



3ML

The Multi-Mission Maximum Likelihood framework

Multi-Mission Maximum Likelihood analysis with 3ML

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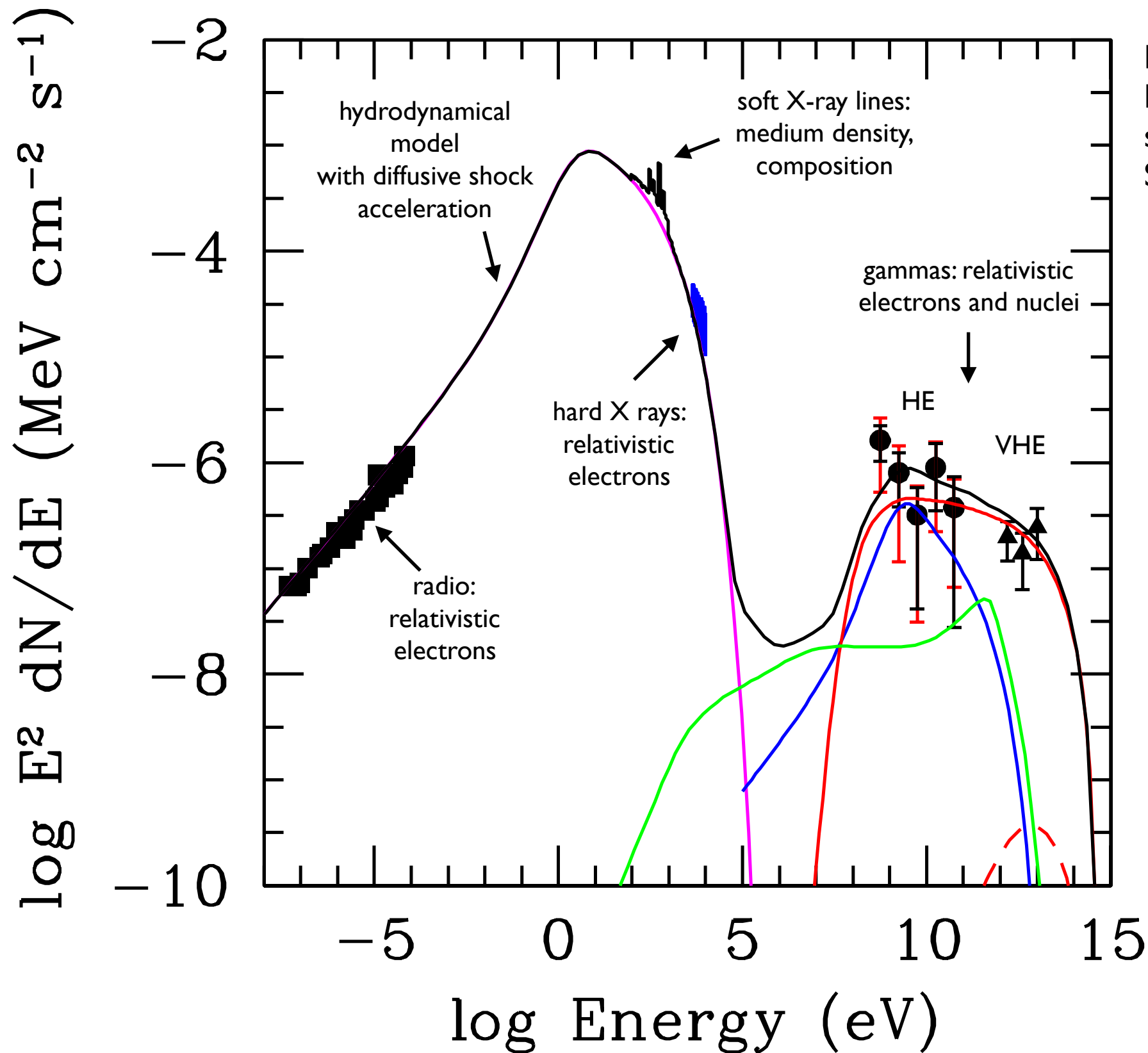
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Python for gamma-ray astronomy
MPI-K, 16 November 2015

