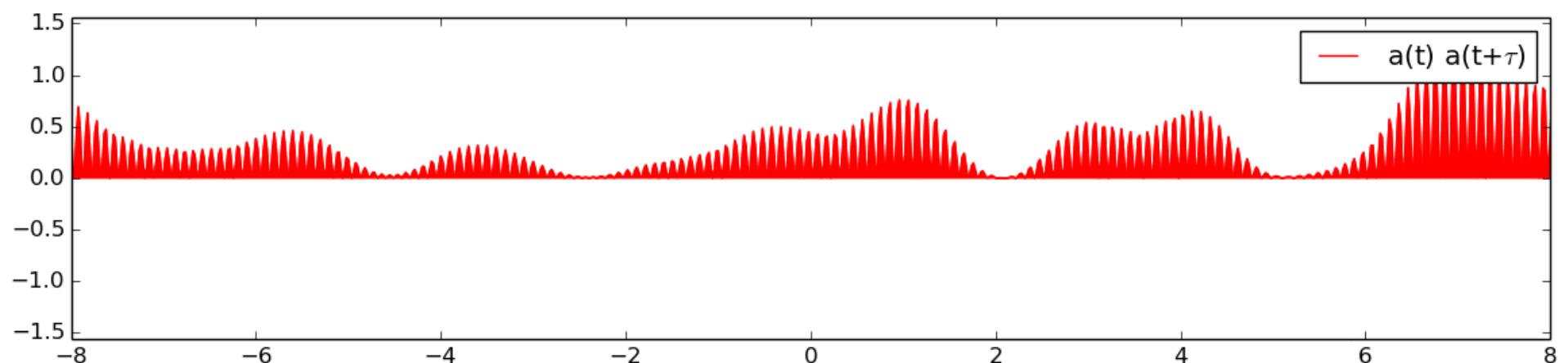
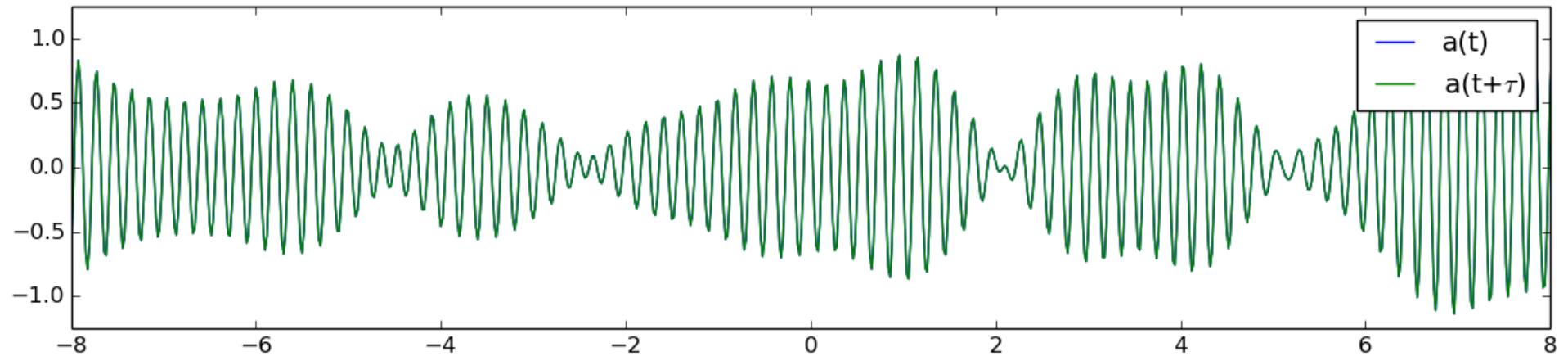


