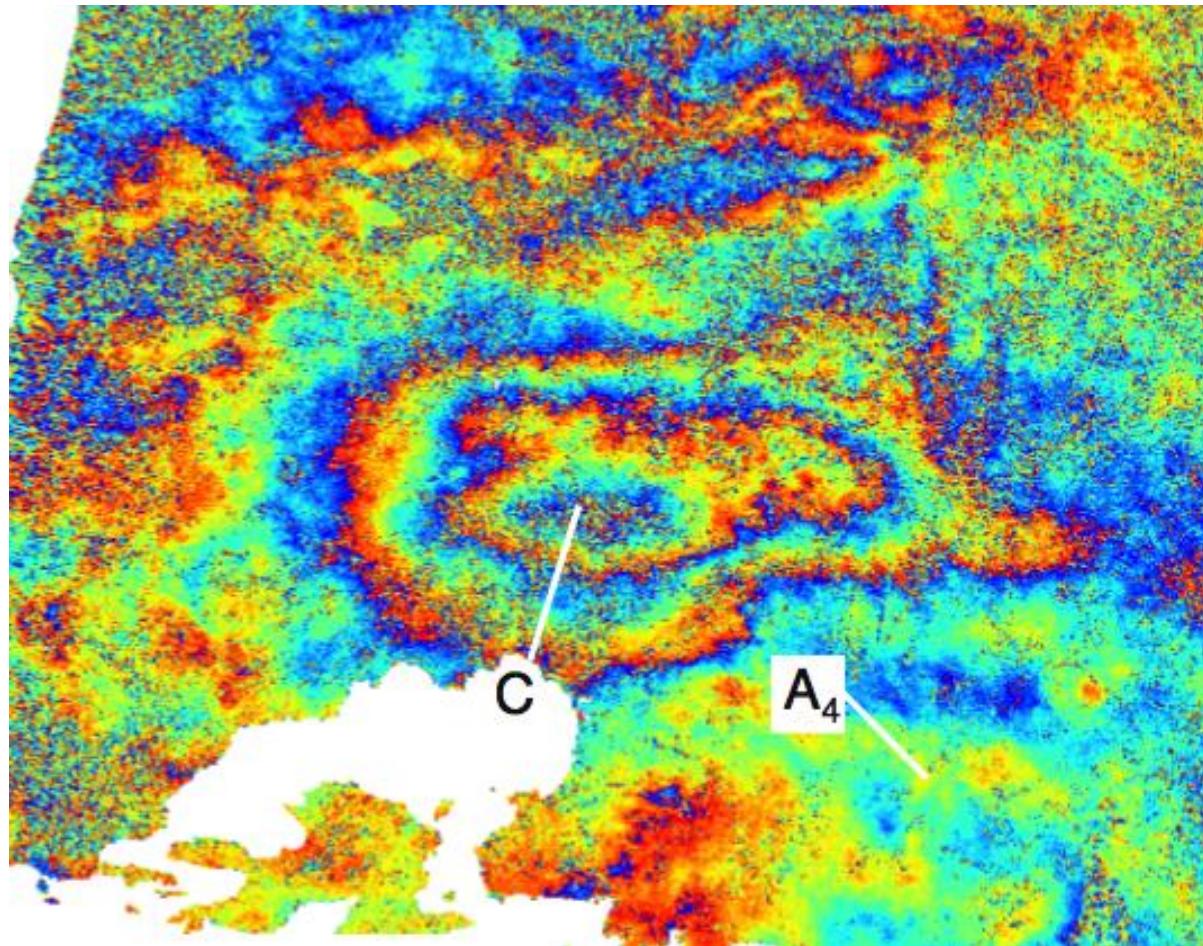


# InSAR training 2024



Contributions to interferometric phase

# Outline

The contributions to interferometric phase

- orbit
- topography
- atmosphere
- in-pixel scatterers
- deformation

# Components of interferometric phase

$$\Delta\phi_{\text{int}} = \underline{\Delta\phi_{\text{orb}}} + \underline{\Delta\phi_{\text{topo}}} + \underline{\Delta\phi_{\text{atm}}} + \underline{\Delta\phi_{\text{pixel}}} + \underline{\Delta\phi_{\text{def}}}$$

The phase of an individual interferogram can be divided into five parts:

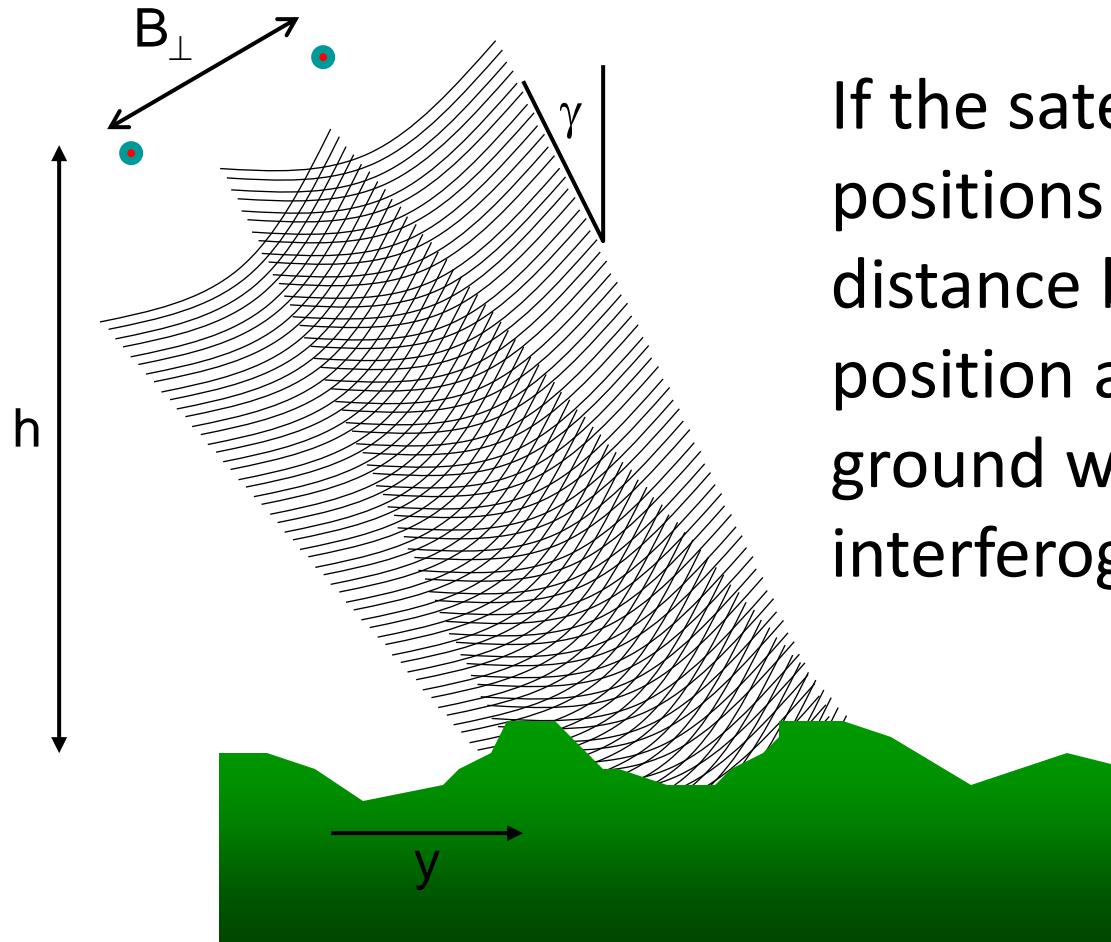
Three are related to the difference in distance between the satellite and the ground

One is related to the difference in the properties of the medium that the radar pulse moves through

One is related to the change in properties of the pixel on the ground

# Orbital phase (1)

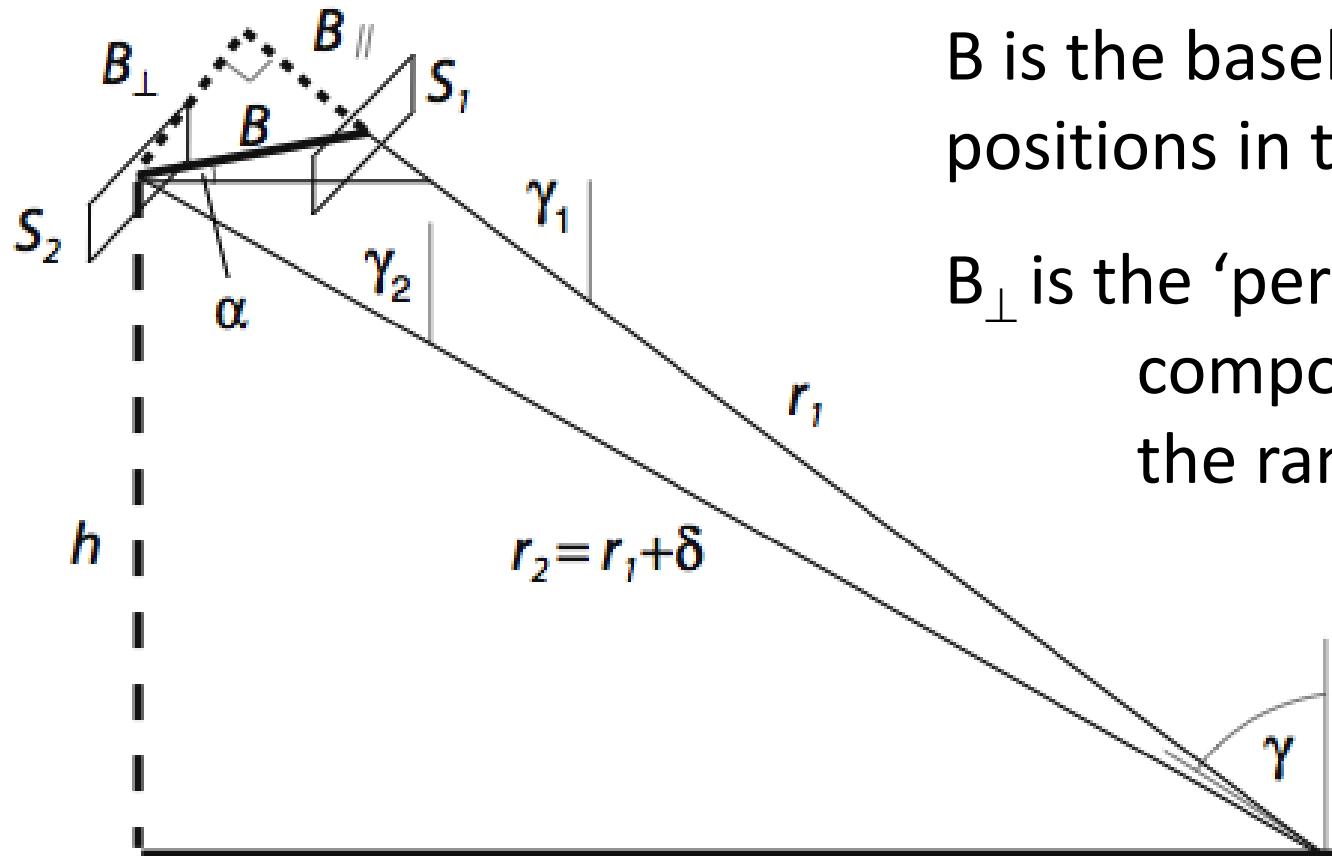
$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$



If the satellite is in different orbital positions for the two images, the distance between each satellite position and successive points on the ground will vary across the whole interferogram

# Orbital phase (2)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$



$B$  is the baseline – the distance between satellite positions in the cross-track plane

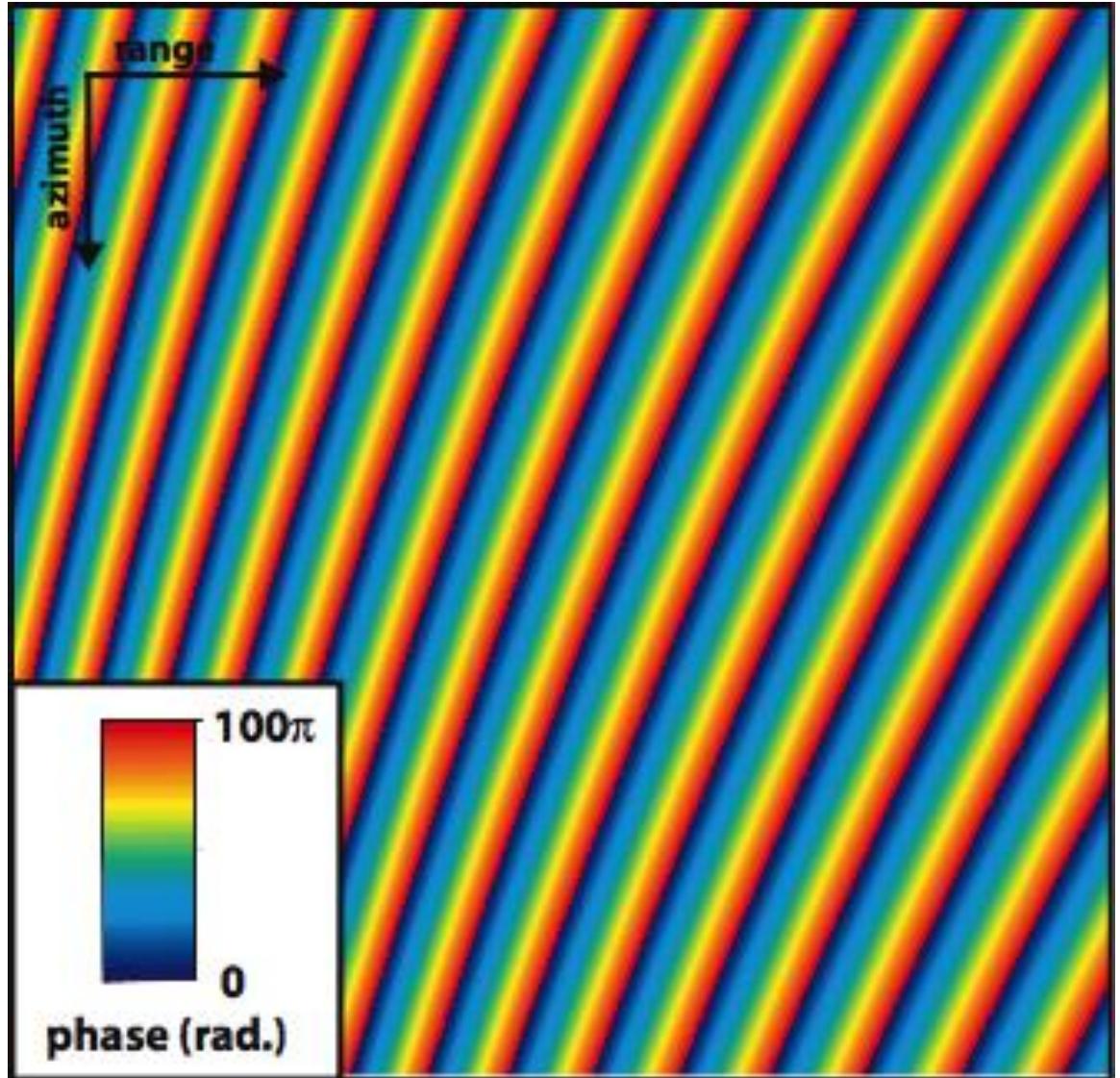
$B_{\perp}$  is the ‘perpendicular baseline’, the component of baseline perpendicular to the range direction

$$\frac{\partial\phi}{\partial y} = \frac{4\pi B_{\perp} \cos^2 \gamma}{h\lambda}$$

