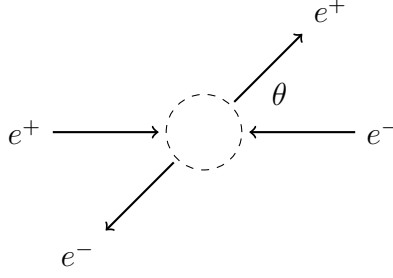
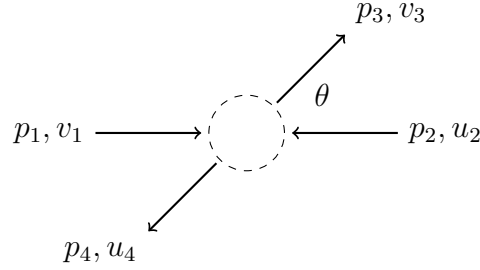


Bhabha scattering is the result of interactions between positrons and electrons. The following diagram represents a collider experiment with collinear electron and positron beams.



Here is the same diagram with momentum and spinor labels.



In a typical collider experiment the momentum vectors are

$$p_1 = \begin{pmatrix} E \\ 0 \\ 0 \\ p \end{pmatrix} \quad p_2 = \begin{pmatrix} E \\ 0 \\ 0 \\ -p \end{pmatrix} \quad p_3 = \begin{pmatrix} E \\ p \sin \theta \cos \phi \\ p \sin \theta \sin \phi \\ p \cos \theta \end{pmatrix} \quad p_4 = \begin{pmatrix} E \\ -p \sin \theta \cos \phi \\ -p \sin \theta \sin \phi \\ -p \cos \theta \end{pmatrix}$$

where  $E = \sqrt{p^2 + m^2}$ . The spinors are

$$\begin{aligned} v_{11} &= \begin{pmatrix} p \\ 0 \\ E + m \\ 0 \end{pmatrix} & u_{21} &= \begin{pmatrix} E + m \\ 0 \\ -p \\ 0 \end{pmatrix} & v_{31} &= \begin{pmatrix} p_3^z \\ p_3^x + ip_3^y \\ E + m \\ 0 \end{pmatrix} & u_{41} &= \begin{pmatrix} E + m \\ 0 \\ p_4^z \\ p_4^x + ip_4^y \end{pmatrix} \\ v_{12} &= \begin{pmatrix} 0 \\ -p \\ 0 \\ E + m \end{pmatrix} & u_{22} &= \begin{pmatrix} 0 \\ E + m \\ 0 \\ p \end{pmatrix} & v_{32} &= \begin{pmatrix} p_3^x - ip_3^y \\ -p_3^z \\ 0 \\ E + m \end{pmatrix} & u_{42} &= \begin{pmatrix} 0 \\ E + m \\ p_4^x - ip_4^y \\ -p_4^z \end{pmatrix} \end{aligned}$$

The spinors shown above are not individually normalized. Instead, a combined spinor normalization constant  $N = (E + m)^4$  will be used.

This is the probability density for Bhabha scattering.

$$|\mathcal{M}(s_1, s_2, s_3, s_4)|^2 = \frac{e^4}{N} \left| -\frac{1}{t} (\bar{v}_1 \gamma^\mu v_3) (\bar{u}_4 \gamma_\mu u_2) + \frac{1}{s} (\bar{v}_1 \gamma^\nu u_2) (\bar{u}_4 \gamma_\nu v_3) \right|^2$$

Symbol  $s_j$  selects the spin (up or down) of spinor  $j$ . Symbol  $e$  is electron charge. Symbols  $s$  and  $t$  are Mandelstam variables  $s = (p_1 + p_2)^2$  and  $t = (p_1 - p_3)^2$ .

