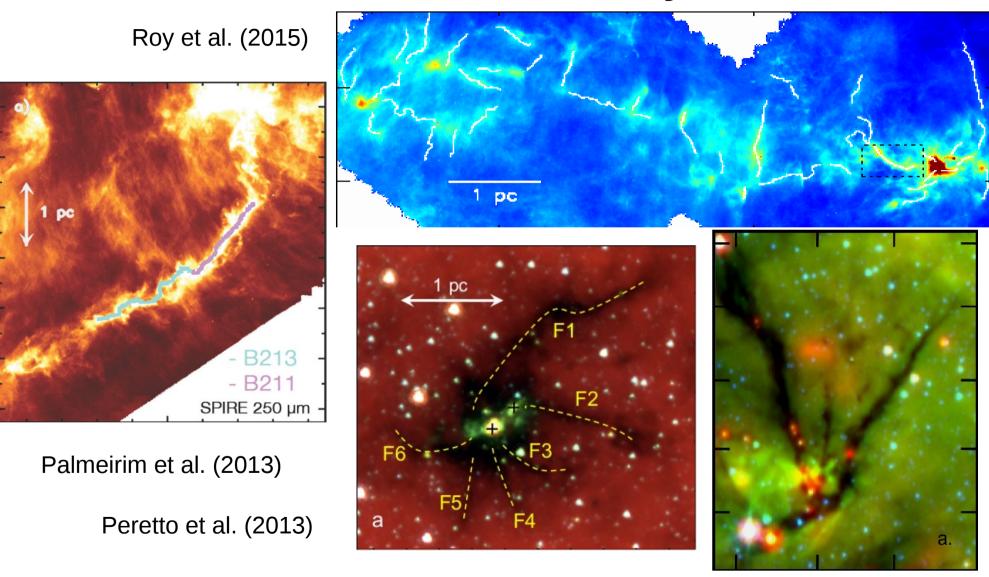
Investigating the global collapse of filaments using SPH

S. D. Clarke & A. P. Whitworth



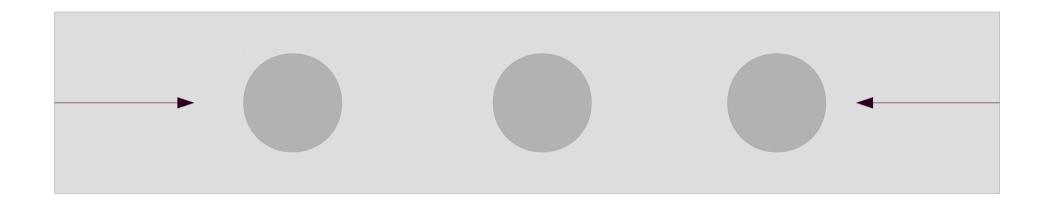
Filaments are everywhere



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Peretto & Fuller (2010)

Global collapse and fragmentation



Bastien (1983) Burkert & Hartmann (2004) $t_{1D} > t_{3D} = \left(\frac{3\pi}{32 G \Omega}\right)^{1/2}$ Pon et al. (2012) Bastien (1983) Pon et al. (2012)

$$t_{1D} > t_{3D} = \left(\frac{3\pi}{32G\rho}\right)^{1/2}$$

Nagasawa (1987) Inutsuka & Miyama (1992) Pon et al. (2011)

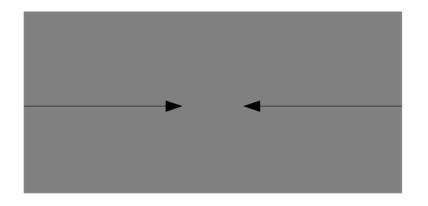
Homologous collapse



Homologous collapse



Homologous collapse



$$t_{1D} \approx \frac{0.44 \, A}{\sqrt{G \, \rho}}$$

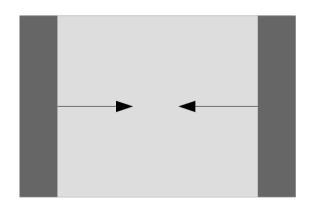
End-dominated collapse



End-dominated collapse



End-dominated collapse



$$t_{1D} \approx 0.98 \sqrt{\frac{A}{G\rho}}$$

Numerical set-up

Uniform density
Cylindrically symmetric
Stationary
Isothermal

- SPH code GANDALF
- Grad-h SPH
- Gravity and hydrodynamics
- T = 1 K as we wish to approximate free-fall
- Radial motion suppressed, $v_x = v_y = 0$
- 100,000 200,000 particles used

