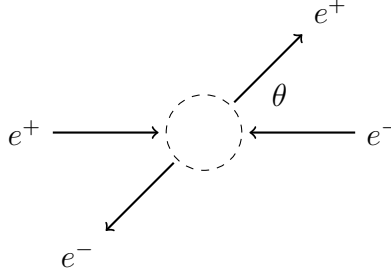
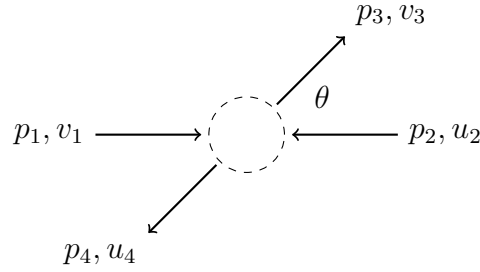


Bhabha scattering is the result of interactions between positrons and electrons. The following diagram represents a collider experiment with collinear electron and positron beams.



Here is the same diagram with momentum and spinor labels.



In a typical collider experiment the momentum vectors are

$$\begin{array}{cccc}
 p_1 = \begin{pmatrix} E \\ 0 \\ 0 \\ p \end{pmatrix} & p_2 = \begin{pmatrix} E \\ 0 \\ 0 \\ -p \end{pmatrix} & p_3 = \begin{pmatrix} E \\ p \sin \theta \cos \phi \\ p \sin \theta \sin \phi \\ p \cos \theta \end{pmatrix} & p_4 = \begin{pmatrix} E \\ -p \sin \theta \cos \phi \\ -p \sin \theta \sin \phi \\ -p \cos \theta \end{pmatrix} \\
 \text{inbound positron} & \text{inbound electron} & \text{outbound positron} & \text{outbound electron}
 \end{array}$$

Symbol  $p$  is incident momentum,  $E$  is total energy  $E = \sqrt{p^2 + m^2}$ , and  $m$  is electron mass. Polar angle  $\theta$  is the observed scattering angle. Azimuth angle  $\phi$  cancels out in scattering calculations.

The spinors are

$$\begin{array}{cccc}
 v_{11} = \begin{pmatrix} p \\ 0 \\ E + m \\ 0 \end{pmatrix} & v_{12} = \begin{pmatrix} 0 \\ -p \\ 0 \\ E + m \end{pmatrix} & u_{21} = \begin{pmatrix} E + m \\ 0 \\ -p \\ 0 \end{pmatrix} & u_{22} = \begin{pmatrix} 0 \\ E + m \\ 0 \\ p \end{pmatrix} \\
 \text{inbound positron, spin up} & \text{inbound positron, spin down} & \text{inbound electron, spin up} & \text{inbound electron, spin down} \\
 v_{31} = \begin{pmatrix} p_3^z \\ p_3^x + ip_3^y \\ E + m \\ 0 \end{pmatrix} & v_{32} = \begin{pmatrix} p_3^x - ip_3^y \\ -p_3^z \\ 0 \\ E + m \end{pmatrix} & u_{41} = \begin{pmatrix} E + m \\ 0 \\ p_4^z \\ p_4^x + ip_4^y \end{pmatrix} & u_{42} = \begin{pmatrix} 0 \\ E + m \\ p_4^x - ip_4^y \\ -p_4^z \end{pmatrix} \\
 \text{outbound positron, spin up} & \text{outbound positron, spin down} & \text{outbound electron, spin up} & \text{outbound electron, spin down}
 \end{array}$$

The spinors shown above are not individually normalized. Instead, a combined spinor normalization constant  $N = (E + m)^4$  will be used.

