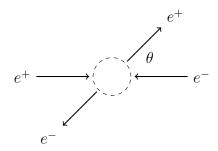
# Bhabha scattering

Bhabha scattering is the interaction  $e^- + e^+ \rightarrow e^- + e^+$ .



Define the following momentum vectors and spinors. Symbol p is incident momentum. Symbol E is total energy  $E = \sqrt{p^2 + m^2}$  where m is electron mass. Polar angle  $\theta$  is the observed scattering angle. Azimuth angle  $\phi$  cancels out in scattering calculations.

$$p_{1} = \begin{pmatrix} E \\ 0 \\ 0 \\ p \end{pmatrix} \qquad p_{2} = \begin{pmatrix} E \\ 0 \\ 0 \\ -p \end{pmatrix} \qquad p_{3} = \begin{pmatrix} E \\ p \sin \theta \cos \phi \\ p \sin \theta \sin \phi \\ p \cos \theta \end{pmatrix} \qquad p_{4} = \begin{pmatrix} E \\ -p \sin \theta \cos \phi \\ -p \sin \theta \sin \phi \\ -p \cos \theta \end{pmatrix}$$

$$v_{11} = \begin{pmatrix} p \\ 0 \\ E+m \\ 0 \\ spin up \end{pmatrix} \qquad u_{21} = \begin{pmatrix} E+m \\ 0 \\ -p \\ 0 \\ spin up \end{pmatrix} \qquad v_{31} = \begin{pmatrix} p_{3}^{z} \\ p_{3}^{x}+ip_{3}^{y} \\ E+m \\ 0 \\ spin up \end{pmatrix} \qquad u_{41} = \begin{pmatrix} E+m \\ 0 \\ p_{4}^{z} \\ p_{4}^{y}+ip_{4}^{y} \\ spin up \end{pmatrix}$$

$$v_{12} = \begin{pmatrix} 0 \\ -p \\ 0 \\ E+m \end{pmatrix} \qquad u_{22} = \begin{pmatrix} 0 \\ E+m \\ 0 \\ p \end{pmatrix} \qquad v_{32} = \begin{pmatrix} p_{3}^{x}-ip_{3}^{y} \\ -p_{3}^{z} \\ 0 \\ E+m \end{pmatrix} \qquad u_{42} = \begin{pmatrix} 0 \\ E+m \\ p_{4}^{x}-ip_{4}^{y} \\ -p_{4}^{z} \\ -p_{4}^{z} \\ -p_{4}^{z} \end{pmatrix}$$
spin down

The spinors are not individually normalized. Instead, a combined spinor normalization constant  $N = (E + m)^4$  will be used.

This is the probability density for spin state *abcd*. The formula is derived from Feynman diagrams for Bhabha scattering.

$$|\mathcal{M}_{abcd}|^2 = \frac{e^4}{N} \left| -\frac{1}{t} (\bar{v}_{1a} \gamma^{\mu} v_{3c}) (\bar{u}_{4d} \gamma_{\mu} u_{2b}) + \frac{1}{s} (\bar{v}_{1a} \gamma^{\nu} u_{2b}) (\bar{u}_{4d} \gamma_{\nu} v_{3c}) \right|^2$$

Symbol e is electron charge and

$$s = (p_1 + p_2)^2 = 4E^2$$
  

$$t = (p_1 - p_3)^2 = (p_1 - p_3)^{\mu} g_{\mu\nu} (p_1 - p_3)^{\nu}$$

Let

$$a_1 = (\bar{v}_{1a}\gamma^{\mu}v_{3c})(\bar{u}_{4d}\gamma_{\mu}u_{2b}), \quad a_2 = (\bar{v}_{1a}\gamma^{\nu}u_{2b})(\bar{u}_{4d}\gamma_{\nu}v_{3c})$$