Let

$$a_1 = (\bar{v}_1 \gamma^\mu v_3)(\bar{u}_4 \gamma_\mu u_2) \quad a_2 = (\bar{v}_1 \gamma^\nu u_2)(\bar{u}_4 \gamma_\nu v_3)$$

Then

$$\begin{aligned} |\mathcal{M}(s_1, s_2, s_3, s_4)|^2 &= \frac{e^4}{N} \left| -\frac{a_1}{t} + \frac{a_2}{s} \right|^2 \\ &= \frac{e^4}{N} \left( -\frac{a_1}{t} + \frac{a_2}{s} \right) \left( -\frac{a_1}{t} + \frac{a_2}{s} \right)^* \\ &= \frac{e^4}{N} \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right) \end{aligned}$$

The expected probability density  $\langle |\mathcal{M}|^2 \rangle$  is computed by summing  $|\mathcal{M}|^2$  over all spin states and dividing by the number of inbound states. There are four inbound states.

$$\begin{aligned} \langle |\mathcal{M}|^2 \rangle &= \frac{1}{4} \sum_{s_1=1}^2 \sum_{s_2=1}^2 \sum_{s_3=1}^2 \sum_{s_4=1}^2 |\mathcal{M}(s_1, s_2, s_3, s_4)|^2 \\ &= \frac{e^4}{4} \sum_{s_1=1}^2 \sum_{s_2=1}^2 \sum_{s_3=1}^2 \sum_{s_4=1}^2 \frac{1}{N} \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right) \end{aligned}$$

Use the Casimir trick to replace sums over spins with matrix products.

$$\begin{aligned} f_{11} &= \frac{1}{N} \sum_{\text{spins}} a_1 a_1^* = \text{Tr} \left( (\not{p}_1 - m) \gamma^\mu (\not{p}_3 - m) \gamma^\nu \right) \text{Tr} \left( (\not{p}_4 + m) \gamma_\mu (\not{p}_2 + m) \gamma_\nu \right) \\ f_{12} &= \frac{1}{N} \sum_{\text{spins}} a_1 a_2^* = \text{Tr} \left( (\not{p}_1 - m) \gamma^\mu (\not{p}_2 + m) \gamma^\nu (\not{p}_4 + m) \gamma_\mu (\not{p}_3 - m) \gamma_\nu \right) \\ f_{22} &= \frac{1}{N} \sum_{\text{spins}} a_2 a_2^* = \text{Tr} \left( (\not{p}_1 - m) \gamma^\mu (\not{p}_2 + m) \gamma^\nu \right) \text{Tr} \left( (\not{p}_4 + m) \gamma_\mu (\not{p}_3 - m) \gamma_\nu \right) \end{aligned}$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left( \frac{f_{11}}{t^2} - \frac{f_{12}}{st} - \frac{f_{12}^*}{st} + \frac{f_{22}}{s^2} \right)$$

Run “bhabha-scattering-1.txt” to verify the Casimir trick.

The following momentum formulas are equivalent to the Casimir trick. (Recall that  $a \cdot b = a^\mu g_{\mu\nu} b^\nu$ )

$$\begin{aligned} f_{11} &= 32(p_1 \cdot p_2)(p_3 \cdot p_4) + 32(p_1 \cdot p_4)(p_2 \cdot p_3) - 32m^2(p_1 \cdot p_3) - 32m^2(p_2 \cdot p_4) + 64m^4 \\ f_{12} &= -32(p_1 \cdot p_4)(p_2 \cdot p_3) - 16m^2(p_1 \cdot p_2) + 16m^2(p_1 \cdot p_3) - 16m^2(p_1 \cdot p_4) \\ &\quad - 16m^2(p_2 \cdot p_3) + 16m^2(p_2 \cdot p_4) - 16m^2(p_3 \cdot p_4) - 32m^4 \\ f_{22} &= 32(p_1 \cdot p_3)(p_2 \cdot p_4) + 32(p_1 \cdot p_4)(p_2 \cdot p_3) + 32m^2(p_1 \cdot p_2) + 32m^2(p_3 \cdot p_4) + 64m^4 \end{aligned}$$

In Mandelstam variables  $s = (p_1 + p_2)^2$ ,  $t = (p_1 - p_3)^2$ ,  $u = (p_1 - p_4)^2$  the formulas are

$$\begin{aligned} f_{11} &= 8s^2 + 8u^2 - 64sm^2 - 64um^2 + 192m^4 \\ f_{12} &= -8u^2 + 64um^2 - 96m^4 \\ f_{22} &= 8t^2 + 8u^2 - 64tm^2 - 64um^2 + 192m^4 \end{aligned}$$