The following formula computes a probability density  $|\mathcal{M}_{abcd}|^2$  for Bhabha scattering where  $abcd$  are spin states.

$$|\mathcal{M}_{abcd}|^2 = \frac{e^4}{N} \left| -\frac{1}{t} (\bar{v}_{1a} \gamma^\mu v_{3c}) (\bar{u}_{4d} \gamma_\mu u_{2b}) + \frac{1}{s} (\bar{v}_{1a} \gamma^\nu u_{2b}) (\bar{u}_{4d} \gamma_\nu v_{3c}) \right|^2$$

Symbol  $e$  is electron charge. Symbols  $s$  and  $t$  are Mandelstam variables

$$\begin{aligned} s &= (p_1 + p_2)^2 \\ t &= (p_1 - p_3)^2 \end{aligned}$$

Let

$$a_1 = (\bar{v}_{1a} \gamma^\mu v_{3c}) (\bar{u}_{4d} \gamma_\mu u_{2b}) \quad a_2 = (\bar{v}_{1a} \gamma^\nu u_{2b}) (\bar{u}_{4d} \gamma_\nu v_{3c})$$

Then

$$\begin{aligned} |\mathcal{M}_{abcd}|^2 &= \frac{e^4}{N} \left| -\frac{a_1}{t} + \frac{a_2}{s} \right|^2 \\ &= \frac{e^4}{N} \left( -\frac{a_1}{t} + \frac{a_2}{s} \right) \left( -\frac{a_1}{t} + \frac{a_2}{s} \right)^* \\ &= \frac{e^4}{N} \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right) \end{aligned}$$

The expected probability density  $\langle |\mathcal{M}|^2 \rangle$  is computed by summing  $|\mathcal{M}_{abcd}|^2$  over all spin states and then dividing by the number of inbound states. There are four inbound states.

$$\begin{aligned} \langle |\mathcal{M}|^2 \rangle &= \frac{1}{4} \sum_{a=1}^2 \sum_{b=1}^2 \sum_{c=1}^2 \sum_{d=1}^2 |\mathcal{M}_{abcd}|^2 \\ &= \frac{e^4}{4N} \sum_{a=1}^2 \sum_{b=1}^2 \sum_{c=1}^2 \sum_{d=1}^2 \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right) \end{aligned}$$

Use the Casimir trick to replace sums over spins with matrix products.

$$\begin{aligned} f_{11} &= \frac{1}{N} \sum_{\text{spins}} a_1 a_1^* = \text{Tr} \left( (\not{p}_1 - m) \gamma^\mu (\not{p}_3 - m) \gamma_\mu \right) \text{Tr} \left( (\not{p}_4 + m) \gamma_\mu (\not{p}_2 + m) \gamma_\nu \right) \\ f_{12} &= \frac{1}{N} \sum_{\text{spins}} a_1 a_2^* = \text{Tr} \left( (\not{p}_1 - m) \gamma^\mu (\not{p}_2 + m) \gamma^\nu (\not{p}_4 + m) \gamma_\mu (\not{p}_3 - m) \gamma_\nu \right) \\ f_{22} &= \frac{1}{N} \sum_{\text{spins}} a_2 a_2^* = \text{Tr} \left( (\not{p}_1 - m) \gamma^\mu (\not{p}_2 + m) \gamma^\nu \right) \text{Tr} \left( (\not{p}_4 + m) \gamma_\mu (\not{p}_3 - m) \gamma_\nu \right) \end{aligned}$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left( \frac{f_{11}}{t^2} - \frac{f_{12}}{st} - \frac{f_{12}^*}{st} + \frac{f_{22}}{s^2} \right)$$

Run “bhabha-scattering-1.txt” to verify the Casimir trick.

The following formulas are equivalent to the Casimir trick. (Recall that  $a \cdot b = a^\mu g_{\mu\nu} b^\nu$ )

$$\begin{aligned} f_{11} &= 32(p_1 \cdot p_2)^2 + 32(p_1 \cdot p_4)^2 - 64m^2(p_1 \cdot p_3) + 64m^4 \\ f_{12} &= -32(p_1 \cdot p_4)^2 - 32m^2(p_1 \cdot p_2) + 32m^2(p_1 \cdot p_3) - 32m^2(p_1 \cdot p_4) - 32m^4 \\ f_{22} &= 32(p_1 \cdot p_3)^2 + 32(p_1 \cdot p_4)^2 + 64m^2(p_1 \cdot p_2) + 64m^4 \end{aligned}$$