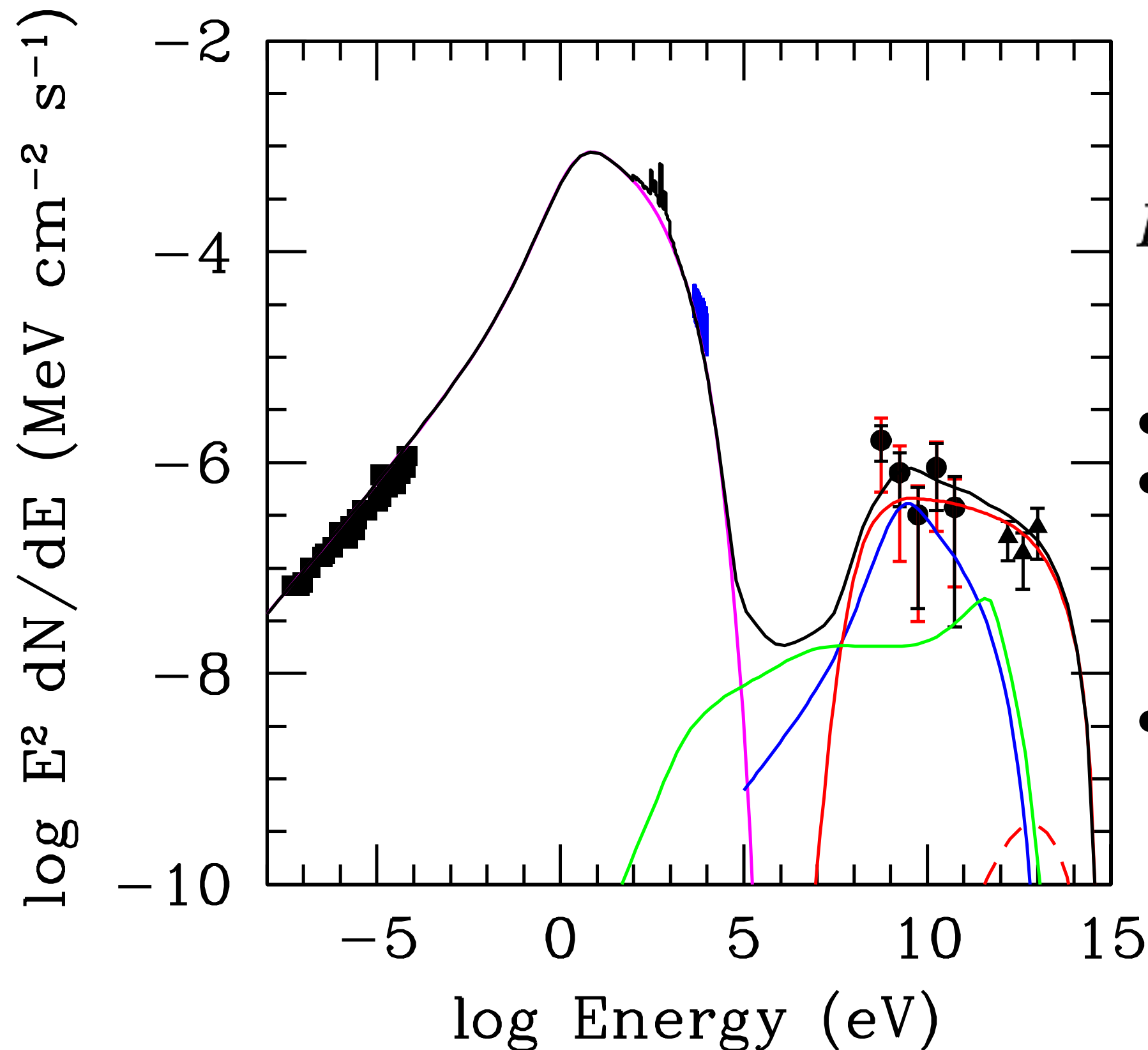
Why multi-mission studies?



