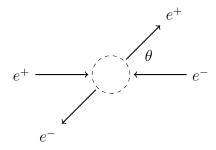
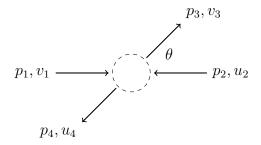
Bhabha scattering is the result of interactions between positrons and electrons. The following diagram represents a collider experiment with collinear electron and positron beams.



Here is the same diagram with momentum and spinor labels.



In a typical collider experiment the momentum vectors are

$$p_{1} = \begin{pmatrix} E \\ 0 \\ 0 \\ p \end{pmatrix} \qquad p_{2} = \begin{pmatrix} E \\ 0 \\ 0 \\ -p \end{pmatrix} \qquad p_{3} = \begin{pmatrix} E \\ p\sin\theta\cos\phi \\ p\sin\theta\sin\phi \\ p\cos\theta \end{pmatrix} \qquad p_{4} = \begin{pmatrix} E \\ -p\sin\theta\cos\phi \\ -p\sin\theta\sin\phi \\ -p\cos\theta \end{pmatrix}$$

where  $E = \sqrt{p^2 + m^2}$ . The spinors are

$$v_{11} = \begin{pmatrix} p \\ 0 \\ E+m \\ 0 \end{pmatrix} \quad u_{21} = \begin{pmatrix} E+m \\ 0 \\ -p \\ 0 \end{pmatrix} \quad v_{31} = \begin{pmatrix} p_3^z \\ p_3^x + ip_3^y \\ E+m \\ 0 \end{pmatrix} \quad u_{41} = \begin{pmatrix} E+m \\ 0 \\ p_4^z \\ p_4^x + ip_4^y \end{pmatrix}$$

$$v_{12} = \begin{pmatrix} 0 \\ -p \\ 0 \\ E+m \end{pmatrix} \quad u_{22} = \begin{pmatrix} 0 \\ E+m \\ 0 \\ p \end{pmatrix} \quad v_{32} = \begin{pmatrix} p_3^x - ip_3^y \\ -p_3^z \\ 0 \\ E+m \end{pmatrix} \quad u_{42} = \begin{pmatrix} 0 \\ E+m \\ p_4^x - ip_4^y \\ -p_4^z \end{pmatrix}$$

The spinors shown above are not individually normalized. Instead, a combined spinor normalization constant  $N = (E + m)^4$  will be used.

This is the probability density for Bhabha scattering.

$$|\mathcal{M}(s_1, s_2, s_3, s_4)|^2 = \frac{e^4}{N} \left| -\frac{1}{t} (\bar{v}_1 \gamma^{\mu} v_3) (\bar{u}_4 \gamma_{\mu} u_2) + \frac{1}{s} (\bar{v}_1 \gamma^{\nu} u_2) (\bar{u}_4 \gamma_{\nu} v_3) \right|^2$$

Symbol  $s_j$  selects the spin (up or down) of spinor j. Symbol e is electron charge. Symbols s and t are Mandelstam variables  $s = (p_1 + p_2)^2$  and  $t = (p_1 - p_3)^2$ .

Let

$$a_1 = (\bar{v}_1 \gamma^{\mu} v_3)(\bar{u}_4 \gamma_{\mu} u_2)$$
  $a_2 = (\bar{v}_1 \gamma^{\nu} u_2)(\bar{u}_4 \gamma_{\nu} v_3)$ 

Then

$$|\mathcal{M}(s_1, s_2, s_3, s_4)|^2 = \frac{e^4}{N} \left| -\frac{a_1}{t} + \frac{a_2}{s} \right|^2$$

$$= \frac{e^4}{N} \left( -\frac{a_1}{t} + \frac{a_2}{s} \right) \left( -\frac{a_1}{t} + \frac{a_2}{s} \right)^*$$

$$= \frac{e^4}{N} \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right)$$

The expected probability density  $\langle |\mathcal{M}|^2 \rangle$  is computed by summing  $|\mathcal{M}|^2$  over all spin states and dividing by the number of inbound states. There are four inbound states.