Using Mandelstam variables

$$\begin{aligned}s &= (p_1 + p_2)^2 \\ t &= (p_1 - p_3)^2 \\ u &= (p_1 - p_4)^2\end{aligned}$$

the formulas are

$$\begin{aligned}f_{11} &= 8s^2 + 8u^2 - 64sm^2 - 64um^2 + 192m^4 \\ f_{12} &= -8u^2 + 64um^2 - 96m^4 \\ f_{22} &= 8t^2 + 8u^2 - 64tm^2 - 64um^2 + 192m^4\end{aligned}$$

## High energy approximation

When  $E \gg m$  a useful approximation is to set  $m = 0$  and obtain

$$\begin{aligned}f_{11} &= 8s^2 + 8u^2 \\ f_{12} &= -8u^2 \\ f_{22} &= 8t^2 + 8u^2\end{aligned}$$

Hence

$$\begin{aligned}\langle |\mathcal{M}|^2 \rangle &= \frac{e^4}{4} \left( \frac{8s^2 + 8u^2}{t^2} - \frac{-8u^2}{st} - \frac{-8u^2}{st} + \frac{8t^2 + 8u^2}{s^2} \right) \\ &= 2e^4 \left( \frac{s^2 + u^2}{t^2} + \frac{2u^2}{st} + \frac{t^2 + u^2}{s^2} \right)\end{aligned}$$

In the following steps we require a common denominator.

$$\langle |\mathcal{M}|^2 \rangle = 2e^4 \left( \frac{s^2(s^2 + u^2) + 2stu^2 + t^2(t^2 + u^2)}{s^2t^2} \right)$$

For  $m = 0$  the Mandelstam variables are

$$\begin{aligned}s &= 4E^2 \\ t &= 2E^2(\cos \theta - 1) \\ u &= -2E^2(\cos \theta + 1)\end{aligned}$$

Hence

$$\begin{aligned}\langle |\mathcal{M}|^2 \rangle &= 2e^4 \left( \frac{32E^8 \cos^4 \theta + 192E^8 \cos^2 \theta + 288E^8}{64E^8(\cos \theta - 1)^2} \right) \\ &= e^4 \left( \frac{\cos^4 \theta + 6 \cos^2 \theta + 9}{(\cos \theta - 1)^2} \right) \\ &= e^4 \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2\end{aligned}$$

Run “bhabha-scattering-2.txt” to verify.

## Cross section

The differential cross section is

$$\frac{d\sigma}{d\Omega} = \frac{\langle |\mathcal{M}|^2 \rangle}{64\pi^2 s} = \frac{e^4}{64\pi^2 s} \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2, \quad s \gg m$$

Substituting  $e^4 = 16\pi^2 \alpha^2$  yields

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2}{4s} \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2$$

We can integrate  $d\sigma$  to obtain a cumulative distribution function. Recall that

$$d\Omega = \sin \theta \, d\theta \, d\phi$$

Hence

$$d\sigma = \frac{\alpha^2}{4s} \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2 \sin \theta \, d\theta \, d\phi$$

Let  $I(\theta)$  be the following integral of  $d\sigma$ .

$$\begin{aligned} I(\theta) &= \frac{4s}{2\pi\alpha^2} \int_0^{2\pi} \int d\sigma \\ &= \int \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2 \sin \theta \, d\theta, \quad a \leq \theta \leq \pi \end{aligned}$$

Angular support is limited to an arbitrary  $a > 0$  because  $I(0)$  is undefined. Assume that  $I(\theta) - I(a)$  is computable given  $\theta$  by either symbolic or numerical integration.

Let  $C$  be the normalization constant

$$C = I(\pi) - I(a)$$

Then the cumulative distribution function  $F(\theta)$  is

$$F(\theta) = \frac{I(\theta) - I(a)}{C}, \quad a \leq \theta \leq \pi$$

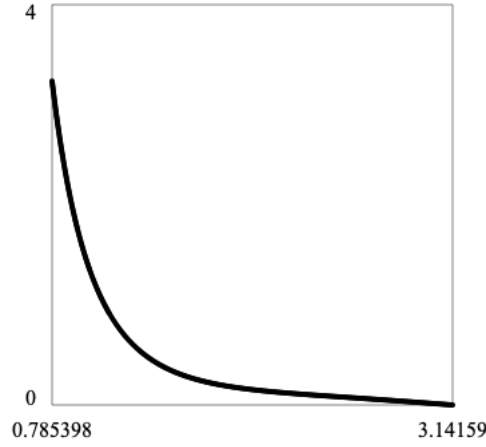
The probability of observing scattering events in the interval  $\theta_1$  to  $\theta_2$  can now be computed.