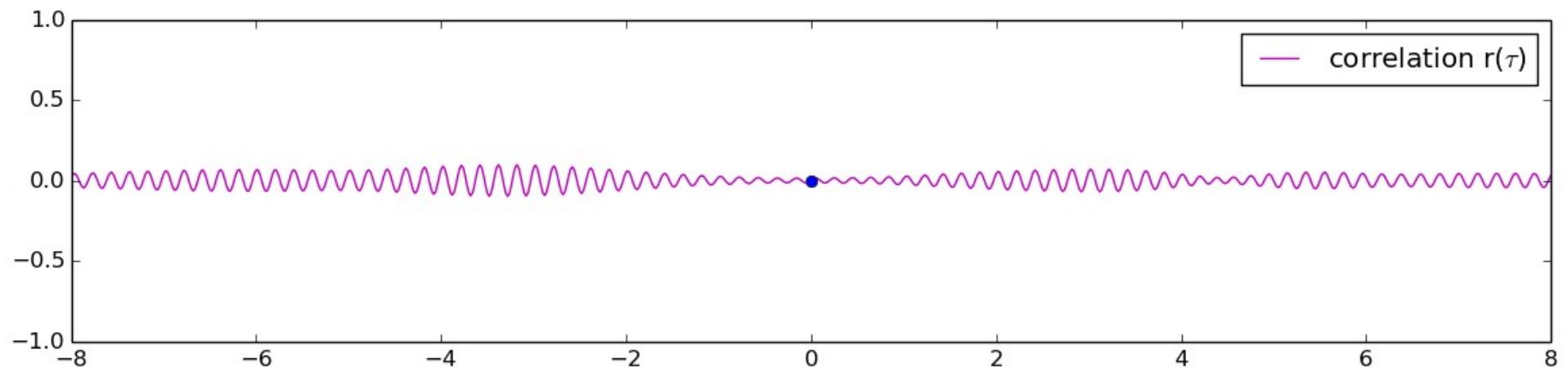
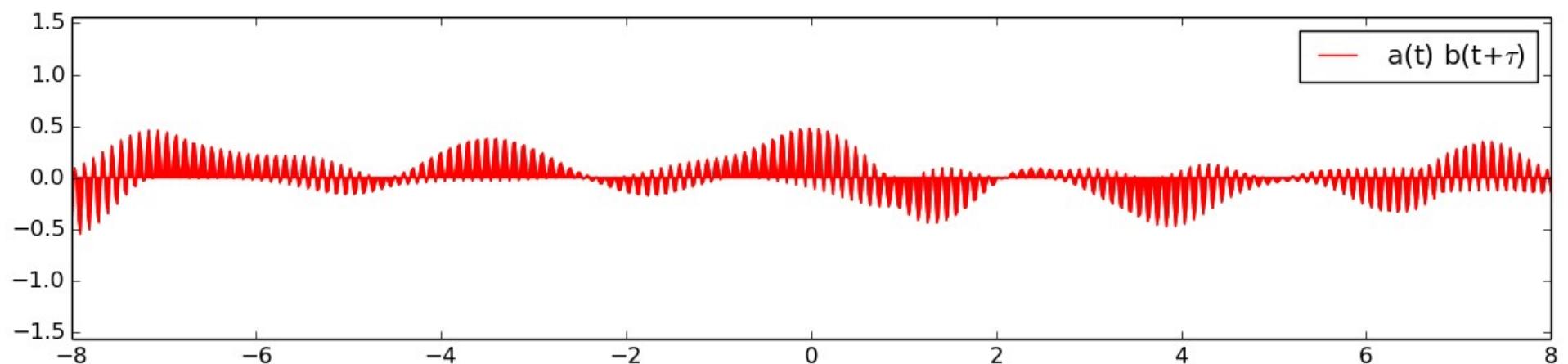
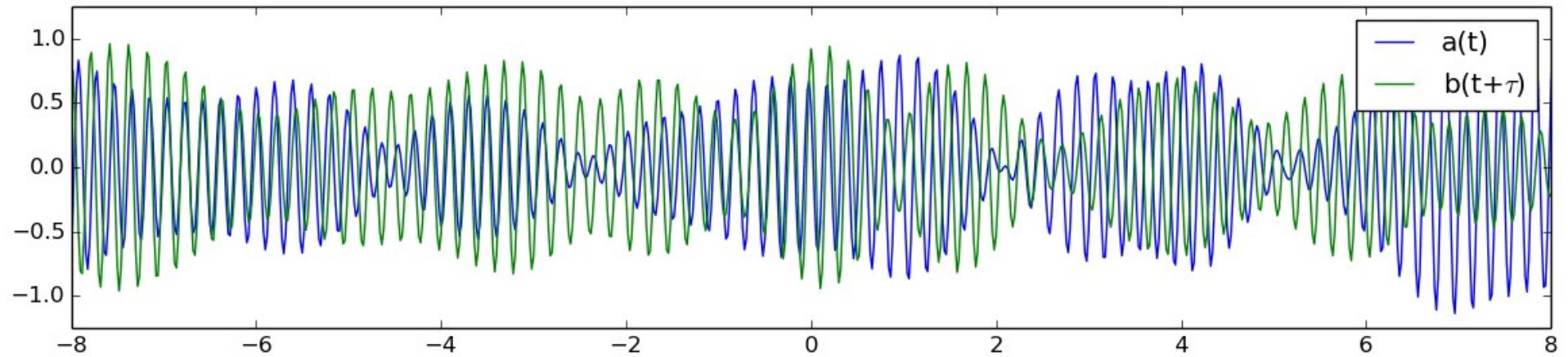
# Narrow-Band Noise