# Orbital phase (3)

The orbital phase component is usually the largest contributor to an interferogram

ERS-1, baseline varies between 121 m at the top and 136 m at the bottom

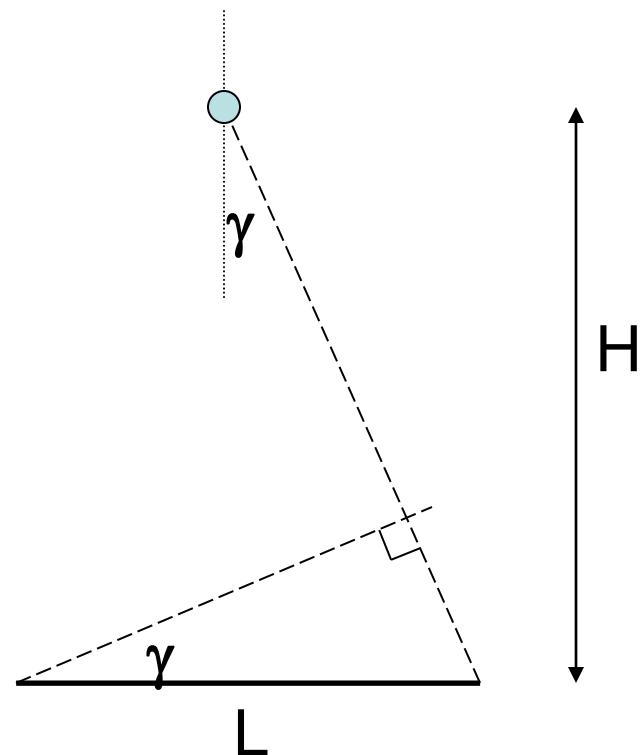


# The critical baseline

In order to identify the phase ramp between pixels, it cannot exceed a shift of  $2\pi$  (or a round-trip distance of  $\lambda$ ) per pixel

Difference in round trip  
distance at ends of a pixel with  
length  $L$ :

$$2L \sin \gamma$$



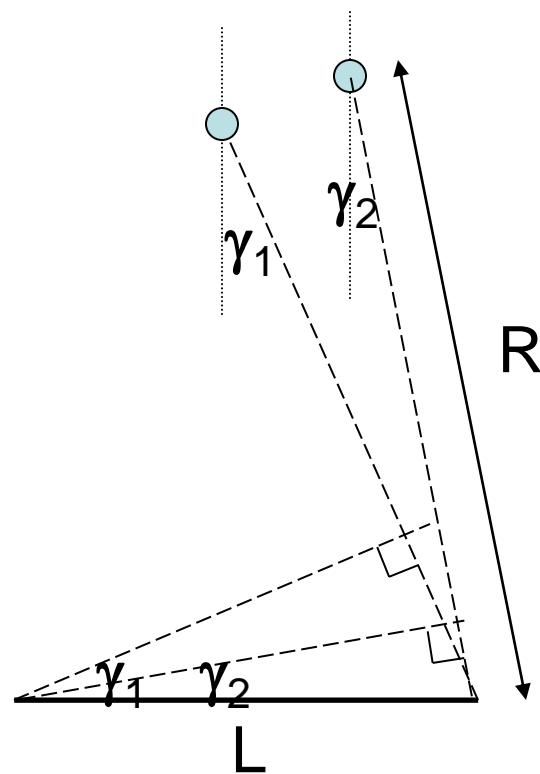
# The critical baseline

In order to identify the phase ramp between pixels, it cannot exceed a shift of  $2\pi$  (or a round-trip distance of  $\lambda$ ) per pixel

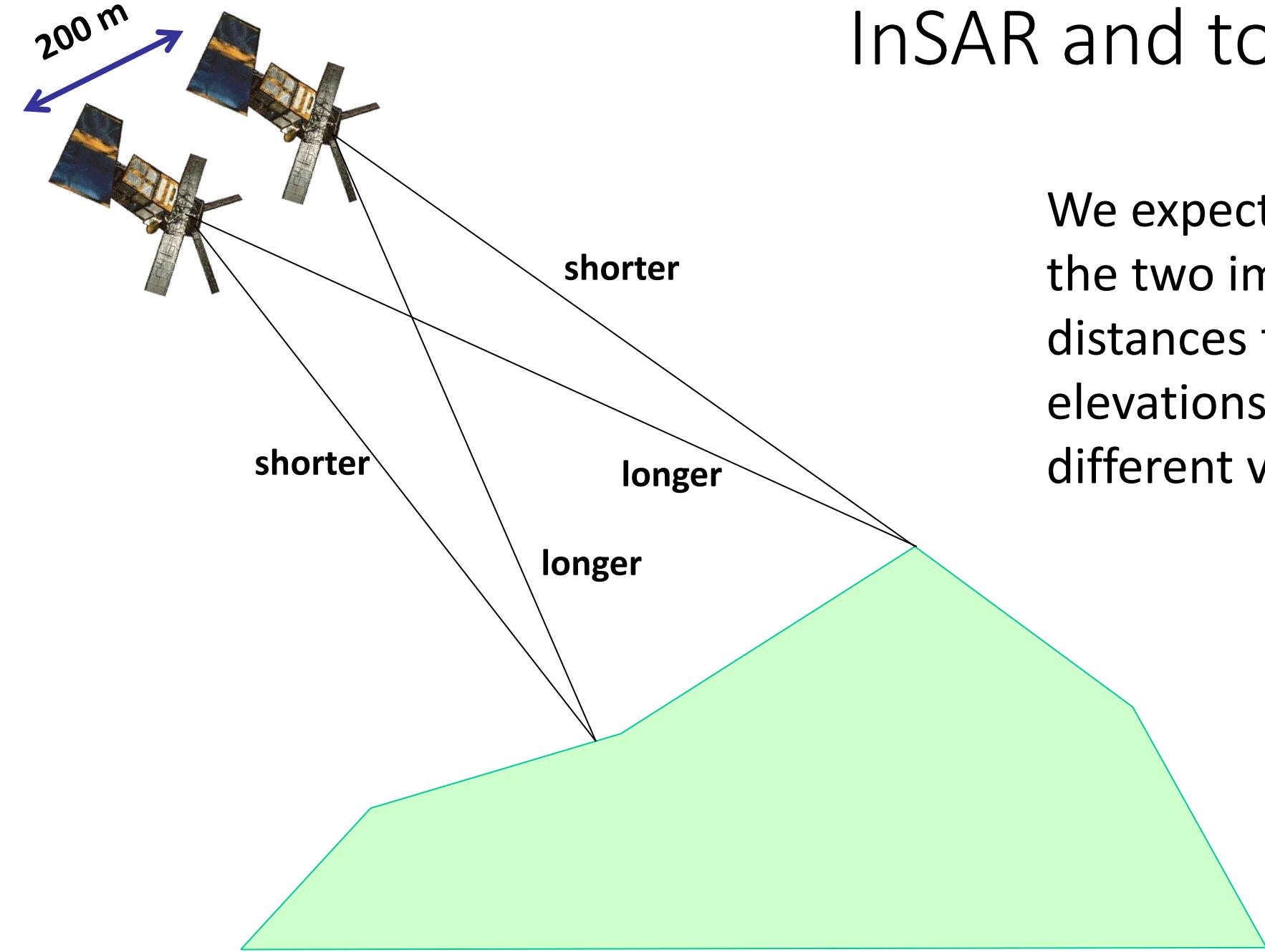
Difference in difference in round trip distance at ends of a pixel with length  $L$  must be less than  $\lambda$ :

$$2L (\sin \gamma_1 - \sin \gamma_2) < \lambda$$

If  $L = 20$  m,  $R = 850$  km,  $\gamma_1 \approx \gamma_2 \approx 23^\circ$  and  $\lambda = 0.056$  m, what is the critical baseline?



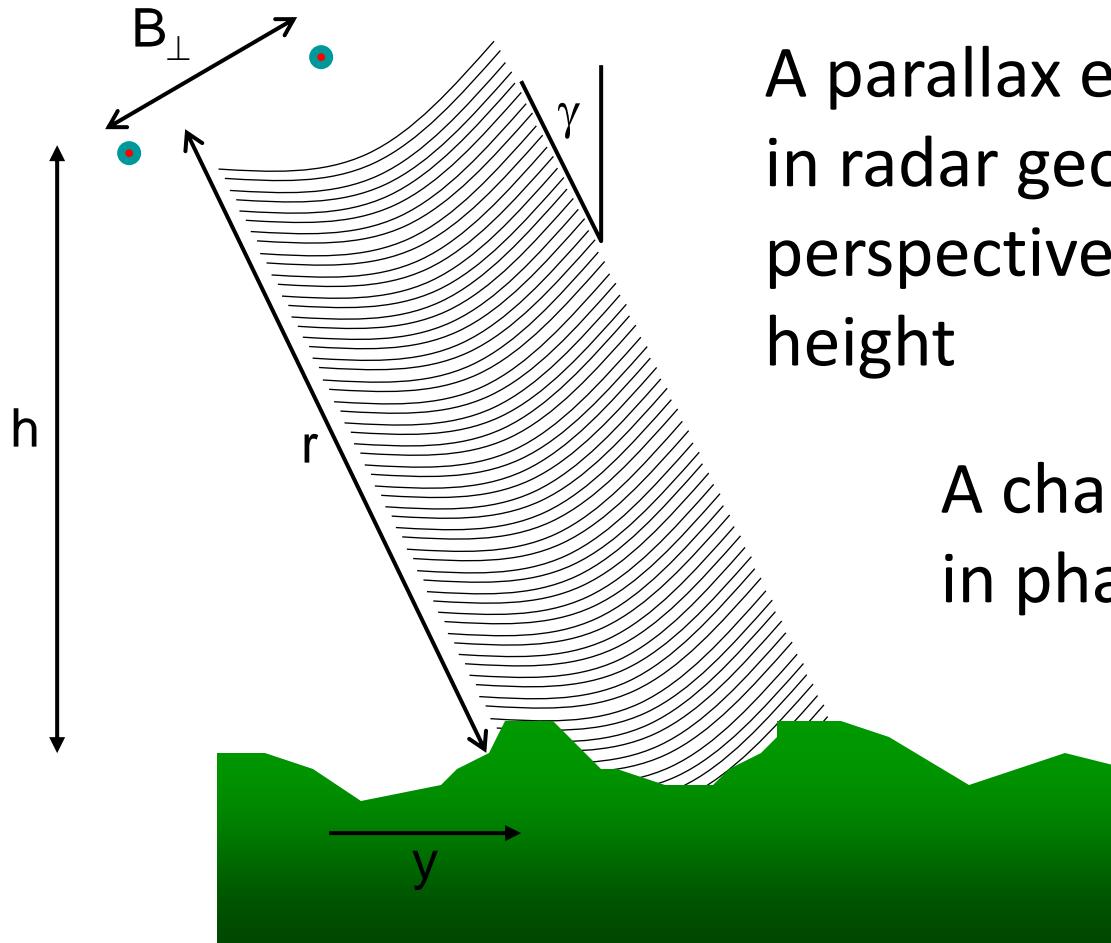
# InSAR and topography



We expect phase shifts between the two images due to different distances from different elevations of the ground at different viewing positions

# Topographic phase (1)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$



A parallax effect (like stereoscopy, but in radar geometry) gives different perspective on changes in topographic height

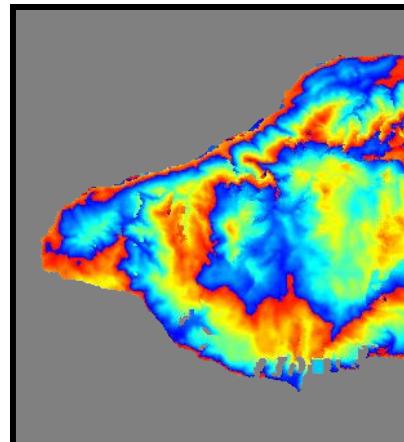
A change in height,  $h_a$ , equals  $2\pi$  in phase change

$$h_a = \frac{r\lambda \sin \gamma}{2B_{\perp}}$$

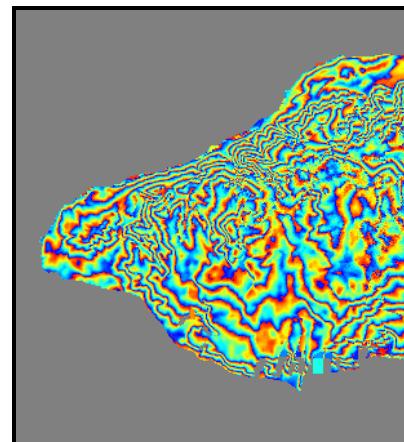
# Topographic phase (2)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

$h_a = 500 \text{ m}$   
 $B_\perp = 20 \text{ m}$



$h_a = 100 \text{ m}$   
 $B_\perp = 100 \text{ m}$



- Stereoscopic effect  $\Rightarrow$  topographic fringes
- 1 fringe for each change in elevation  $h_a$

$$h_a = \frac{r\lambda \sin \gamma}{2B_\perp} \approx \frac{10,000}{B_\perp} \text{ (for ERS)}$$

- Remove  $Df_{\text{topo}}$  using a pre-existing DEM
- Quality of DEM required  $f(h_a)$

Height error =  $\varepsilon_h$

$\Rightarrow$  phase error =  $2\pi(e_h / h_a)$

# Atmospheric phase (1)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

A foggy morning, near ancient  
Mycenae, Greece



# Atmospheric phase (2)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

Atmosphere contributes to measured phase:

$$\text{path\_delay} = \int_0^{\text{atm thickness}} (n_1(h) - n_2(h)) dh$$

n = refractive index,

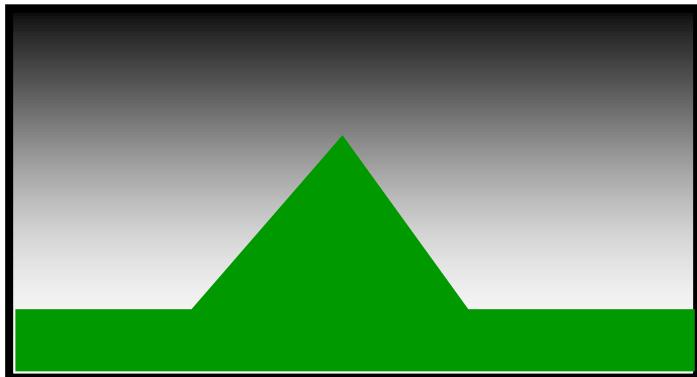
= f (**humidity**, temperature, pressure)

- varies by very small amount
- BUT, integrated over whole atmosphere
  - ⇒ differential path delays of up to 10 cm in extreme cases.