Then

$$|\mathcal{M}_{abcd}|^2 = \frac{e^4}{N} \left| -\frac{a_1}{t} + \frac{a_2}{s} \right|^2$$

$$= \frac{e^4}{N} \left( -\frac{a_1}{t} + \frac{a_2}{s} \right) \left( -\frac{a_1}{t} + \frac{a_2}{s} \right)^*$$

$$= \frac{e^4}{N} \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right)$$

The expected probability density  $\langle |\mathcal{M}|^2 \rangle$  is computed by summing  $|\mathcal{M}_{abcd}|^2$  over all spin states and then dividing by the number of inbound states. There are four inbound states.

$$\langle |\mathcal{M}|^2 \rangle = \frac{1}{4} \sum_{a=1}^2 \sum_{b=1}^2 \sum_{c=1}^2 \sum_{d=1}^2 |\mathcal{M}_{abcd}|^2$$

$$= \frac{e^4}{4N} \sum_{a=1}^2 \sum_{b=1}^2 \sum_{c=1}^2 \sum_{d=1}^2 \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right)$$

The Casimir trick uses matrix arithmetic to compute sums.

$$f_{11} = \frac{1}{N} \sum_{abcd} a_1 a_1^* = \operatorname{Tr} \left( (\not p_1 - m) \gamma^{\mu} (\not p_3 - m) \gamma^{\nu} \right) \operatorname{Tr} \left( (\not p_4 + m) \gamma_{\mu} (\not p_2 + m) \gamma_{\nu} \right)$$

$$f_{12} = \frac{1}{N} \sum_{abcd} a_1 a_2^* = \operatorname{Tr} \left( (\not p_1 - m) \gamma^{\mu} (\not p_2 + m) \gamma^{\nu} (\not p_4 + m) \gamma_{\mu} (\not p_3 - m) \gamma_{\nu} \right)$$

$$f_{22} = \frac{1}{N} \sum_{abcd} a_2 a_2^* = \operatorname{Tr} \left( (\not p_1 - m) \gamma^{\mu} (\not p_2 + m) \gamma^{\nu} \right) \operatorname{Tr} \left( (\not p_4 + m) \gamma_{\mu} (\not p_3 - m) \gamma_{\nu} \right)$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left( \frac{f_{11}}{t^2} - \frac{f_{12}}{st} - \frac{f_{12}^*}{st} + \frac{f_{22}}{s^2} \right)$$

The following formulas are equivalent to the Casimir trick. (Recall that  $a \cdot b = a^{\mu}g_{\mu\nu}b^{\nu}$ )

$$f_{11} = 32(p_1 \cdot p_2)^2 + 32(p_1 \cdot p_4)^2 - 64m^2(p_1 \cdot p_3) + 64m^4$$

$$f_{12} = -32(p_1 \cdot p_4)^2 - 32m^2(p_1 \cdot p_2) + 32m^2(p_1 \cdot p_3) - 32m^2(p_1 \cdot p_4) - 32m^4$$

$$f_{22} = 32(p_1 \cdot p_3)^2 + 32(p_1 \cdot p_4)^2 + 64m^2(p_1 \cdot p_2) + 64m^4$$

For Mandelstam variables

$$s = (p_1 + p_2)^2$$
  

$$t = (p_1 - p_3)^2$$
  

$$u = (p_1 - p_4)^2$$

the formulas are

$$f_{11} = 8s^2 + 8u^2 - 64sm^2 - 64um^2 + 192m^4$$
  

$$f_{12} = -8u^2 + 64um^2 - 96m^4$$
  

$$f_{22} = 8t^2 + 8u^2 - 64tm^2 - 64um^2 + 192m^4$$

### High energy approximation

For high energy experiments  $E\gg m$  a useful approximation is to set m=0 and obtain

$$f_{11} = 8s^2 + 8u^2$$
$$f_{12} = -8u^2$$
$$f_{22} = 8t^2 + 8u^2$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left( \frac{f_{11}}{t^2} - \frac{f_{12}}{st} - \frac{f_{12}^*}{st} + \frac{f_{22}}{s^2} \right)$$

$$= \frac{e^4}{4} \left( \frac{8s^2 + 8u^2}{t^2} - \frac{-8u^2}{st} - \frac{-8u^2}{st} + \frac{8t^2 + 8u^2}{s^2} \right)$$

$$= 2e^4 \left( \frac{s^2 + u^2}{t^2} + \frac{2u^2}{st} + \frac{t^2 + u^2}{s^2} \right)$$