Example:
MWL spectrum of Tycho
supernova remnant
Slane et al. ApJ 783, 33 (2014)

more data
↓
more constraints
on models
↓
more physics!

The SED approach



idealised case

$$F(E) = E \cdot \frac{N}{\Delta t A}$$

- narrow energy bins
- counts integrated in PSF-like region, background subtracted
- measured energy \sim real energy

Deriving a SED: the real deal

- Complications:
 - limited counting statistics require large-ish bins
 - energy variable PSF and backgrounds
 - energy dispersion

$$N = \int dt dE d\Omega \left[\overset{\text{flux normalisation}}{\downarrow} k \cdot \overset{\text{spectral model}}{\downarrow} S(\vec{p}', E' | \vec{\alpha}) \otimes \overset{\text{instrument response}}{\downarrow} \text{IRF}(\vec{p}', E' | \vec{p}, E; t) + \overset{\text{background model}}{\downarrow} \text{BG} \right]$$

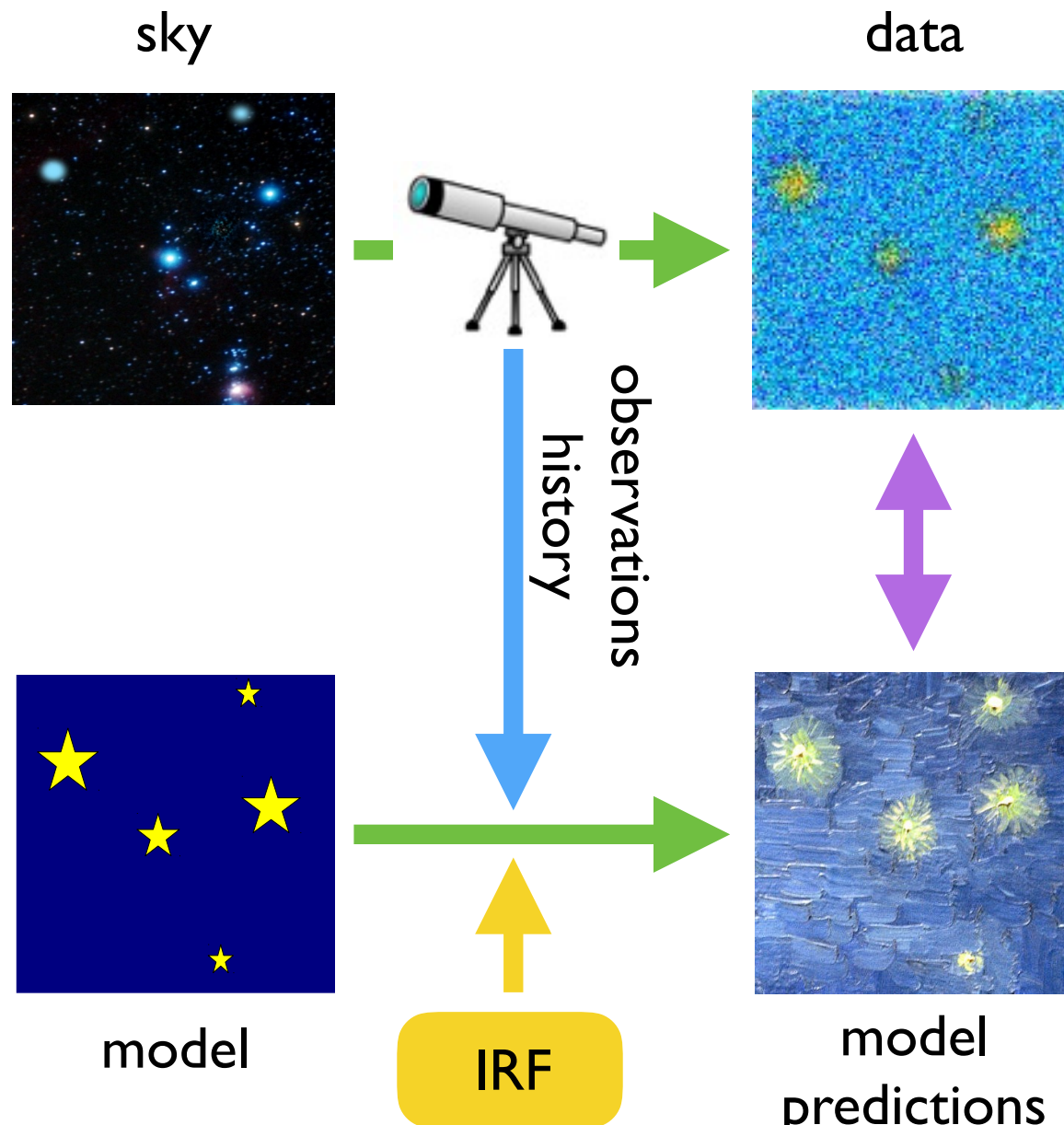
$$\overset{\text{from fit in specific energy bin}}{\uparrow} F(\langle E \rangle_{12}) = \tilde{k}_{12} \int_{E1}^{E2} \int_{\Delta\Omega} dE d\Omega \quad \overset{\text{from broadband fit}}{\uparrow} E S(\vec{p}, E | \vec{\alpha})$$

