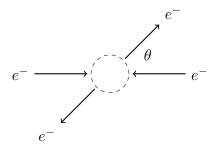
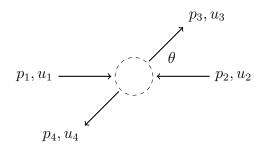
Moller (or electron) scattering is the result of interactions between electrons. The following diagram shows the geometry of a typical collider experiment that generates Moller scattering data.



Here is the same diagram with momentum and spinor labels.



In center of mass coordinates the momentum vectors are

$$p_{1} = \begin{pmatrix} E \\ 0 \\ 0 \\ p \end{pmatrix} \quad p_{2} = \begin{pmatrix} E \\ 0 \\ 0 \\ -p \end{pmatrix} \quad p_{3} = \begin{pmatrix} E \\ p\sin\theta\cos\phi \\ p\sin\theta\sin\phi \\ p\cos\theta \end{pmatrix} \quad p_{4} = \begin{pmatrix} E \\ -p\sin\theta\cos\phi \\ -p\sin\theta\sin\phi \\ -p\cos\theta \end{pmatrix}$$
inbound electron inbound electron outbound electron outbound electron

Symbol p is incident momentum, E is total energy $E = \sqrt{p^2 + m^2}$, and m is electron mass. Polar angle θ is the observed scattering angle. Azimuth angle ϕ cancels out in scattering calculations.

The spinors are

$$u_{11} = \begin{pmatrix} E+m \\ 0 \\ p \\ 0 \end{pmatrix} \qquad u_{12} = \begin{pmatrix} 0 \\ E+m \\ 0 \\ -p \end{pmatrix} \qquad u_{21} = \begin{pmatrix} E+m \\ 0 \\ -p \\ 0 \end{pmatrix} \qquad u_{22} = \begin{pmatrix} 0 \\ E+m \\ 0 \\ p \end{pmatrix}$$
 inbound electron, spin up inbound electron, spin down inbound electron, spin up inbound electron, spin down
$$u_{31} = \begin{pmatrix} E+m \\ 0 \\ p_{3z} \\ p_{3z} + ip_{3y} \end{pmatrix} \qquad u_{32} = \begin{pmatrix} 0 \\ E+m \\ p_{3x} - ip_{3y} \\ -p_{3z} \end{pmatrix} \qquad u_{41} = \begin{pmatrix} E+m \\ 0 \\ p_{4z} \\ p_{4x} + ip_{4y} \end{pmatrix} \qquad u_{42} = \begin{pmatrix} 0 \\ E+m \\ p_{4x} - ip_{4y} \\ -p_{4z} \end{pmatrix}$$
 outbound electron, spin up outbound electron, spin down outbound electron, spin up outbound electron, spin down

The spinors shown above are not individually normalized. Instead, a combined spinor normalization constant $N = (E + m)^4$ will be used.

The following formula computes a probability density $|\mathcal{M}_{abcd}|^2$ for Moller scattering where the subscripts abcd are the spin states of the electrons.

$$|\mathcal{M}_{abcd}|^2 = \frac{e^4}{N} \begin{vmatrix} \frac{1}{t} (\bar{u}_{3c} \gamma^{\mu} u_{1a}) (\bar{u}_{4d} \gamma_{\mu} u_{2b}) - \frac{1}{u} (\bar{u}_{4d} \gamma^{\nu} u_{1a}) (\bar{u}_{3c} \gamma_{\nu} u_{2b}) \\ \text{from Feynman diagram with} \\ \text{photon exchange} \\ \text{no electron interchange} \end{vmatrix}$$
 from Feynman diagram with photon exchange electron interchange

Symbol e is electron charge. Symbols t and u are Mandelstam variables

$$t = (p_1 - p_3)^2$$
$$u = (p_1 - p_4)^2$$

Let

$$a_1 = (\bar{u}_{3c}\gamma^{\mu}u_{1a})(\bar{u}_{4d}\gamma_{\mu}u_{2b})$$
 $a_2 = (\bar{u}_{4d}\gamma^{\nu}u_{1a})(\bar{u}_{3c}\gamma_{\nu}u_{2b})$

Then

$$|\mathcal{M}_{abcd}|^2 = \frac{e^4}{N} \left| \frac{a_1}{t} - \frac{a_2}{u} \right|^2$$

$$= \frac{e^4}{N} \left(\frac{a_1}{t} - \frac{a_2}{u} \right) \left(\frac{a_1}{t} - \frac{a_2}{u} \right)^*$$

$$= \frac{e^4}{N} \left(\frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{tu} - \frac{a_1^* a_2}{tu} + \frac{a_2 a_2^*}{u^2} \right)$$