$$\langle |\mathcal{M}|^2 \rangle = \frac{1}{4} \sum_{s_1=1}^2 \sum_{s_2=1}^2 \sum_{s_3=1}^2 \sum_{s_4=1}^2 |\mathcal{M}(s_1, s_2, s_3, s_4)|^2$$

$$= \frac{e^4}{4} \sum_{s_1=1}^2 \sum_{s_2=1}^2 \sum_{s_3=1}^2 \sum_{s_4=1}^2 \frac{1}{N} \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right)$$

Use the Casimir trick to replace sums over spins with matrix products.

$$f_{11} = \frac{1}{N} \sum_{\text{spins}} a_1 a_1^* = \text{Tr}\left( (\not p_1 - m) \gamma^{\mu} (\not p_3 - m) \gamma^{\nu} \right) \text{Tr}\left( (\not p_4 + m) \gamma_{\mu} (\not p_2 + m) \gamma_{\nu} \right)$$

$$f_{12} = \frac{1}{N} \sum_{\text{spins}} a_1 a_2^* = \text{Tr}\left( (\not p_1 - m) \gamma^{\mu} (\not p_2 + m) \gamma^{\nu} (\not p_4 + m) \gamma_{\mu} (\not p_3 - m) \gamma_{\nu} \right)$$

$$f_{22} = \frac{1}{N} \sum_{\text{spins}} a_2 a_2^* = \text{Tr}\left( (\not p_1 - m) \gamma^{\mu} (\not p_2 + m) \gamma^{\nu} \right) \text{Tr}\left( (\not p_4 + m) \gamma_{\mu} (\not p_3 - m) \gamma_{\nu} \right)$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left( \frac{f_{11}}{t^2} - \frac{f_{12}}{st} - \frac{f_{12}^*}{st} + \frac{f_{22}}{s^2} \right)$$

Run "bhabha-scattering-1.txt" to verify the Casimir trick.

The following momentum formulas are equivalent to the Casimir trick. (Recall that  $a \cdot b = a^{\mu} g_{\mu\nu} b^{\nu}$ )

$$f_{11} = 32(p_1 \cdot p_2)(p_3 \cdot p_4) + 32(p_1 \cdot p_4)(p_2 \cdot p_3) - 32m^2(p_1 \cdot p_3) - 32m^2(p_2 \cdot p_4) + 64m^4$$

$$f_{12} = -32(p_1 \cdot p_4)(p_2 \cdot p_3) - 16m^2(p_1 \cdot p_2) + 16m^2(p_1 \cdot p_3) - 16m^2(p_1 \cdot p_4)$$

$$- 16m^2(p_2 \cdot p_3) + 16m^2(p_2 \cdot p_4) - 16m^2(p_3 \cdot p_4) - 32m^4$$

$$f_{22} = 32(p_1 \cdot p_3)(p_2 \cdot p_4) + 32(p_1 \cdot p_4)(p_2 \cdot p_3) + 32m^2(p_1 \cdot p_2) + 32m^2(p_3 \cdot p_4) + 64m^4$$

In Mandelstam variables  $s = (p_1 + p_2)^2$ ,  $t = (p_1 - p_3)^2$ ,  $u = (p_1 - p_4)^2$  the formulas are

$$f_{11} = 8s^2 + 8u^2 - 64sm^2 - 64um^2 + 192m^4$$
  

$$f_{12} = -8u^2 + 64um^2 - 96m^4$$
  

$$f_{22} = 8t^2 + 8u^2 - 64tm^2 - 64um^2 + 192m^4$$

# High energy approximation

When  $E \gg m$  a useful approximation is to set m=0 and obtain

$$f_{11} = 8s^2 + 8u^2$$
$$f_{12} = -8u^2$$
$$f_{22} = 8t^2 + 8u^2$$

For m=0 the Mandelstam variables are

$$s = 4E^{2}$$

$$t = -2E^{2}(1 - \cos \theta) = -4E^{2}\sin^{2}(\theta/2)$$

$$u = -2E^{2}(1 + \cos \theta) = -4E^{2}\cos^{2}(\theta/2)$$

It follows that

$$s^2 t^2 = 256 E^8 \sin^4(\theta/2)$$

The corresponding expected probability density is

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left( \frac{f_{11}}{t^2} - \frac{f_{12}}{st} - \frac{f_{12}^*}{st} + \frac{f_{22}}{s^2} \right)$$