$$P(\theta_1 \leq \theta \leq \theta_2) = F(\theta_2) - F(\theta_1)$$

Probability density function  $f(\theta)$  is the derivative of  $F(\theta)$ .

$$f(\theta) = \frac{dF(\theta)}{d\theta} = \frac{1}{C} \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2 \sin \theta$$

Run “bhabha-scattering-3.txt” to draw  $f(\theta)$  for  $a = \pi/4 = 45^\circ$ .



Probability distribution for  $45^\circ$  bins ( $a = 45^\circ$ ).

$\theta_1$	$\theta_2$	$P(\theta_1 \leq \theta \leq \theta_2)$
$0^\circ$	$45^\circ$	—
$45^\circ$	$90^\circ$	0.83
$90^\circ$	$135^\circ$	0.13
$135^\circ$	$180^\circ$	0.04

## Data from SLAC SPEAR experiment

The following Bhabha scattering data is adapted from SLAC-PUB-1501.

Bin	$x_k, x_{k+1}$	$y$
(Smallest $\theta$ ) $k = 1$	0.6, 0.5	4432
2	0.5, 0.4	2841
3	0.4, 0.3	2045
4	0.3, 0.2	1420
5	0.2, 0.1	1136
6	0.1, 0.0	852
7	0.0, -0.1	656
8	-0.1, -0.2	625
9	-0.2, -0.3	511
10	-0.3, -0.4	455
11	-0.4, -0.5	402
(Largest $\theta$ ) 12	-0.5, -0.6	398

Data column  $x$  is  $\cos \theta$  of scattering angle  $\theta$ , data column  $y$  is number of scattering events.

To compute predicted values  $\hat{y}$ , start by integrating the probability density function over each bin.

$$I_k = \int_{\arccos(x_k)}^{\arccos(x_{k+1})} \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2 \sin \theta d\theta$$

Normalize  $I_k$  to obtain scattering probability  $P_k$ .

$$P_k = \frac{I_k}{N}, \quad N = \sum_{j=1}^{12} I_j$$

Multiply  $P_k$  by total counts to obtain  $\hat{y}_k$ .

$$\hat{y}_k = P_k T, \quad T = \sum_{j=1}^{12} y_j = 15773$$

Bin	$x_k, x_{k+1}$	$y$	$\hat{y}$
1	0.6, 0.5	4432	4598
2	0.5, 0.4	2841	2880
3	0.4, 0.3	2045	1955
4	0.3, 0.2	1420	1410
5	0.2, 0.1	1136	1068
6	0.1, 0.0	852	843
7	0.0, -0.1	656	689
8	-0.1, -0.2	625	582
9	-0.2, -0.3	511	505
10	-0.3, -0.4	455	450
11	-0.4, -0.5	402	411
12	-0.5, -0.6	398	382

The coefficient of determination  $R^2$  measures how well predicted values fit the real data.

$$R^2 = 1 - \frac{\sum (y - \hat{y})^2}{\sum (y - \bar{y})^2} = 0.997$$

The result indicates that the model  $d\sigma$  explains 99.7% of the variance in the data.

Run “bhabha-scattering-4.txt” to verify.

## Data from DESY PETRA experiment

See [www.hepdata.net/record/ins191231](http://www.hepdata.net/record/ins191231), Table 3, 14.0 GeV.