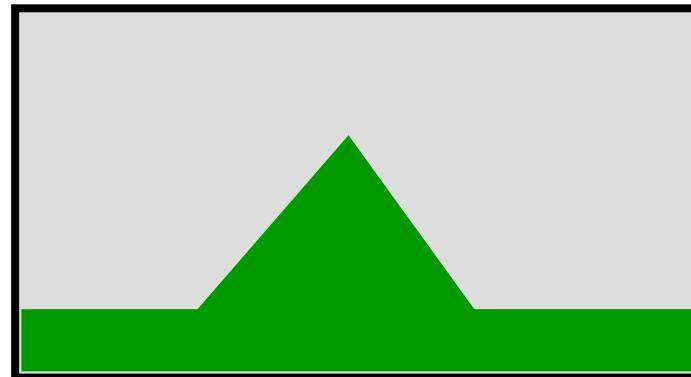
# Atmospheric phase (3)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

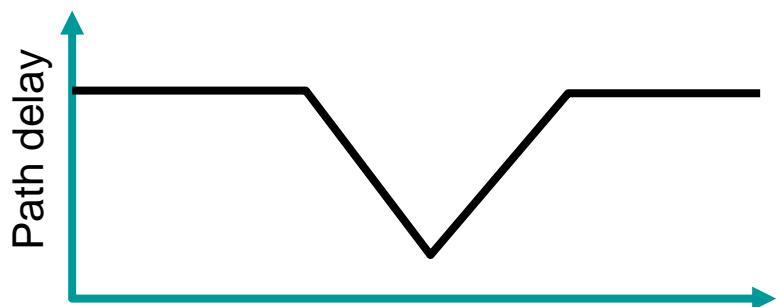
*Layered atmosphere*



Pass 1



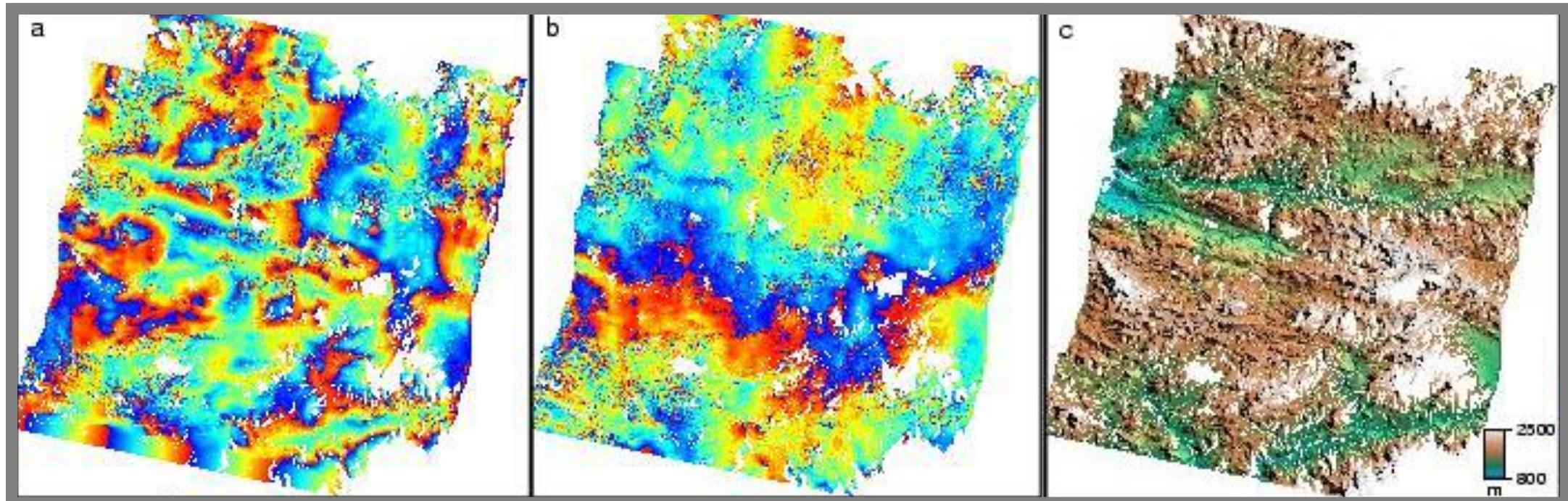
Pass 2



# Atmospheric phase (4)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

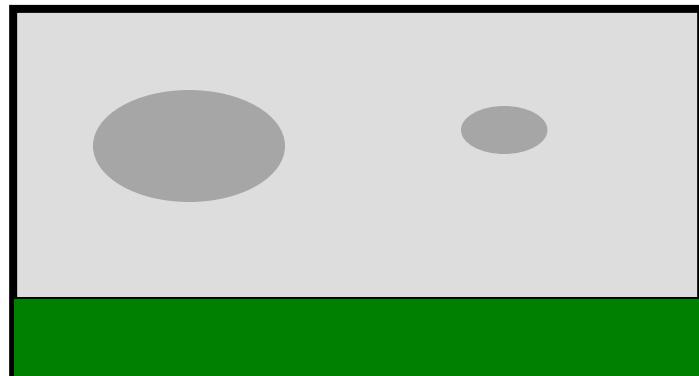
*Layered atmosphere*



# Atmospheric phase (5)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

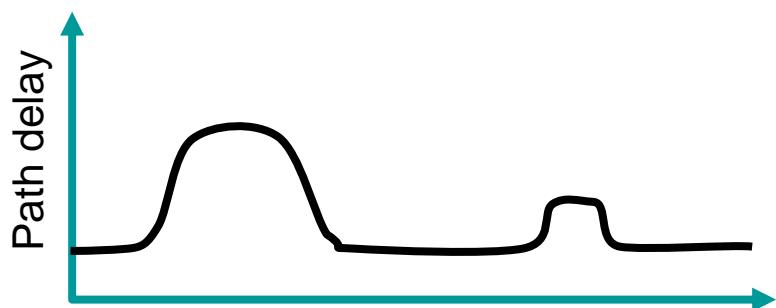
*Turbulent atmosphere*



Pass 1



Pass 2

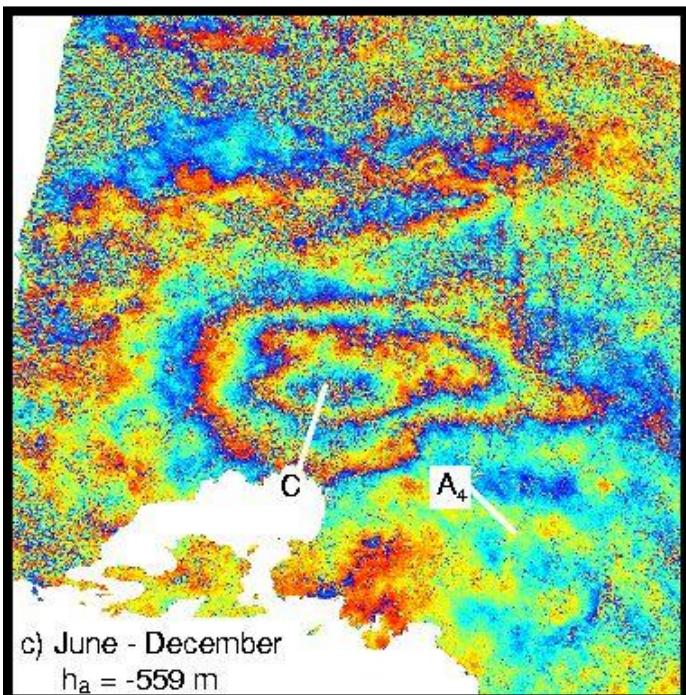


# Atmospheric phase (6)

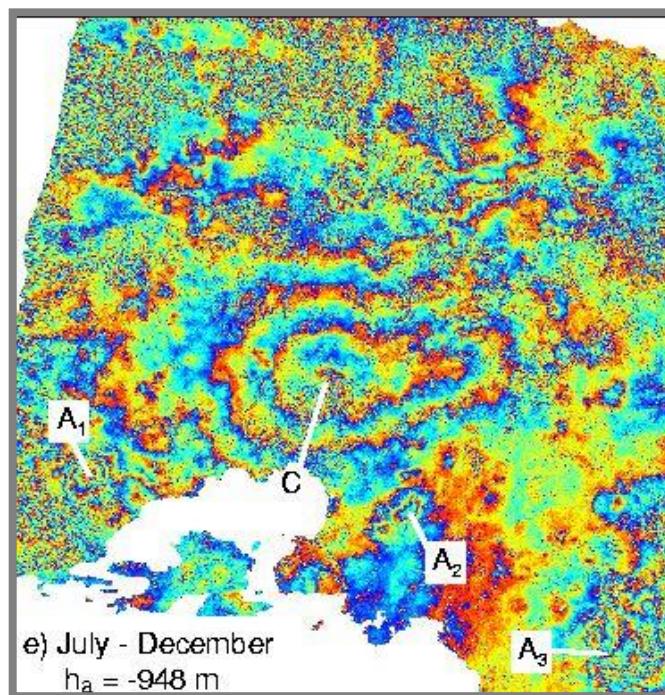
$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