$$= \frac{e^4}{4s^2t^2} \left( s^2 f_{11} - st f_{12} - st f_{12}^* + t^2 f_{22} \right)$$

$$= \frac{e^4}{4s^2t^2} \left( s^2 \left( 8s^2 + 8u^2 \right) + 16stu^2 + t^2 \left( 8t^2 + 8u^2 \right) \right)$$

$$= \frac{e^4}{1024E^8 \sin^4(\theta/2)} \left( 256E^8 \cos^4 \theta + 1536E^8 \cos^2 \theta + 2304E^8 \right)$$

$$= \frac{e^4}{4\sin^4(\theta/2)} \left( \cos^4 \theta + 6\cos^2 \theta + 9 \right)$$

$$= \frac{e^4}{4\sin^4(\theta/2)} \left( \cos^4 \theta + 3 \right)^2$$

$$= \frac{e^4}{4\sin^4(\theta/2)} \left( \sin^4 \theta + 3 \right)^2$$

Run "bhabha-scattering-2.txt" to verify.

#### Cross section

This is the differential cross section for Bhabha scattering.

$$\frac{d\sigma}{d\Omega} = \frac{\langle |\mathcal{M}|^2 \rangle}{64\pi^2 s} = \frac{e^4}{1024\pi^2 E^2} \frac{\left(\cos^2 \theta + 3\right)^2}{\sin^4(\theta/2)}$$

Substituting  $e^4 = 16\pi^2\alpha^2$  yields

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2}{64E^2} \frac{\left(\cos^2\theta + 3\right)^2}{\sin^4(\theta/2)}$$

We can integrate  $d\sigma$  to obtain a cumulative distribution function. Recall that

$$d\Omega = \sin\theta \, d\theta \, d\phi$$

Hence

$$d\sigma = \frac{\alpha^2}{64E^2} \frac{\left(\cos^2\theta + 3\right)^2}{\sin^4(\theta/2)} \sin\theta \, d\theta \, d\phi$$

Let  $I(\theta)$  be the following integral.

$$I(\theta) = \left(\frac{64E^2}{\alpha^2}\right) \frac{1}{2\pi} \int_0^{2\pi} \int d\sigma$$
$$= \int \frac{\left(\cos^2 \theta + 3\right)^2}{\sin^4(\theta/2)} \sin \theta \, d\theta, \quad a \le \xi \le \pi$$

Angular support is limited to an arbitrary a > 0 because I(0) is undefined.

Let C be the normalization constant  $C = I(\pi) - I(a)$ . Then the cumulative distribution function  $F(\theta)$  is

$$F(\theta) = \frac{I(\theta) - I(a)}{C}, \quad a \le \theta \le \pi$$

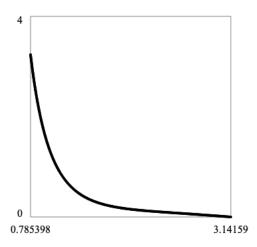
The probability of observing scattering events in the interval  $\theta_1$  to  $\theta_2$  can now be computed.

$$P(\theta_1 \le \theta \le \theta_2) = F(\theta_2) - F(\theta_1)$$

Probability density function  $f(\theta)$  is the derivative of  $F(\theta)$ .

$$f(\theta) = \frac{dF(\theta)}{d\theta} = C^{-1} \frac{dI(\theta)}{d\theta} = C^{-1} \frac{(\cos^2 \theta + 3)^2}{\sin^4(\theta/2)} \sin \theta$$

Run "bhabha-scattering-3.txt" to draw  $f(\theta)$  for  $a = \pi/4 = 45^{\circ}$ .