Pitfalls of the SED approach

$$F(\langle E \rangle_{12}) = \tilde{k}_{12} \int_{E_1}^{E_2} \int_{\Delta\Omega} dE d\Omega \quad ES(\vec{p}, E | \vec{\alpha})$$

- depends on the spectral model chosen and on datasets used to determine the spectral parameters
- consistency of spectral models/parameters in combining different instruments and comparing to model
- impossible to take into account source morphology: spatial information not used/SED points from inconsistent regions
- difficult to incorporate other messengers (neutrinos)

The forward-folding approach



Likelihood:

- estimate model parameters
- compare different models

$$\mathcal{L}(O|M)$$

multiple independent
observations (instruments)

$$\mathcal{L}(\vec{O}|M) = \mathcal{L}(O_1|M) \times \mathcal{L}(O_2|M) \times \dots$$

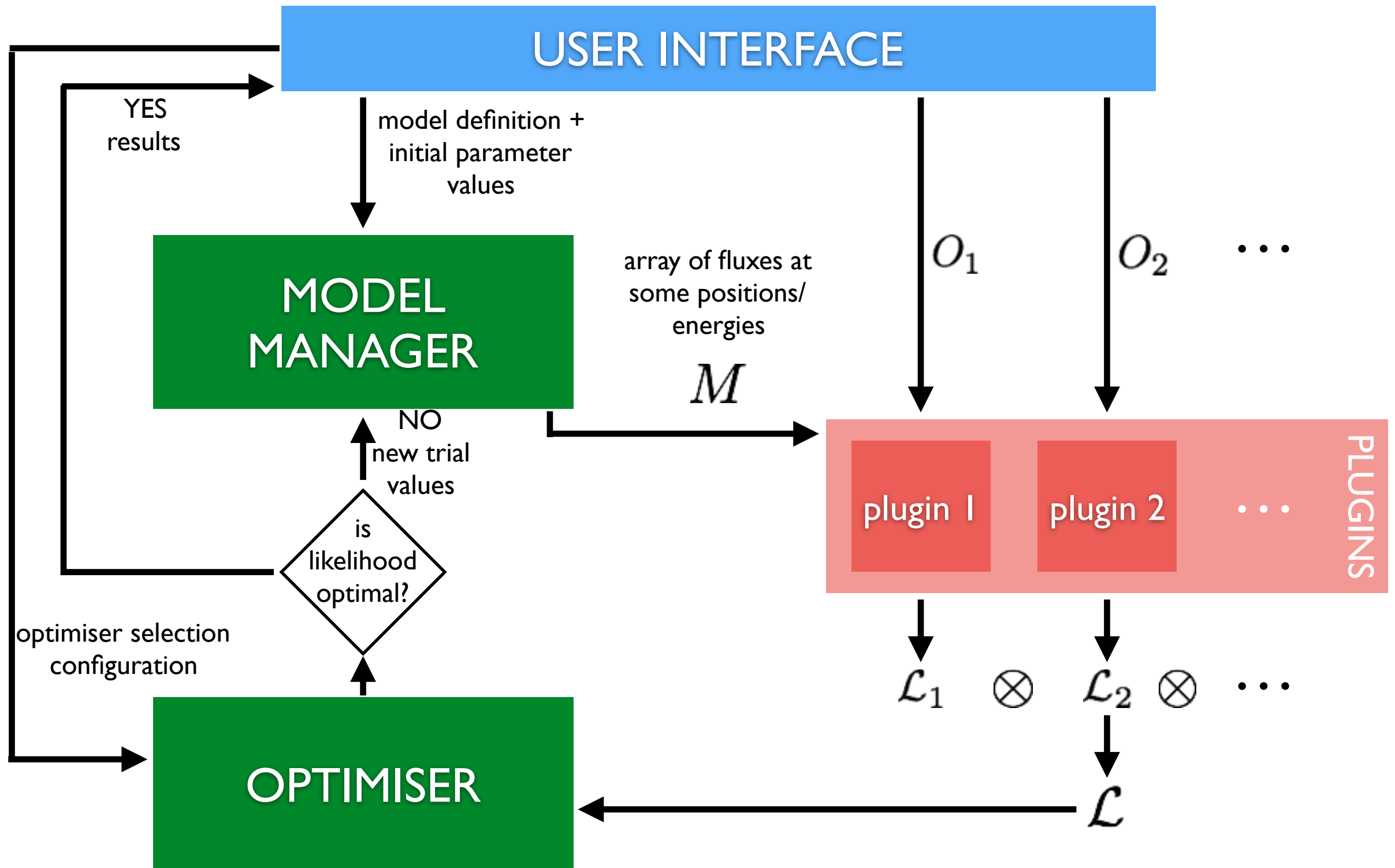


The 3ML project

- lead developer: Giacomo Vianello (Stanford U.); co-developers: J.M. Burgess, R. Lauer, N. Omodei, L. Tibaldo, P. Yunk
- python project
 - user friendly
 - relies on open packages (scipy, astropy)
 - very limited C++ code (under the hood)
- website: <https://threeml.stanford.edu/>
- git repository: <https://github.com/giacomov/3ML>
- ICRC 2015 proceedings paper: <http://arxiv.org/abs/1507.08343>