*Turbulent atmosphere*

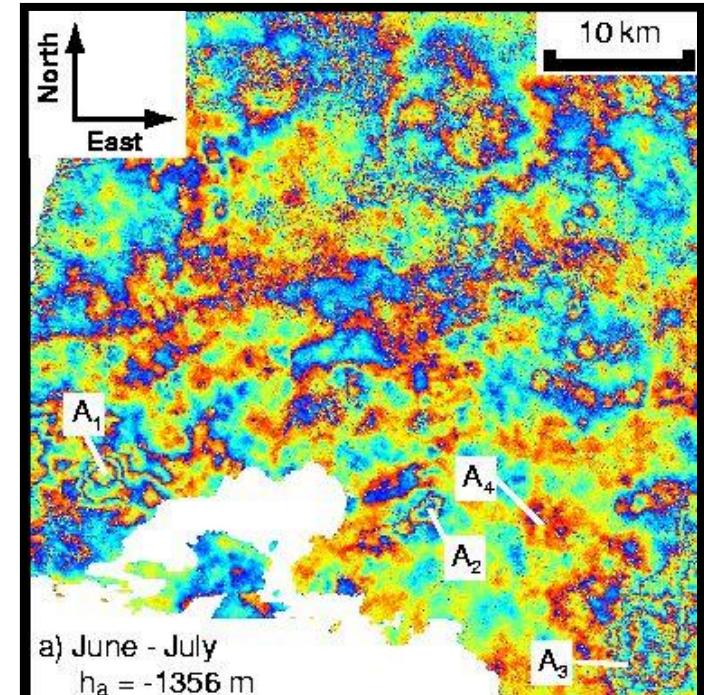
Athens Earthquake – September 1999



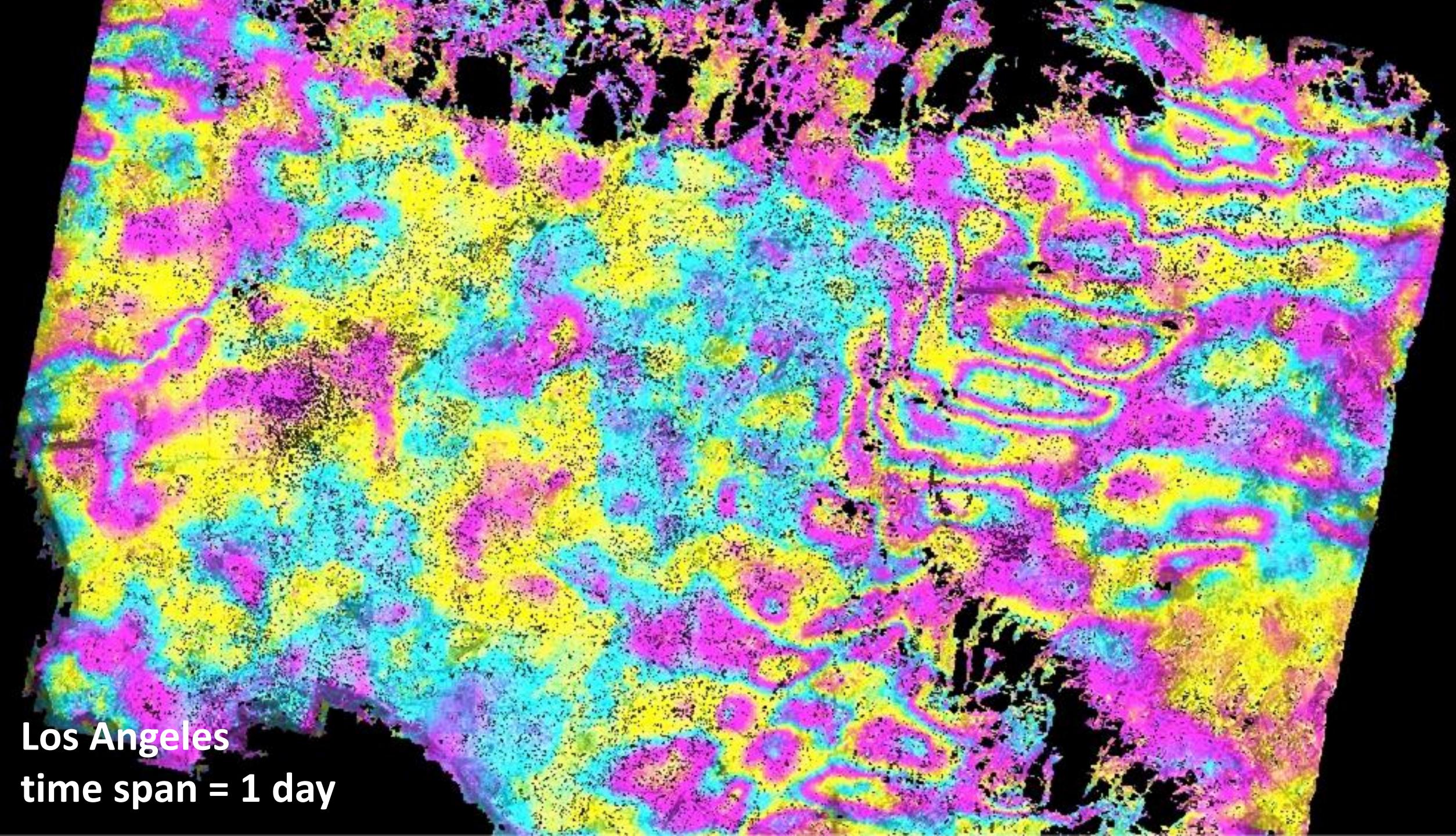
June to December



July to December



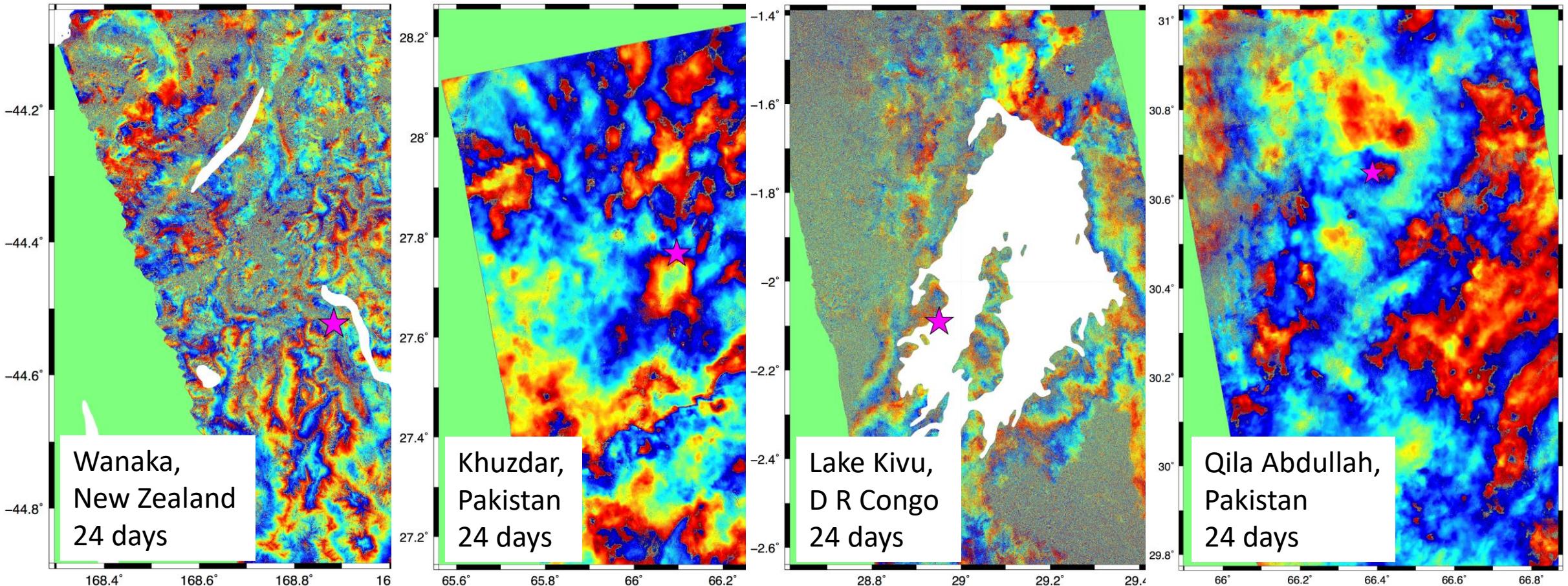
June to July



**Los Angeles**  
**time span = 1 day**

# Which blob?

Shortest Sentinel-1 pairs for shallow M>5.5 earthquakes



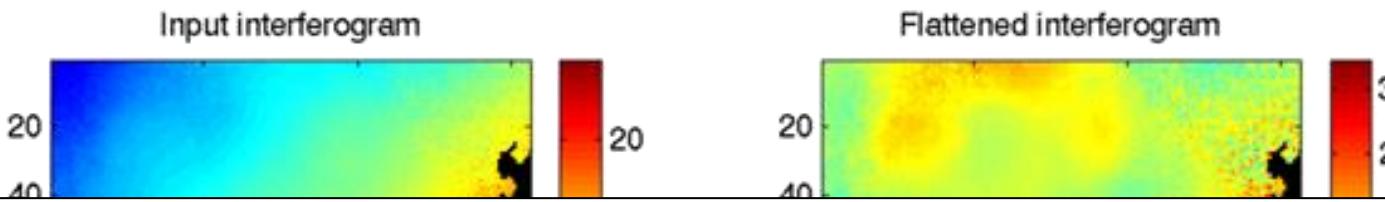
# Atmospheric phase (7)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

Size of  $Df_{\text{atm}}$  (at sea level)  $\sim$  0 to 3 fringes (90 mm)

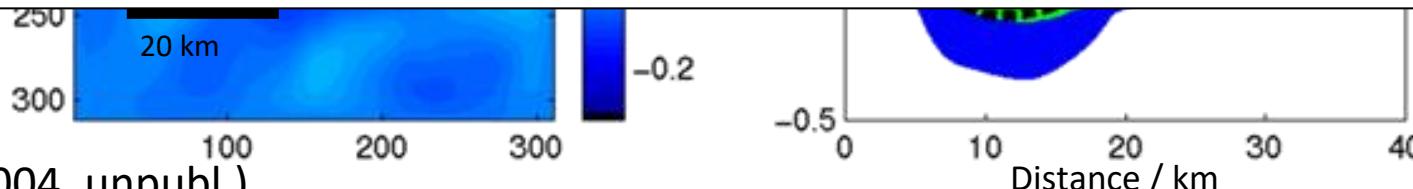
Methods for dealing with  $Df_{\text{atm}}$

- Ignore (most common)
- Exclude
- Quantify
- Model based on other observations (e.g. GPS, meteorology...)
- Increase SNR by stacking

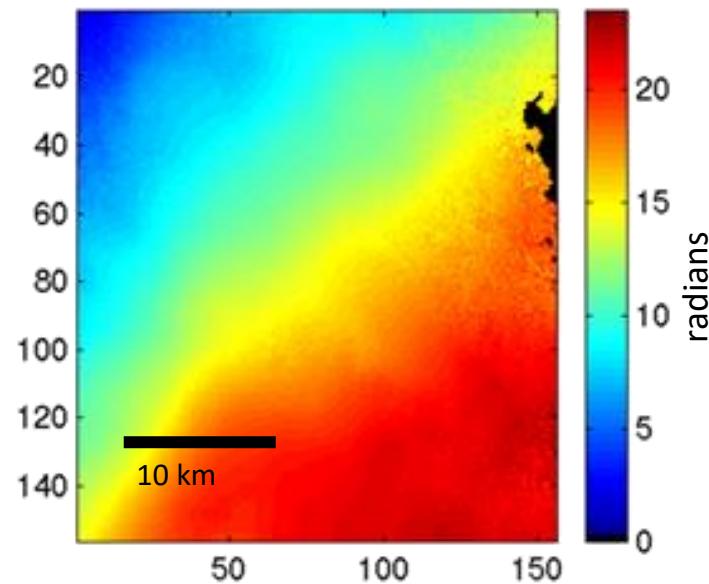


Several studies have **quantified** the tropospheric noise in interferograms (e.g. Hanssen, 2002; Wright et al., 2003; Funning et al., 2005; Lohman and Simons, 2005)

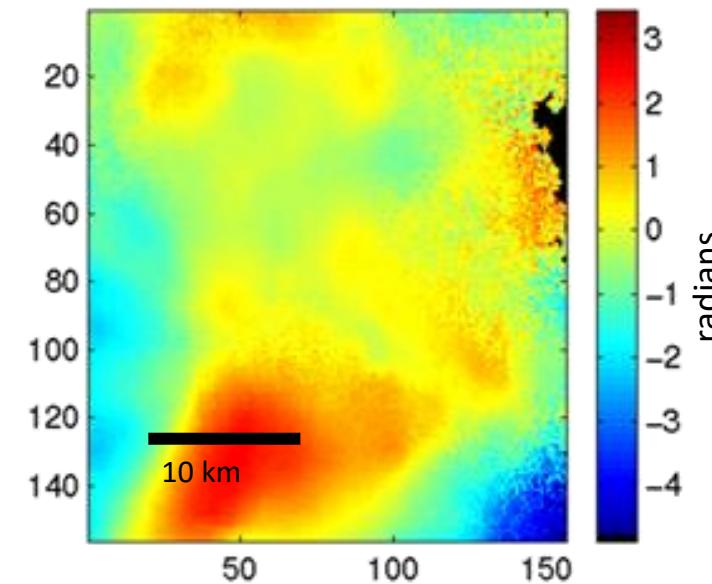
- Remove a best-fitting ramp from data
- Compute autocorrelation (can be achieved in Fourier domain) or semi-variogram
- Calculate the radial average and fit a function; its e-folding wavelength is an estimate of the correlation length scale of the noise; its amplitude an estimate of the variance of the noise



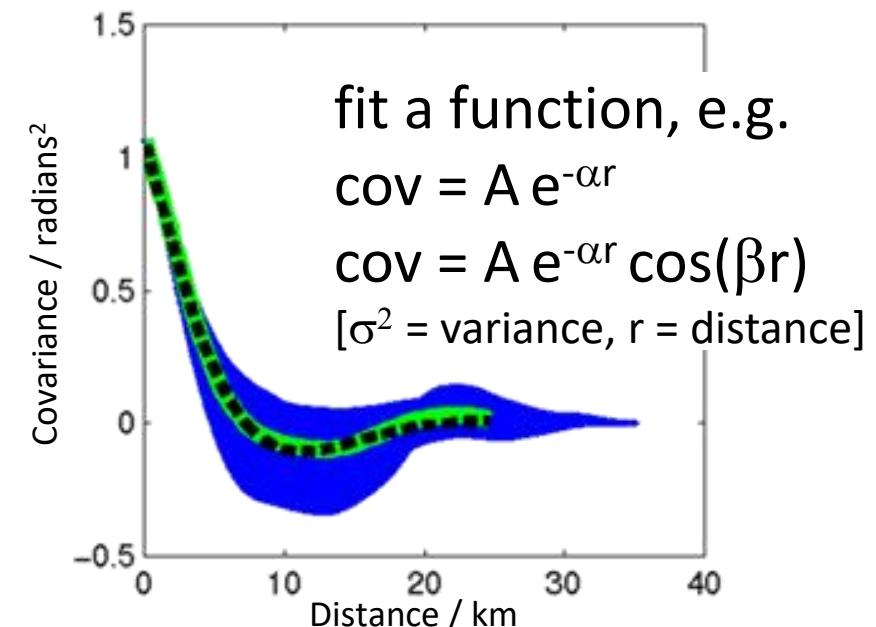
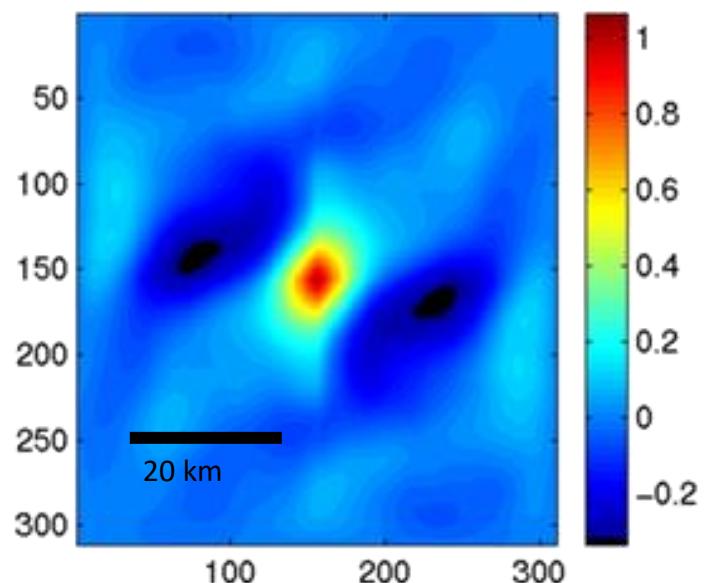
Input interferogram