## High energy approximation

When  $E \gg m$  a useful approximation is to set  $m = 0$  and obtain

$$\begin{aligned}f_{11} &= 8s^2 + 8u^2 \\f_{12} &= -8u^2 \\f_{22} &= 8t^2 + 8u^2\end{aligned}$$

For  $m = 0$  the Mandelstam variables are

$$\begin{aligned}s &= 4E^2 \\t &= -2E^2(1 - \cos \theta) = -4E^2 \sin^2(\theta/2) \\u &= -2E^2(1 + \cos \theta) = -4E^2 \cos^2(\theta/2)\end{aligned}$$

It follows that

$$s^2 t^2 = 256 E^8 \sin^4(\theta/2)$$

The corresponding expected probability density is

$$\begin{aligned}\langle |\mathcal{M}|^2 \rangle &= \frac{e^4}{4} \left( \frac{f_{11}}{t^2} - \frac{f_{12}}{st} - \frac{f_{12}^*}{st} + \frac{f_{22}}{s^2} \right) \\&= \frac{e^4}{4s^2 t^2} (s^2 f_{11} - st f_{12} - st f_{12}^* + t^2 f_{22}) \\&= \frac{e^4}{4s^2 t^2} (s^2 (8s^2 + 8u^2) + 16stu^2 + t^2 (8t^2 + 8u^2)) \\&= \frac{e^4}{1024 E^8 \sin^4(\theta/2)} (256 E^8 \cos^4 \theta + 1536 E^8 \cos^2 \theta + 2304 E^8) \\&= \frac{e^4}{4 \sin^4(\theta/2)} (\cos^4 \theta + 6 \cos^2 \theta + 9) \\&= \frac{e^4}{4} \frac{(\cos^2 \theta + 3)^2}{\sin^4(\theta/2)}\end{aligned}$$

Run “bhabha-scattering-2.txt” to verify.

## Cross section

This is the differential cross section for Bhabha scattering.

$$\frac{d\sigma}{d\Omega} = \frac{\langle |\mathcal{M}|^2 \rangle}{64\pi^2 s} = \frac{e^4}{1024\pi^2 E^2} \frac{(\cos^2 \theta + 3)^2}{\sin^4(\theta/2)}$$

Substituting  $e^4 = 16\pi^2 \alpha^2$  yields

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2}{64 E^2} \frac{(\cos^2 \theta + 3)^2}{\sin^4(\theta/2)}$$

We can integrate  $d\sigma$  to obtain a cumulative distribution function. Recall that

$$d\Omega = \sin \theta d\theta d\phi$$

Hence

$$d\sigma = \frac{\alpha^2}{64E^2} \frac{(\cos^2 \theta + 3)^2}{\sin^4(\theta/2)} \sin \theta d\theta d\phi$$

Let  $I(\xi)$  be the following definite integral.

$$\begin{aligned} I(\xi) &= \frac{64E^2}{2\pi\alpha^2} \int_0^{2\pi} \int_a^\xi d\sigma \sin \theta d\theta d\phi \\ &= \int_a^\xi \frac{(\cos^2 \theta + 3)^2}{\sin^4(\theta/2)} \sin \theta d\theta, \quad a \leq \xi \leq \pi \end{aligned}$$

Angular support is limited to  $a > 0$  because  $I(0)$  is undefined.

Let  $C$  be the normalization constant  $C = I(\pi)$ . Then the cumulative distribution function  $F(\theta)$  is

$$F(\theta) = C^{-1} I(\theta), \quad a \leq \theta \leq \pi$$

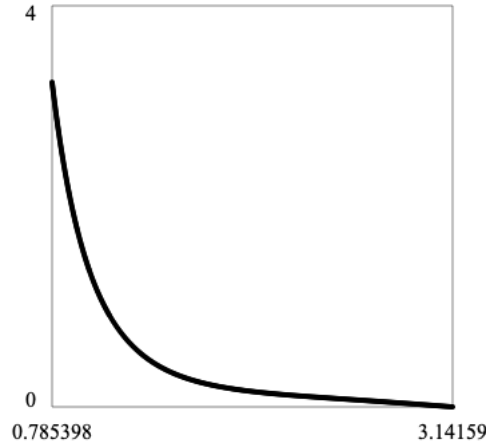
The probability of observing scattering events in the interval  $\theta_1$  to  $\theta_2$  can now be computed.