The following table shows the corresponding probability distribution for three bins.

| $\theta_1$ | $\theta_2$ | $P(\theta_1 \le \theta \le \theta_2)$ |
|------------|------------|---------------------------------------|
| 0°         | 45°        | _                                     |
| 45°        | 90°        | 0.83                                  |
| 90°        | 135°       | 0.13                                  |
| 135°       | 180°       | 0.04                                  |

## **SLAC** data

The following Bhabha scattering data is adapted from SLAC-PUB-1501.

|                      | Bin | $x_k, x_{k+1}$ | y    |
|----------------------|-----|----------------|------|
| (Smallest $\theta$ ) | 1   | 0.6, 0.5       | 4432 |
|                      | 2   | 0.5, 0.4       | 2841 |
|                      | 3   | 0.4, 0.3       | 2045 |
|                      | 4   | 0.3, 0.2       | 1420 |
|                      | 5   | 0.2, 0.1       | 1136 |
|                      | 6   | 0.1, 0.0       | 852  |
|                      | 7   | 0.0, -0.1      | 656  |
|                      | 8   | -0.1, -0.2     | 625  |
|                      | 9   | -0.2, -0.3     | 511  |
|                      | 10  | -0.3, -0.4     | 455  |
|                      | 11  | -0.4, -0.5     | 402  |
| (Largest $\theta$ )  | 12  | -0.5, -0.6     | 398  |

Data column y is the observed number of scattering events per bin.

To compute predicted counts  $\hat{y}$ , integrate the probability density function over each bin and multiply by the total number of observed counts.

$$P_k = C^{-1} \int_{\arccos(x_k)}^{\arccos(x_{k+1})} \frac{(\cos^2 \theta + 3)^2}{\sin^4(\theta/2)} \sin \theta$$
$$\hat{y}_k = P_k \sum y$$

| Bin | $x_k, x_{k+1}$ | y    | $\hat{y}$ |
|-----|----------------|------|-----------|
| 1   | 0.6, 0.5       | 4432 | 4598      |
| 2   | 0.5, 0.4       | 2841 | 2880      |
| 3   | 0.4, 0.3       | 2045 | 1955      |
| 4   | 0.3, 0.2       | 1420 | 1410      |
| 5   | 0.2, 0.1       | 1136 | 1068      |
| 6   | 0.1, 0.0       | 852  | 843       |
| 7   | 0.0, -0.1      | 656  | 689       |
| 8   | -0.1, -0.2     | 625  | 582       |
| 9   | -0.2, -0.3     | 511  | 505       |
| 10  | -0.3, -0.4     | 455  | 450       |
| 11  | -0.4, -0.5     | 402  | 411       |
| 12  | -0.5, -0.6     | 398  | 382       |

The coefficient of determination  $\mathbb{R}^2$  measures how well predicted values fit the real data. Let y be observed counts per bin and let  $\hat{y}$  be predicted counts per bin. Then

$$R^{2} = 1 - \frac{\sum (y - \hat{y})^{2}}{\sum (y - \bar{y})^{2}} = 0.997$$

The result indicates that the model  $\langle |\mathcal{M}|^2 \rangle$  explains 99.7% of the variance in the data.

Run "bhabha-scattering-4.txt" to verify.

### **DESY** data

The following table shows DESY-PETRA Bhabha scattering data obtained from HEP Data.<sup>1</sup>

| x       | y       |
|---------|---------|
| -0.73   | 0.10115 |
| -0.6495 | 0.12235 |
| -0.5495 | 0.11258 |
| -0.4494 | 0.09968 |
| -0.3493 | 0.14749 |
| -0.2491 | 0.14017 |
| -0.149  | 0.1819  |
| -0.0488 | 0.22964 |
| 0.0514  | 0.25312 |
| 0.1516  | 0.30998 |
| 0.252   | 0.40898 |
| 0.3524  | 0.62695 |
| 0.4529  | 0.91803 |
| 0.5537  | 1.51743 |
| 0.6548  | 2.56714 |
| 0.7323  | 4.30279 |