Flattened interferogram



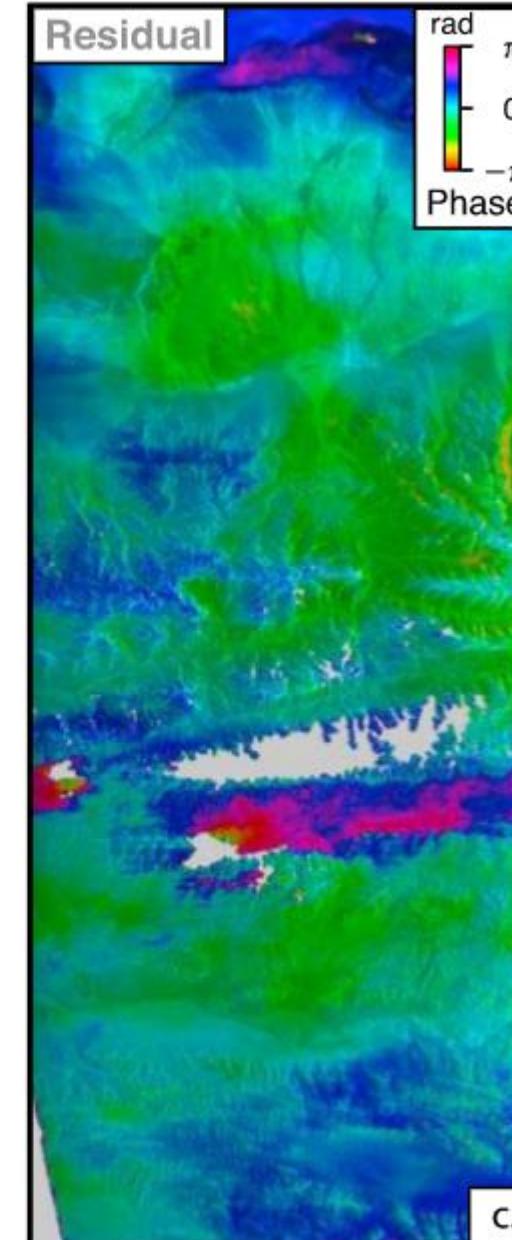
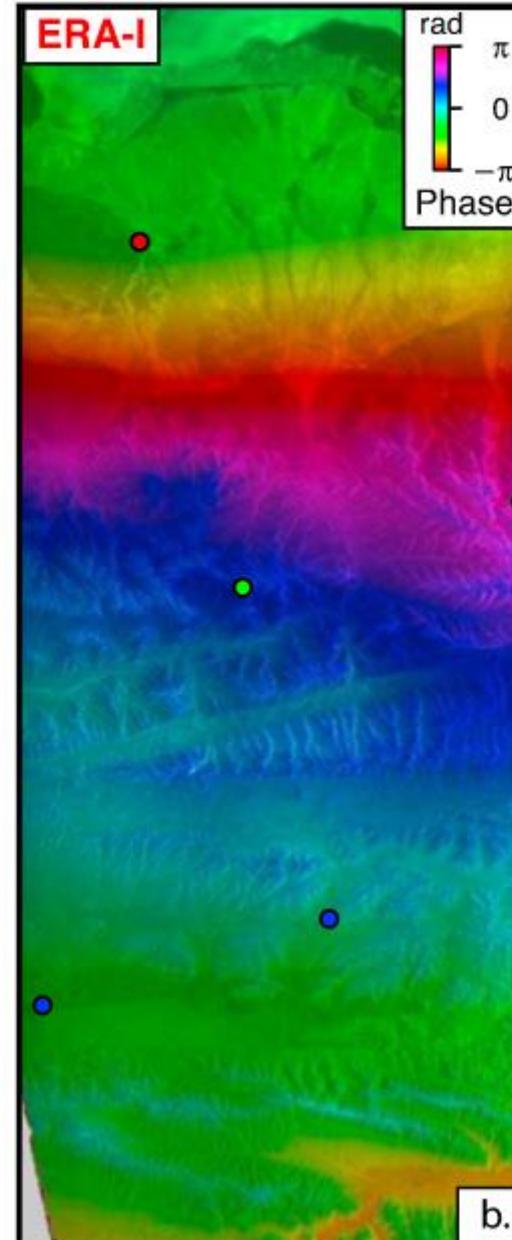
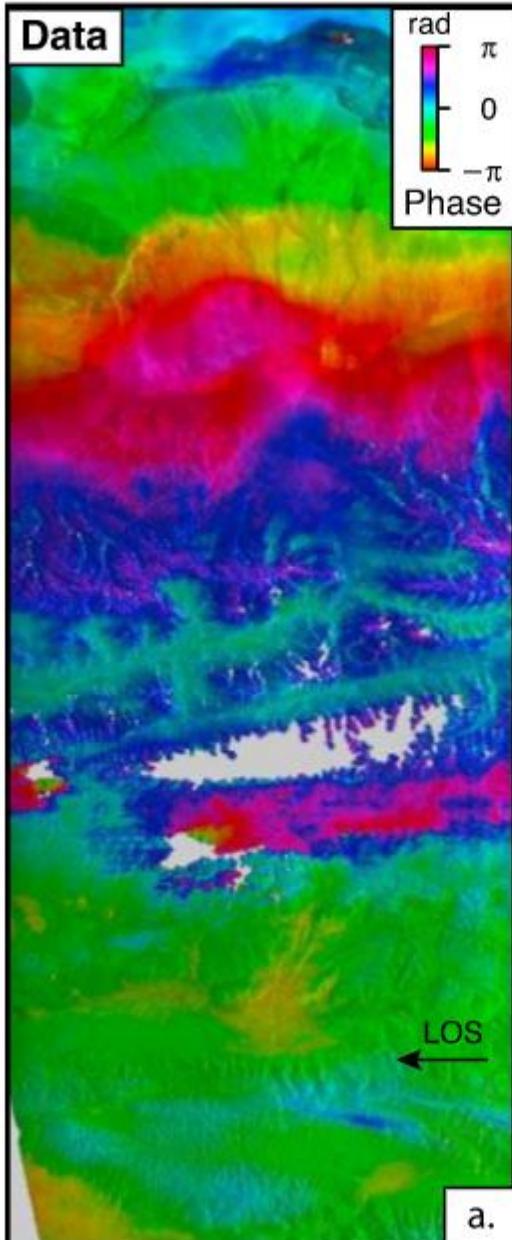
Autocorrelation grid



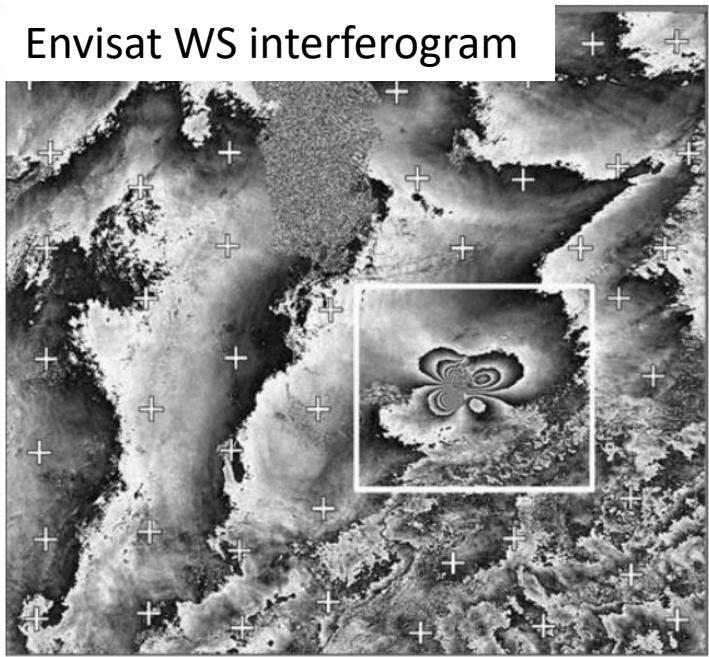
# Modeling the troposphere

The static, layered portion of the troposphere depends on temperature, pressure, humidity and elevation, and can be estimated using weather/atmospheric reanalysis models such as ERA-I/ERA5

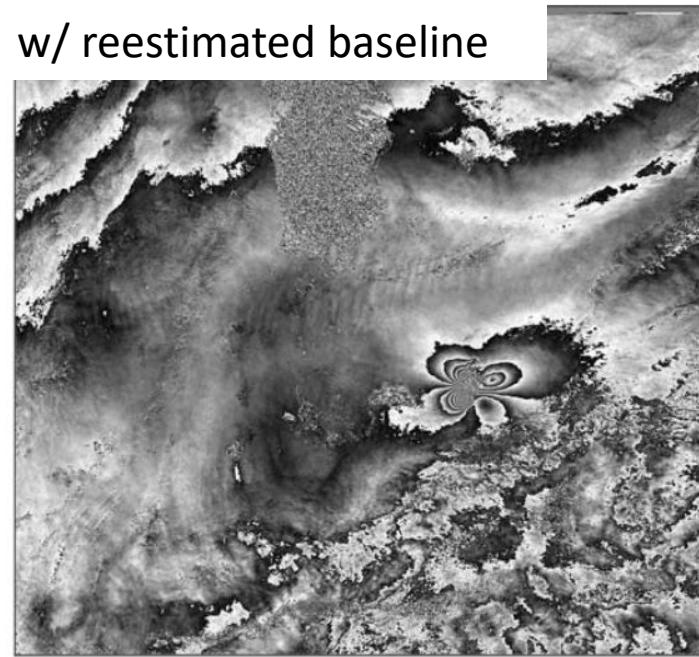
The turbulent portion is a bit more difficult, requiring contemporaneous measurements of the troposphere by other means (e.g. GPS, optical imagery). These corrections are not (yet) routine...



