

### By Dr. Vishal Chauhan

### Contributions

### 17<sup>th</sup> Century



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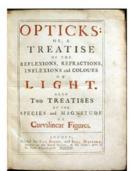
The laws of the arpropaga





Principle of wavefront

#### 18th Century



By Issac Newton





Coulomb

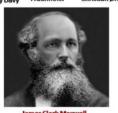
19th Century











**Wave Particle Duality** 

#### Particle Nature of light:

An object has a definite position in space which cannot be simultaneously occupied by another particle and indentifiable by their distinct properties such as mass, momentum, kinetic engergy, spin and electric charge.

Examples: Photoelectric effect, emission and absroption of radiation by substances, black body radiation etc.

#### Wave Nature of light:

A wave means periodically repeated pattern in space which is specified by its wavelength, frequency, amplitude of disturbance, intensity, energy and momentum.

Examples: Interference, Diffraction, Polarization etc.



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1900: Max Planck suggests that radiation is quantized (it comes in discrete amounts.)

1905: Albert Einstein, one of the few scientists to take Planck's ideas seriously, proposes a quantum of light (the photon) which behaves like a particle. Einstein's other theories explained the equivalence of mass and energy, the particle-wave duality of photons, the equivalence principle, and special relativity.

1913: Niels Bohr succeeds in constructing a theory of atomic structure based on quantum ideas.

1919: Ernest Rutherford finds the first evidence for a proton.

1921: James Chadwick and E.S. Bieler conclude that some strong force holds the nucleus together.

1923: Arthur Compton discovers the quantum (particle) nature of x rays, thus confirming photons as particles.

1924: Louis de Broglie proposes that matter has wave properties.

1925 (Jan): Wolfgang Pauli formulates the exclusion principle for electrons in an atom.

1926: Erwin Schroedinger develops wave mechanics, which describes the behavior of quantum systems for bosons. Max Born gives a probability interpretation of quantum mechanics. G.N. Lewis proposes the name "photon" for a light quantum.

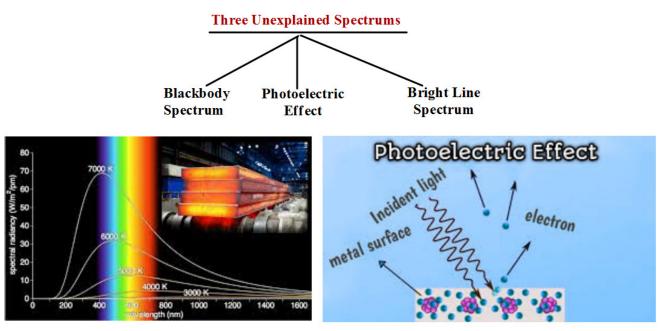
1927: Werner Heisenberg formulates the uncertainty principle: the more you know about a particle's energy, the less you know about the time of the energy (and vice versa.) The same uncertainty applies to momenta and coordinates.

1928: Paul Dirac combines quantum mechanics and special relativity to describe the electron.

1930: Quantum mechanics and special relativity are well established. There are just three fundamental particles: protons, electrons, and photons.





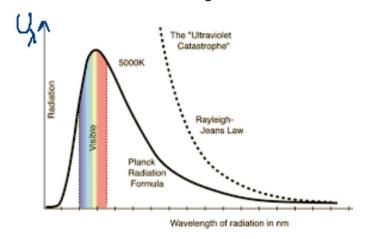


https://phet.colorado.edu/sims/html/blackbody-spectrum/latest/blackbody-spectrum en.html



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Rayleigh- Tears Low Ford for larger wavelengths  $U_{\lambda} = \frac{8\pi}{\lambda^{4}} R.T$ Bad for smaller R = Boltzman Constant wavelengths



#### **Entry of Max Planck**

- \* Max Planck made a big advancement in quantum theory when he put forward a model saying that energy of any oscillation can be absorbed or emitted only in units of a basic energy E which is proportional to the frequency of oscillation.
- \* In case of light, this means the energy of various modes of stationary waves in the enclosure can be E, 2E, 3E......etc. with  $E=h \mathcal{V} = h \mathcal{C}/\lambda$ , h is a constant. Using this hypothesis he dervied the following equation for spectral distribution;

$$U_{\lambda} = \frac{8\pi hc}{\lambda^{5}} \cdot \frac{1}{e^{hc/\lambda kT} - 1}$$

\* This equation matches extremely well with experimental results in the entire range of wavelength, when the value of h is given 6.6 x 10-34 J.S

Planck's hypothesis was a great revolution and the constant h is rightly called 'Planck's Constant'.