$x$	$y$
-0.73	0.10115
-0.6495	0.12235
-0.5495	0.11258
-0.4494	0.09968
-0.3493	0.14749
-0.2491	0.14017
-0.149	0.1819
-0.0488	0.22964
0.0514	0.25312
0.1516	0.30998
0.252	0.40898
0.3524	0.62695
0.4529	0.91803
0.5537	1.51743
0.6548	2.56714
0.7323	4.30279

Data  $x$  and  $y$  have the following relationship with the cross section model.

$$x = \cos \theta \quad y = \frac{d\sigma}{d\Omega}$$

The cross section formula for Bhabha scattering is

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2}{4s} \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2$$

To compute predicted values  $\hat{y}$ , substitute  $x$  for  $\cos \theta$ , multiply by  $(\hbar c)^2$  to convert to SI, and multiply by  $10^{37}$  to convert square meters to nanobarns.

$$\hat{y} = \frac{\alpha^2}{4s} \left( \frac{x^2 + 3}{x - 1} \right)^2 \times (\hbar c)^2 \times 10^{37}$$

The following table shows  $\hat{y}$  for  $s = (14.0 \text{ GeV})^2$ .

$x$	$y$	$\hat{y}$
-0.73	0.10115	0.110296
-0.6495	0.12235	0.113816
-0.5495	0.11258	0.120101
-0.4494	0.09968	0.129075
-0.3493	0.14749	0.141592
-0.2491	0.14017	0.158934
-0.149	0.1819	0.182976
-0.0488	0.22964	0.216737
0.0514	0.25312	0.264989
0.1516	0.30998	0.335782
0.252	0.40898	0.44363
0.3524	0.62695	0.615528
0.4529	0.91803	0.9077
0.5537	1.51743	1.45175
0.6548	2.56714	2.60928
0.7323	4.30279	4.61509

The coefficient of determination  $R^2$  measures how well predicted values fit the real data.

$$R^2 = 1 - \frac{\sum (y - \hat{y})^2}{\sum (y - \bar{y})^2} = 0.995$$

The result indicates that the model  $d\sigma$  explains 99.5% of the variance in the data.

Run “bhabha-scattering-5.txt” to verify.

## Notes

Here are a few notes about how the Eigenmath scripts work. In component notation the trace operators of the Casimir trick become sums over the repeated index  $\alpha$ .

$$\begin{aligned} f_{11} &= \left( (\not{p}_1 - m)^\alpha {}_\beta \gamma^{\mu\beta} {}_\rho (\not{p}_3 - m)^\rho {}_\sigma \gamma^{\nu\sigma} {}_\alpha \right) \left( (\not{p}_4 + m)^\alpha {}_\beta \gamma_\mu {}^\beta {}_\rho (\not{p}_2 + m)^\rho {}_\sigma \gamma_\nu {}^\sigma {}_\alpha \right) \\ f_{12} &= (\not{p}_1 - m)^\alpha {}_\beta \gamma^{\mu\beta} {}_\rho (\not{p}_2 + m)^\rho {}_\sigma \gamma^{\nu\sigma} {}_\tau (\not{p}_4 + m)^\tau {}_\delta \gamma_\mu {}^\delta {}_\eta (\not{p}_3 - m)^\eta {}_\xi \gamma_\nu {}^\xi {}_\alpha \\ f_{22} &= \left( (\not{p}_1 - m)^\alpha {}_\beta \gamma^{\mu\beta} {}_\rho (\not{p}_2 + m)^\rho {}_\sigma \gamma^{\nu\sigma} {}_\alpha \right) \left( (\not{p}_4 + m)^\alpha {}_\beta \gamma_\mu {}^\beta {}_\rho (\not{p}_3 - m)^\rho {}_\sigma \gamma_\nu {}^\sigma {}_\alpha \right) \end{aligned}$$

To convert the above formulas to Eigenmath code, the  $\gamma$  tensors need to be transposed so that repeated indices are adjacent to each other. Also, multiply  $\gamma^\mu$  by the metric tensor to lower the index.

$$\begin{aligned} \gamma^{\beta\mu} {}_\rho &\rightarrow \text{gammaT} = \text{transpose}(\text{gamma}) \\ \gamma^\beta {}_{\mu\rho} &\rightarrow \text{gammaL} = \text{transpose}(\text{dot}(\text{gmunu}, \text{gamma})) \end{aligned}$$

Define the following  $4 \times 4$  matrices.