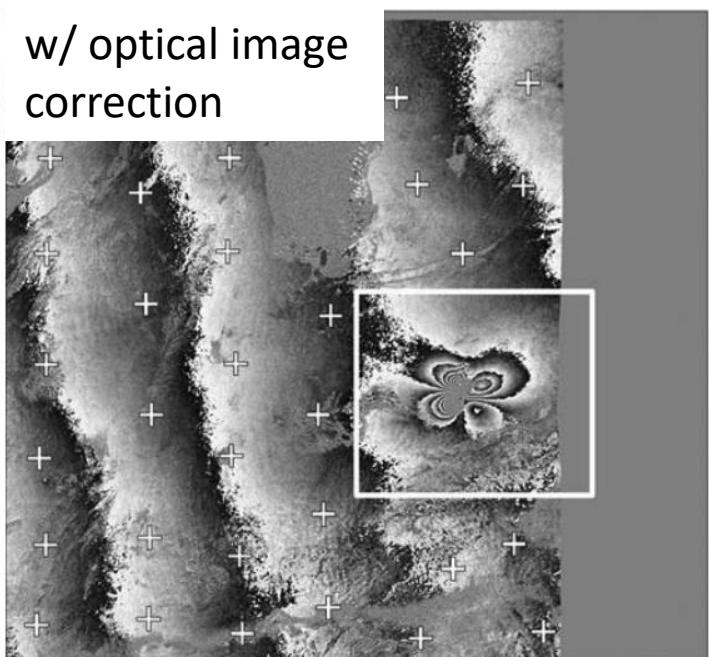
Envisat WS interferogram



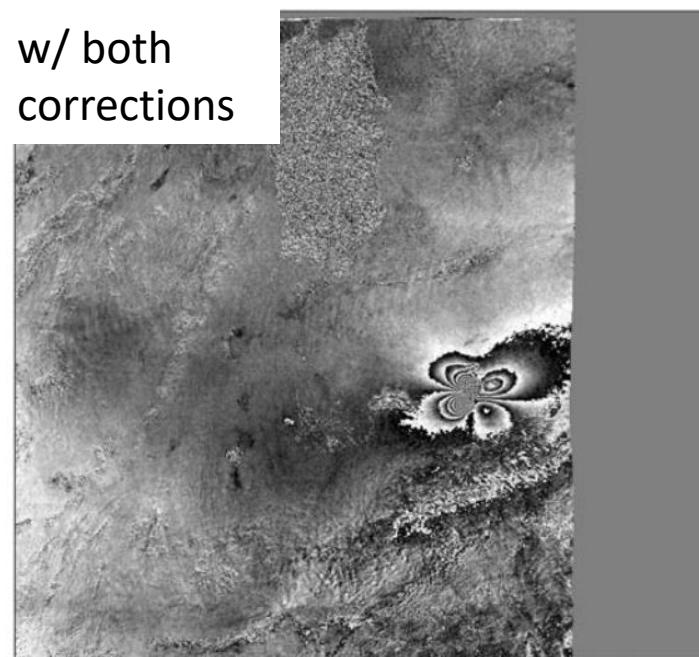
w/ reestimated baseline



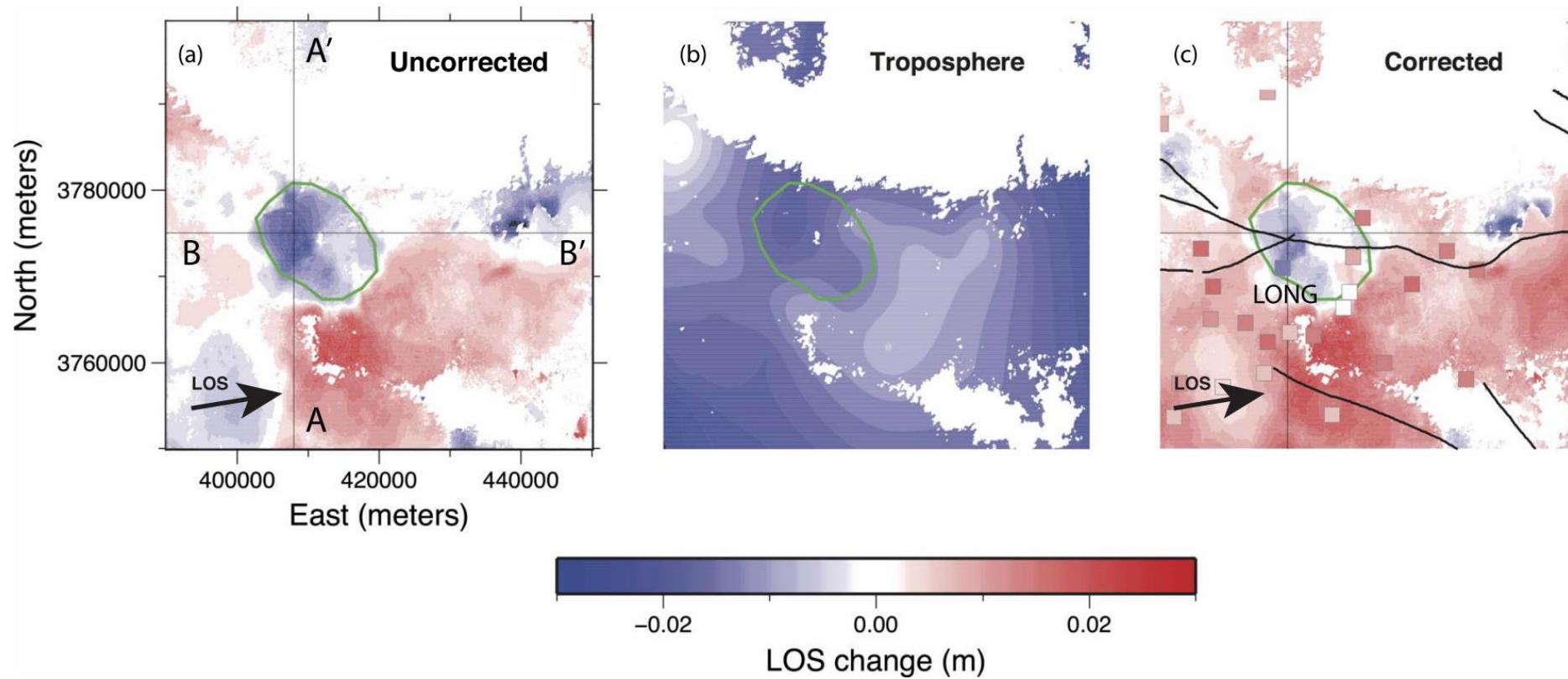
w/ optical image  
correction

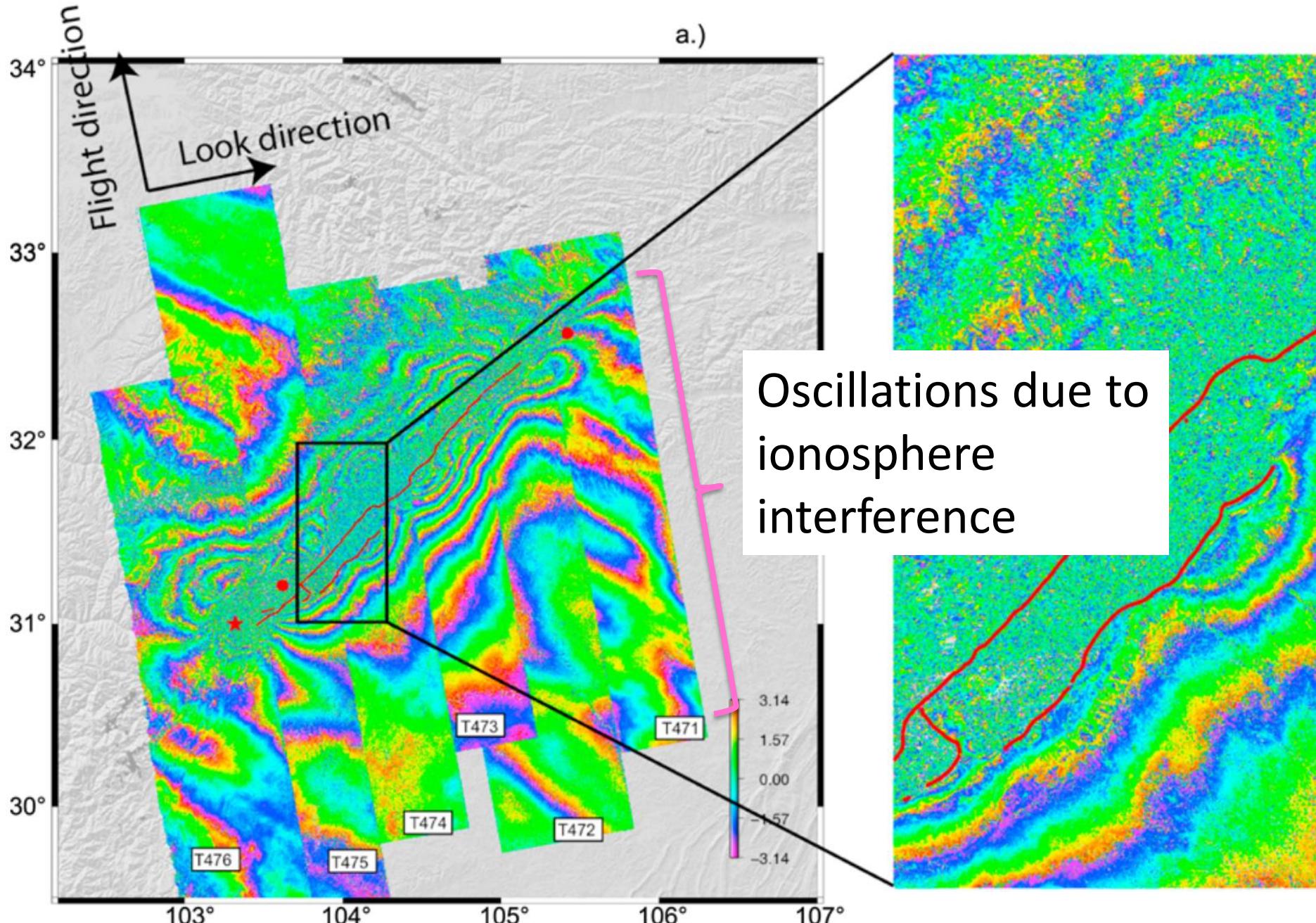


w/ both  
corrections



In regions of dense GPS, the zenith delay correction (a troposphere delay estimate) can be used to correct the troposphere





# Ionosphere distortions

The ionosphere also adds a distortion to radar waves. It is dispersive (frequency-dependent) and affects long wavelengths (e.g. L-band missions such as ALOS and ALOS-2) more than shorter wavelengths

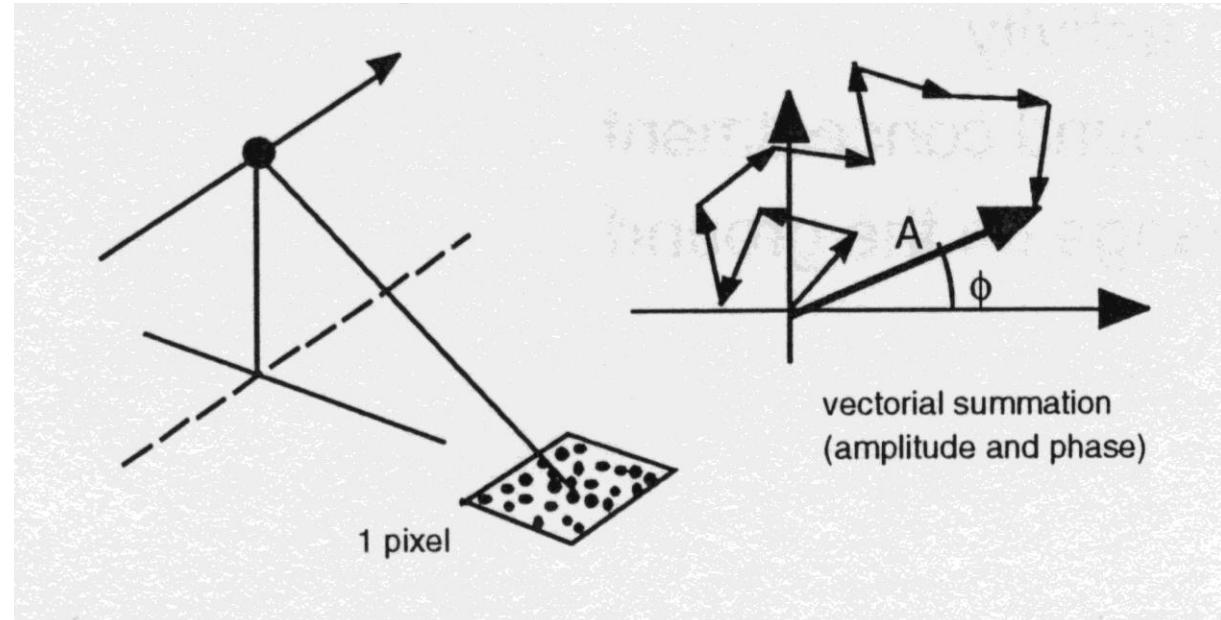
Since it is frequency dependent, this property can potentially be used to correct for it, as it is in GPS – split the bandwidth in processing, and form a higher frequency and lower frequency interferograms

This technique is known as 'enhanced spectral diversity'

# Pixel phase (1)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

- Each pixel on the ground scatters radar in a unique way; there is a phase shift associated with the configuration of objects within it
- If the pixel is undisturbed, this phase shift cancels out in the interferogram; if not, the phase is effectively random



## Pixel phase (2)

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

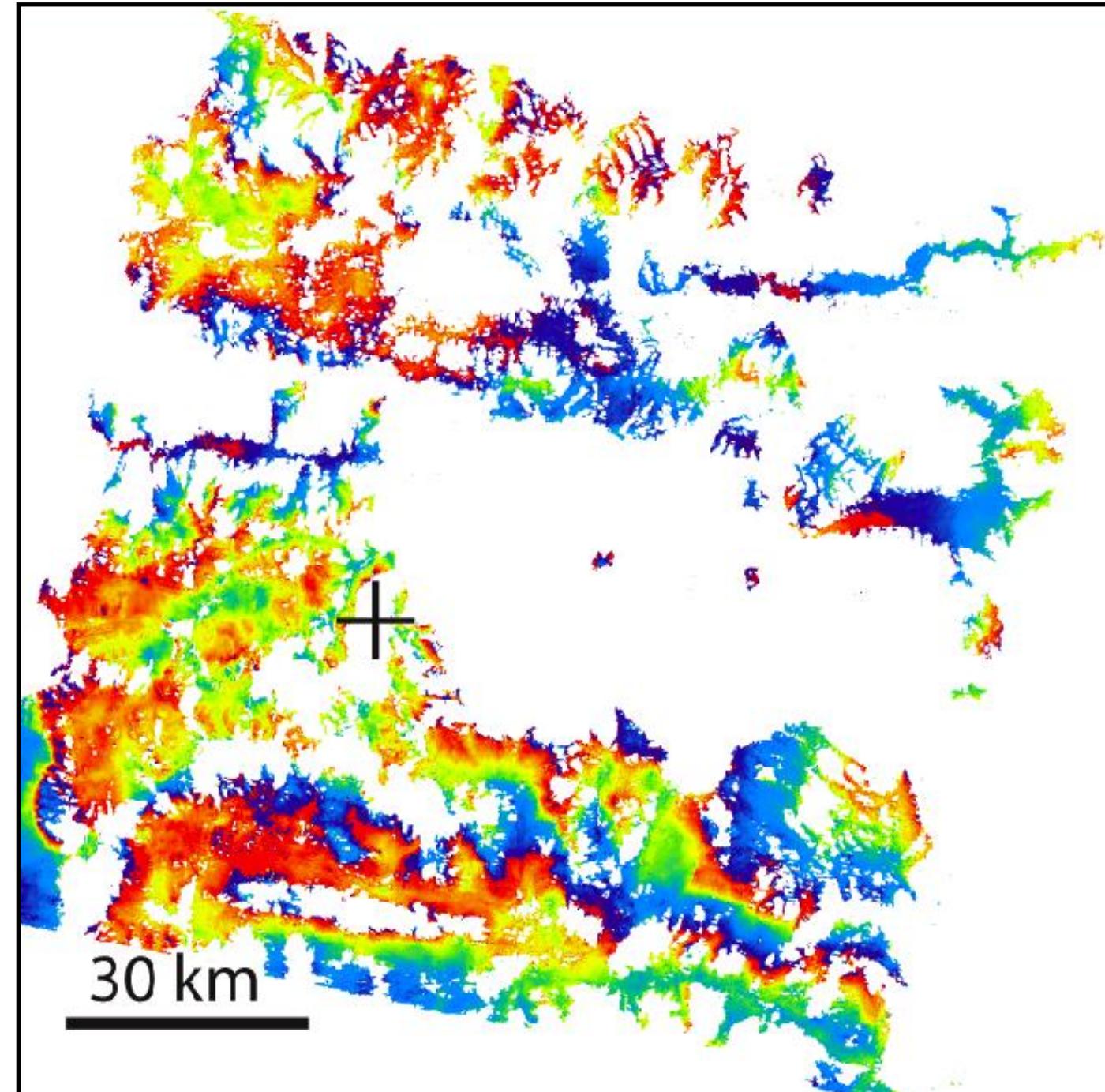
- The effect of this effectively random signal is termed *decorrelation* (or *incoherence*), and the deformation signal cannot be recovered.
- Can use *multilooking* (spatial averaging) to pull out coherent signal from larger effective pixel sizes (at cost of reduced spatial resolution).

# Decorrelation

If the scattering properties of the pixel change between acquisitions, the signal is lost

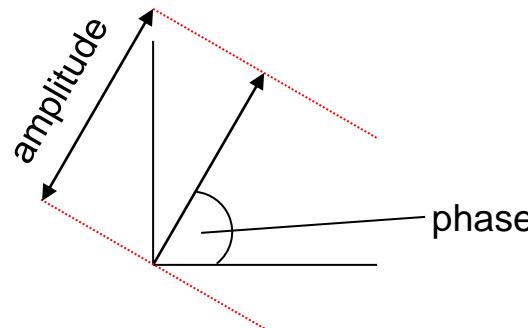
The probability of this occurring increases with time

ERS-1, Ngamring  
County, Tibet, July  
1992–Feb 1993

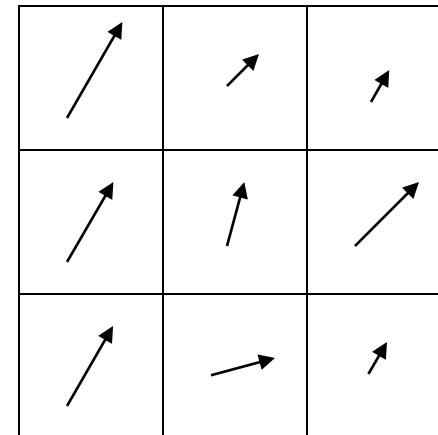


# Interferometric correlation

Interferometric correlation, C, compares the phase of a pixel to its neighbors over a specified window

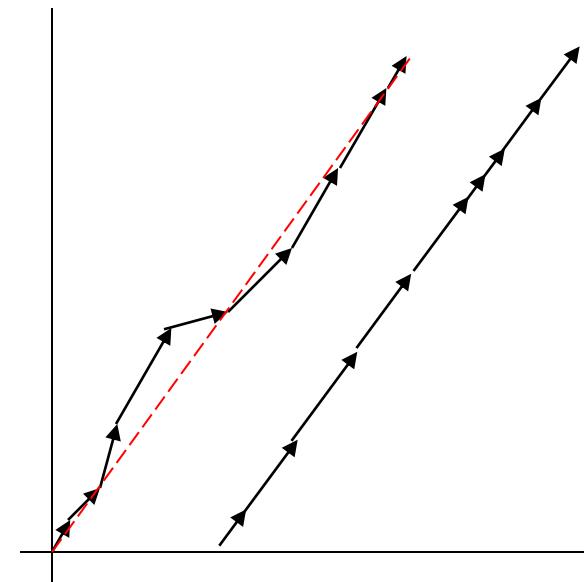


$$C = \frac{\left| \sum_{i=1}^n I_i \right|}{\sqrt{\sum_{i=1}^n |I_i|^2}}$$



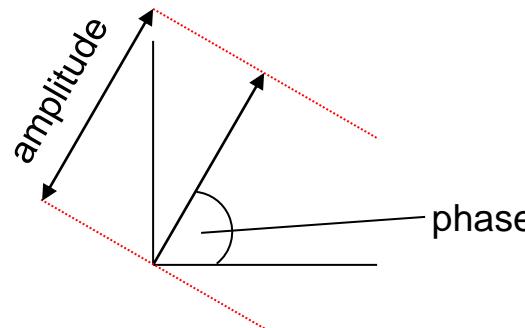
If the phases are similar, the correlation is high, and vice-versa

High correlation

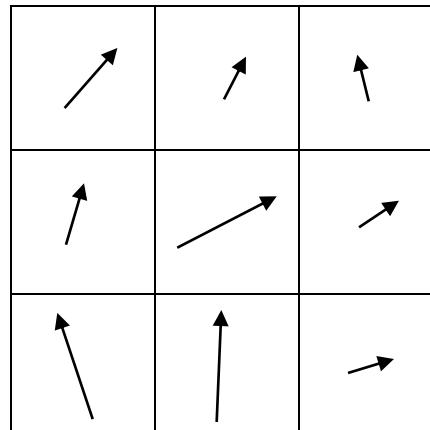


# Interferometric correlation

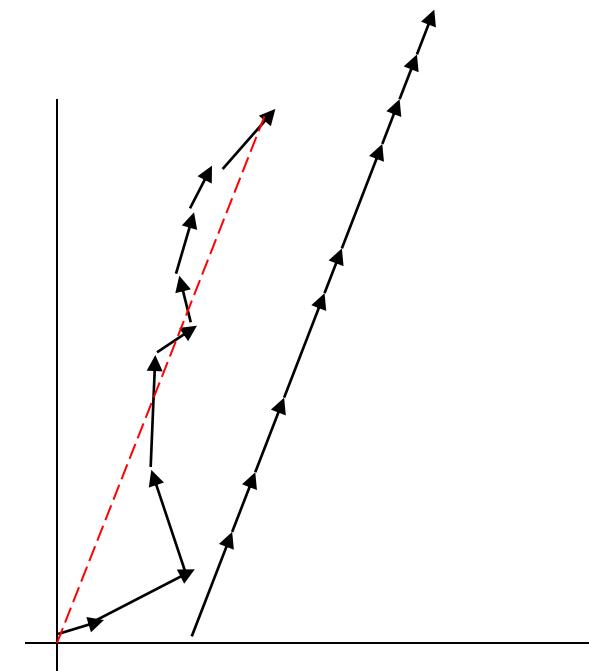
Interferometric correlation, C, compares the phase of a pixel to its neighbors over a specified window



$$C = \frac{\left| \sum_{i=1}^n I_i \right|}{\left| \sum_{i=1}^n |I_i| \right|}$$