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### Equation of monochromatic light

$$\vec{E} = \vec{E}_a \cos (\kappa z - \omega t)$$

$$(i) \lambda = \frac{2\pi}{K}, K = \frac{2\pi}{\lambda}$$

(iii) frequency 
$$(v) = \frac{\omega}{2\pi}$$
  
 $\omega = 2\pi v$  (Angular frequency)



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If wave is a photon, then each photon has an energy (E)

$$F = h\nu$$

$$= h\nu \times \frac{2\pi}{2\pi}$$

$$= \frac{h}{2\pi} \cdot 2\pi\nu \quad \therefore h = \frac{h}{2\pi}$$

 $E = h.\omega$  &  $\omega = 2\pi v$ For EM Radiation (No matter)



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### Relation between momentum and wavelength

$$P = \frac{h}{\lambda} \quad (Muldiply & Divide by 2\pi)$$

$$= \frac{h}{\lambda} \times \frac{2\pi}{2\pi}$$

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$$= \frac{h}{2\pi} \times \frac{2\pi}{\lambda}$$

$$= \frac{h}{2\pi} \cdot \frac{2\pi}{\lambda} \quad here \frac{h}{2\pi} = \frac{h}{h}$$

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When does quantum mechanics apply?

Not a simple question

But \* when angular momentum ~ T \* When uncertainties APAR 25 AEAt ~t

of when any action Sat

1 = 1.05457148 X10-34 kgm2/5

### Electron in hydrogen atom

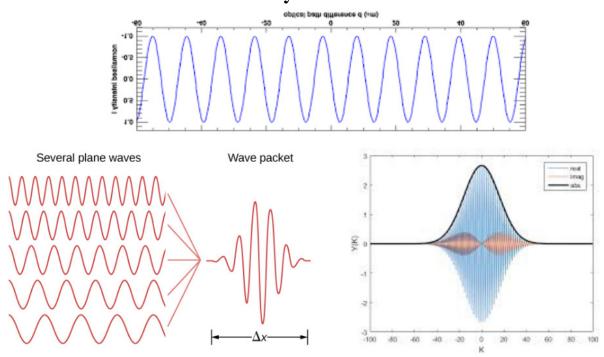
energy = 10eV = 102 kg·m/s Size of atom ~ 10-10m APARNITATION KONT Quantum domain

### Speck of dust

mass ~ 10-6 Hg relocity~ 1-m/s Size ~ 10-5 m momentum p=10 kgm/s AP = 10 8 Kgm/s position uncertainty ۵2 م اه- د س DP Dx = 10-14 Kgmt/s ~10th Classical



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Many waves in a wave packet can be defined by a single differential equation.

The general wave equation

$$\frac{3x^2}{3^2} = \frac{y^2}{1} \cdot \frac{3t^2}{3^2}$$

In this equation velocity (v) is considered constant but it is not practical in case of material waves therefore Schrodinger provided a solution for it which is known as Schrodinger wave equation.

$$-\frac{h^2}{2\pi}\frac{3x^2}{3^2\Psi} + V(x)\Psi = i\hbar \frac{3\Psi}{3t}$$



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Let a common wave is defined as

$$y = Af(x-vt)$$

$$\frac{dy}{dx} = A\frac{df}{d(x-vt)} \cdot \frac{d}{dx}(x-vt)$$
with  $v$ -velocity.
$$v(fixed) = const.$$

O First differentiate w.r. to z, t-fixed

$$\frac{\partial y}{\partial x} = Af'$$

$$\frac{\partial^2 y}{\partial x^2} = A \frac{df'}{d(x-yt)} \cdot \frac{d(x-yt)}{dx}$$

$$\frac{\partial^2 y}{\partial x^2} = A f''$$

Now differentiate w.r. to t, x-fixed

$$\frac{\partial t}{\partial y} = Af' \cdot \frac{d}{dt}(x - vt)$$
$$= Af' \cdot (-v)$$

$$\frac{\partial^{2}y}{\partial t^{2}} = Af''(-y)(-y)$$

$$= Af''y^{2}$$

$$\frac{\partial^{2}y}{\partial t^{2}} = y^{2}(Af'') \qquad "Af'' = \frac{\partial^{2}y}{\partial x^{2}}$$

$$\therefore \frac{3^2y}{3t^2} = y^2 \frac{3^2y}{3x^2} \text{ or } \frac{3^2y}{3x^2} = \frac{1}{1} \cdot \frac{3^2y}{3t^2}$$
General wave equation



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### Time Dependent Schrodinger wave equation

In general wave equation velocity (v) is considered constant but it is not practical in case of material waves threfore Schrödinger provided a solution for it which is known as Schrodinger wave equation.