Combine terms so  $\langle |\mathcal{M}|^2 \rangle$  has a common denominator.

$$\langle |\mathcal{M}|^2 \rangle = 2e^4 \left( \frac{s^2 (s^2 + u^2) + 2stu^2 + t^2 (t^2 + u^2)}{s^2 t^2} \right)$$

For m = 0 the Mandelstam variables are

$$s = 4E^{2}$$
  

$$t = 2E^{2}(\cos \theta - 1)$$
  

$$u = -2E^{2}(\cos \theta + 1)$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = 2e^4 \left( \frac{32E^8 \cos^4 \theta + 192E^8 \cos^2 \theta + 288E^8}{64E^8 (\cos \theta - 1)^2} \right)$$
$$= e^4 \left( \frac{\cos^4 \theta + 6 \cos^2 \theta + 9}{(\cos \theta - 1)^2} \right)$$
$$= e^4 \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2$$

### Cross section

The differential cross section is

$$\frac{d\sigma}{d\Omega} = \frac{\langle |\mathcal{M}|^2 \rangle}{4(4\pi\varepsilon_0)^2 s}, \quad s = (p_1 + p_2)^2 = 4E^2$$

For high energy experiments we have

$$\langle |\mathcal{M}|^2 \rangle = e^4 \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2$$

Substitute for  $\langle |\mathcal{M}|^2 \rangle$ .

$$\frac{d\sigma}{d\Omega} = \frac{e^4}{4(4\pi\varepsilon_0)^2 s} \left(\frac{\cos^2\theta + 3}{\cos\theta - 1}\right)^2$$

Noting that

$$e^2 = 4\pi\varepsilon_0 \alpha \hbar c$$

we can also write

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2 (\hbar c)^2}{4s} \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2$$

We can integrate  $d\sigma$  to obtain a cumulative distribution function. Let  $I(\theta)$  be the following integral of  $d\sigma$ . (The  $\sin\theta$  is from  $d\Omega = \sin\theta \, d\theta \, d\phi$ .)

$$I(\theta) = \int \left(\frac{\cos^2 \theta + 3}{\cos \theta - 1}\right)^2 \sin \theta \, d\theta$$

The result is

$$I(\theta) = \frac{16}{\cos \theta - 1} - \frac{\cos^3 \theta}{3} - \cos^2 \theta - 9\cos \theta - 16\log(1 - \cos \theta)$$

The cumulative distribution function is

$$F(\theta) = \frac{I(\theta) - I(a)}{I(\pi) - I(a)}, \quad a \le \theta \le \pi$$

Angular support is reduced by an arbitrary angle a > 0 because I(0) is undefined.

The probability of observing scattering events in the interval  $\theta_1$  to  $\theta_2$  is

$$P(\theta_1 \le \theta \le \theta_2) = F(\theta_2) - F(\theta_1)$$

Let N be the number of scattering events from an experiment. Then the number of scattering events in the interval  $\theta_1$  to  $\theta_2$  is predicted to be

$$N\left(F(\theta_2) - F(\theta_1)\right)$$

The probability density function is

$$f(\theta) = \frac{dF(\theta)}{d\theta} = \frac{1}{I(\pi) - I(a)} \left(\frac{\cos^2 \theta + 3}{\cos \theta - 1}\right)^2 \sin \theta$$

Note that if we had carried through the  $\alpha^2(\hbar c)^2/4s$  in  $I(\theta)$ , it would have cancelled out in  $F(\theta)$ .