The expected probability density $\langle |\mathcal{M}|^2 \rangle$ is computed by summing $|\mathcal{M}_{abcd}|^2$ over all spin states and dividing by the number of inbound states. There are four inbound states.

$$\langle |\mathcal{M}|^2 \rangle = \frac{1}{4} \sum_{a=1}^2 \sum_{b=1}^2 \sum_{c=1}^2 \sum_{d=1}^2 |\mathcal{M}_{abcd}|^2$$

$$= \frac{e^4}{4N} \sum_{a=1}^2 \sum_{b=1}^2 \sum_{c=1}^2 \sum_{d=1}^2 \left(\frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{tu} - \frac{a_1^* a_2}{tu} + \frac{a_2 a_2^*}{u^2} \right)$$

Use the Casimir trick to replace sums over spins with matrix products.

$$f_{11} = \frac{1}{N} \sum_{abcd} a_1 a_1^* = \operatorname{Tr} \left((\not p_3 + m) \gamma^{\mu} (\not p_1 + m) \gamma^{\nu} \right) \operatorname{Tr} \left((\not p_4 + m) \gamma_{\mu} (\not p_2 + m) \gamma_{\nu} \right)$$

$$f_{12} = \frac{1}{N} \sum_{abcd} a_1 a_2^* = \operatorname{Tr} \left((\not p_3 + m) \gamma^{\mu} (\not p_1 + m) \gamma^{\nu} (\not p_4 + m) \gamma_{\mu} (\not p_2 + m) \gamma_{\nu} \right)$$

$$f_{22} = \frac{1}{N} \sum_{abcd} a_2 a_2^* = \operatorname{Tr} \left((\not p_4 + m) \gamma^{\mu} (\not p_1 + m) \gamma^{\nu} \right) \operatorname{Tr} \left((\not p_3 + m) \gamma_{\mu} (\not p_2 + m) \gamma_{\nu} \right)$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left(\frac{f_{11}}{t^2} - \frac{f_{12}}{tu} - \frac{f_{12}^*}{tu} + \frac{f_{22}}{u^2} \right)$$

Run "moller-scattering-1.txt" to verify the Casimir trick.

The following formulas are equivalent to the Casimir trick. (Recall that $a \cdot b = a^{\mu}g_{\mu\nu}b^{\nu}$)

$$f_{11} = 32(p_1 \cdot p_2)^2 + 32(p_1 \cdot p_4)^2 - 64m^2(p_1 \cdot p_3) + 64m^4$$

$$f_{12} = -32(p_1 \cdot p_2)^2 + 32m^2(p_1 \cdot p_2) + 32m^2(p_1 \cdot p_3) + 32m^2(p_1 \cdot p_4) - 32m^4$$

$$f_{22} = 32(p_1 \cdot p_2)^2 + 32(p_1 \cdot p_3)^2 - 64m^2(p_1 \cdot p_4) + 64m^4$$

In Mandelstam variables

$$s = (p_1 + p_2)^2$$

$$t = (p_1 - p_3)^2$$

$$u = (p_1 - p_4)^2$$

the formulas are

$$f_{11} = 8s^2 + 8u^2 - 64sm^2 - 64um^2 + 192m^4$$

$$f_{12} = -8s^2 + 64sm^2 - 96m^4$$

$$f_{22} = 8s^2 + 8t^2 - 64sm^2 - 64tm^2 + 192m^4$$

Run "moller-scattering-2.txt" to verify.

High energy approximation

When $E \gg m$ a useful approximation is to set m=0 and obtain

$$f_{11} = 8s^2 + 8u^2$$
$$f_{12} = -8s^2$$
$$f_{22} = 8s^2 + 8t^2$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left(\frac{f_{11}}{t^2} - \frac{f_{12}}{tu} - \frac{f_{12}^*}{tu} + \frac{f_{22}}{u^2} \right)$$

$$= \frac{e^4}{4} \left(\frac{8s^2 + 8u^2}{t^2} - \frac{-8s^2}{tu} - \frac{-8s^2}{tu} + \frac{8s^2 + 8t^2}{u^2} \right)$$

$$= 2e^4 \left(\frac{s^2 + u^2}{t^2} + \frac{2s^2}{tu} + \frac{s^2 + t^2}{u^2} \right)$$