3ML architecture

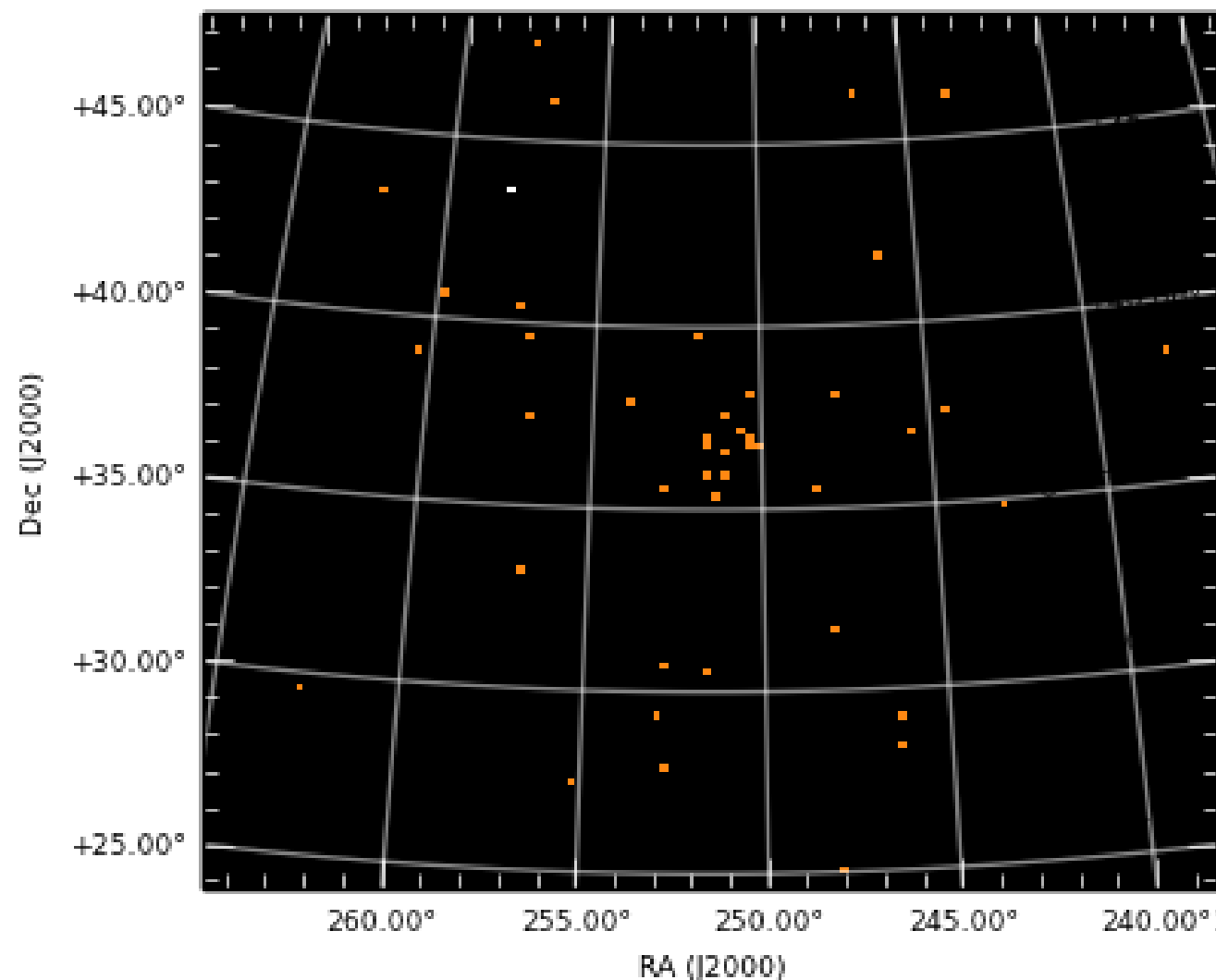


What's special about 3ML? Plugins

- instrument-specific part of the analysis handled through plugins using official/existing software (or at least preserve the existing/official data and instrument response format)
 - no need to re-develop software specific to each instrument
 - no need to export data to a different format
 - flexible: extendable to any instrument (multi-messenger)
- plugins under development
 - working: *Fermi* LAT, *Fermi* GBM, OGIP (*Swift* XRT), HAWC (data/software not public)
 - planned/possible: VERITAS, HESS, GammaLib, anything you want to implement!

Get the most out of each dataset

- software like XSPEC or ISIS retains only spectral information (OGIP format)