$$\begin{aligned} (\not{p}_1 - m) &\rightarrow \text{X1} = \text{pslash1} - \text{m I} \\ (\not{p}_2 + m) &\rightarrow \text{X2} = \text{pslash2} + \text{m I} \\ (\not{p}_3 - m) &\rightarrow \text{X3} = \text{pslash3} - \text{m I} \\ (\not{p}_4 + m) &\rightarrow \text{X4} = \text{pslash4} + \text{m I} \end{aligned}$$

Then for  $f_{11}$  we have the following Eigenmath code. The contract function sums over  $\alpha$ .

$$\begin{aligned} (\not{p}_1 - m)^\alpha {}_\beta \gamma^{\mu\beta} {}_\rho (\not{p}_3 - m)^\rho {}_\sigma \gamma^{\nu\sigma} {}_\alpha &\rightarrow \text{T1} = \text{contract}(\text{dot}(\text{X1}, \text{gammaT}, \text{X3}, \text{gammaT}), 1, 4) \\ (\not{p}_4 + m)^\alpha {}_\beta \gamma_\mu {}^\beta {}_\rho (\not{p}_2 + m)^\rho {}_\sigma \gamma_\nu {}^\sigma {}_\alpha &\rightarrow \text{T2} = \text{contract}(\text{dot}(\text{X4}, \text{gammaL}, \text{X2}, \text{gammaL}), 1, 4) \end{aligned}$$

Next, multiply then sum over repeated indices. The dot function sums over  $\nu$  then the contract function sums over  $\mu$ . The transpose makes the  $\nu$  indices adjacent as required by the dot function.

$$f_{11} = \text{Tr}(\cdots \gamma^\mu \cdots \gamma^\nu) \text{Tr}(\cdots \gamma_\mu \cdots \gamma_\nu) \rightarrow \text{f11} = \text{contract}(\text{dot}(\text{T1}, \text{transpose}(\text{T2})))$$

Follow suit for  $f_{22}$ .

$$\begin{aligned} (\not{p}_1 - m)^\alpha {}_\beta \gamma^{\mu\beta} {}_\rho (\not{p}_2 + m)^\rho {}_\sigma \gamma^{\nu\sigma} {}_\alpha &\rightarrow \text{T1} = \text{contract}(\text{dot}(\text{X1}, \text{gammaT}, \text{X2}, \text{gammaT}), 1, 4) \\ (\not{p}_4 + m)^\alpha {}_\beta \gamma_\mu {}^\beta {}_\rho (\not{p}_3 - m)^\rho {}_\sigma \gamma_\nu {}^\sigma {}_\alpha &\rightarrow \text{T2} = \text{contract}(\text{dot}(\text{X4}, \text{gammaL}, \text{X3}, \text{gammaL}), 1, 4) \end{aligned}$$

Hence

$$f_{22} = \text{Tr}(\cdots \gamma^\mu \cdots \gamma^\nu) \text{Tr}(\cdots \gamma_\mu \cdots \gamma_\nu) \rightarrow \text{f22} = \text{contract}(\text{dot}(\text{T1}, \text{transpose}(\text{T2})))$$

The calculation of  $f_{12}$  begins with

$$\begin{aligned} (\not{p}_1 - m)^\alpha {}_\beta \gamma^{\mu\beta} {}_\rho (\not{p}_2 + m)^\rho {}_\sigma \gamma^{\nu\sigma} {}_\tau (\not{p}_4 + m)^\tau {}_\delta \gamma_\mu {}^\delta {}_\eta (\not{p}_3 - m)^\eta {}_\xi \gamma_\nu {}^\xi {}_\alpha \\ \rightarrow \text{T} = \text{contract}(\text{dot}(\text{X1}, \text{gammaT}, \text{X2}, \text{gammaT}, \text{X4}, \text{gammaL}, \text{X3}, \text{gammaL}), 1, 6) \end{aligned}$$

Then sum over repeated indices  $\mu$  and  $\nu$ .

$$f_{12} = \text{Tr}(\cdots \gamma^\mu \cdots \gamma^\nu \cdots \gamma_\mu \cdots \gamma_\nu) \rightarrow \text{f12} = \text{contract}(\text{contract}(\text{T}, 1, 3))$$