Combine terms so $\langle |\mathcal{M}|^2 \rangle$ has a common denominator.

$$\langle |\mathcal{M}|^2 \rangle = 2e^4 \left(\frac{u^2(s^2 + u^2) + 2s^2tu + t^2(s^2 + t^2)}{t^2u^2} \right)$$

For m = 0 the Mandelstam variables are

$$s = 4E^{2}$$

$$t = 2E^{2}(\cos \theta - 1)$$

$$u = -2E^{2}(\cos \theta + 1)$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = 2e^4 \left(\frac{32E^8 \cos^4 \theta + 192E^8 \cos^2 \theta + 288E^8}{16E^8 (\cos \theta - 1)^2 (\cos \theta + 1)^2} \right)$$
$$= 4e^4 \frac{(\cos^2 \theta + 3)^2}{(\cos \theta - 1)^2 (\cos \theta + 1)^2}$$
$$= 4e^4 \left(\frac{\cos^2 \theta + 3}{\cos^2 \theta - 1} \right)^2$$

The following equivalent formula can also be used.

$$\langle |\mathcal{M}|^2 \rangle = 2e^4 \left(\frac{s^2 + u^2}{t^2} + \frac{2s^2}{tu} + \frac{s^2 + t^2}{u^2} \right)$$

$$= 2e^4 \left(\frac{1 + \cos^4(\theta/2)}{\sin^4(\theta/2)} + \frac{2}{\sin^2(\theta/2)\cos^2(\theta/2)} + \frac{1 + \sin^4(\theta/2)}{\cos^4(\theta/2)} \right)$$
from Feynman diagram with photon exchange no electron interchange

For example, see A. Zee, p. 134.

Cross section

The differential cross section is

$$\frac{d\sigma}{d\Omega} = \frac{\langle |\mathcal{M}|^2 \rangle}{64\pi^2 s} = \frac{e^4}{16\pi^2 s} \left(\frac{\cos^2 \theta + 3}{\cos^2 \theta - 1}\right)^2, \quad s \gg m$$

Substituting $e^4 = 16\pi^2\alpha^2$ yields

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2}{s} \left(\frac{\cos^2 \theta + 3}{\cos^2 \theta - 1} \right)^2$$

We can integrate $d\sigma$ to obtain a cumulative distribution function. Recall that

$$d\Omega = \sin\theta \, d\theta \, d\phi$$

Hence

$$d\sigma = \frac{\alpha^2}{s} \left(\frac{\cos^2 \theta + 3}{\cos^2 \theta - 1} \right)^2 \sin \theta \, d\theta \, d\phi$$

Let $I(\theta)$ be the following integral of $d\sigma$.

$$I(\theta) = \frac{s}{2\pi\alpha^2} \int_0^{2\pi} \int d\sigma$$
$$= \int \left(\frac{\cos^2 \theta + 3}{\cos^2 \theta - 1}\right)^2 \sin \theta \, d\theta, \quad a \le \theta \le \pi - a$$

Angular support is limited to an arbitrary a > 0 because I(0) and $I(\pi)$ are undefined. Assume that $I(\theta) - I(a)$ is computable given θ by either symbolic or numerical integration.

Let C be the normalization constant

$$C = I(\pi - a) - I(a)$$

Then the cumulative distribution function $F(\theta)$ is

$$F(\theta) = \frac{I(\theta) - I(a)}{C}, \quad a \le \theta \le \pi - a$$

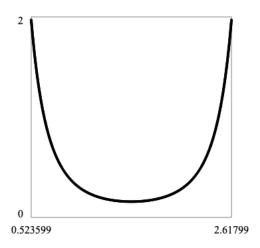
The probability of observing scattering events in the interval θ_1 to θ_2 can now be computed.

$$P(\theta_1 \le \theta \le \theta_2) = F(\theta_2) - F(\theta_1)$$

Probability density function $f(\theta)$ is the derivative of $F(\theta)$.

$$f(\theta) = \frac{dF(\theta)}{d\theta} = \frac{1}{C} \left(\frac{\cos^2 \theta + 3}{\cos^2 \theta - 1} \right)^2 \sin \theta$$

Run "moller-scattering-3.txt" to draw a graph of $f(\theta)$ for $a = \pi/6 = 30^{\circ}$.



Probability distribution for 30° bins $(a = 30^{\circ})$.