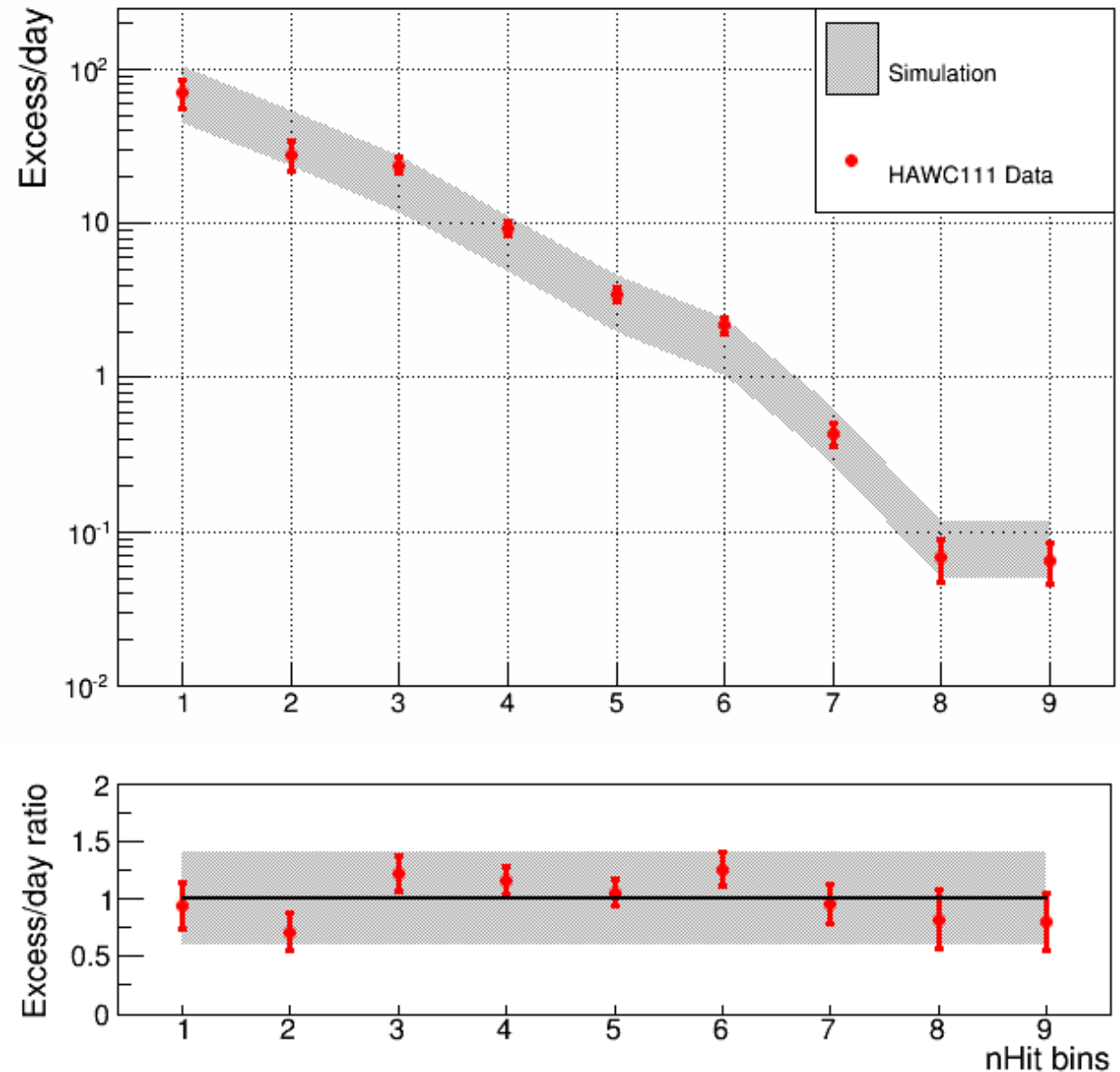


simulation of faint GRB with *Fermi* LAT

- 45 background + 15 signal
- Poisson probability of 60 photons for 45 background expected: 2%, significance $\sim 2\sigma$
- significance from *Fermi* LAT Science Tools (considering spatial information): 7σ

Flexibility

- HAWC does not have (yet) a good energy estimator
 - nHit (integrated number of p.e. in a shower)
 - not a good energy estimator (strong dependence on core location, zenith angle)
 - MC can reliably predict nHit from energy
 - 3ML makes possible analyses in nHit space (rather than energy)



nHit “spectrum” of the Crab nebula seen by HAWC
A.J. Smith for the HAWC collaboration, ICRC 2015

Models

- Spectral
 - for the moment several analytical functions and tabulated
 - anything that you can program in python (or fetch from python)
- Spatial
 - several analytical functions
 - FITS maps in WCS coordinates and HEALPix coming soon

Optimisers

- classical likelihood optimisation with Minuit (iMinuit)
- under development: Bayesian sampling of likelihood profile with emcee
- possible to implement other optimisation strategies

Final remarks

- 3ML provides a framework for multi-mission analysis
 - based on forward folding approach: statistically/
methodologically robust
 - based on **plugins** that exploit existing software for each
instrument and offer maximum flexibility
 - python based: user friendly and can rely on many
community-driven packages
- young project: lots of development required
 - feedback and contributions are welcome
- tutorial will be given on Wed at 11 (unusual room:
Multimedia, first floor)