HW4

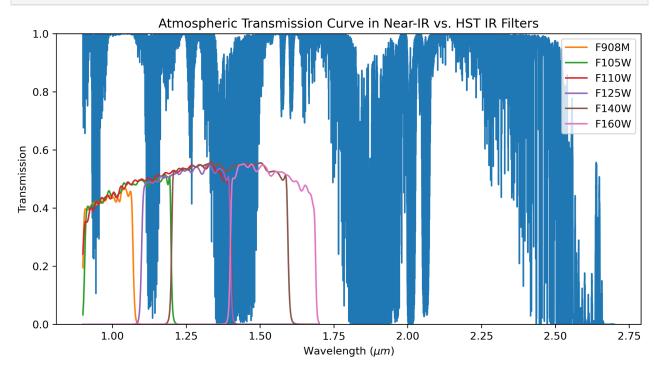
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1)

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In []: import pandas as pd
import numpy as np
import matplotlib.pyplot as plt
```

```
In [ ]: nir_transmission = pd.read_csv("mktrans_zm_10_10.dat", sep = ' ', header =
        nir transmission = nir transmission.apply(pd.to numeric)
        nir transmission.columns = ['wavelength', 'transmission']
        f98 = pd.read_csv("filters/f098m.IR.tab.txt", sep = ' ', header = None, ski
        f98 = f98.drop(0, axis=1)
        f98.columns = ["wavelength", "transmission"]
        f98.wavelength /= 1e4
        f105 = pd.read_csv("filters/f105w.IR.tab.txt", sep = ' ', header = None, sk
        f105 = f105.drop(0, axis=1)
        f105.columns = ["wavelength", "transmission"]
        f105.wavelength /= 1e4
        f110 = pd.read_csv("filters/f110w.IR.tab.txt", sep = ' ', header = None, sk
        f110 = f110.drop(0, axis=1)
        f110.columns = ["wavelength", "transmission"]
        f110.wavelength /= 1e4
        f125 = pd.read_csv("filters/f125w.IR.tab.txt", sep = ' ', header = None, sk
        f125 = f125.drop(0, axis=1)
        f125.columns = ["wavelength", "transmission"]
        f125.wavelength /= 1e4
        f140 = pd.read_csv("filters/f140w.IR.tab.txt", sep = ' ', header = None, sk
        f140 = f140.drop(0, axis=1)
        f140.columns = ["wavelength", "transmission"]
        f140.wavelength /= 1e4
        f160 = pd.read_csv("filters/f160w.IR.tab.txt", sep = ' ', header = None, sk
        f160 = f160.drop(0, axis=1)
        f160.columns = ["wavelength", "transmission"]
        f160.wavelength /= 1e4
        filters = [f98, f105, f110, f125, f140, f160]
        labels = ['F908M', 'F105W', 'F110W', 'F125W', 'F140W', 'F160W']
```

```
In []: fig = plt.figure(figsize=(10, 5), dpi= 300)
        ax = fig.gca()
        mask = np.where((nir_transmission.wavelength < 2.7) & (nir_transmission.wave
        x = nir_transmission.wavelength.to_numpy()[mask]
        y = nir_transmission.transmission.to_numpy()[mask]
        ax.plot(x, y)
        for i, filter in enumerate(filters):
            mask = np.where((filter.wavelength <= 2.7) & (filter.wavelength >= .9))
            x = filter.wavelength.to_numpy()[mask]
            y = filter.transmission.to_numpy()[mask]
            ax.plot(x, y, label = labels[i])
        ax.set_ylim(0, 1)
        ax.set xlabel("Wavelength ($\mu m$)")
        ax.set_ylabel("Transmission")
        ax.legend(loc = "upper right")
        ax.set_title("Atmospheric Transmission Curve in Near-IR vs. HST IR Filters")
        plt.show()
```



$$s = \frac{206265}{6} = 3.44 \times 10^4 s \cdot m^{-1} = .0344 s \cdot \mu m^{-1}$$

pixel scale = $18 \mu m imes .0344 s \cdot \mu m^{-1} = .62 s$

Frame width = 1024 * .62 = 634s = 10.6min

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$$F = 1 \times 10^{-17} Wm^{-2}$$

$$\delta = 15s$$

$$\sigma = \frac{4F}{\pi \delta^2}$$

$$d_p=25\mu m$$

$$s_p = rac{206265*d_p}{f}$$

Albert:

$$D_A = .3m$$

$$f_A=2.4m$$

$$F_A = \sigma rac{\pi D^2}{4}$$

$$E = 100 s_p^2 F_A = 100 (rac{206265 d_p}{f_A})^2 rac{4F}{\pi \delta^2} rac{\pi D_A^2}{4}$$

```
In []: F = 1e-7
    delta = 15
    D = .3
    dp = 25e-6
    f = 2.4

sigma = 4*F / (np.pi * delta**2)
Fa = sigma * np.pi * D**2 / 4
    s = 206265 * dp / f

print("Pixel energy collected for Albert:", np.round(100 * s**2 *Fa, 11), "J
```

Pixel energy collected for Albert: 1.847e-08 J

Bertha:

$$D_B = 30m$$

$$f_B = 120m$$

$$F_B = \sigma rac{\pi D^2}{4}$$

$$E=100s_p^2F_B$$

```
In []: F = 1e-7
    delta = 15
    D = 30
    dp = 25e-6
    f = 120

sigma = 4*F / (np.pi * delta**2)
Fb = sigma * np.pi * D**2 / 4
    s = 206265 * dp / f

print("Pixel energy collected for Bertha:", np.round(100 * s**2 * Fb, 11), "
```

Pixel energy collected for Bertha: 7.386e-08 J

9)

For 300 nm:

$$\theta = \frac{1.22 \cdot 300 \times 10^{-9}}{2.4m}$$

In []: diffraction_limit = 1.22 * 300e-9 / 2.4 * 206265
print("The diffraction limit for Hubble at 300 nm is:", np.round(diffraction

The diffraction limit for Hubble at 300 nm is: 0.031 arc seconds

At 2 μm :

$$heta=rac{1.22\cdot2 imes10^{-6}}{2.4m}$$

In []: diffraction_limit = 1.22 * 2e-6 / 2.4 * 206265
print("The diffraction limit for Hubble at 2 micrometers is:", np.round(diff

The diffraction limit for Hubble at 2 micrometers is: 0.21 arc seconds

For an 8m telescope at 2 μm :

$$heta=rac{1.22\cdot2 imes10^{-6}}{8m}$$

In []: diffraction_limit = 1.22 * 2e-6 / 8 * 206265 print("The diffraction limit for the 8 meter scope at 2 micrometers is:", ng

The diffraction limit for the 8 meter scope at 2 micrometers is: 0.063 arc seconds