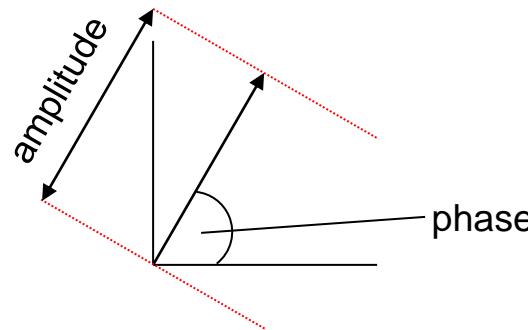
If the phases are similar, the correlation is high, and vice-versa



Lower correlation

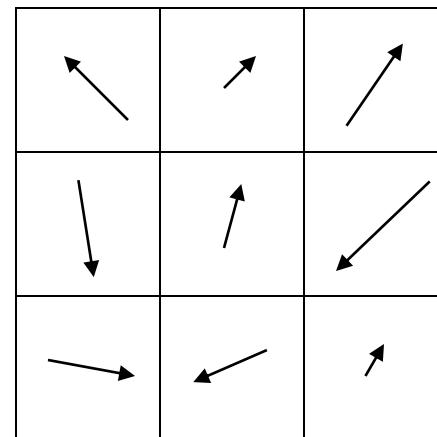
# Interferometric correlation

Interferometric correlation, C, compares the phase of a pixel to its neighbors over a specified window

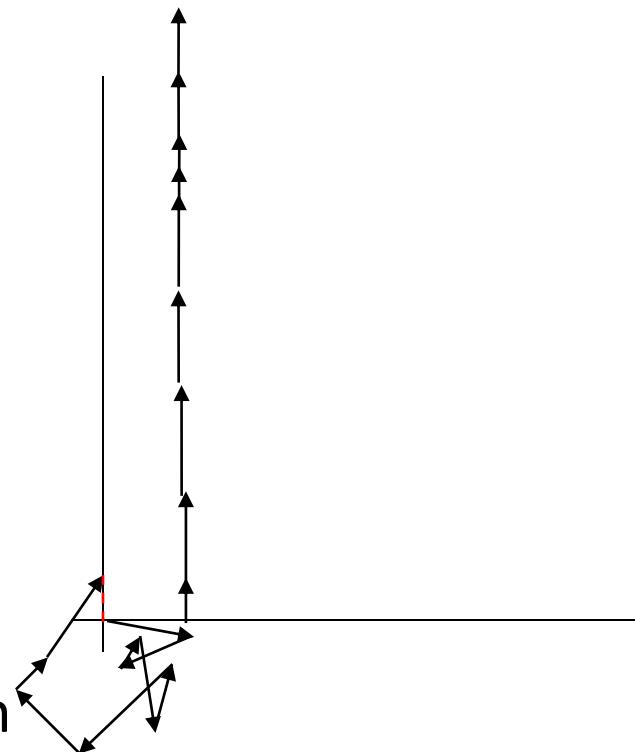


$$C = \frac{\left| \sum_{i=1}^n I_i \right|}{\sum_{i=1}^n |I_i|}$$

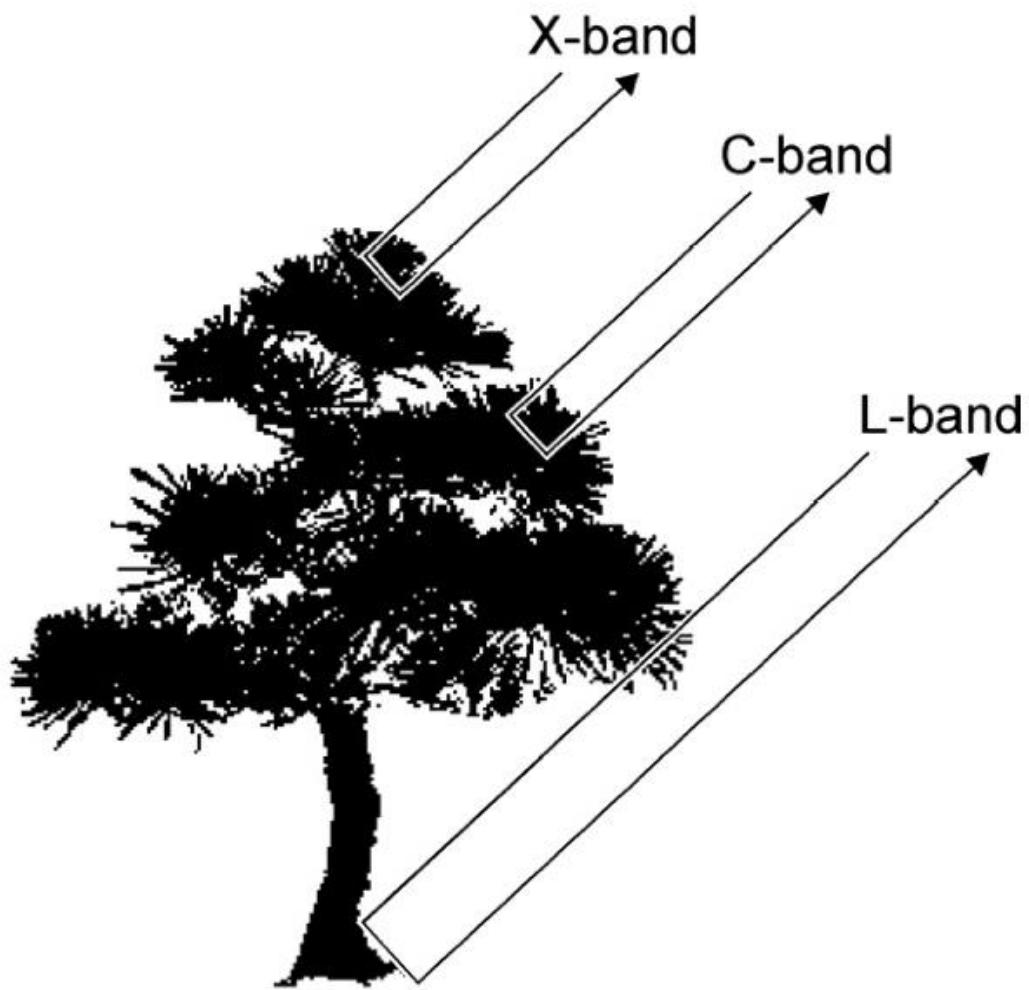
If the phases are similar, the correlation is high, and vice-versa



Low correlation



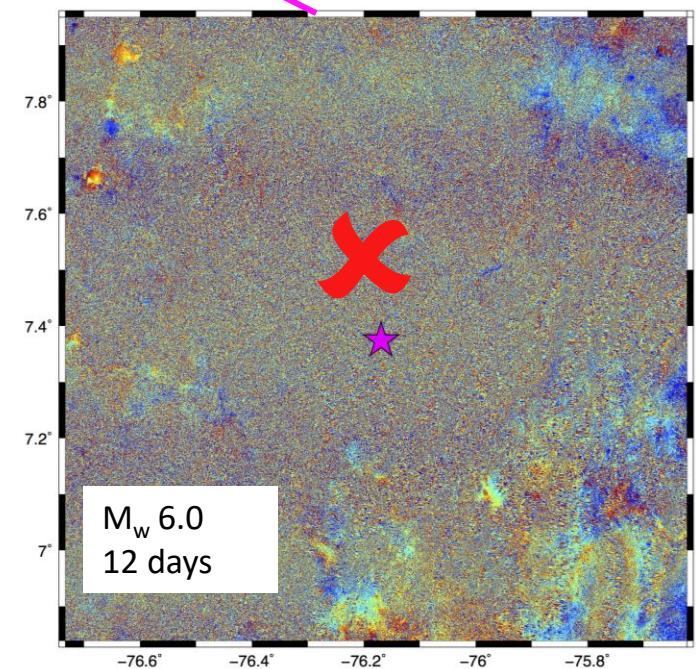
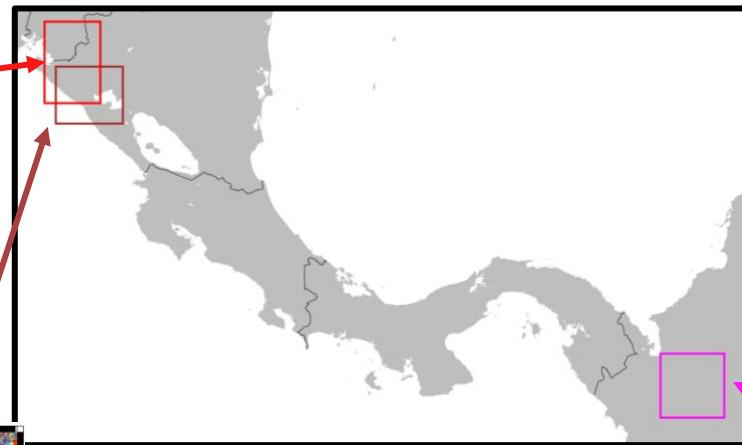
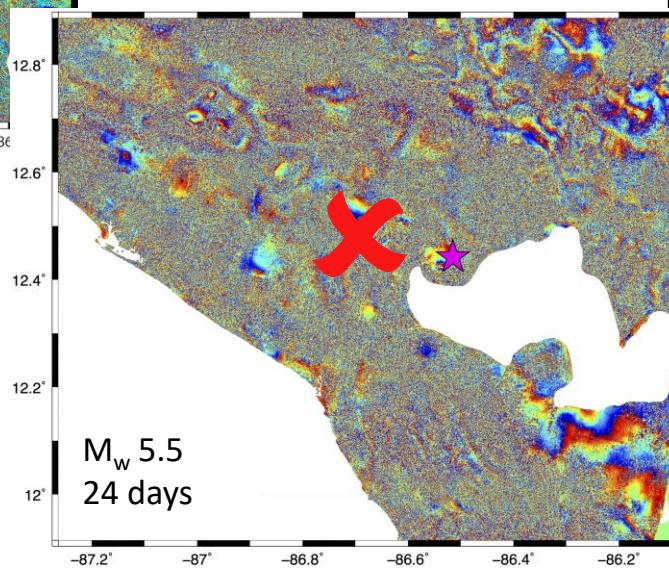
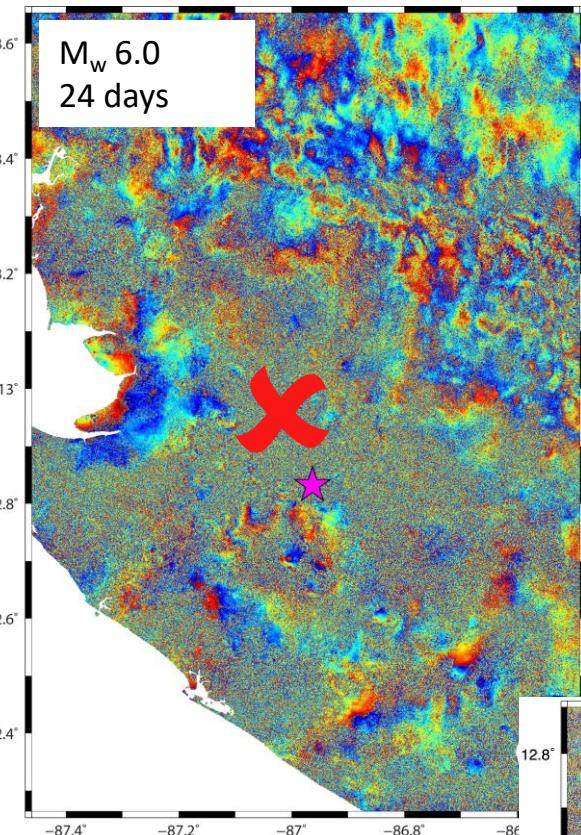
# Volume scattering of radar



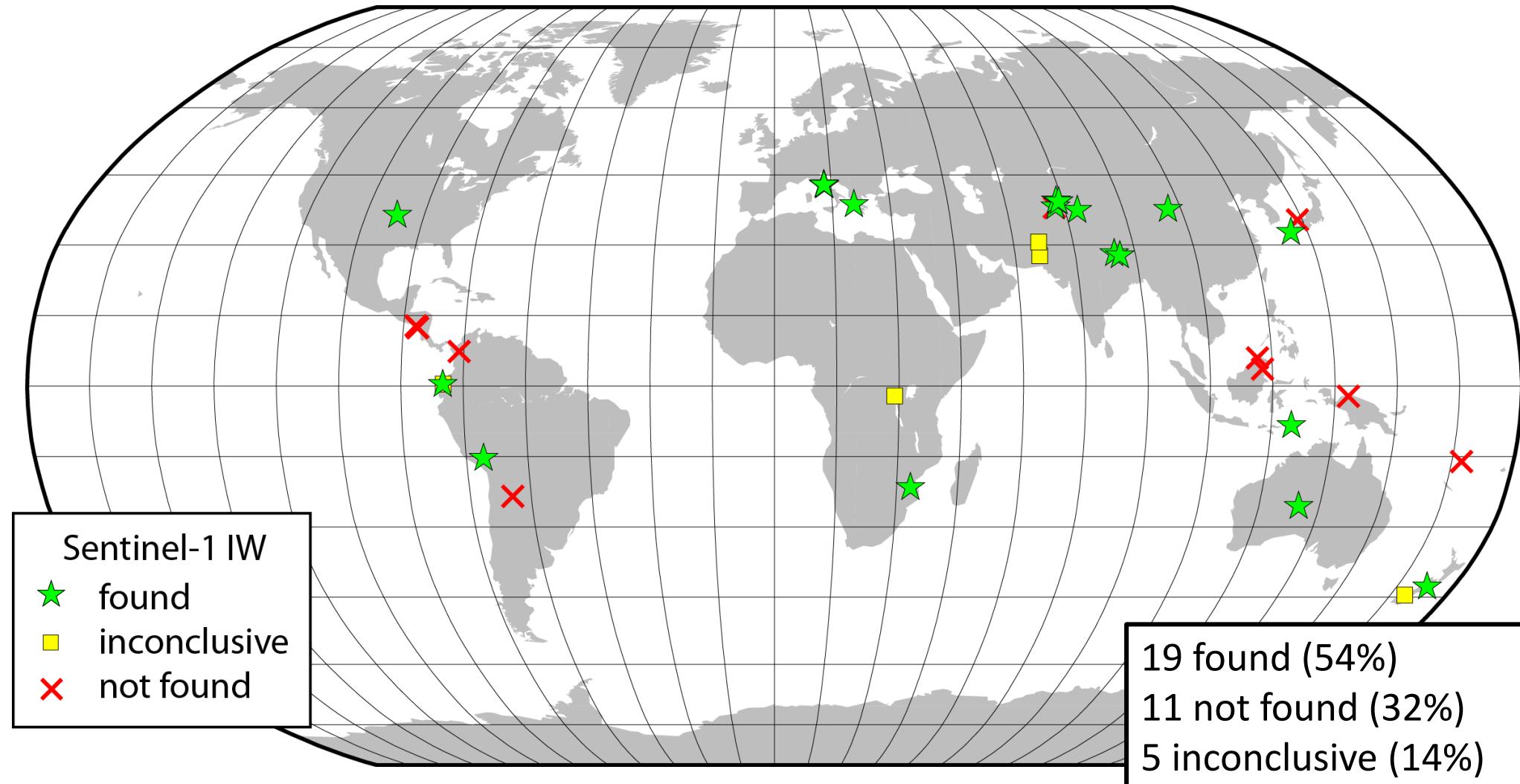
- Vegetation is especially problematic for InSAR
- C-band radar (ERS, Envisat, RADARSAT) scatters off the canopies of trees
- If the tree grows/ branches blow in the wind/ leaves drop, the data will be decorrelated
- Longer wavelengths (e.g. L-band, as on ALOS) penetrate deeper into the tree, and perform better for InSAR