- Average period is  $\sim 1/f$
- Amplitude constant for  $t \sim 1/\text{bandwidth}$
- Similar for phase.

# Narrow-Band Correlations

- Auto-correlation
- Cross-correlation

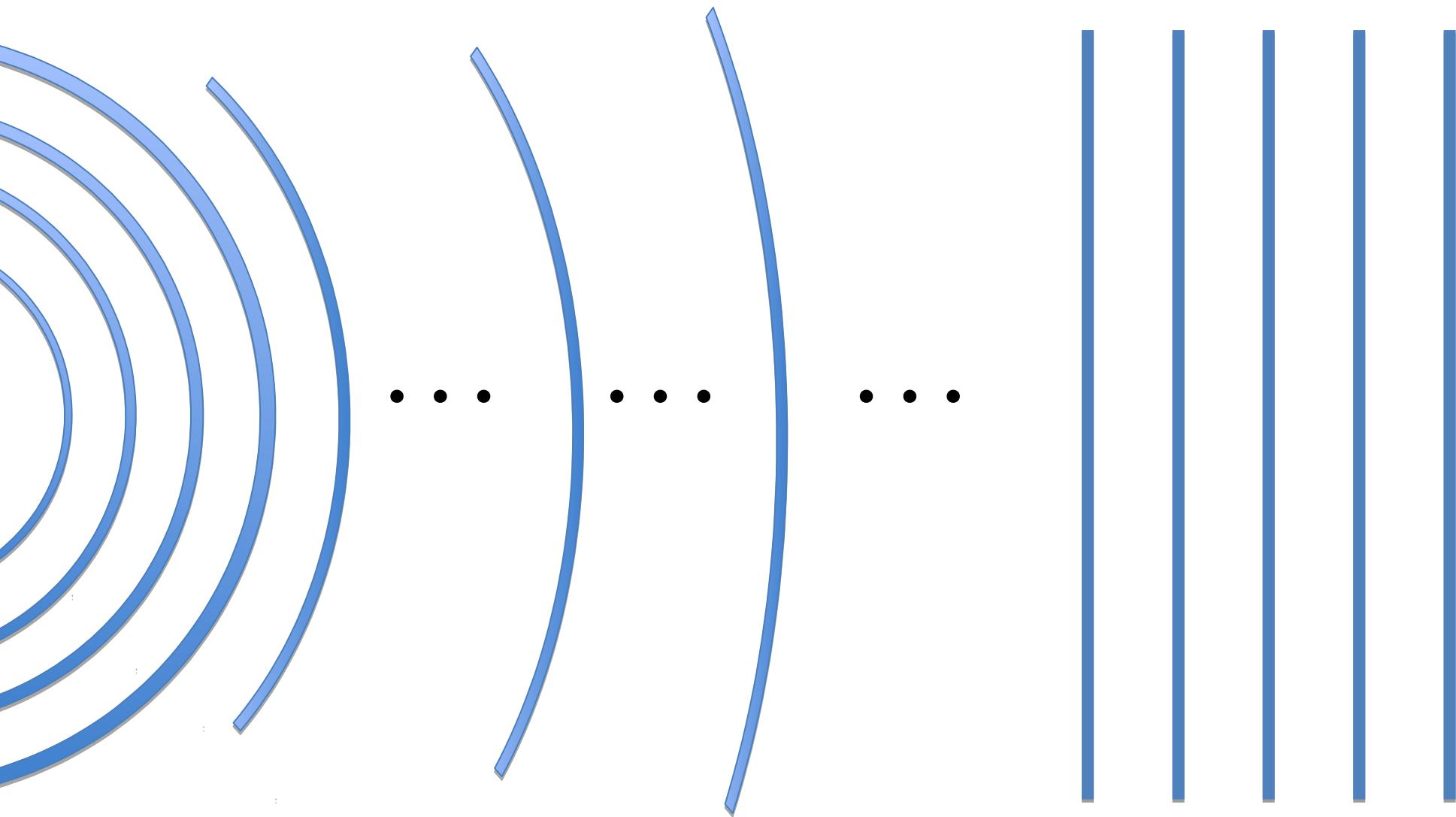




# Review

- Coherence
  - *Spatially incoherent* sources make coherent fields far away.