$$P(\theta_1 \leq \theta \leq \theta_2) = F(\theta_2) - F(\theta_1)$$

Probability density function  $f(\theta)$  is the derivative of  $F(\theta)$ .

$$f(\theta) = \frac{dF(\theta)}{d\theta} = C^{-1} \frac{(\cos^2 \theta + 3)^2}{\sin^4(\theta/2)} \sin \theta$$

Run “bhabha-scattering-3.txt” to draw  $f(\theta)$  for  $a = \pi/4 = 45^\circ$ .



The following table shows the corresponding probability distribution for three bins.

$\theta_1$	$\theta_2$	$P(\theta_1 \leq \theta \leq \theta_2)$
$0^\circ$	$45^\circ$	—
$45^\circ$	$90^\circ$	0.83
$90^\circ$	$135^\circ$	0.13
$135^\circ$	$180^\circ$	0.04

## SLAC data

The following Bhabha scattering data is adapted from SLAC-PUB-1501.

	Bin	$\cos \theta$ (interval)	Count
(Smallest $\theta$ )	1	0.6, 0.5	4432
	2	0.5, 0.4	2841
	3	0.4, 0.3	2045
	4	0.3, 0.2	1420
	5	0.2, 0.1	1136
	6	0.1, 0.0	852
	7	0.0, -0.1	656
	8	-0.1, -0.2	625
	9	-0.2, -0.3	511
	10	-0.3, -0.4	455
	11	-0.4, -0.5	402
(Largest $\theta$ )	12	-0.5, -0.6	398

“Count” is the number of Bhabha scattering events observed per bin. Let us see if the density function  $\langle |\mathcal{M}|^2 \rangle$  explains the distribution of counts in the table. Start by integrating  $\langle |\mathcal{M}|^2 \rangle$  over all the bins to obtain a normalization constant.

$$\int_{\text{bins}} \langle |\mathcal{M}|^2 \rangle d\Omega = \int_0^{2\pi} \int_{\arccos 0.6}^{\arccos -0.6} \langle |\mathcal{M}|^2 \rangle \sin \theta d\theta d\phi = 2\pi \times 9.3817 \times 2e^4$$

Let

$$f(\theta) = \frac{\langle |\mathcal{M}|^2 \rangle}{2\pi \times 9.3817 \times 2e^4} = \frac{1}{2\pi \times 9.3817} \left( \frac{1 + \cos^4(\theta/2)}{\sin^4(\theta/2)} - \frac{2 \cos^4(\theta/2)}{\sin^2(\theta/2)} + \frac{1 + \cos^2 \theta}{2} \right)$$

The probability of a scattering event occurring in an interval  $\theta_1$  to  $\theta_2$  is obtained by integrating  $f(\theta)$  over that interval.

$$P(\theta_1 < \theta < \theta_2) = \int_0^{2\pi} \int_{\theta_1}^{\theta_2} f(\theta) \sin \theta d\theta d\phi = 2\pi \int_{\theta_1}^{\theta_2} f(\theta) \sin \theta d\theta$$

The total number of counts in the table is 15773. To obtain a predicted distribution, multiply 15773 times the probability for each bin. For example, for the first bin we have

$$P(\arccos 0.6 < \theta < \arccos 0.5) \times 15773 = 4598$$

Repeat for all bins to obtain the following predicted distribution.