Numerical set-up

Aspect ratio, A Line-mass, µ Radius, R

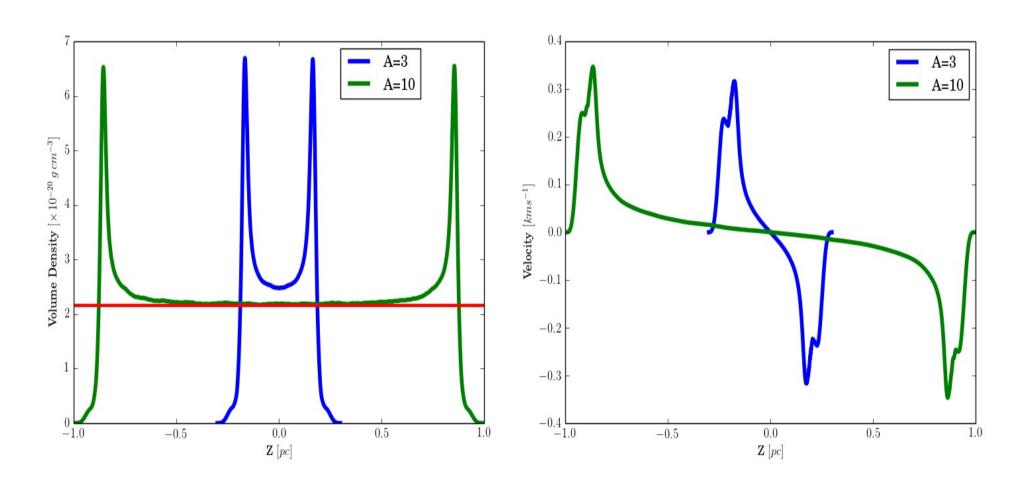
$$2 \le A \le 20$$
 $2 \text{ M pc}^{-1} \le \mu \le 50 \text{ M pc}^{-1}$
 $0.05 \text{ pc} \le R \le 0.15 \text{ pc}$

Fiducial case:

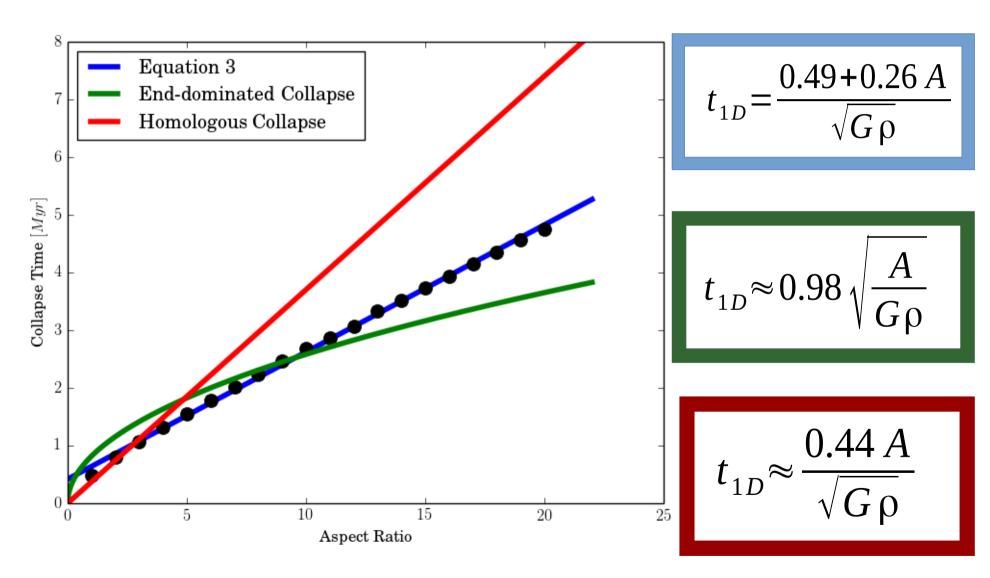
$$A = 10, \mu = 10 \text{ M pc}^{-1},$$

 $R = 0.1 \text{ pc}$

No homologous collapse

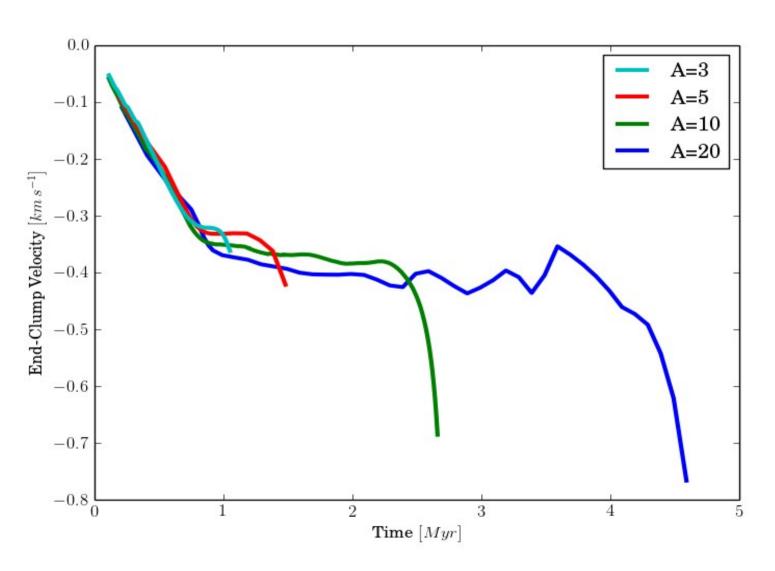


Collapse time



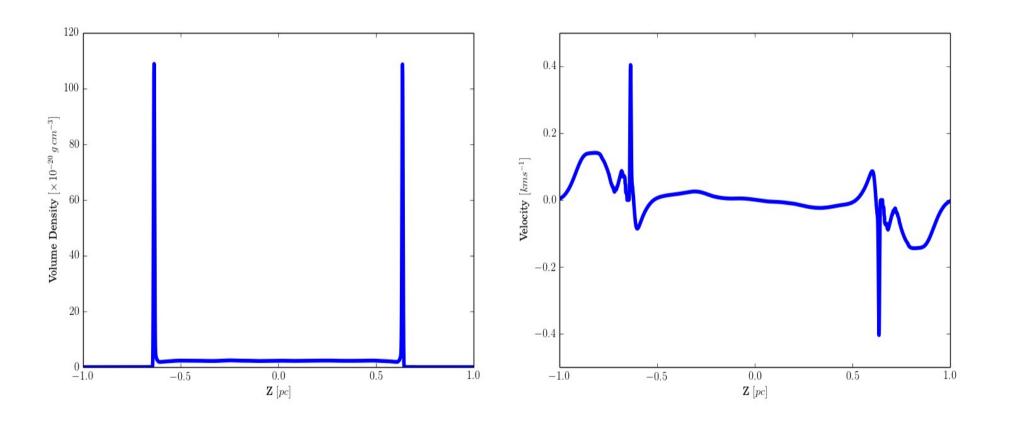
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End-clump velocity