θ_1	θ_2	$P(\theta_1 \le \theta \le \theta_2)$
0°	30°	_
30°	60°	0.40
60°	90°	0.10
90°	120°	0.10
120°	150°	0.40
150°	180°	_

Notes

In component notation, the trace operators of the Casimir trick become sums over a repeated index, in this case α .

$$f_{11} = \left((\not p_3 + m)^{\alpha}{}_{\beta} \gamma^{\mu\beta}{}_{\rho} (\not p_1 + m)^{\rho}{}_{\sigma} \gamma^{\nu\sigma}{}_{\alpha} \right) \left((\not p_4 + m)^{\alpha}{}_{\beta} \gamma_{\mu}{}^{\beta}{}_{\rho} (\not p_2 + m)^{\rho}{}_{\sigma} \gamma_{\nu}{}^{\sigma}{}_{\alpha} \right)$$

$$f_{12} = (\not p_3 + m)^{\alpha}{}_{\beta} \gamma^{\mu\beta}{}_{\rho} (\not p_1 + m)^{\rho}{}_{\sigma} \gamma^{\nu\sigma}{}_{\tau} (\not p_4 + m)^{\tau}{}_{\delta} \gamma_{\mu}{}^{\delta}{}_{\eta} (\not p_2 + m)^{\eta}{}_{\xi} \gamma_{\nu}{}^{\xi}{}_{\alpha}$$

$$f_{22} = \left((\not p_4 + m)^{\alpha}{}_{\beta} \gamma^{\mu\beta}{}_{\rho} (\not p_1 + m)^{\rho}{}_{\sigma} \gamma^{\nu\sigma}{}_{\alpha} \right) \left((\not p_3 + m)^{\alpha}{}_{\beta} \gamma_{\mu}{}^{\beta}{}_{\rho} (\not p_2 + m)^{\rho}{}_{\sigma} \gamma_{\nu}{}^{\sigma}{}_{\alpha} \right)$$

To convert the above formulas to Eigenmath code, the γ tensors need to be transposed so that repeated indices are adjacent to each other. Also, multiply γ^{μ} by the metric tensor to lower the index.

$$\gamma^{\beta\mu}_{\ \rho} \rightarrow {\rm gammaT = transpose(gamma)}$$
 $\gamma^{\beta}_{\ \mu\rho} \rightarrow {\rm gammaL = transpose(dot(gmunu,gamma))}$

Define the following 4×4 matrices.

Then for f_{11} we have the following Eigenmath code. The contract function sums over α .

$$(\not\!p_3 + m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_1 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \quad \rightarrow \quad \text{T1 = contract(dot(X3,gammaT,X1,gammaT),1,4)}$$

$$(\not\!p_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \quad \rightarrow \quad \text{T2 = contract(dot(X4,gammaL,X2,gammaL),1,4)}$$

Next, multiply then sum over repeated indices. The dot function sums over ν then the contract function sums over μ . The transpose makes the ν indices adjacent as required by the dot function.

$$f_{11} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu}) \operatorname{Tr}(\cdots \gamma_{\mu} \cdots \gamma_{\nu}) \quad \rightarrow \quad \operatorname{contract(dot(T1, transpose(T2)))}$$

Follow suit for f_{22} .

$$(\not\!p_4 + m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_1 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \quad \rightarrow \quad \text{T1 = contract(dot(X4,gammaT,X1,gammaT),1,4)}$$

$$(\not\!p_3 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \quad \rightarrow \quad \text{T2 = contract(dot(X3,gammaL,X2,gammaL),1,4)}$$

Then

$$f_{22} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu}) \operatorname{Tr}(\cdots \gamma_{\mu} \cdots \gamma_{\nu}) \quad o \quad \operatorname{contract(dot(T1,transpose(T2)))}$$

The calculation of f_{12} begins with

$$\begin{split} (\not\!p_3 + m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_1 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\tau (\not\!p_4 + m)^\tau{}_\delta \gamma_\mu{}^\delta{}_\eta (\not\!p_2 + m)^\eta{}_\xi \gamma_\nu{}^\xi{}_\alpha \\ \rightarrow \quad \mathsf{T} = \mathsf{contract}(\mathsf{dot}(\mathsf{X3},\mathsf{gammaT},\mathsf{X1},\mathsf{gammaT},\mathsf{X4},\mathsf{gammaL},\mathsf{X2},\mathsf{gammaL}),\mathsf{1},\mathsf{6}) \end{split}$$

Then sum over repeated indices μ and ν .

$$f_{12} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu} \cdots \gamma_{\mu} \cdots \gamma_{\nu}) \quad o \quad \operatorname{contract(contract(T,1,3))}$$