Bin	$\cos \theta$ (interval)	Count	Predicted
1	0.6, 0.5	4432	4598
2	0.5, 0.4	2841	2880
3	0.4, 0.3	2045	1955
4	0.3, 0.2	1420	1410
5	0.2, 0.1	1136	1068
6	0.1, 0.0	852	843
7	0.0, -0.1	656	689
8	-0.1, -0.2	625	582
9	-0.2, -0.3	511	505
10	-0.3, -0.4	455	450
11	-0.4, -0.5	402	411
12	-0.5, -0.6	398	382

The coefficient of determination  $R^2$  measures how well predicted values fit the real data. Let  $y$  be observed counts per bin and let  $\hat{y}$  be predicted counts per bin. Then

$$R^2 = 1 - \frac{\sum (y - \hat{y})^2}{\sum (y - \bar{y})^2} = 0.997$$

The result indicates that the model  $\langle |\mathcal{M}|^2 \rangle$  explains 99.7% of the variance in the data.

Run “bhabha-scattering-4.txt” to verify.

## DESY data

The following table shows DESY-PETRA Bhabha scattering data obtained from HEP Data.<sup>1</sup>

$x$	$y$
-0.73	0.10115
-0.6495	0.12235
-0.5495	0.11258
-0.4494	0.09968
-0.3493	0.14749
-0.2491	0.14017
-0.149	0.1819
-0.0488	0.22964
0.0514	0.25312
0.1516	0.30998
0.252	0.40898
0.3524	0.62695
0.4529	0.91803
0.5537	1.51743
0.6548	2.56714
0.7323	4.30279

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<sup>1</sup>[www.hepdata.net/record/ins191231](http://www.hepdata.net/record/ins191231) (Table 3, 14.0 GeV)

Data  $x$  and  $y$  have the following relationship with the cross section model.

$$x = \cos \theta \quad y = \frac{d\sigma}{d\Omega}$$

The differential cross section for Bhabha scattering is

$$\frac{d\sigma}{d\Omega} = \frac{\langle |\mathcal{M}|^2 \rangle}{64\pi^2 s} = \frac{\alpha^2}{2s} \left( \frac{s^2 + u^2}{t^2} + \frac{2u^2}{st} + \frac{t^2 + u^2}{s^2} \right)$$

The predicted cross section  $\hat{y}$  is computed from data  $x$  and beam energy  $E$  as

$$\hat{y} = \frac{\alpha^2}{2s} \left( \frac{s^2 + u^2}{t^2} + \frac{2u^2}{st} + \frac{t^2 + u^2}{s^2} \right) \times (\hbar c)^2 \times 10^{37}$$

where

$$\begin{aligned} s &= 4E^2 \\ t &= -2E^2(1 - x) \\ u &= -2E^2(1 + x) \end{aligned}$$

Factor  $(\hbar c)^2$  converts the result to SI and factor  $10^{37}$  converts square meters to nanobarns.

The following table shows  $\hat{y}$  for  $E = 7.0$  GeV.

$x$	$y$	$\hat{y}$
-0.73	0.10115	0.110296
-0.6495	0.12235	0.113816
-0.5495	0.11258	0.120101
-0.4494	0.09968	0.129075
-0.3493	0.14749	0.141592
-0.2491	0.14017	0.158934
-0.149	0.1819	0.182976
-0.0488	0.22964	0.216737
0.0514	0.25312	0.264989
0.1516	0.30998	0.335782
0.252	0.40898	0.44363
0.3524	0.62695	0.615528
0.4529	0.91803	0.9077
0.5537	1.51743	1.45175
0.6548	2.56714	2.60928
0.7323	4.30279	4.61509

The coefficient of determination  $R^2$  measures how well predicted values fit the real data.

$$R^2 = 1 - \frac{\sum (y - \hat{y})^2}{\sum (y - \bar{y})^2} = 0.995$$

The result indicates that the model  $d\sigma$  explains 99.5% of the variance in the data.

Run “bhabha-scattering-5.txt” to verify.

