

Physics 441/541: Stars and Star Formation
Final Exam Formulas May 5, 2022

In addition to this page, you are allowed both sides of a letter-sized page for your own notes or formulas.

some possibly useful formulas

(for which I am on purpose not providing more details)

$$\begin{aligned}
 \frac{dP}{dr} &= -\rho(r)g(r) = -\frac{Gm(r)\rho(r)}{r^2} & \langle P \rangle &= -\frac{1}{3} \frac{E_{\text{pot}}}{V} & T_c &\sim \frac{GM\mu m_p}{kR} & P_c &\sim \frac{GM^2}{R^4} \\
 B_\lambda(T) &= \frac{2hc^2/\lambda^5}{\exp(hc/\lambda kT) - 1} & F &= \sigma T^4 & L &= 4\pi R^2 \sigma T_{\text{eff}}^4 & f &= L/4\pi d^2 & d &= 1/p \\
 m &= -2.5 \log(f/f_0) & \mu &= m - M = 5 \log(d/10 \text{ pc}) & M_{\text{bol}} &= -2.5 \log(L/L_\odot) + 4.74 \\
 P &= K\rho^\gamma & \gamma &= 1 + 1/n & R &\propto M^{(1-n)/(3-n)} & E_{\text{pot}} &= -\frac{3}{5-n} \frac{GM^2}{R} \\
 P &= \frac{1}{3} aT^4 & P &= nkT = \frac{\rho kT}{\mu m_p} & P &= \frac{h^2}{5m_e} \left[\frac{3}{8\pi} \right]^{2/3} n_e^{5/3} & P &= \frac{hc}{4} \left[\frac{3}{8\pi} \right]^{1/3} n_e^{4/3} \\
 E_n &= -13.6 \text{ eV} \left(\frac{Z^2}{n^2} \right) & g_n &= 2n^2 & \frac{n_m}{n_n} &= \frac{g_m}{g_n} \exp\left(-\frac{E_m - E_n}{kT}\right) \\
 \frac{n_{II}}{n_I} &= \frac{2Z_{II}}{n_e Z_I} \left(\frac{2\pi m_e kT}{h^2} \right)^{3/2} \exp\left(-\frac{\chi_I}{kT}\right) & \left| \frac{dT}{dr} \right| &> \frac{\gamma-1}{\gamma} \frac{T}{P} \left| \frac{dP}{dr} \right| = \frac{\gamma-1}{\gamma} \frac{\mu m_p}{k} g \\
 \ell &= \frac{1}{n\sigma} = \frac{1}{\rho\kappa} & \frac{dT}{dr} &= -\frac{3}{4ac} \frac{\rho\kappa}{T^3} \frac{L}{4\pi r^2} & E &= \Delta mc^2 = \epsilon Mc^2 & E_G &= (\pi\alpha Z_A Z_B)^2 2m_r c^2 \\
 \text{probability} &\approx \exp\left[-\left(\frac{E_G}{E}\right)^{1/2}\right] & E_0 &= \left[\frac{E_G(kT)^2}{4}\right]^{1/3} & \alpha &= \left(\frac{E_G}{4kT}\right)^{1/3} - \frac{2}{3} \\
 R_{AB} &= n_A n_B \left[\frac{8}{\pi m_r} \right]^{1/2} \left[\frac{1}{kT} \right]^{3/2} \int_0^\infty S(E) \exp\left[-\frac{E}{kT} - \left(\frac{E_G}{E}\right)^{1/2}\right] dE & R_{AB} &\propto T^\alpha \\
 \frac{dm}{dr} &= 4\pi r^2 \rho & \frac{dL}{dr} &= 4\pi r^2 \rho \epsilon & \kappa_{\text{es}} &= \frac{n_e \sigma_T}{\rho} \approx 0.2(1+X) \text{ cm}^2 \text{ g}^{-1} & L_{\text{Edd}} &= \frac{4\pi GMm_p c}{\sigma_T} \\
 t_{\text{dyn}} &= \frac{1}{\sqrt{G\rho}} & t_{\text{KH}} &= \frac{GM^2}{RL} & t_{\text{nuc}} &= \frac{E_{\text{nuc}}}{L} \approx \frac{f_M \epsilon Mc^2}{L} & t_{\text{dyn}} &\ll t_{\text{conv}} \ll t_{\text{KH}} \ll t_{\text{nuc}} \\
 M_{Ch} &\approx 2.02 \left(\frac{\sqrt{3\pi}}{2} \right) \left(\frac{1}{\mu_e m_p} \right)^2 \left(\frac{\hbar c}{G} \right)^{3/2} & z &= \frac{\Delta\lambda}{\lambda} = \left[1 - \frac{R_S}{R} \right]^{-1/2} - 1 & R_S &= \frac{2GM}{c^2} \\
 J^2 &= Ga \frac{M_1^2 M_2^2}{M_1 + M_2}
 \end{aligned}$$