Consider a case of free particle which is not interacting and no force is applied.

wave nature > k (wave factor) W (Angular freq.)

Total energy 
$$\Rightarrow E = K \cdot E \cdot = \frac{1}{2}mv^2$$
  
or  $E = \frac{p^2}{2m}$ 

20

$$E = \frac{b^2}{2\pi n}$$
,  $E = h\nu$ ,  $P = hR$  (wave associated with free particle)

It is the real part of

$$e^{ix} = Casx + iSin(x)$$

$$\Psi = A e^{[i(\kappa z - \omega t)]}$$

$$(l=\sqrt{-1})$$

or  $\Psi = Ae[i(kz-\omega t)]$   $(i=\sqrt{-1})$ General wave equation (i=-1)



$$Y = A e^{\left(i\left(kx - \omega t\right)\right)}$$

$$k = \frac{P}{\hbar}, \quad \omega = \frac{E}{\hbar}$$

$$\Psi = A e^{i\left(\frac{P}{\hbar}x - \frac{E}{\hbar}t\right)}$$

$$\Psi = A e^{i\sqrt{\hbar}} \left(\frac{Px - Et}{\hbar}\right)$$

$$(2)$$

- 1. Now differentiate equation (2) w. r. to (x) two times.
- 2. Then differentiate equation (2) w. r. to (t) two times

$$\frac{\partial \psi}{\partial x} = A e^{i/h} (px - Et), \frac{\partial}{\partial x} \left[ \frac{i}{h} (px - Et) \right]$$

$$= A e^{i/h} (px - Et), \frac{i}{h} \frac{\partial}{\partial x} (px - Et)$$

$$= \psi \cdot \frac{i}{h} p$$

$$\frac{\partial \psi}{\partial x} = \frac{i}{h} p \psi \Rightarrow p \psi = \frac{h}{i} \frac{\partial \psi}{\partial x}$$

$$p \psi = -i \frac{\partial \psi}{\partial x}$$



$$\frac{\partial^{2} \Psi}{\partial x^{2}} = A \left( \frac{i}{\hbar} P \right) \left( \frac{i}{\hbar} P \right) e^{i/h} \left( Px - Et \right)$$

$$= -A \frac{P^{2}}{\hbar^{2}} e^{i/h} \left( Px - Et \right)$$

$$= -\frac{P^{2}}{\hbar^{2}} \left[ A e^{i/h} \left( Px - Et \right) \right]$$

$$\frac{\partial^{2} \Psi}{\partial x^{2}} - \frac{P^{2}}{\hbar^{2}} \Psi$$

$$P^{2} \Psi = -\frac{\hbar^{2}}{\hbar^{2}} \frac{\partial^{2} \Psi}{\partial x^{2}}$$

Now diff, war, but 
$$-\frac{3\Psi}{3t} = \left[Ae^{i\lambda(Px-Et)}\right](-\frac{i}{\lambda}E)$$

$$= \left[\Psi\right](-\frac{i}{\lambda}E)$$

$$\frac{3\Psi}{3t} = -\frac{i}{\lambda}E\Psi$$

$$E\Psi = -\frac{t}{\lambda}\frac{3\Psi}{3t}$$

$$E\Psi = i \frac{3\Psi}{3t}$$



For free particle
$$E = \frac{b^2}{2m}$$

$$E\Psi = \frac{p^2}{2m}\Psi$$

$$\frac{1}{2} + \frac{3 \psi}{3 t} = \frac{- t^2 3^2 \psi}{2 m 3 x ^2}$$

$$-\frac{\hbar^2}{3m} \cdot \frac{3^2 \Psi}{3 \pi^2} = i \frac{\hbar^3 \Psi}{3t}$$



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if particle is not free inclusion of potential energy. Here  $E = K \cdot E + P \cdot E \cdot E = \frac{p^2}{2m} + V(x) \psi$   $EV = \frac{p^2 v}{2m} + V(x) \psi$ 

$$-\frac{k^2}{2m}\cdot\frac{3^2\Psi}{3x^2}+V(x)\Psi=it\frac{3\Psi}{3t}$$

This is final time dependent Schrodinger wave equation