# EM Waves

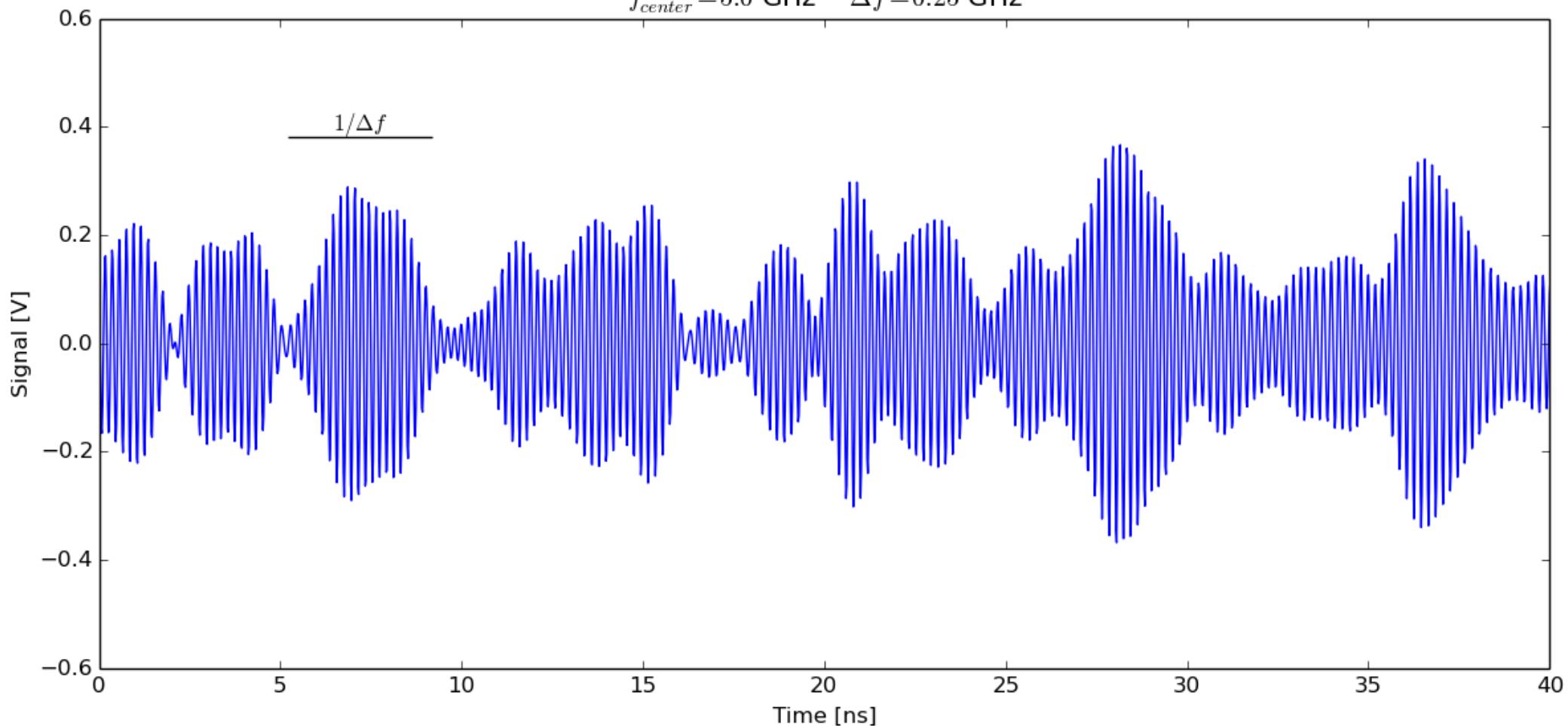


# Review

- Coherence
  - *Spatially incoherent* sources make coherent fields far away
  - *Temporally incoherent* sources become coherent through filtering (even photons!)