# Central America

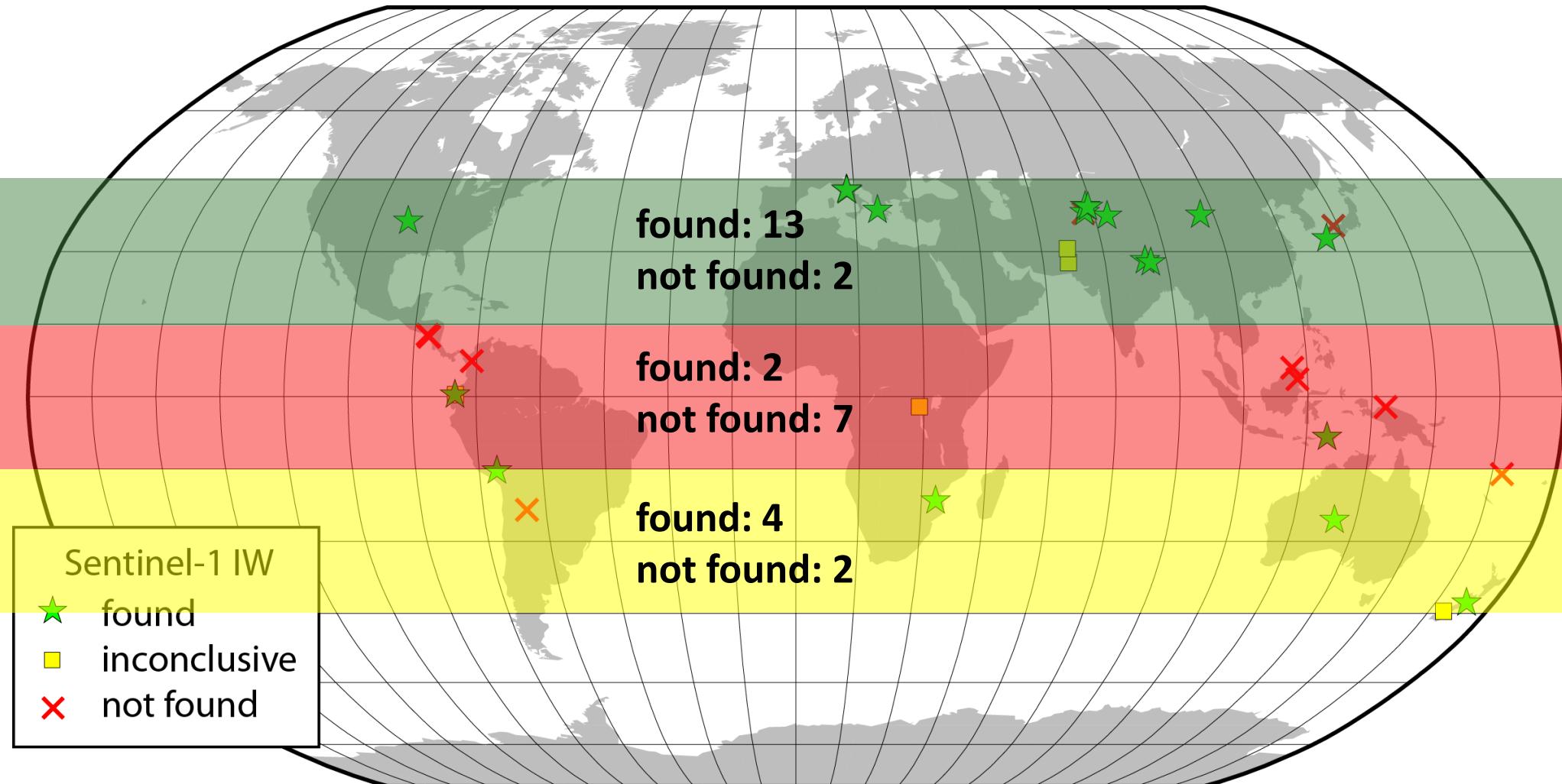
Shortest Sentinel-1 pairs for shallow M>5.5 earthquakes



# 35 M>5.5 earthquakes in 18 months with Sentinel-1

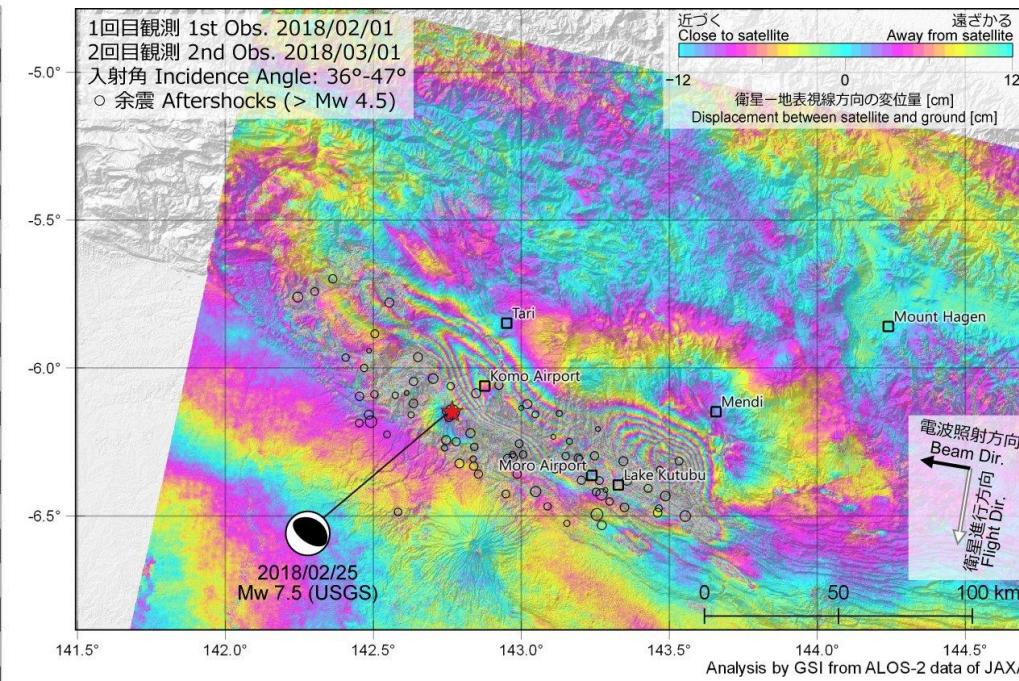
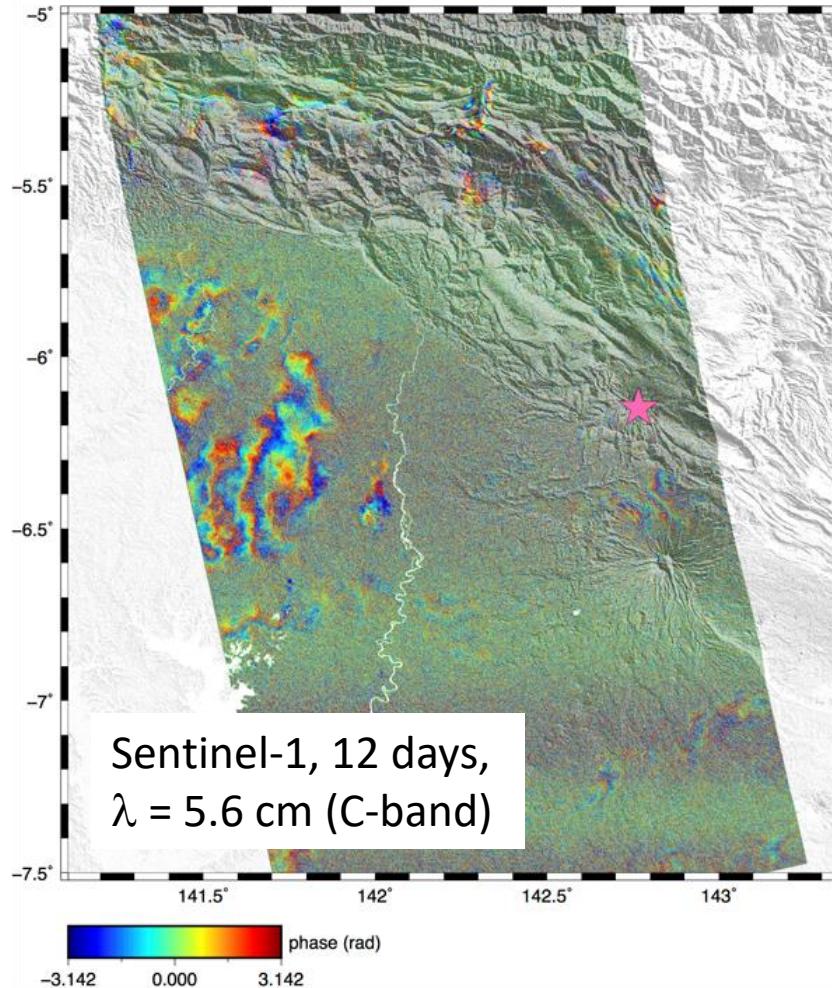


# Most nondetections within 15° of equator – link to tropical climate and vegetation?



# Effect of wavelength on decorrelation

M7.6, Papua New Guinea, February 25th, 2018



ALOS-2, 28 days,  
 $\lambda = 23.8$  cm (L-band)  
[Processed by  
Yu Morishita, GSI, Japan]

# Tips for avoiding decorrelation

- Short interferogram time spans reduce the probability of temporal decorrelation
- Short perpendicular baselines reduce geometric decorrelation
- Longer radar wavelengths are less susceptible to volume decorrelation
- Don't work on forested areas or the tropics?
- Only work in urban areas or the mid-latitudes?



Less of a problem with modern missions



ALOS-2! NISAR!  
(2023?)

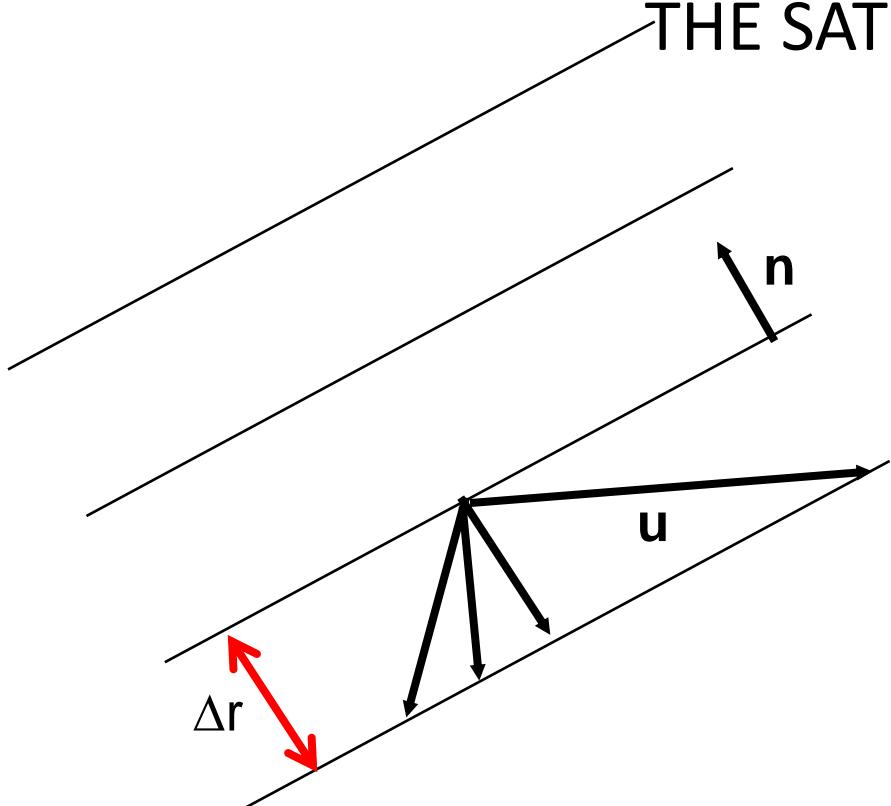


No, but you might need to choose data carefully

# Deformation phase

$$\Delta\phi_{\text{int}} = \Delta\phi_{\text{orb}} + \Delta\phi_{\text{topo}} + \Delta\phi_{\text{atm}} + \Delta\phi_{\text{pixel}} + \Delta\phi_{\text{def}}$$

InSAR ONLY MEASURES THE COMPONENT OF SURFACE DEFORMATION IN  
THE SATELLITE'S LINE OF SIGHT (LOS)



$$\Delta\mathbf{r} = -\mathbf{n} \cdot \mathbf{u}$$

where  $\mathbf{n}$  is a unit vector pointing  
from the ground to the satellite

$$\Delta\phi_{\text{def}} = (4\pi / \lambda) \Delta\mathbf{r}$$

i.e. 1 fringe = 28.3 mm LOS  
deformation for ERS