Data x and y have the following relationship with the cross section model.

$$x = \cos \theta$$
  $y = \frac{d\sigma}{d\Omega}$ 

The differential cross section for Bhabha scattering is

$$\frac{d\sigma}{d\Omega} = \frac{\langle |\mathcal{M}|^2 \rangle}{64\pi^2 s} = \frac{\alpha^2}{2s} \left( \frac{s^2 + u^2}{t^2} + \frac{2u^2}{st} + \frac{t^2 + u^2}{s^2} \right)$$

The predicted cross section  $\hat{y}$  is computed from data x and beam energy E as

$$\hat{y} = \frac{\alpha^2}{2s} \left( \frac{s^2 + u^2}{t^2} + \frac{2u^2}{st} + \frac{t^2 + u^2}{s^2} \right) \times (\hbar c)^2 \times 10^{37}$$

where

$$s = 4E2$$
  

$$t = -2E2(1 - x)$$
  

$$u = -2E2(1 + x)$$

Factor  $(\hbar c)^2$  converts the result to SI and factor  $10^{37}$  converts square meters to nanobarns.

The following table shows  $\hat{y}$  for  $E = 7.0 \,\text{GeV}$ .

<sup>&</sup>lt;sup>1</sup>www.hepdata.net/record/ins191231 (Table 3, 14.0 GeV)

| x       | y       | $\hat{y}$ |
|---------|---------|-----------|
| -0.73   | 0.10115 | 0.110296  |
| -0.6495 | 0.12235 | 0.113816  |
| -0.5495 | 0.11258 | 0.120101  |
| -0.4494 | 0.09968 | 0.129075  |
| -0.3493 | 0.14749 | 0.141592  |
| -0.2491 | 0.14017 | 0.158934  |
| -0.149  | 0.1819  | 0.182976  |
| -0.0488 | 0.22964 | 0.216737  |
| 0.0514  | 0.25312 | 0.264989  |
| 0.1516  | 0.30998 | 0.335782  |
| 0.252   | 0.40898 | 0.44363   |
| 0.3524  | 0.62695 | 0.615528  |
| 0.4529  | 0.91803 | 0.9077    |
| 0.5537  | 1.51743 | 1.45175   |
| 0.6548  | 2.56714 | 2.60928   |
| 0.7323  | 4.30279 | 4.61509   |

The coefficient of determination  $R^2$  measures how well predicted values fit the real data.

$$R^{2} = 1 - \frac{\sum (y - \hat{y})^{2}}{\sum (y - \bar{y})^{2}} = 0.995$$

The result indicates that the model  $d\sigma$  explains 99.5% of the variance in the data.

Run "bhabha-scattering-5.txt" to verify.

## Notes on Eigenmath scripts

Here are a few notes about how the Eigenmath scripts work. In component notation the trace operators of the Casimir trick become sums over the repeated index  $\alpha$ .

$$f_{11} = \left( (\not p_1 - m)^{\alpha}{}_{\beta} \gamma^{\mu\beta}{}_{\rho} (\not p_3 - m)^{\rho}{}_{\sigma} \gamma^{\nu\sigma}{}_{\alpha} \right) \left( (\not p_4 + m)^{\alpha}{}_{\beta} \gamma_{\mu}{}^{\beta}{}_{\rho} (\not p_2 + m)^{\rho}{}_{\sigma} \gamma_{\nu}{}^{\sigma}{}_{\alpha} \right)$$

$$f_{12} = (\not p_1 - m)^{\alpha}{}_{\beta} \gamma^{\mu\beta}{}_{\rho} (\not p_2 + m)^{\rho}{}_{\sigma} \gamma^{\nu\sigma}{}_{\tau} (\not p_4 + m)^{\tau}{}_{\delta} \gamma_{\mu}{}^{\delta}{}_{\eta} (\not p_3 - m)^{\eta}{}_{\xi} \gamma_{\nu}{}^{\xi}{}_{\alpha}$$