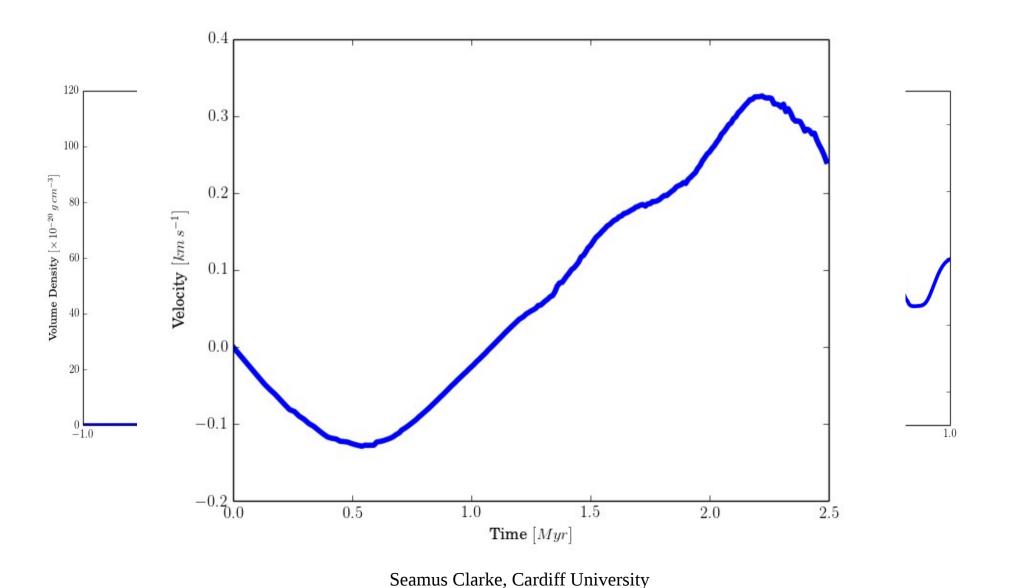


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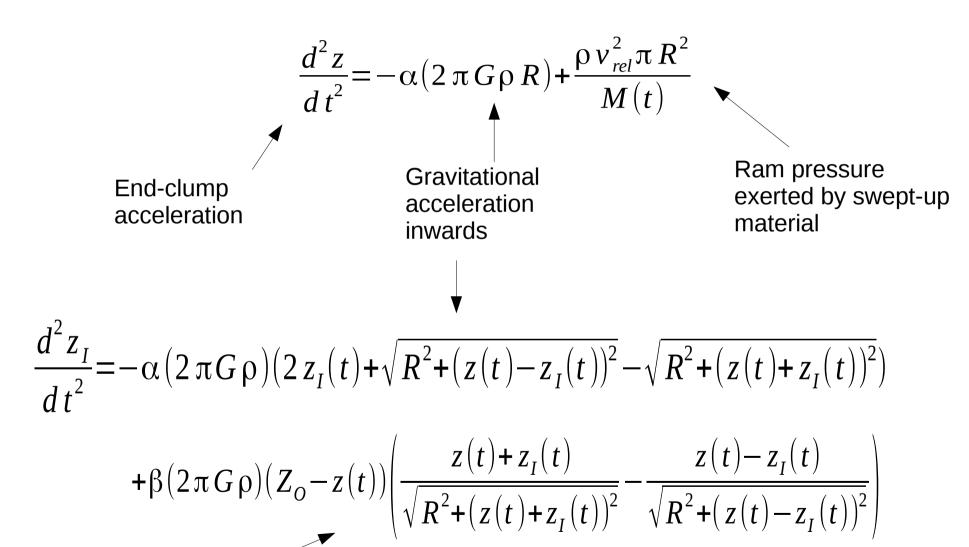
How to explain this terminal velocity