# Data from SLAC SPEAR experiment

The following Bhabha scattering data is adapted from SLAC-PUB-1501.

k	$x_k, x_{k+1}$	y
1	0.6, 0.5	4432
2	0.5, 0.4	2841
3	0.4, 0.3	2045
4	0.3, 0.2	1420
5	0.2, 0.1	1136
6	0.1, 0.0	852
7	0.0, -0.1	656
8	-0.1, -0.2	625
9	-0.2, -0.3	511
10	-0.3, -0.4	455
11	-0.4, -0.5	402
12	-0.5, -0.6	398

Column k is the bin number, column y is the number of scattering events, and

$$x_k = \cos \theta_k$$

The cumulative distribution function for this experiment is

$$F(\theta) = \frac{I(\theta) - I(\theta_1)}{I(\theta_{13}) - I(\theta_1)}$$

where

$$\theta_{13} = \arccos(-0.6), \quad \theta_1 = \arccos(0.6)$$

The scattering probability  $P_k$  is

$$P_k = F\left(\arccos(x_{k+1})\right) - F\left(\arccos(x_k)\right)$$

Multiply  $P_k$  by total scattering events to obtain predicted number of events  $\hat{y}_k$ .

$$\sum y_k = 15773, \quad \hat{y}_k = 15773 \, P_k$$

Bin	$x_k, x_{k+1}$	y	$\hat{y}$
1	0.6, 0.5	4432	4598
2	0.5, 0.4	2841	2880
3	0.4, 0.3	2045	1955
4	0.3, 0.2	1420	1410
5	0.2, 0.1	1136	1068
6	0.1, 0.0	852	843
7	0.0, -0.1	656	689
8	-0.1, -0.2	625	582
9	-0.2, -0.3	511	505
10	-0.3, -0.4	455	450
11	-0.4, -0.5	402	411
12	-0.5, -0.6	398	382

The coefficient of determination  $R^2$  measures how well predicted values fit the data.

$$R^{2} = 1 - \frac{\sum (y - \hat{y})^{2}}{\sum (y - \bar{y})^{2}} = 0.997$$

The result indicates that  $F(\theta)$  explains 99.7% of the variance in the data.

## Data from DESY PETRA experiment

See www.hepdata.net/record/ins191231, Table 3, 14.0 GeV.

x	y
-0.73	0.10115
-0.6495	0.12235
-0.5495	0.11258
-0.4494	0.09968
-0.3493	0.14749
-0.2491	0.14017
-0.149	0.1819
-0.0488	0.22964
0.0514	0.25312
0.1516	0.30998
0.252	0.40898
0.3524	0.62695
0.4529	0.91803
0.5537	1.51743
0.6548	2.56714
0.7323	4.30279

Data x and y have the following relationship with the differential cross section formula.

$$x = \cos \theta$$
,  $y = \frac{d\sigma}{d\Omega}$  in nanobarns

The cross section formula is

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2}{4s} \left( \frac{\cos^2 \theta + 3}{\cos \theta - 1} \right)^2 \times (\hbar c)^2$$

To compute predicted values  $\hat{y}$ , multiply by  $10^{37}$  to convert square meters to nanobarns.

$$\hat{y} = \frac{\alpha^2}{4s} \left( \frac{x^2 + 3}{x - 1} \right)^2 \times (\hbar c)^2 \times 10^{37}$$

The following table shows predicted values  $\hat{y}$  for  $s = (14.0 \,\text{GeV})^2$ .

x	y	$\hat{y}$
-0.73	0.10115	0.110296
-0.6495	0.12235	0.113816
-0.5495	0.11258	0.120101
-0.4494	0.09968	0.129075
-0.3493	0.14749	0.141592
-0.2491	0.14017	0.158934
-0.149	0.1819	0.182976
-0.0488	0.22964	0.216737
0.0514	0.25312	0.264989
0.1516	0.30998	0.335782
0.252	0.40898	0.44363
0.3524	0.62695	0.615528
0.4529	0.91803	0.9077
0.5537	1.51743	1.45175
0.6548	2.56714	2.60928
0.7323	4.30279	4.61509

The coefficient of determination  $R^2$  measures how well predicted values fit the data.

$$R^{2} = 1 - \frac{\sum (y - \hat{y})^{2}}{\sum (y - \bar{y})^{2}} = 0.995$$

The result indicates that the model  $d\sigma$  explains 99.5% of the variance in the data.