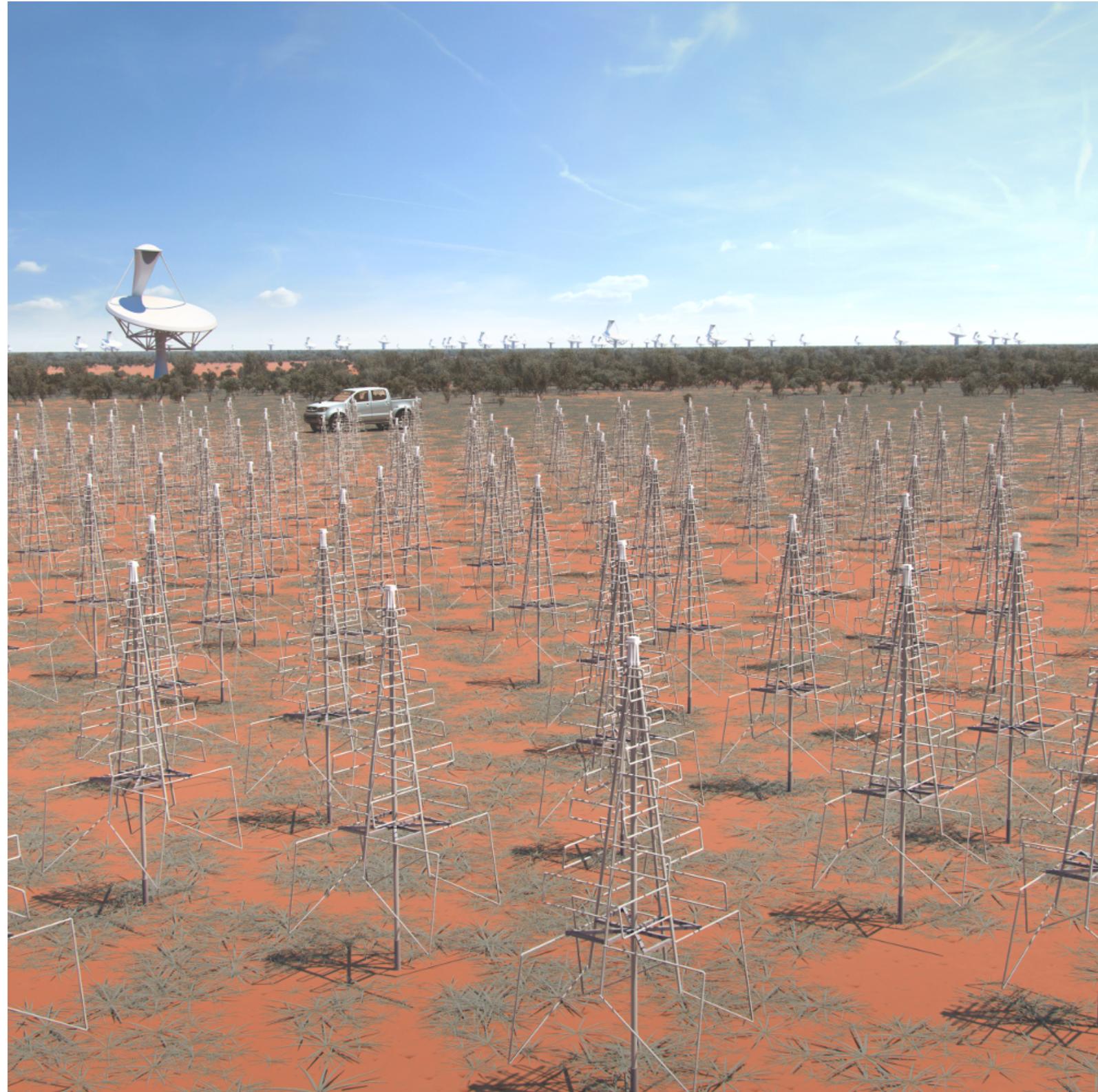
# Temporal Coherence

White noise filtered with bandpass  
 $f_{center} = 5.0 \text{ GHz}$     $\Delta f = 0.25 \text{ GHz}$

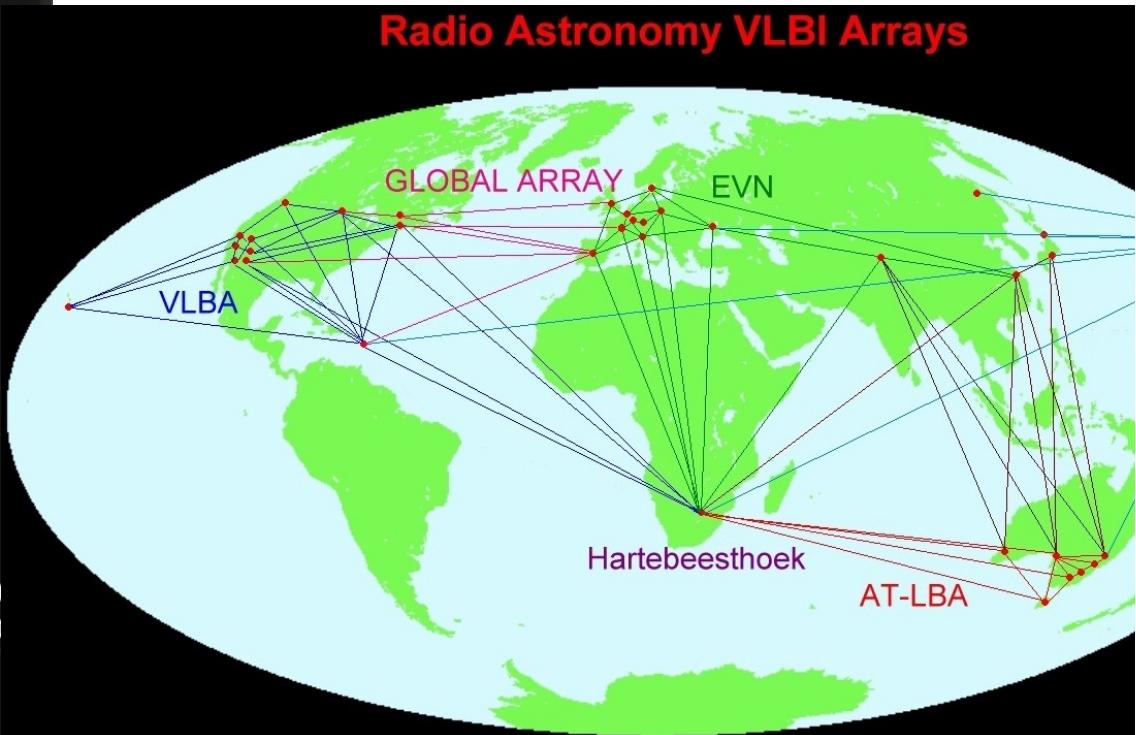
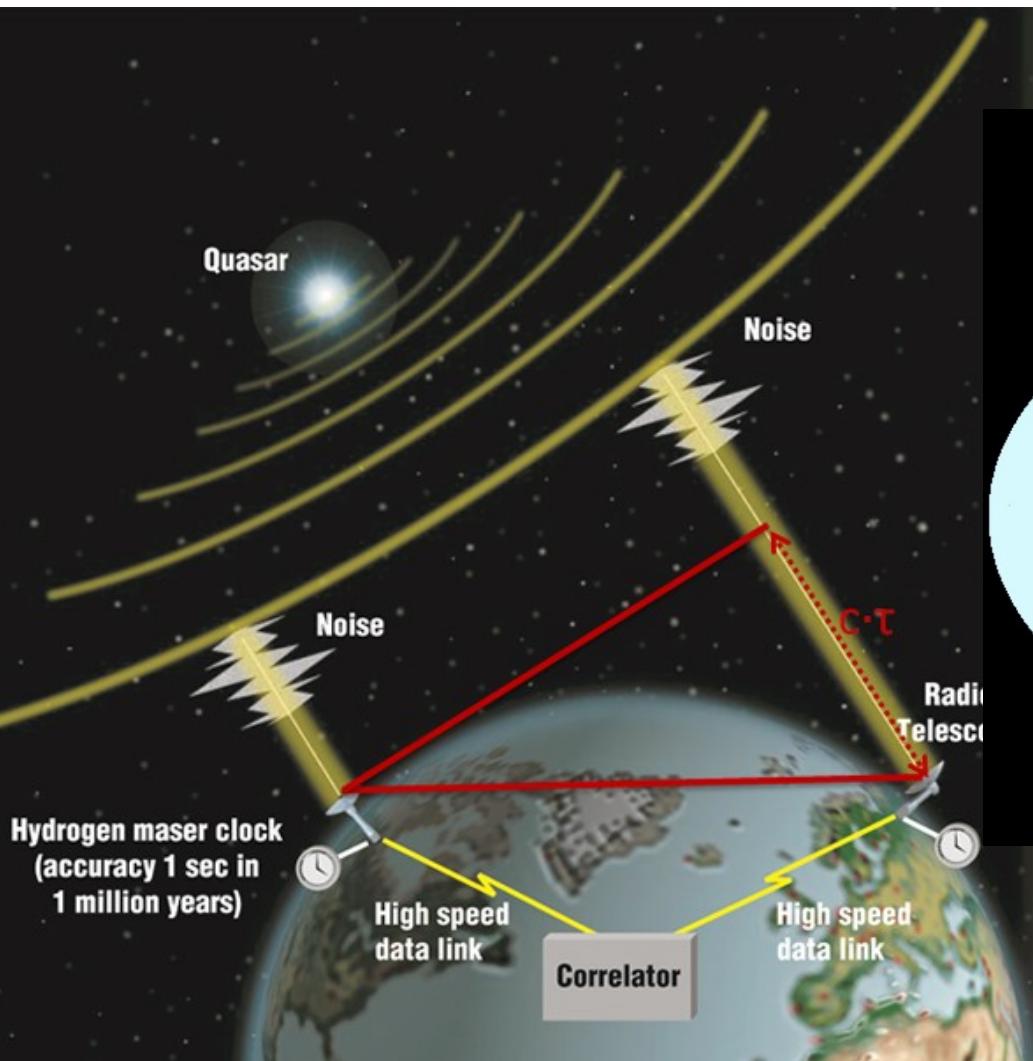


# Review

- Coherence
  - *Spatially incoherent* sources make coherent fields far away
  - *Temporally incoherent* sources become coherent through filtering (even photons)
- Interferometry is a novel and efficient way to make images:
  - No lens or mirror necessary!
  - Measure enough pairwise correlations, Fourier Transform
  - “Van Cittert-Zernike theorem”



# Very Long Baseline Interferometry



$$\begin{aligned} D &= 10,000 \text{ km} \\ \lambda &= 1 \text{ mm} \\ \theta_r &= 20 \text{ } \mu\text{arcsec} \end{aligned}$$

# Review

- Coherence
  - *Spatially incoherent* sources make coherent fields far away
  - *Temporally incoherent* sources become coherent through filtering (even photons)
- Interferometry is a novel and efficient way to make images:
  - No lens or mirror necessary!
  - Measure enough pairwise correlations, Fourier Transform
  - “Van Cittert-Zernike theorem”
- A *limitation* became key to new technique.

# Chem Nobel: Super-Resolved Fluorescence Microscopy