How to explain this terminal velocity

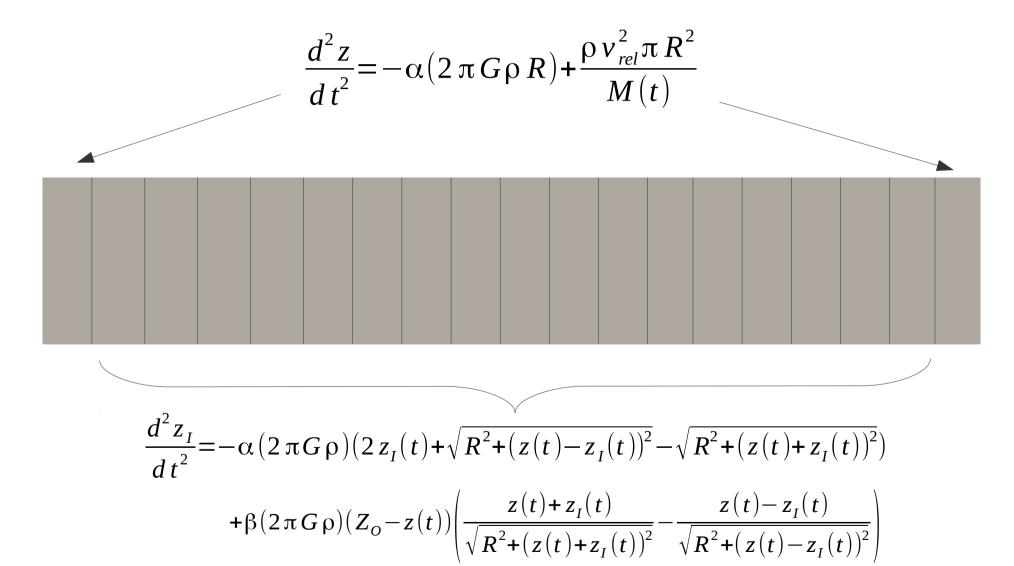


Semi-analytical model



Gravitational acceleration outwards due to the end-clump

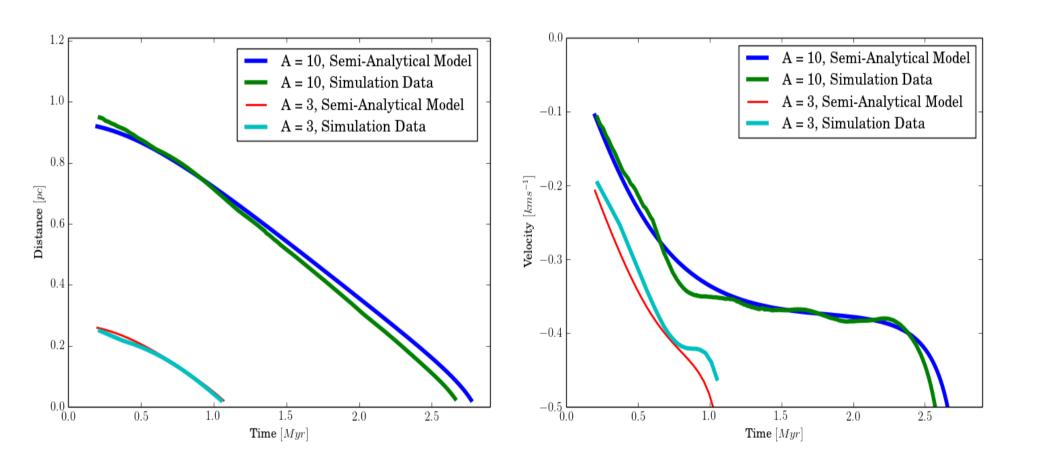
Semi-analytical model



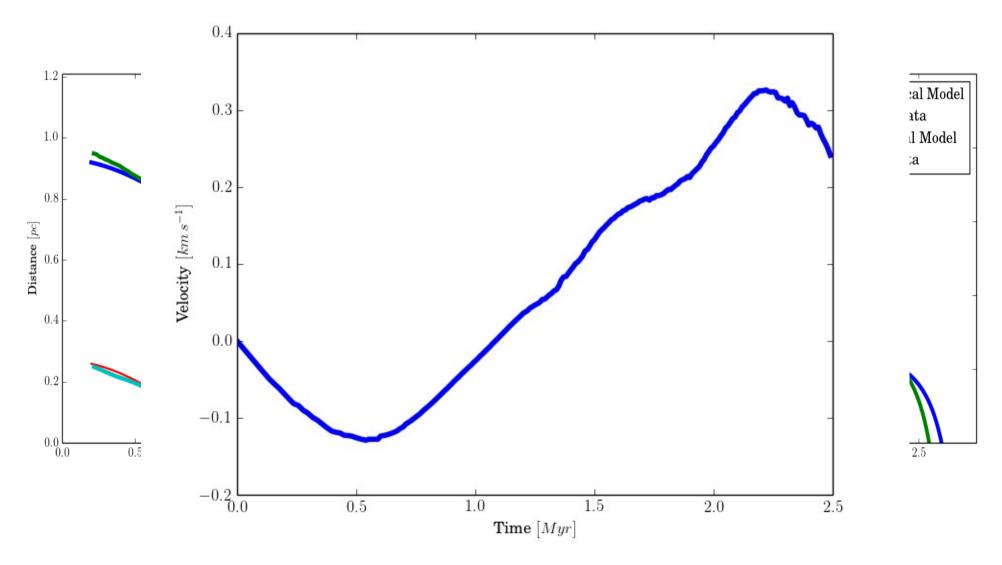
$$\sqrt{R^2 + (z(t) + z_I(t))^2} \sqrt{R^2 + (z(t) - z_I(t))^2}$$