# Notes on Eigenmath scripts

Here are a few notes about how the Eigenmath scripts work. In component notation the trace operators of the Casimir trick become sums over the repeated index  $\alpha$ .

$$\begin{aligned} f_{11} &= \left( (\not{p}_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not{p}_3 - m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \right) \left( (\not{p}_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not{p}_2 + m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \right) \\ f_{12} &= (\not{p}_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not{p}_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\tau (\not{p}_4 + m)^\tau{}_\delta \gamma_\mu{}^\delta{}_\eta (\not{p}_3 - m)^\eta{}_\xi \gamma_\nu{}^\xi{}_\alpha \\ f_{22} &= \left( (\not{p}_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not{p}_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \right) \left( (\not{p}_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not{p}_3 - m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \right) \end{aligned}$$

To convert the above formulas to Eigenmath code, the  $\gamma$  tensors need to be transposed so that repeated indices are adjacent to each other. Also, multiply  $\gamma^\mu$  by the metric tensor to lower the index.

$$\begin{aligned} \gamma^{\beta\mu}{}_\rho &\rightarrow \text{gammaT} = \text{transpose}(\text{gamma}) \\ \gamma^\beta{}_{\mu\rho} &\rightarrow \text{gammaL} = \text{transpose}(\text{dot}(\text{gmunu}, \text{gamma})) \end{aligned}$$

Define the following  $4 \times 4$  matrices.

$$\begin{aligned} (\not{p}_1 - m) &\rightarrow \text{X1} = \text{pslash1} - \text{m I} \\ (\not{p}_2 + m) &\rightarrow \text{X2} = \text{pslash2} + \text{m I} \\ (\not{p}_3 - m) &\rightarrow \text{X3} = \text{pslash3} - \text{m I} \\ (\not{p}_4 + m) &\rightarrow \text{X4} = \text{pslash4} + \text{m I} \end{aligned}$$

Then for  $f_{11}$  we have the following Eigenmath code. The contract function sums over  $\alpha$ .

$$\begin{aligned} (\not{p}_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not{p}_3 - m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha &\rightarrow \text{T1} = \text{contract}(\text{dot}(\text{X1}, \text{gammaT}, \text{X3}, \text{gammaT}), 1, 4) \\ (\not{p}_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not{p}_2 + m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha &\rightarrow \text{T2} = \text{contract}(\text{dot}(\text{X4}, \text{gammaL}, \text{X2}, \text{gammaL}), 1, 4) \end{aligned}$$

Next, multiply then sum over repeated indices. The dot function sums over  $\nu$  then the contract function sums over  $\mu$ . The transpose makes the  $\nu$  indices adjacent as required by the dot function.

$$f_{11} = \text{Tr}(\dots \gamma^\mu \dots \gamma^\nu) \text{Tr}(\dots \gamma_\mu \dots \gamma_\nu) \rightarrow \text{f11} = \text{contract}(\text{dot}(\text{T1}, \text{transpose}(\text{T2})))$$

Follow suit for  $f_{22}$ .

$$\begin{aligned} (\not{p}_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not{p}_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha &\rightarrow \text{T1} = \text{contract}(\text{dot}(\text{X1}, \text{gammaT}, \text{X2}, \text{gammaT}), 1, 4) \\ (\not{p}_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not{p}_3 - m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha &\rightarrow \text{T2} = \text{contract}(\text{dot}(\text{X4}, \text{gammaL}, \text{X3}, \text{gammaL}), 1, 4) \end{aligned}$$

Hence

$$f_{22} = \text{Tr}(\dots \gamma^\mu \dots \gamma^\nu) \text{Tr}(\dots \gamma_\mu \dots \gamma_\nu) \rightarrow \text{f22} = \text{contract}(\text{dot}(\text{T1}, \text{transpose}(\text{T2})))$$

The calculation of  $f_{12}$  begins with

$$\begin{aligned} (\not{p}_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not{p}_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\tau (\not{p}_4 + m)^\tau{}_\delta \gamma_\mu{}^\delta{}_\eta (\not{p}_3 - m)^\eta{}_\xi \gamma_\nu{}^\xi{}_\alpha \\ \rightarrow \text{T} = \text{contract}(\text{dot}(\text{X1}, \text{gammaT}, \text{X2}, \text{gammaT}, \text{X4}, \text{gammaL}, \text{X3}, \text{gammaL}), 1, 6) \end{aligned}$$

Then sum over repeated indices  $\mu$  and  $\nu$ .

$$f_{12} = \text{Tr}(\dots \gamma^\mu \dots \gamma^\nu \dots \gamma_\mu \dots \gamma_\nu) \rightarrow \text{f12} = \text{contract}(\text{contract}(\text{T}, 1, 3))$$