#### Notes

Here are a few notes about how the Eigenmath scripts work. In component notation the trace operators of the Casimir trick become sums over the repeated index  $\alpha$ .

$$\begin{split} f_{11} &= \left( (\not\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_3 - m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \right) \left( (\not\!p_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \right) \\ f_{12} &= (\not\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\tau (\not\!p_4 + m)^\tau{}_\delta \gamma_\mu{}^\delta{}_\eta (\not\!p_3 - m)^\eta{}_\xi \gamma_\nu{}^\xi{}_\alpha \\ f_{22} &= \left( (\not\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \right) \left( (\not\!p_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not\!p_3 - m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \right) \end{split}$$

To convert the above formulas to Eigenmath code, the  $\gamma$  tensors need to be transposed so that repeated indices are adjacent to each other. Also, multiply  $\gamma^{\mu}$  by the metric tensor to lower the index.

$$\gamma^{\beta\mu}{}_{\rho}$$
  $\to$  gammaT = transpose(gamma)  $\gamma^{\beta}{}_{\mu\rho}$   $\to$  gammaL = transpose(dot(gmunu,gamma))

Define the following  $4 \times 4$  matrices.

Then for  $f_{11}$  we have the following Eigenmath code. The contract function sums over  $\alpha$ .

$$(\not\!\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!\!p_3 - m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \quad \rightarrow \quad \text{T1 = contract(dot(X1,gammaT,X3,gammaT),1,4)}$$
 
$$(\not\!\!p_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not\!\!p_2 + m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \quad \rightarrow \quad \text{T2 = contract(dot(X4,gammaL,X2,gammaL),1,4)}$$

Next, multiply then sum over repeated indices. The dot function sums over  $\nu$  then the contract function sums over  $\mu$ . The transpose makes the  $\nu$  indices adjacent as required by the dot function.

$$f_{11} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu}) \operatorname{Tr}(\cdots \gamma_{\mu} \cdots \gamma_{\nu}) \rightarrow \operatorname{fll} = \operatorname{contract}(\operatorname{dot}(\operatorname{T1,transpose}(\operatorname{T2})))$$

Follow suit for  $f_{22}$ .

$$(\not\!p_1-m)^\alpha{}_\beta\gamma^{\mu\beta}{}_\rho(\not\!p_2+m)^\rho{}_\sigma\gamma^{\nu\sigma}{}_\alpha \quad \rightarrow \quad \text{T1 = contract(dot(X1,gammaT,X2,gammaT),1,4)}$$
 
$$(\not\!p_4+m)^\alpha{}_\beta\gamma_\mu{}^\beta{}_\rho(\not\!p_3-m)^\rho{}_\sigma\gamma_\nu{}^\sigma{}_\alpha \quad \rightarrow \quad \text{T2 = contract(dot(X4,gammaL,X3,gammaL),1,4)}$$

Hence

$$f_{22} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu}) \operatorname{Tr}(\cdots \gamma_{\mu} \cdots \gamma_{\nu}) \quad o \quad exttt{f22} = ext{contract(dot(T1,transpose(T2)))}$$

The calculation of  $f_{12}$  begins with

$$\begin{split} (\not\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\tau (\not\!p_4 + m)^\tau{}_\delta \gamma_\mu{}^\delta{}_\eta (\not\!p_3 - m)^\eta{}_\xi \gamma_\nu{}^\xi{}_\alpha \\ & \to \quad \text{T = contract(dot(X1,gammaT,X2,gammaT,X4,gammaL,X3,gammaL),1,6)} \end{split}$$

Then sum over repeated indices  $\mu$  and  $\nu$ .

$$f_{12} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu} \cdots \gamma_{\mu} \cdots \gamma_{\nu}) \quad o \quad ext{f12} = ext{contract(Contract(T,1,3))}$$