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Testing the model

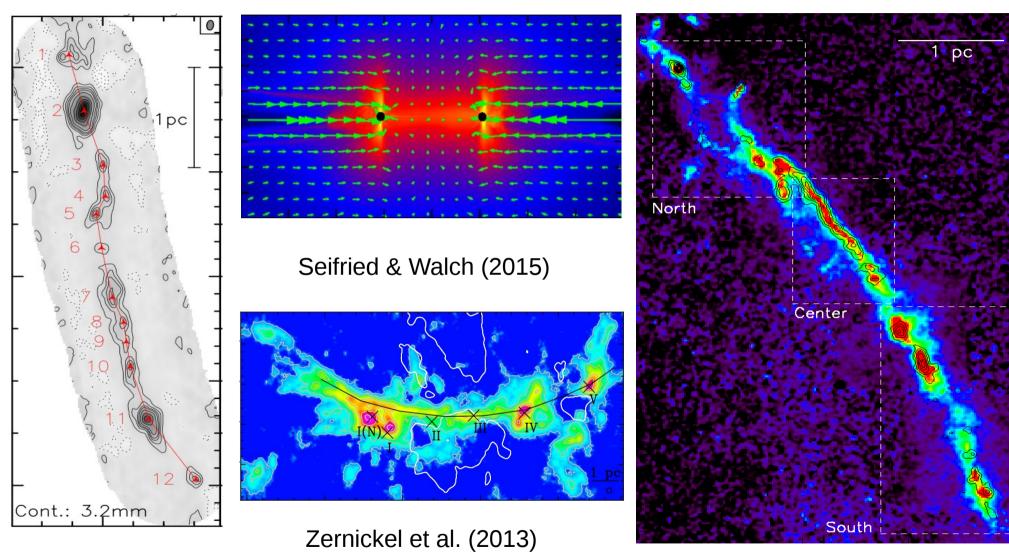


Testing the model



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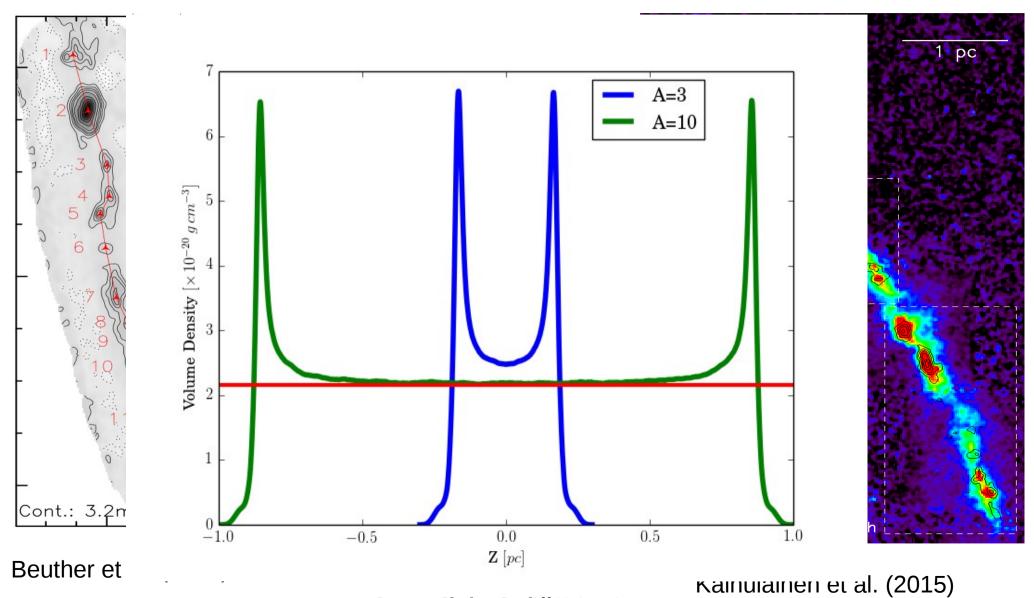
Observational evidence



Beuther et al. (2015)

Kainulainen et al. (2015)

Observational evidence



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Conclusions

- Filaments collapse and fragment differently than higher dimensional objects.
- Global collapse occurs via the end-dominated mode, which produces high density end-clumps.
- The free-fall time of filaments is given by the equation:

$$t_{1D} = \frac{0.49 + 0.26 A}{\sqrt{G \rho}}$$

- Global collapse is much longer than in 3D, this makes filaments perfect sites for fragmentation.
- Observational examples of end-clumps in isolated filaments have been found.