$$f_{22} = \left( (\not p_1 - m)^{\alpha}{}_{\beta} \gamma^{\mu\beta}{}_{\rho} (\not p_2 + m)^{\rho}{}_{\sigma} \gamma^{\nu\sigma}{}_{\alpha} \right) \left( (\not p_4 + m)^{\alpha}{}_{\beta} \gamma_{\mu}{}^{\beta}{}_{\rho} (\not p_3 - m)^{\rho}{}_{\sigma} \gamma_{\nu}{}^{\sigma}{}_{\alpha} \right)$$

To convert the above formulas to Eigenmath code, the  $\gamma$  tensors need to be transposed so that repeated indices are adjacent to each other. Also, multiply  $\gamma^{\mu}$  by the metric tensor to lower the index.

$$\gamma^{eta\mu}_{\phantom{\mu}
ho}$$
  $ightarrow$  gammaT = transpose(gamma)  $\gamma^{eta}_{\phantom{\mu}\mu
ho}$   $ightarrow$  gammaL = transpose(dot(gmunu,gamma))

Define the following  $4 \times 4$  matrices.

Then for  $f_{11}$  we have the following Eigenmath code. The contract function sums over  $\alpha$ .

$$(\not\!\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!\!p_3 - m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \quad \rightarrow \quad \text{T1 = contract(dot(X1,gammaT,X3,gammaT),1,4)}$$
 
$$(\not\!\!p_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not\!\!p_2 + m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \quad \rightarrow \quad \text{T2 = contract(dot(X4,gammaL,X2,gammaL),1,4)}$$

Next, multiply then sum over repeated indices. The dot function sums over  $\nu$  then the contract function sums over  $\mu$ . The transpose makes the  $\nu$  indices adjacent as required by the dot function.

$$f_{11} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu}) \operatorname{Tr}(\cdots \gamma_{\mu} \cdots \gamma_{\nu}) \rightarrow \operatorname{fll} = \operatorname{contract}(\operatorname{dot}(\operatorname{T1,transpose}(\operatorname{T2})))$$

Follow suit for  $f_{22}$ .

$$(\not\!p_1-m)^\alpha{}_\beta\gamma^{\mu\beta}{}_\rho(\not\!p_2+m)^\rho{}_\sigma\gamma^{\nu\sigma}{}_\alpha \quad \rightarrow \quad \text{T1 = contract(dot(X1,gammaT,X2,gammaT),1,4)}$$
 
$$(\not\!p_4+m)^\alpha{}_\beta\gamma_\mu{}^\beta{}_\rho(\not\!p_3-m)^\rho{}_\sigma\gamma_\nu{}^\sigma{}_\alpha \quad \rightarrow \quad \text{T2 = contract(dot(X4,gammaL,X3,gammaL),1,4)}$$

Hence

$$f_{22} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu}) \operatorname{Tr}(\cdots \gamma_{\mu} \cdots \gamma_{\nu}) \rightarrow \text{f22} = \operatorname{contract}(\operatorname{dot}(\mathtt{T1}, \operatorname{transpose}(\mathtt{T2})))$$

The calculation of  $f_{12}$  begins with

$$(\not\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\tau (\not\!p_4 + m)^\tau{}_\delta \gamma_\mu{}^\delta{}_\eta (\not\!p_3 - m)^\eta{}_\xi \gamma_\nu{}^\xi{}_\alpha$$

$$\rightarrow \quad T = \text{contract(dot(X1,gammaT,X2,gammaT,X4,gammaL,X3,gammaL),1,6)}$$

Then sum over repeated indices  $\mu$  and  $\nu$ .

$$f_{12}=\mathrm{Tr}(\cdots\gamma^{\mu}\cdots\gamma^{\nu}\cdots\gamma_{\mu}\cdots\gamma_{\nu})$$
  $ightarrow$  f12 = contract(contract(T,1,3))